

Special Relativity Introduction for General Relativity 2

Monday 01.30.2017

Review: Relativity ρ functions Two famous ones Extremes and plot vs. ρ
Doppler jeopardy Geometric mean and Relativistic hyperbolas
Animation of $e^{\rho}=2$ spacetime and per-spacetime plots

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity
Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry
“Occams Sword” and summary of 16 parameter functions of ρ and σ
Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room
Relativistic action and Lagrangian-Hamiltonian relations
Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization “photon” number N and 1st-quantization wavenumber κ

Relativity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes g -acceleration grid

Analysis of constant- g grid compared to zero- g Minkowski grid

Animation of mechanics and metrology of constant- g grid

Lecture 31

Thur. 12.08.2016

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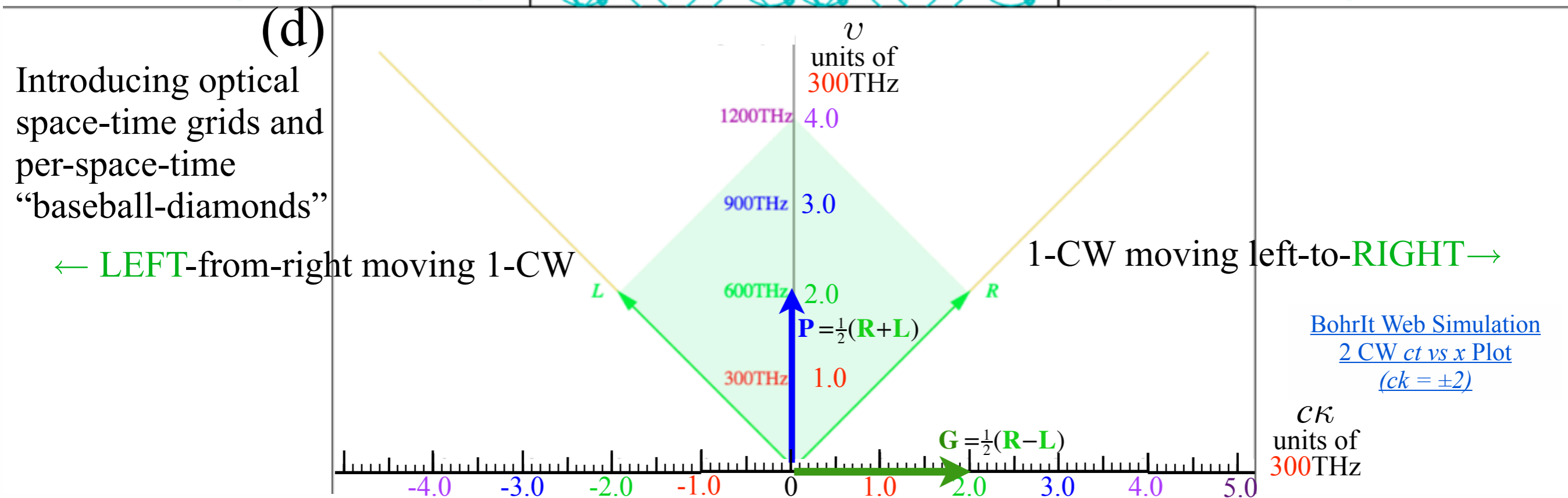
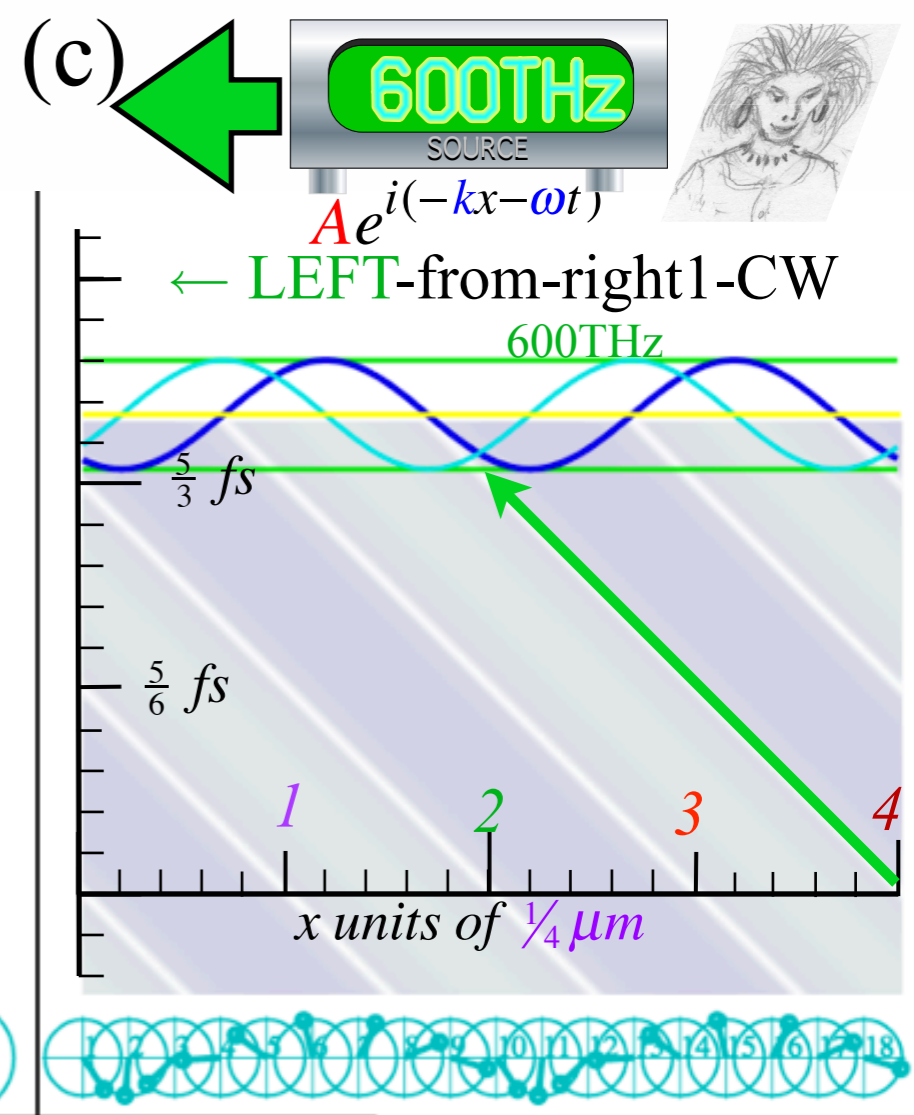
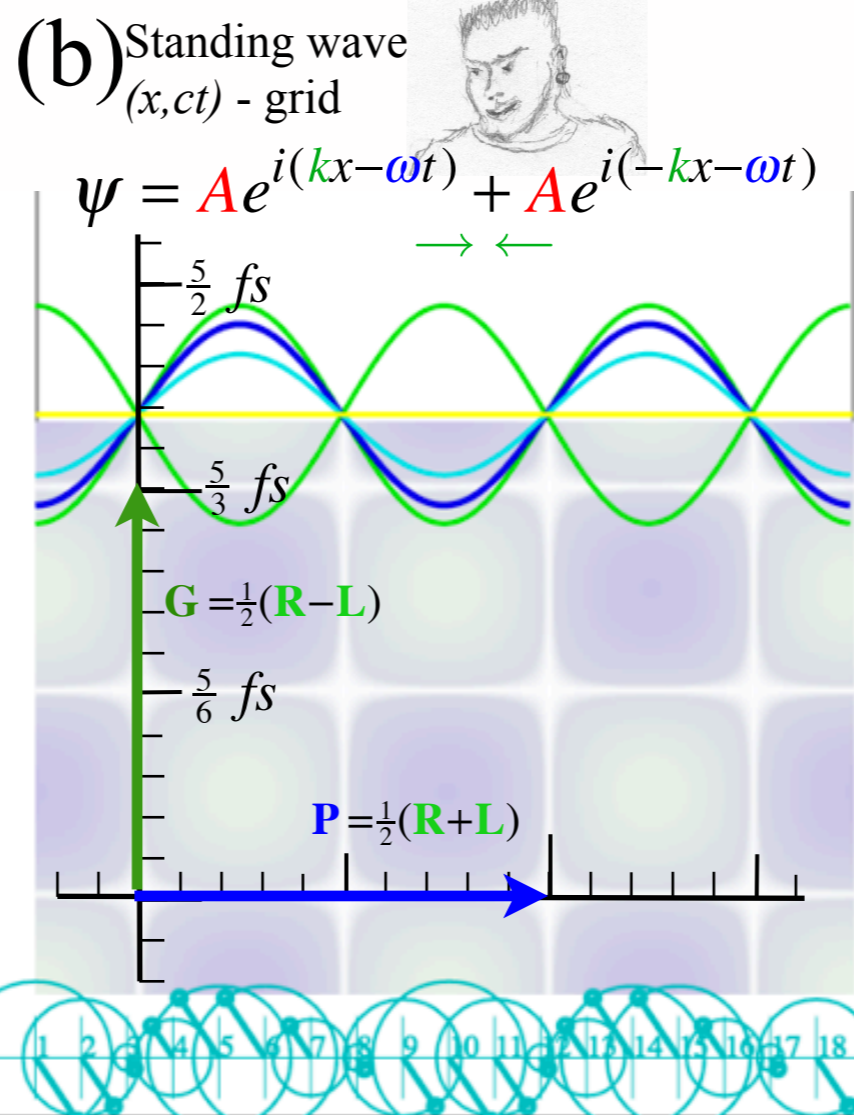
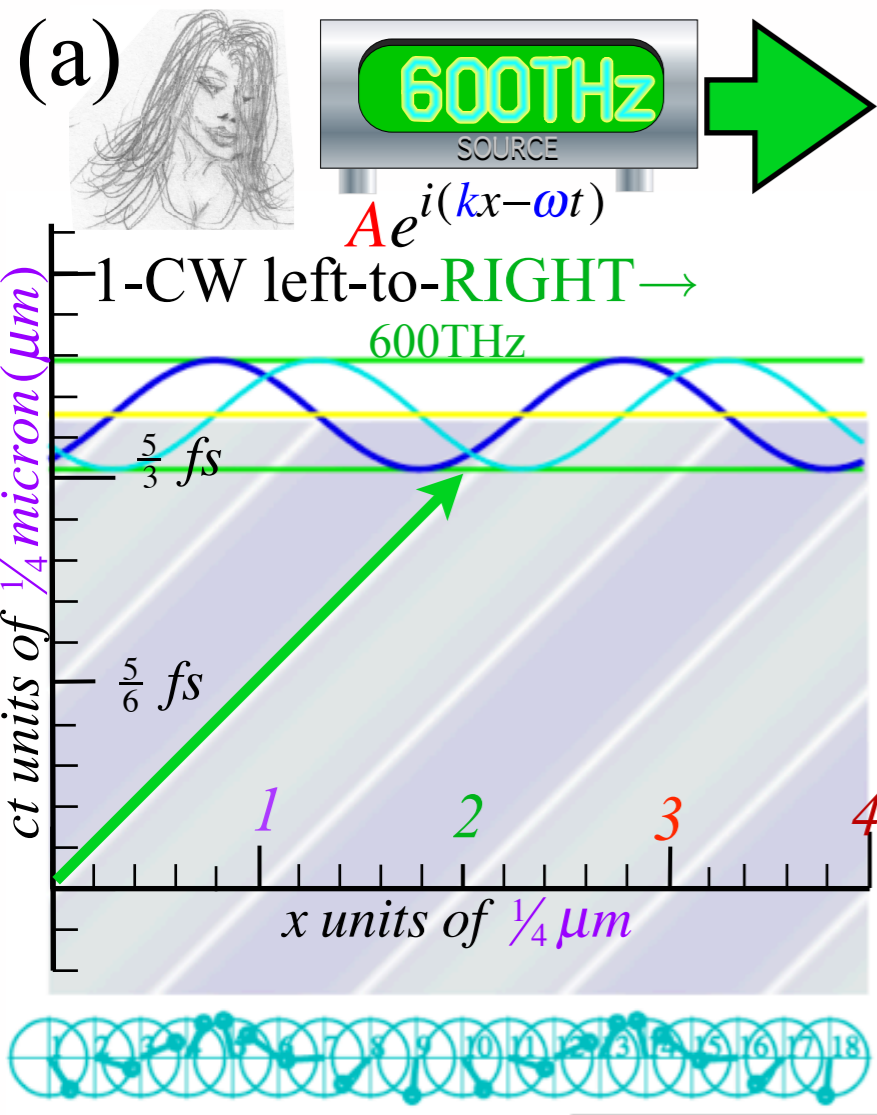
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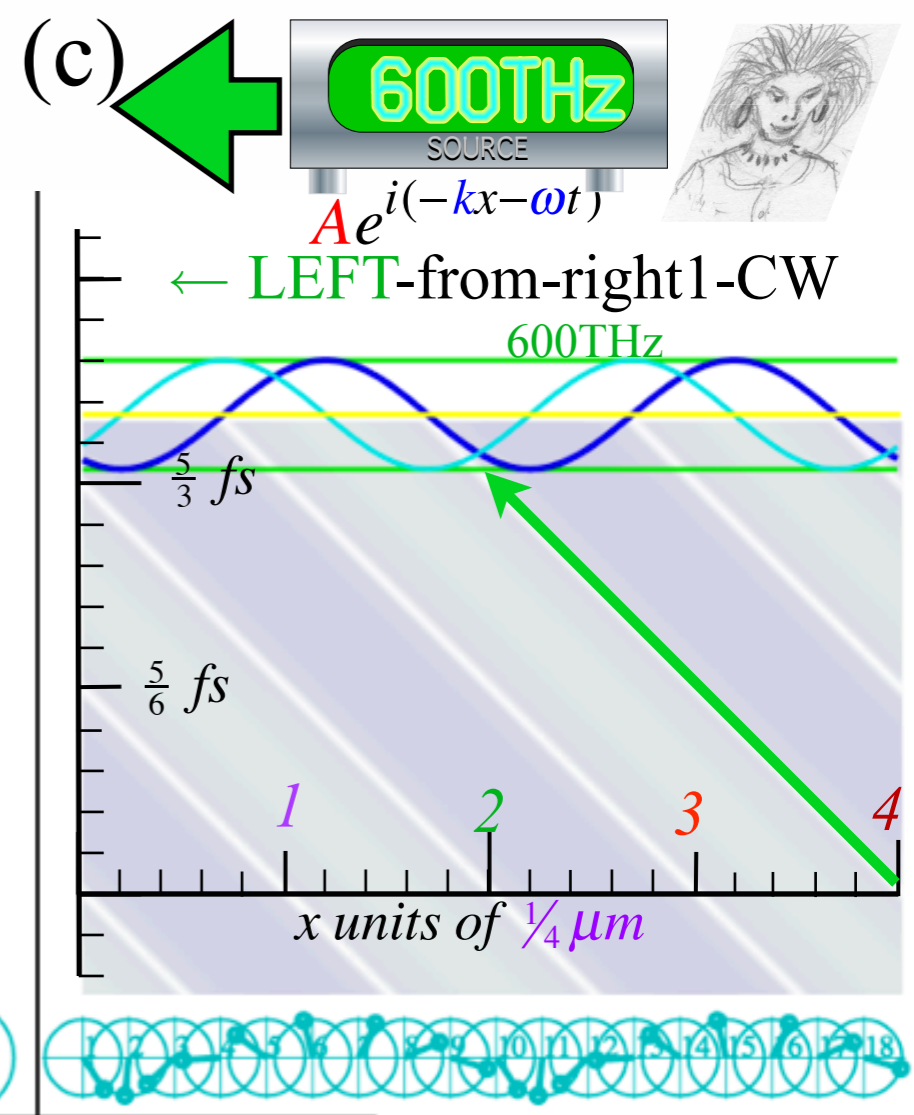
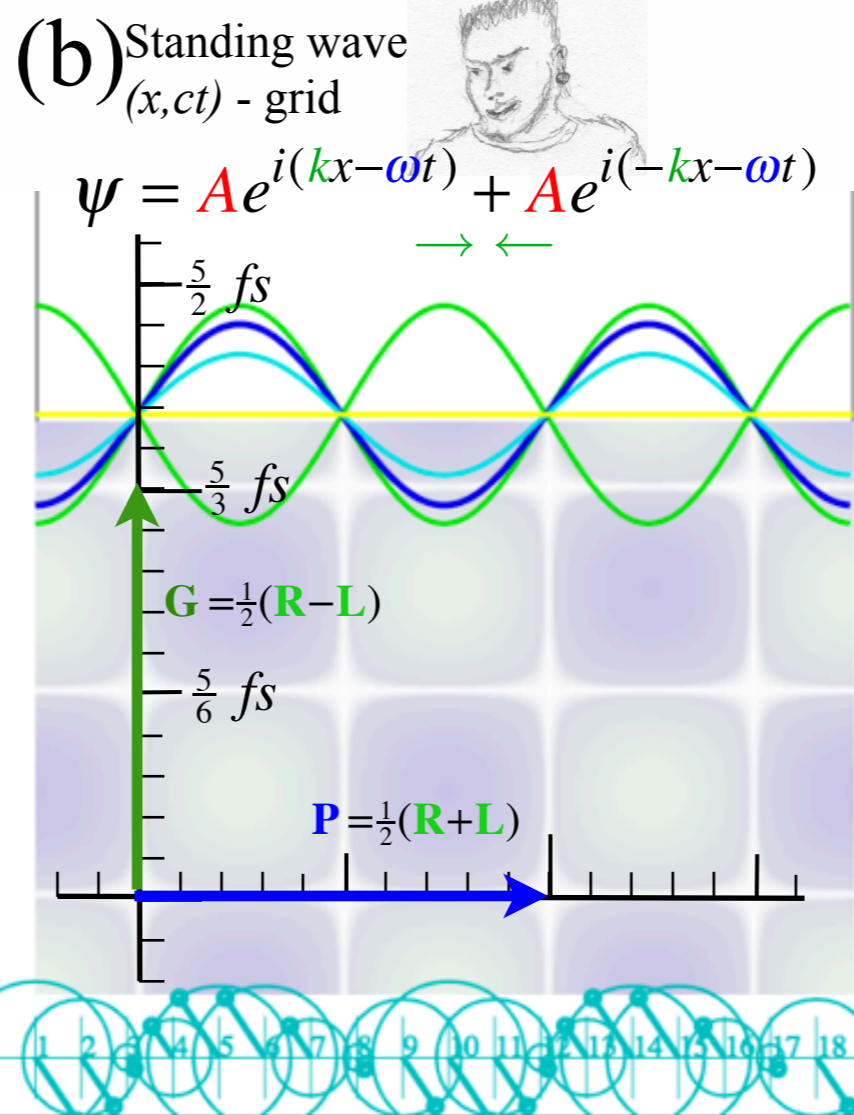
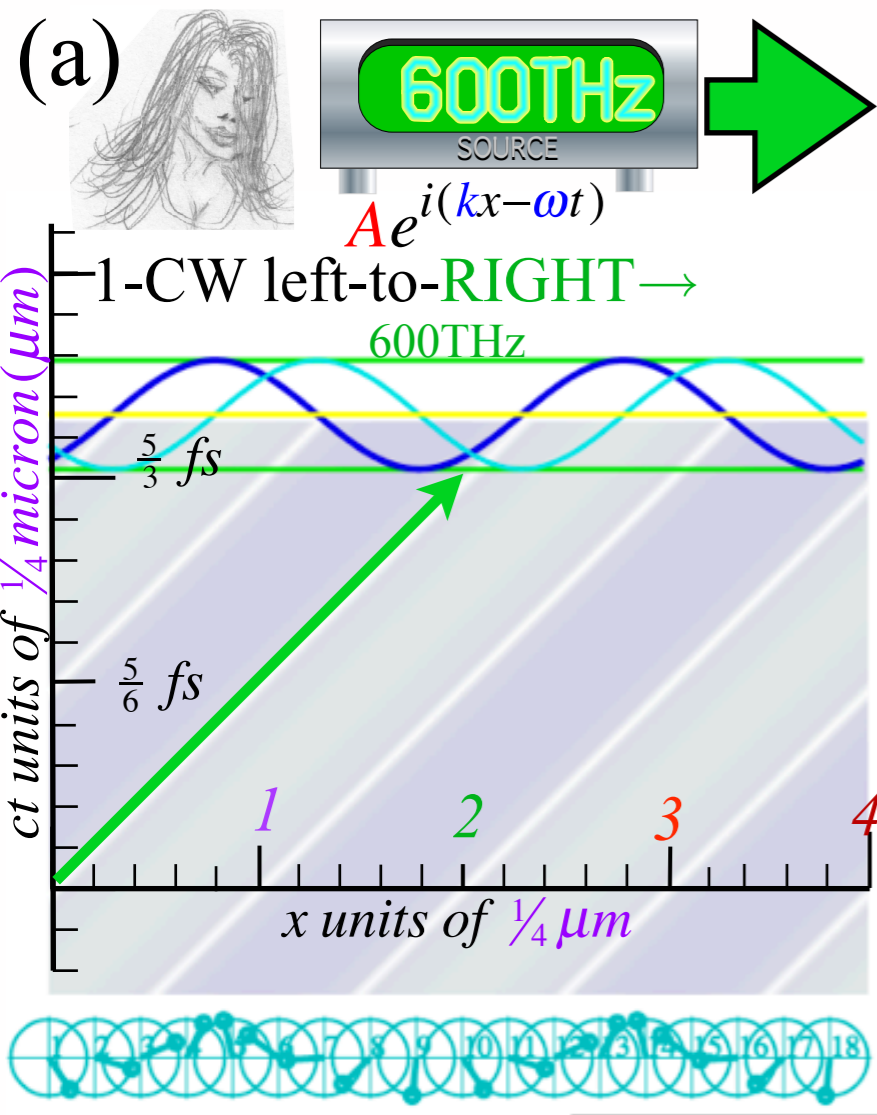
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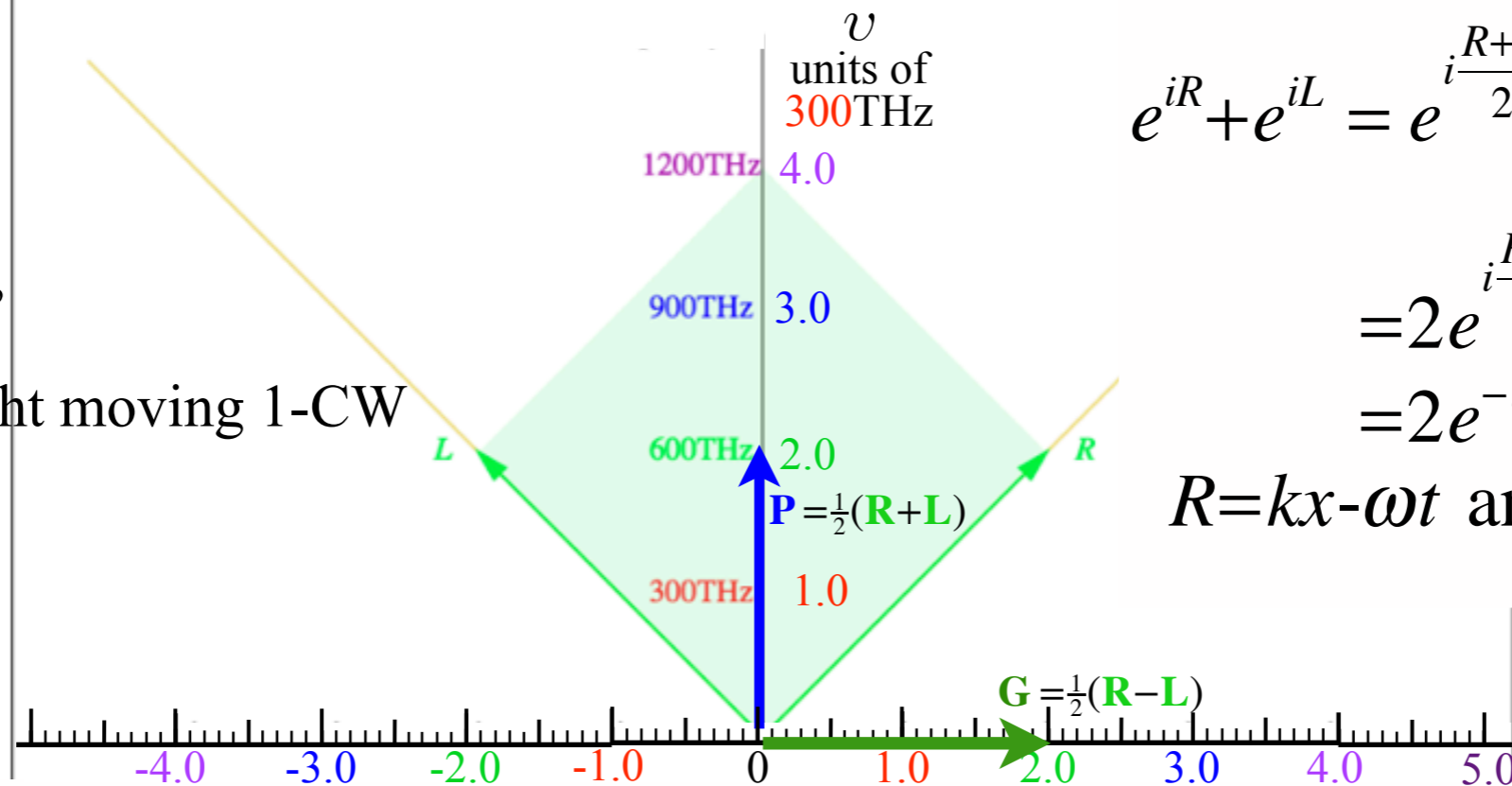
Animation of mechanics and metrology of constant- g grid





(d) Introducing optical space-time grids and per-space-time “baseball-diamonds”

\leftarrow LEFT-from-right moving 1-CW



$$e^{iR} + e^{iL} = e^{i\frac{R+L}{2}} \left(e^{i\frac{R-L}{2}} + e^{-i\frac{R-L}{2}} \right)$$

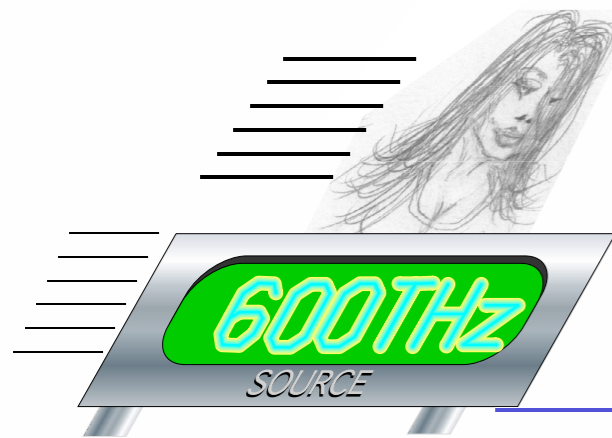

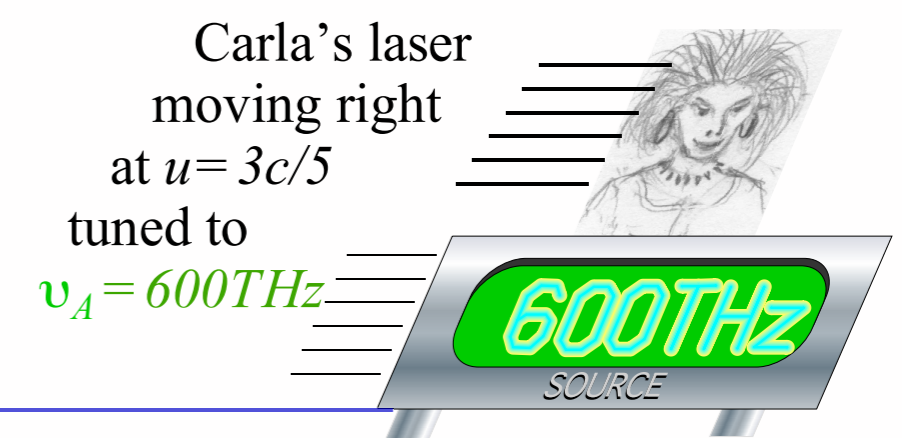
$$= 2e^{i\frac{R+L}{2}} \cos \frac{R-L}{2}$$



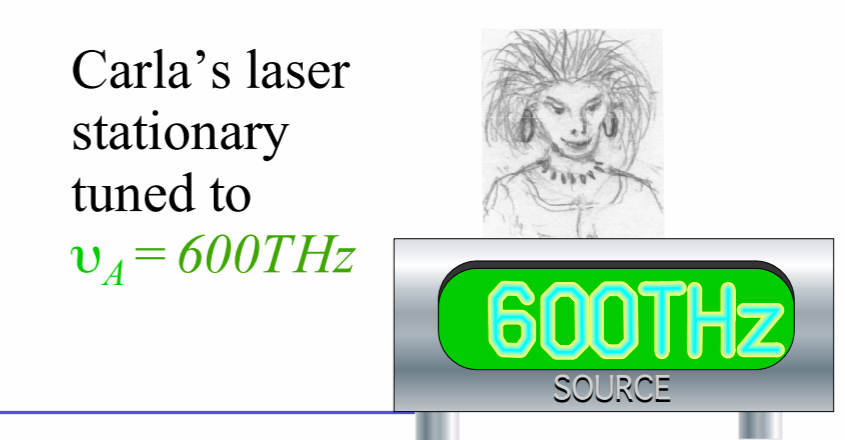
$$= 2e^{-i\omega t} \cos kx$$

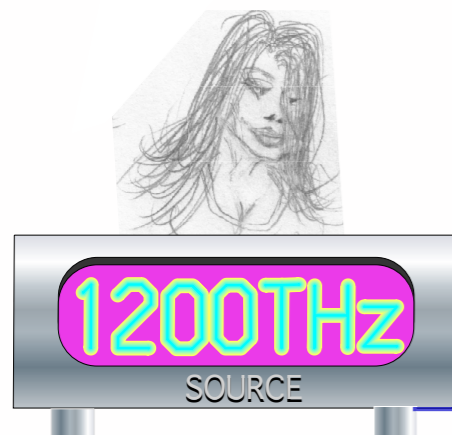

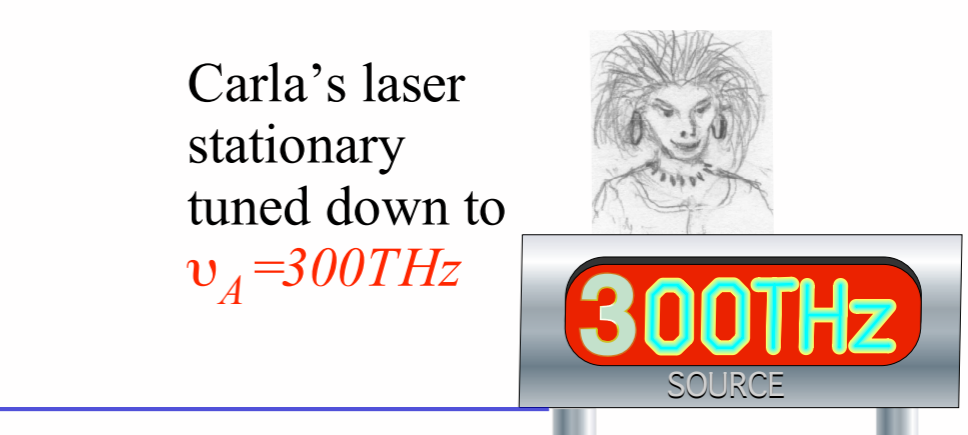
$$R = kx - \omega t \text{ and: } L = -kx - \omega t$$

ck units of 300THz

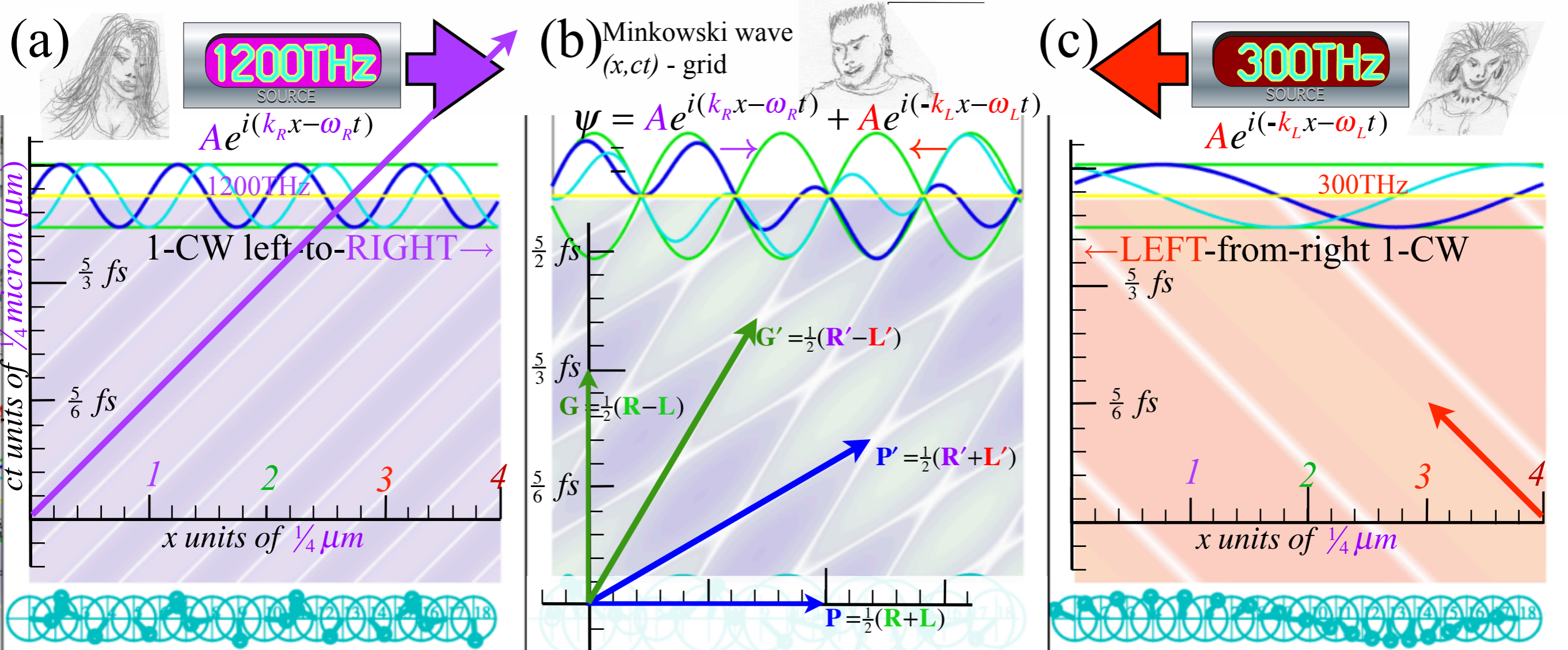
Three scenarios that look the same to Bob

	<p>Alice's laser moving right at $u=3c/5$ tuned to $\nu_A=600\text{THz}$</p>	<p>Bob stationary</p> 	<p>Carla's laser moving right at $u=3c/5$ tuned to $\nu_A=600\text{THz}$</p> 
----------------------------------------------------------------------------------	----------------------------------------------------------------------------------------------------	-----------------------------------------------------------------------------------------------------------	----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------

	<p>Alice's laser stationary tuned to $\nu_A=600\text{THz}$</p>	<p>Bob moving left at $u=-3c/5$</p> 	<p>Carla's laser stationary tuned to $\nu_A=600\text{THz}$</p> 
------------------------------------------------------------------------------------	---------------------------------------------------------------------------	-------------------------------------------------------------------------------------------------------------------------------------	----------------------------------------------------------------------------------------------------------------------------------------------------------------

	<p>Alice's laser stationary tuned up to $\nu_A=1200\text{THz}$</p>	<p>Bob stationary</p> 	<p>Carla's laser stationary tuned down to $\nu_A=300\text{THz}$</p> 
-------------------------------------------------------------------------------------	-------------------------------------------------------------------------------	-------------------------------------------------------------------------------------------------------------	----------------------------------------------------------------------------------------------------------------------------------------------------------------------

Much cheaper to do this one!\$!



(d)

$$e^{iR'} + e^{iL'} = e^{i\frac{R'+L'}{2}} (e^{i\frac{R'-L'}{2}} + e^{-i\frac{R'-L'}{2}})$$

$$= e^{i\frac{R'+L'}{2}} 2 \cos \frac{R'-L'}{2}$$

$$= \psi'_{\text{phase}} \psi'_{\text{group}}$$

$$R' = k_R x - \omega_R t \text{ and: } L' = -k_L x - \omega_L t$$

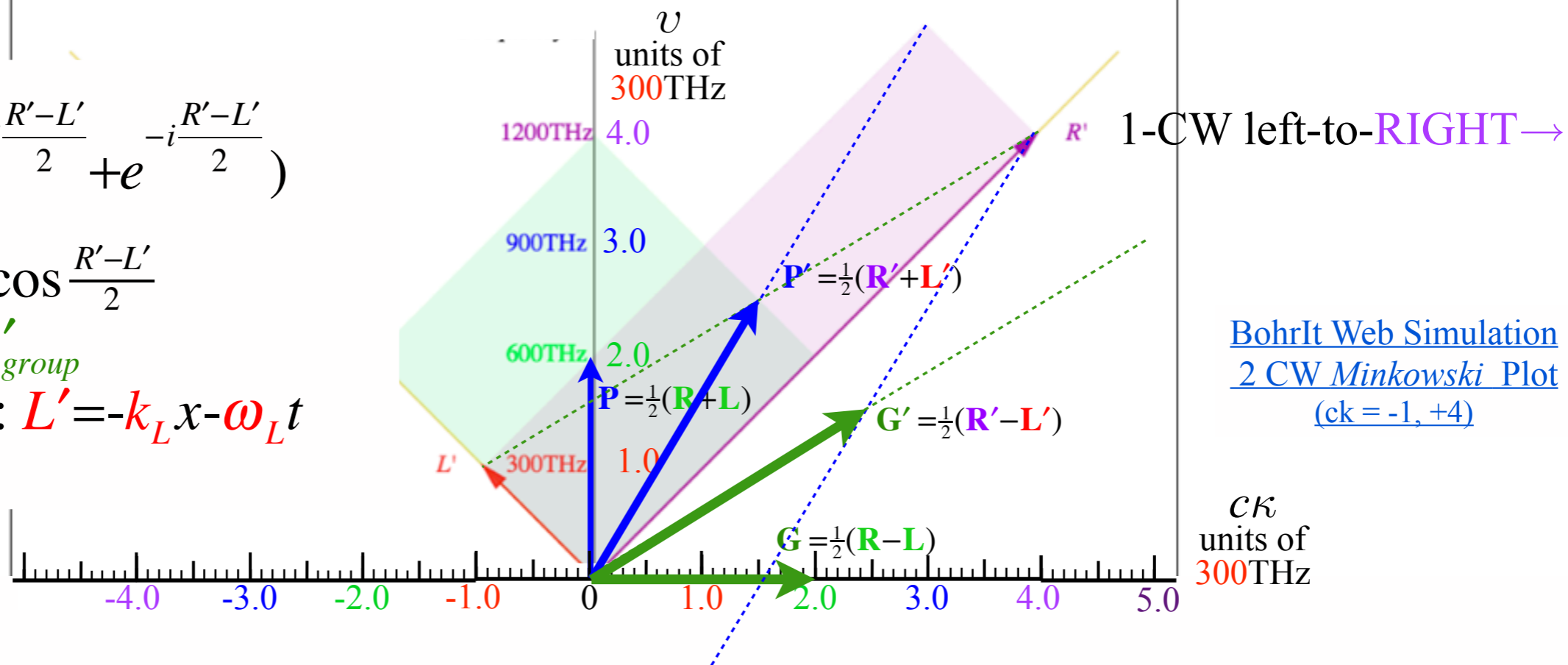


Fig. 10 in text
 Relativity...

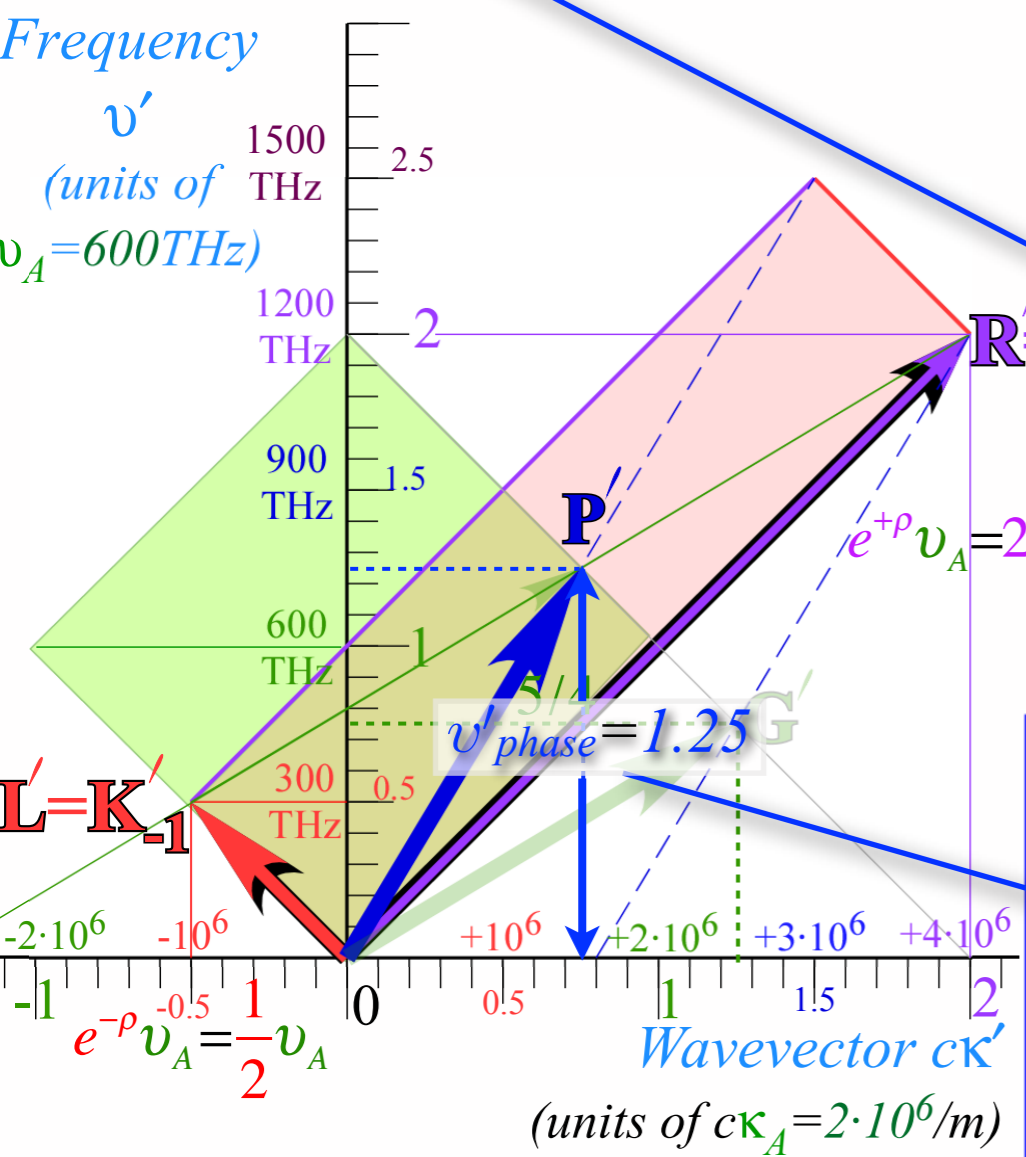
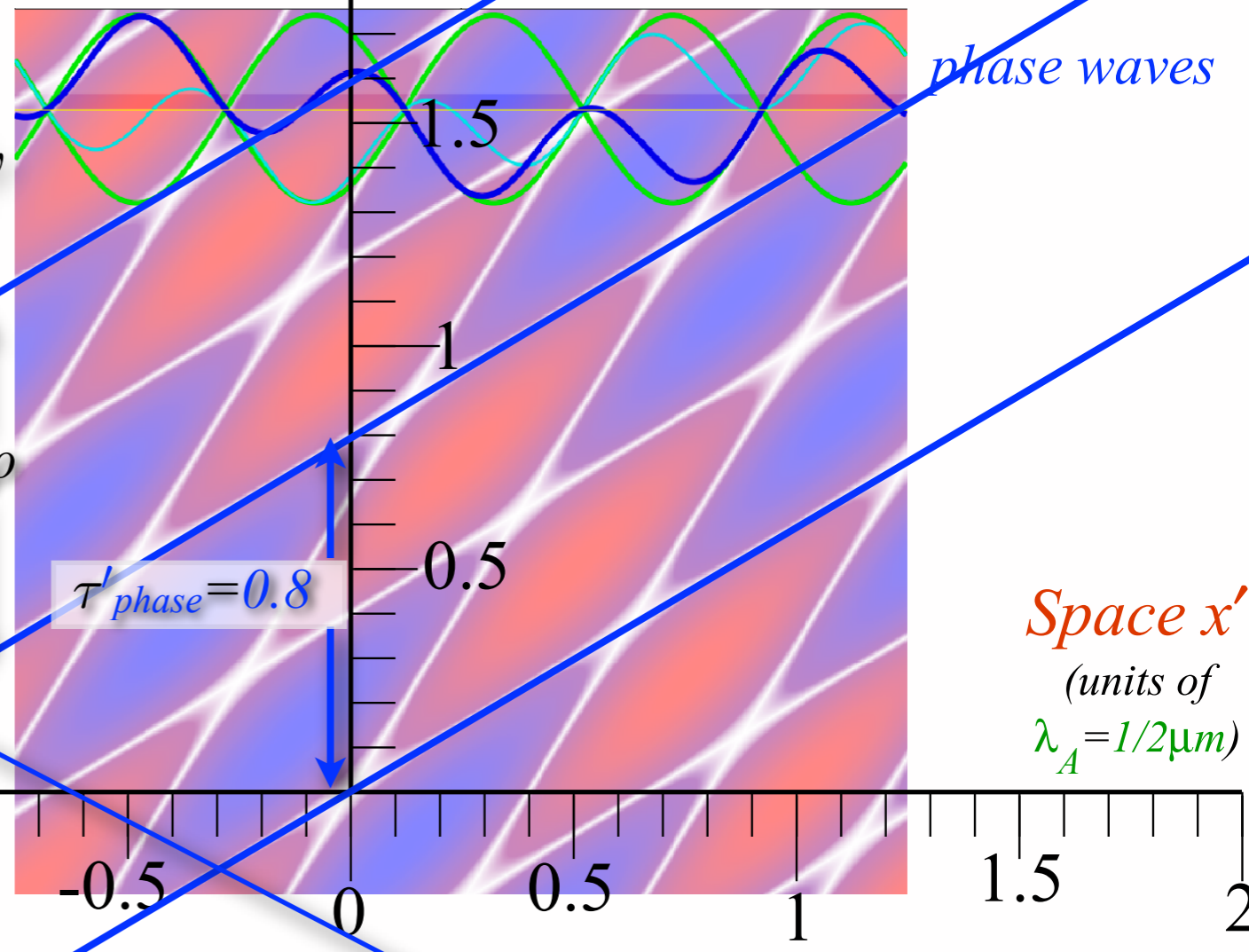
The 16 dimensions of 2CW interference

Time ct'
(units of $\lambda_A = 1/2\mu\text{m}$)

Start with the *Dopplers*
...then do the *phase waves*

$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

Phase frequency $v'_{phase} = v_A \cosh \rho = 5/4 = 1.25$
 flips to Phase period $\tau'_{phase} = \tau_A \text{sech} \rho = 4/5 = 0.8$



phase	$b_{\text{Doppler RED}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{\kappa_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{V_{\text{phase}}}{c}$	$b_{\text{Doppler BLUE}}$
group	1	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{\kappa_{\text{group}}}{\kappa_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{c}{V_{\text{group}}}$	1
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech} \rho$	$\cosh \rho$	$\text{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

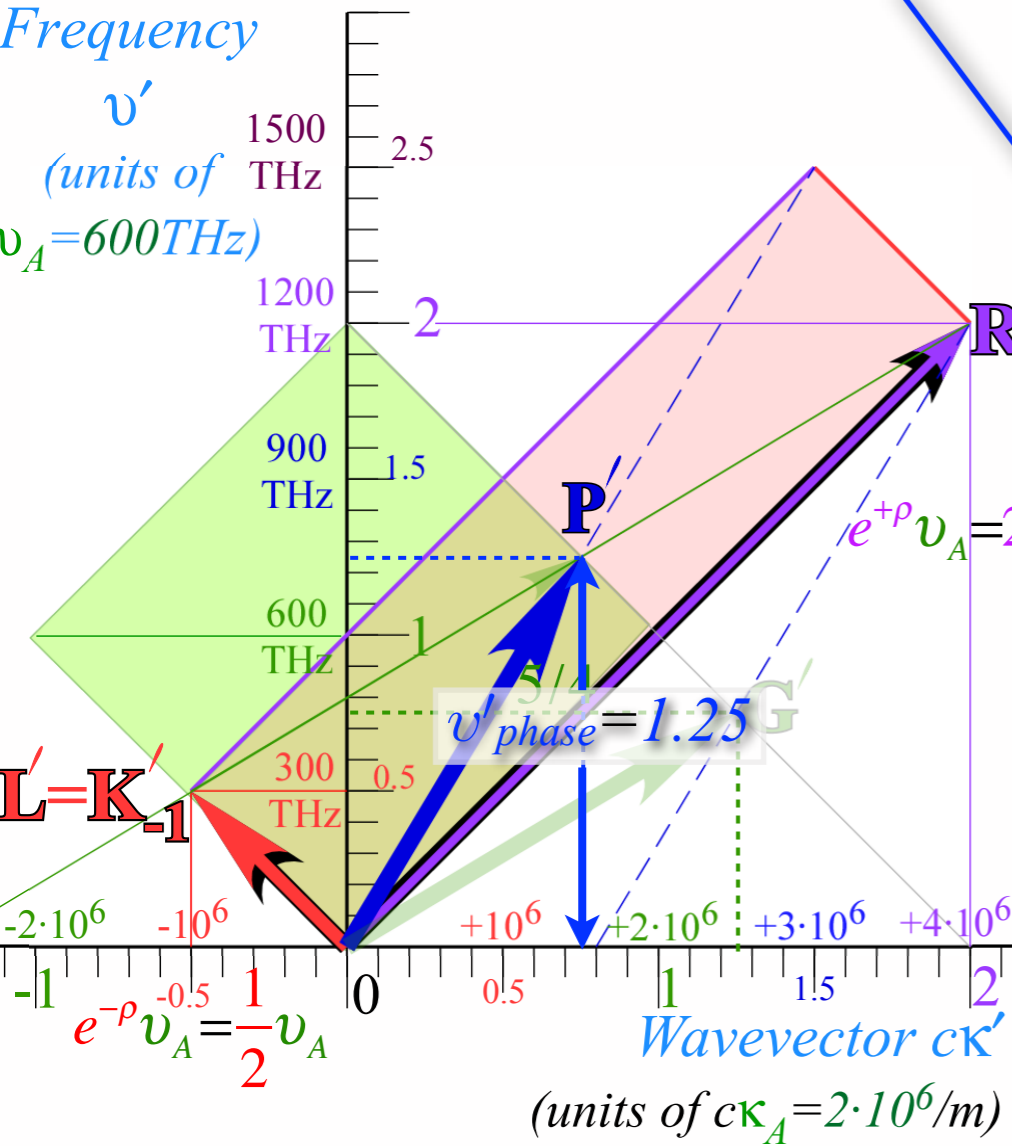
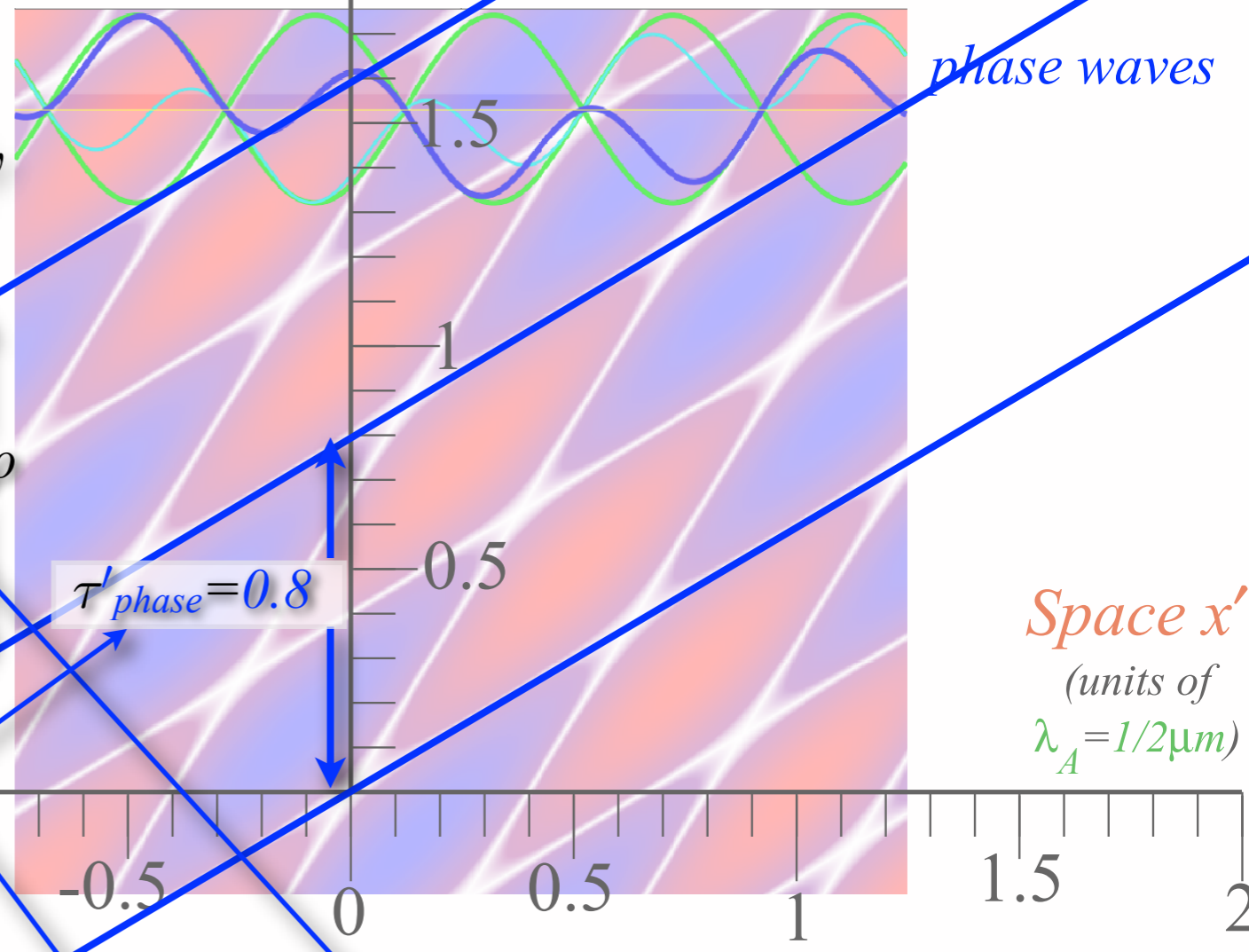
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Time ct'
 (units of $\lambda_A = 1/2 \mu m$)

Start with the Dopplers
 ...then do the phase waves



phase	$b_{Doppler RED}$	$\frac{v_{phase}}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
group	$\frac{1}{b_{Doppler BLUE}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

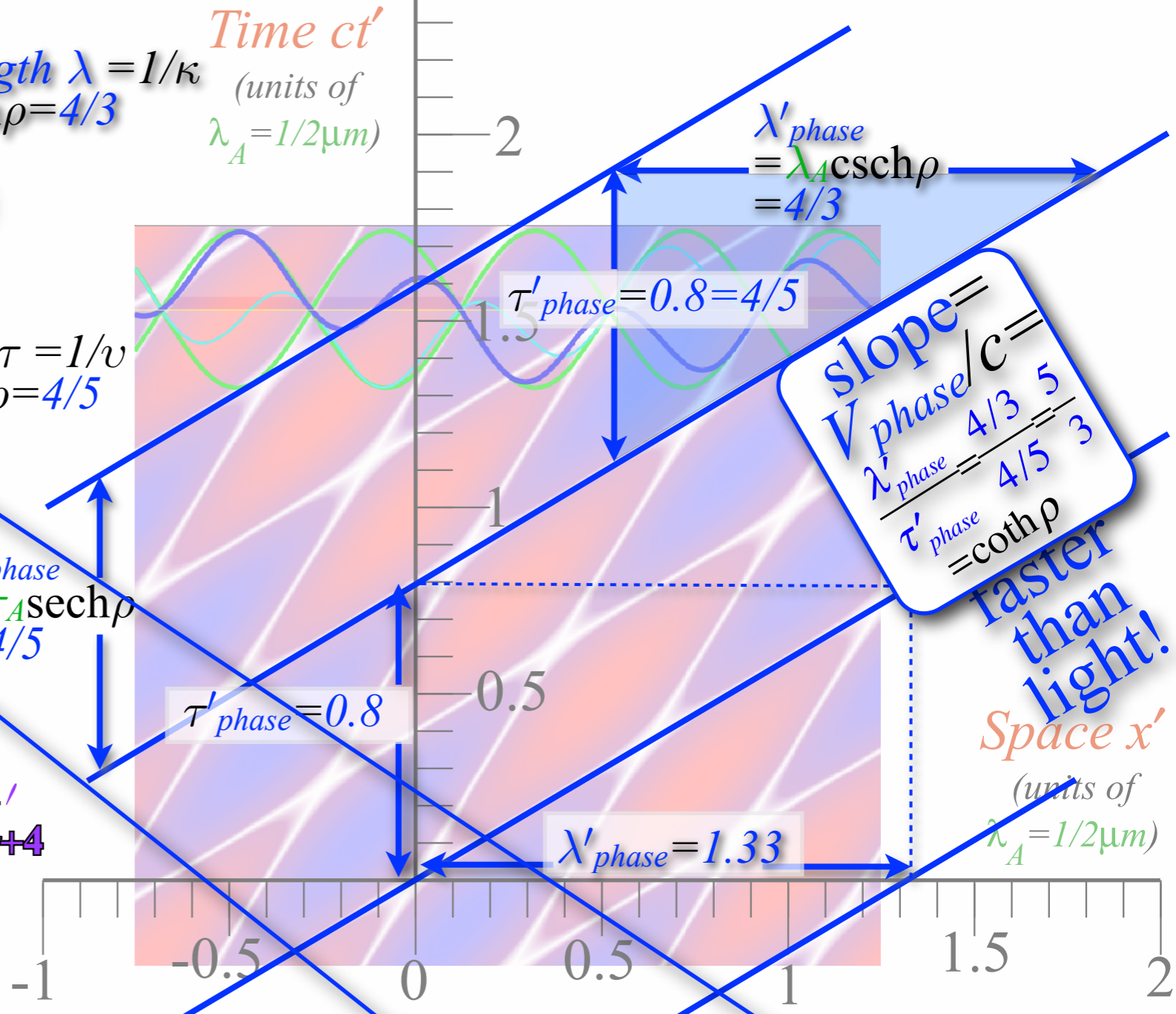
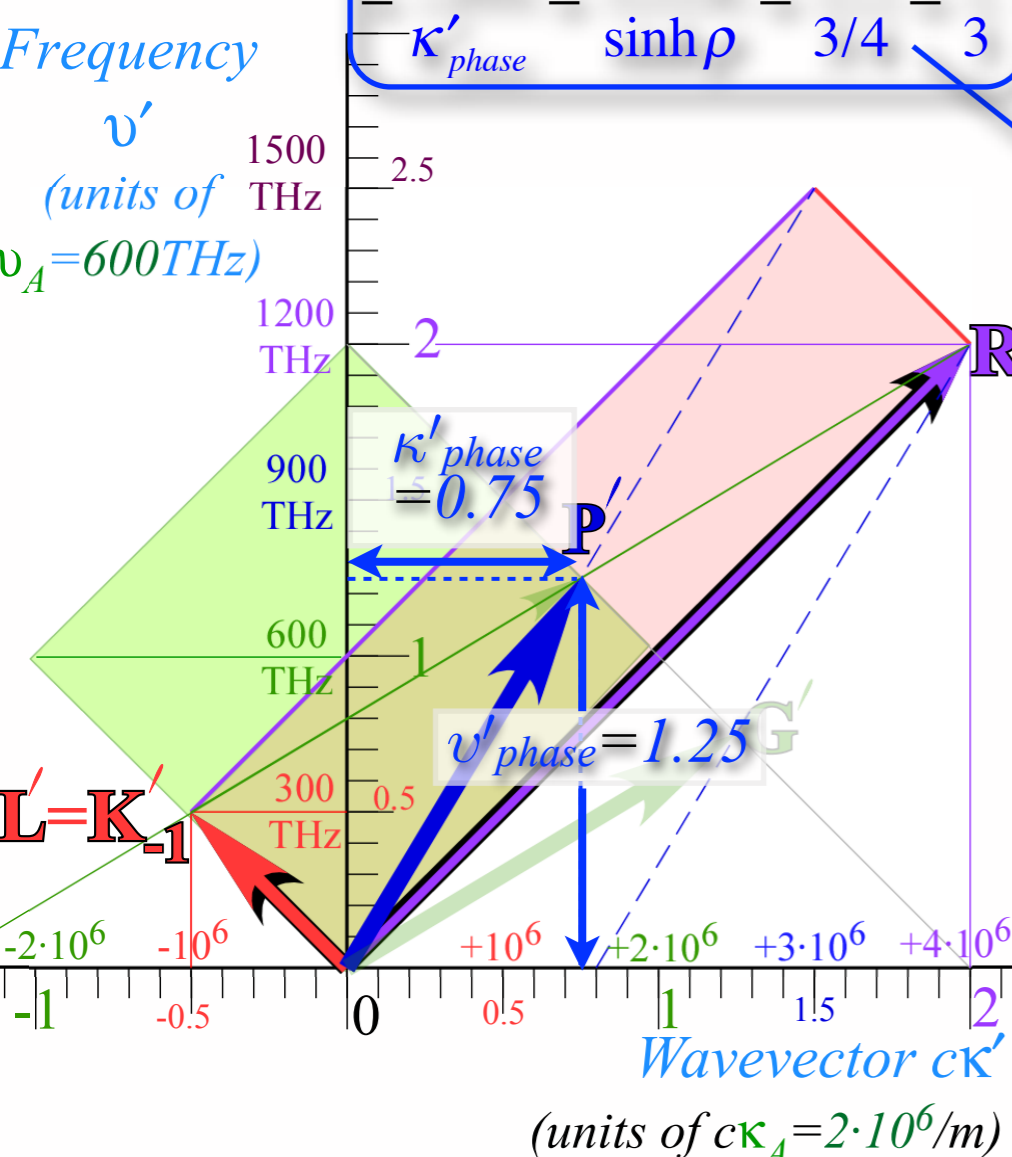
Phase wavenumber $\kappa'_{phase} = \kappa_A \sinh \rho = 3/4$ flips to Phase wavelength $\lambda'_{phase} = \lambda_A \text{csch } \rho = 4/3$ (units of $\lambda_A = 1/2 \mu\text{m}$)

$$\mathbf{P}' = \begin{pmatrix} c\kappa'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

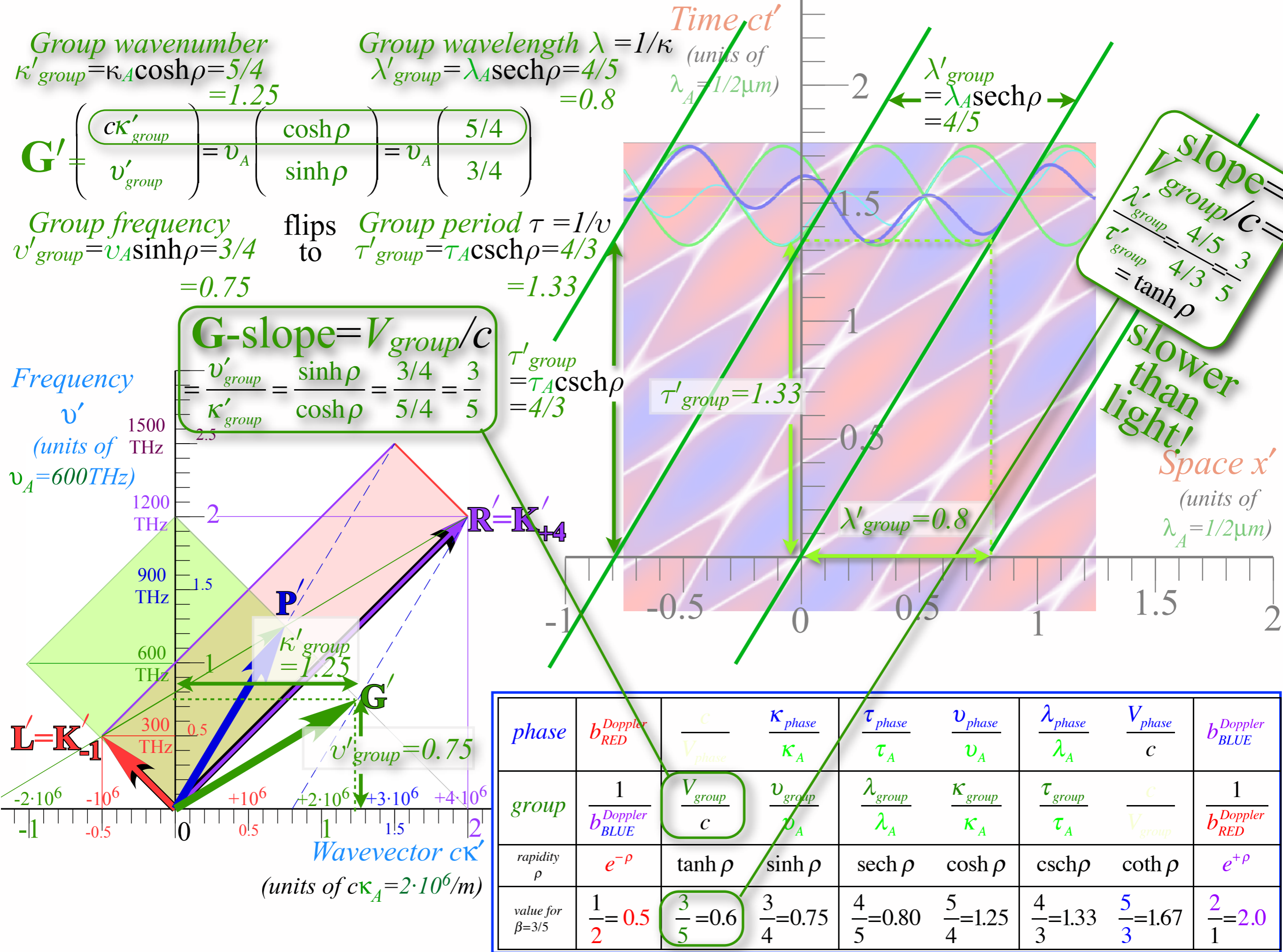
Phase frequency $v'_{phase} = v_A \cosh \rho = 5/4$ flips to Phase period $\tau'_{phase} = \tau_A \text{sech } \rho = 4/5$

P-slope = V_{phase}/c

$$= \frac{v'_{phase}}{\kappa'_{phase}} = \frac{\cosh \rho}{\sinh \rho} = \frac{5/4}{3/4} = \frac{5}{3}$$



phase	$b_{\text{Doppler RED}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{\text{Doppler BLUE}}$
group	$\frac{1}{b_{\text{Doppler BLUE}}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{\text{Doppler RED}}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
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Lorentz transformations...

write \mathbf{G}' and \mathbf{P}' in terms of \mathbf{G} and \mathbf{P} using $\cosh \rho$ and $\sinh \rho$

$$\mathbf{G}' = \begin{pmatrix} c\mathbf{K}'_{group} \\ \mathbf{v}'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

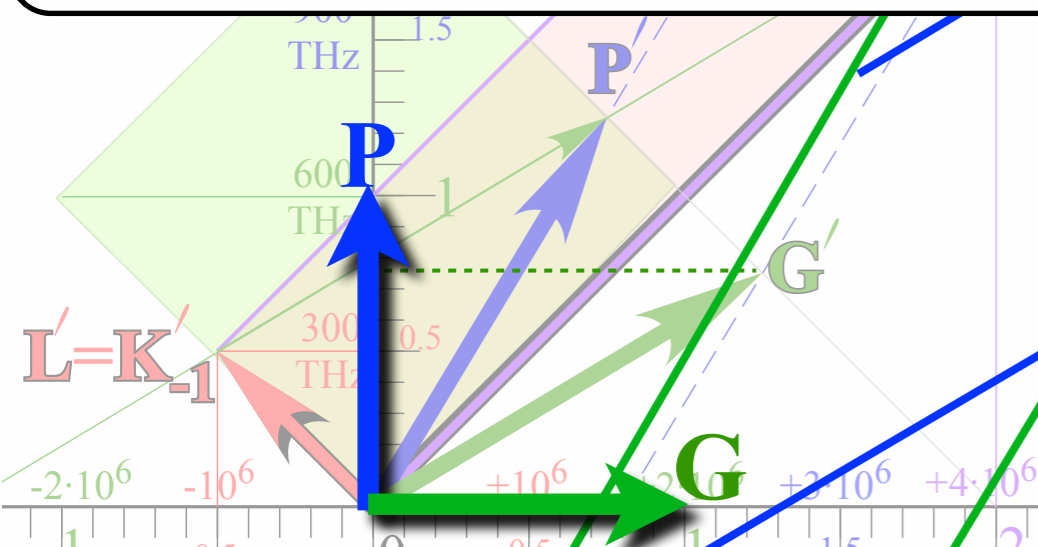
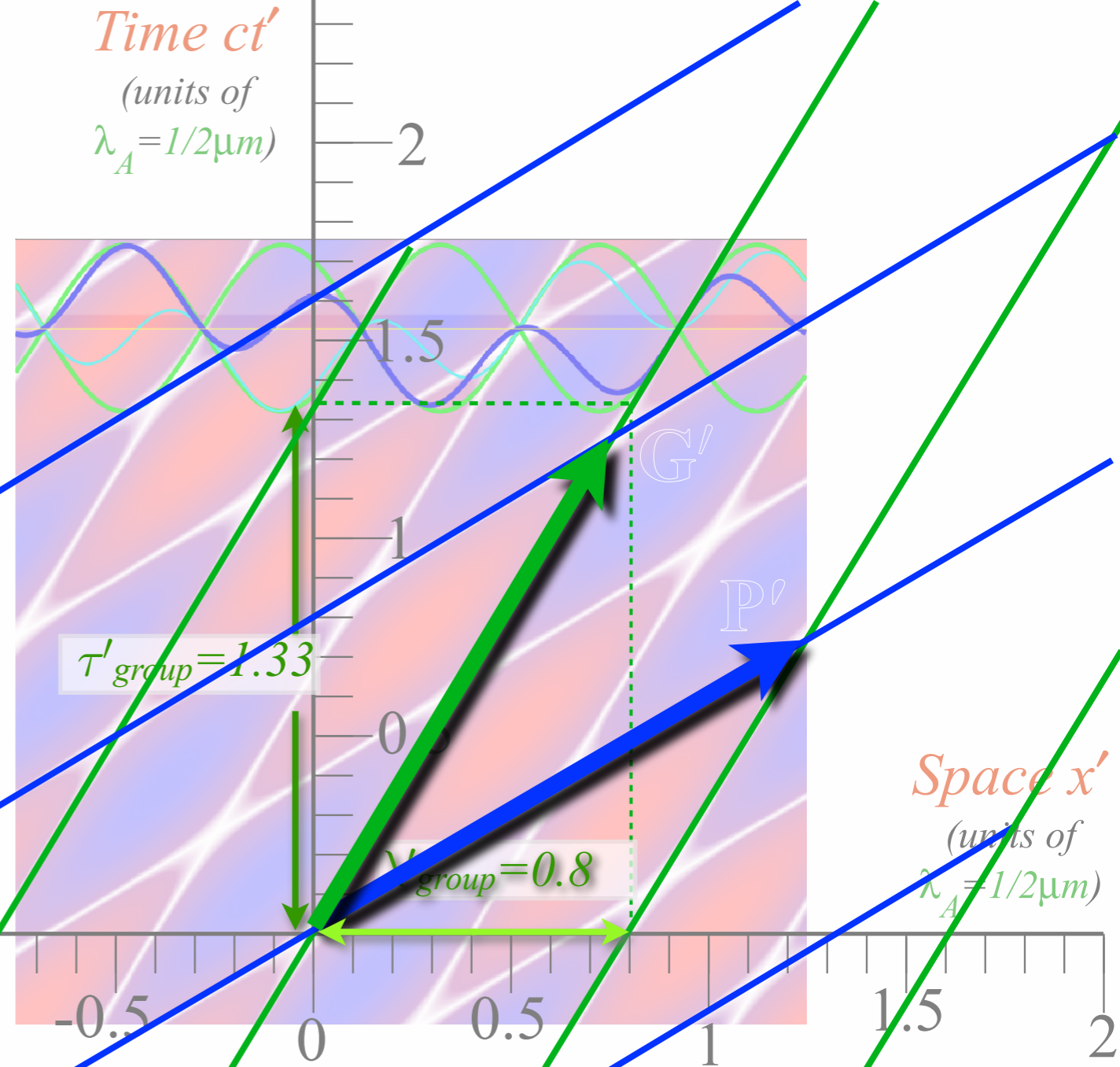
$$= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \cosh \rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \sinh \rho$$

$$\mathbf{G}' = \mathbf{G} \cosh \rho + \mathbf{P} \sinh \rho$$

$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ \mathbf{v}'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

$$= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \sinh \rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \cosh \rho$$

$$\mathbf{P}' = \mathbf{G} \sinh \rho + \mathbf{P} \cosh \rho$$



RelaWavity Web Simulation - 16 Relativity Dimensions

$$\begin{pmatrix} \cosh \rho & \sinh \rho \\ \sinh \rho & \cosh \rho \end{pmatrix} \text{ Lorentz transform matrix}$$

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\mathbf{K}_{phase}}{\mathbf{K}_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{\mathbf{v}_{phase}}{\mathbf{v}_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{\mathbf{v}_{group}}{\mathbf{v}_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\mathbf{K}_{group}}{\mathbf{K}_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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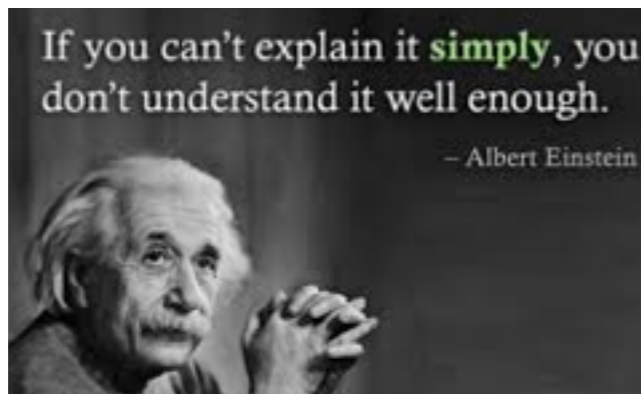
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Two Famous-Name Coefficients

Review of Lect. 30 p.106

Albert Einstein
1859-1955



This number is called an: **Einstein time-dilation** (dilated by 25% here)

This number is called a: **Lorentz length-contraction** (contracted by 20% here)



Hendrik A. Lorentz
1853-1928

Old-Fashioned Notation

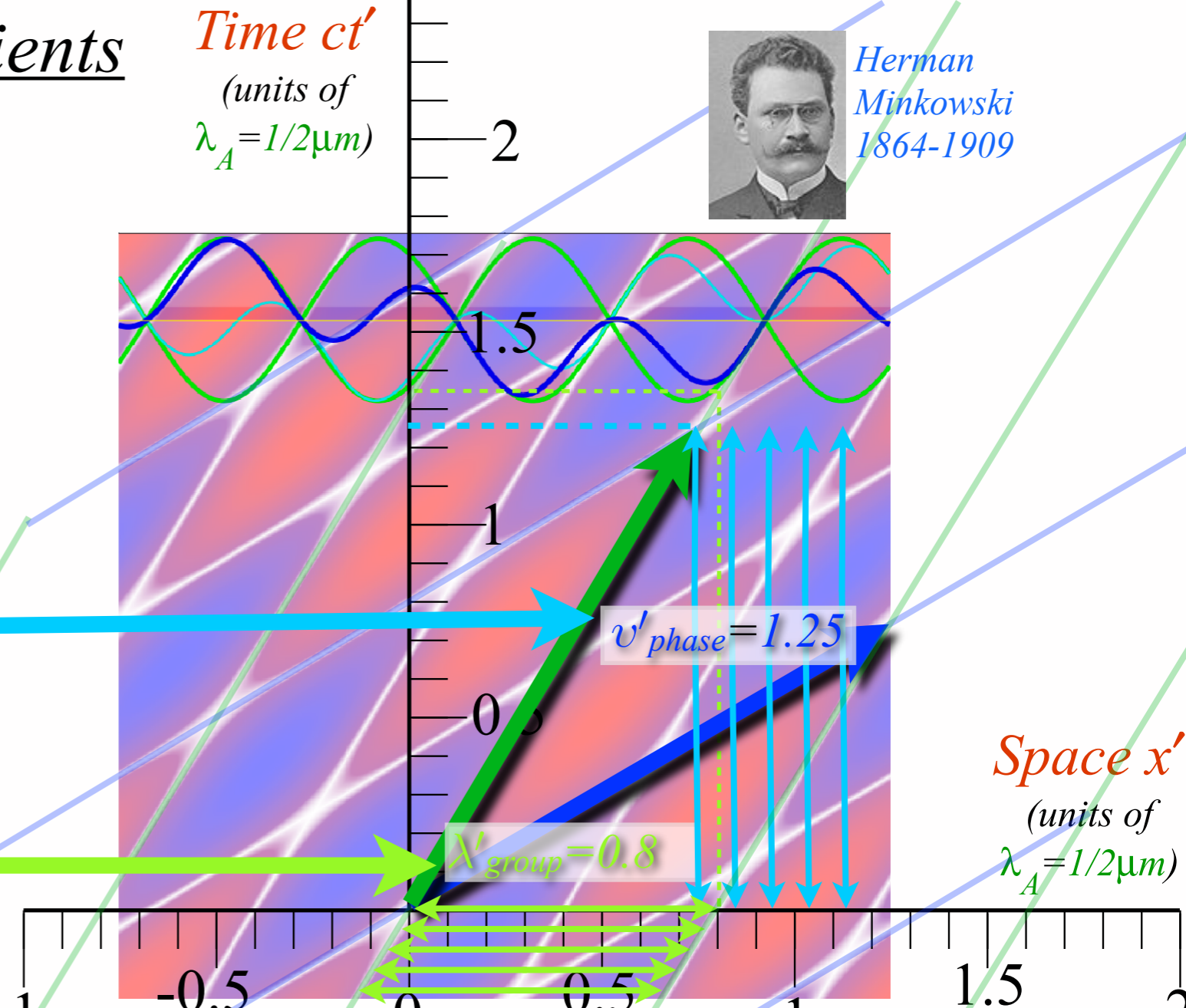
[RelaWavity Web Simulation - Relativistic Terms](#)
(Expanded Table)

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\operatorname{cosh} \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Time ct'
(units of $\lambda_A = 1/2 \mu m$)



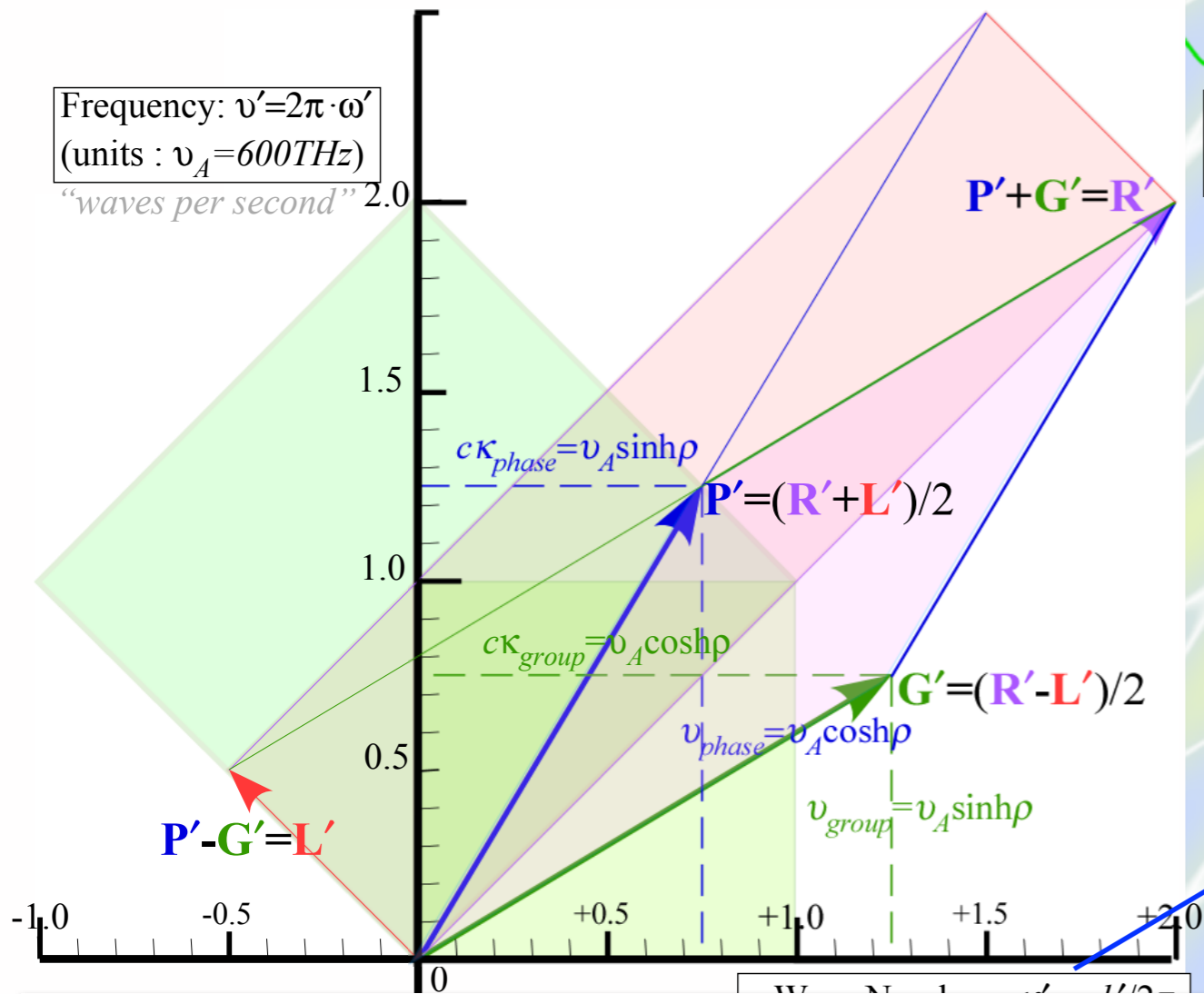
Herman Minkowski
1864-1909



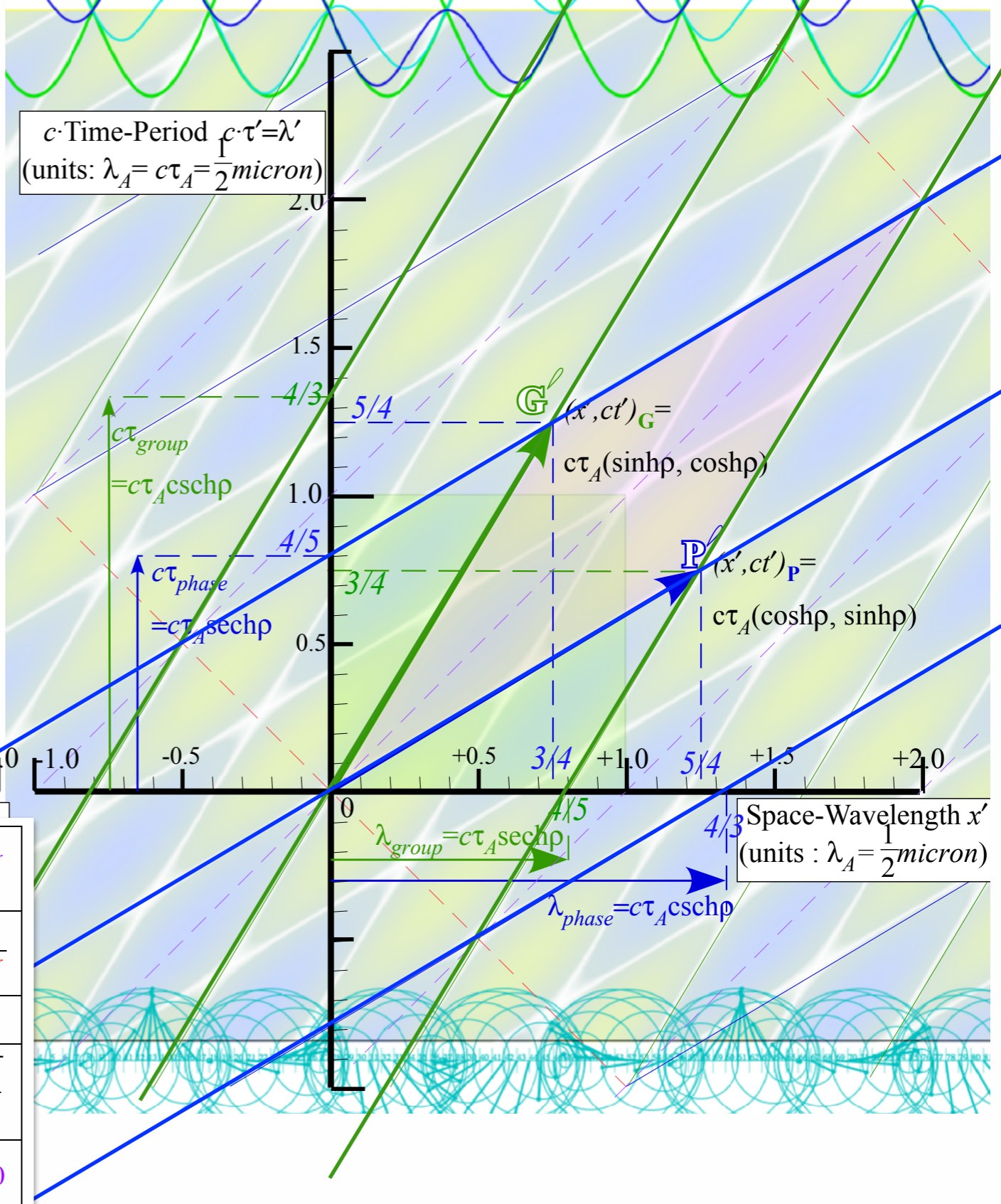
Space x'
(units of $\lambda_A = 1/2 \mu m$)

Fig. 11 in text Relativity...

(a) Per-space-time $(v', c\kappa')$ geometry of 2-CW vectors



(b) Space-time $(c\tau', x')$ geometry of 2-CW paths



phase	$b_{Doppler RED}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
group	$\frac{1}{b_{Doppler BLUE}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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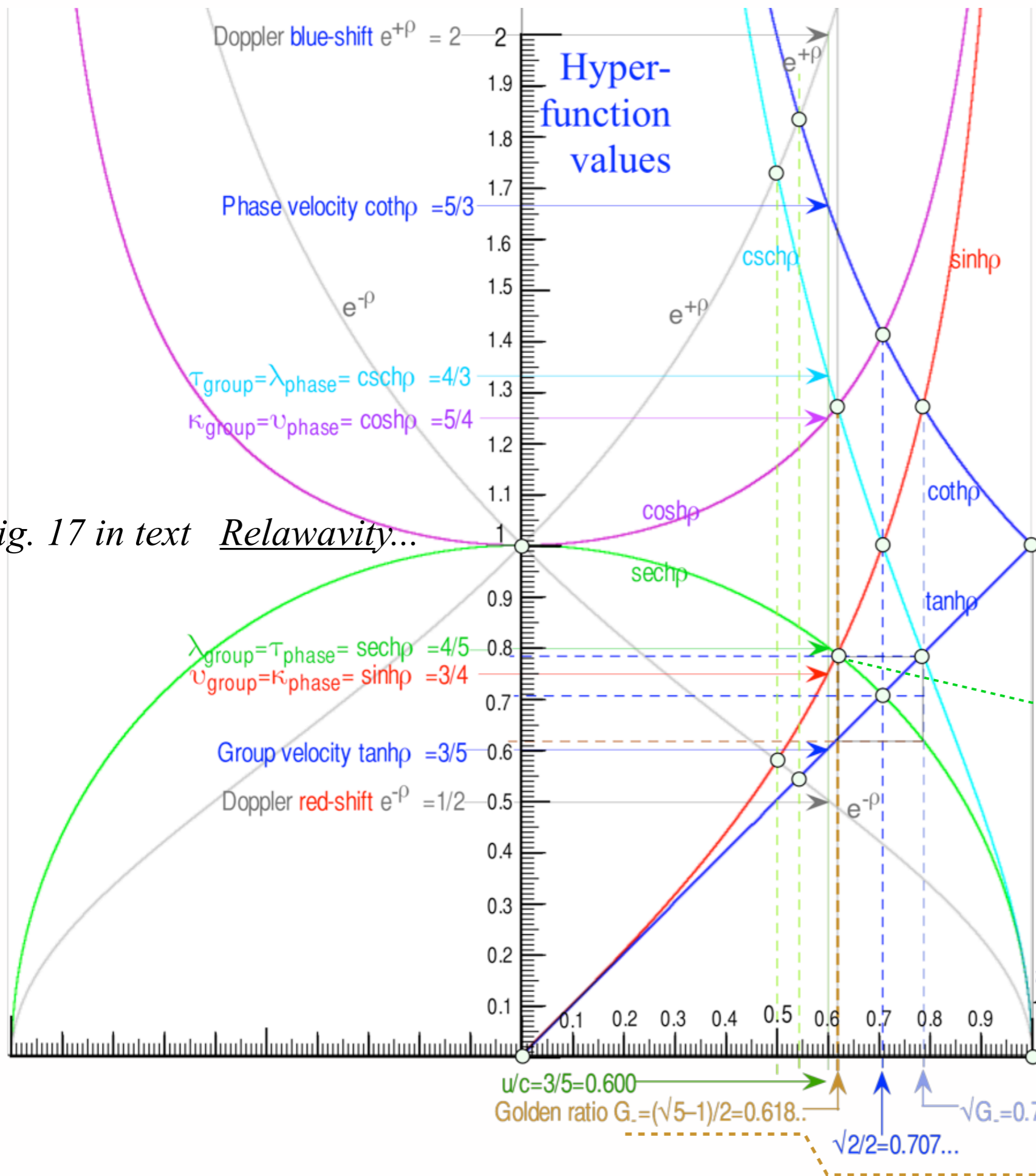
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If $\frac{u}{c} = \tanh \rho = 0.618..$ (Golden-Mean G_-)

two parameters become exactly equal :

$$\frac{ct'_P}{c\tau_A} = \sinh \rho = \frac{\lambda_{\text{group}}}{\lambda_A} = \frac{\tau_{\text{phase}}}{\tau_A} = \text{sech } \rho$$

$$= 0.786.. = \sqrt{G_-} = 0.786..$$

and

$$\frac{x'_P}{\lambda_A} = \cosh \rho = \frac{\lambda_{\text{phase}}}{\lambda_A} = \frac{\tau_{\text{group}}}{\tau_A} = \text{csch } \rho$$

$$= 1.272.. = 1/\sqrt{G_-} = 1.272..$$

Fig. 17 in text Relativity...

Solve :

$$\text{sech } \rho = \sinh \rho$$

or:

$$\sinh \rho \cosh \rho = 1$$

or:

$$\sinh 2\rho = 2$$

$$\rho = \frac{1}{2} \sinh^{-1} 2 = 0.7218..$$

$$\tanh \rho = 0.618.. = \frac{\sqrt{5}-1}{2}$$

Lecture 31

Thur. 12.08.2016

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Doppler Jeopardy

$$\omega_R = 2\pi \nu_R \quad \nu_R = 600 \text{THz}$$



$$\nu_L = 300 \text{THz}$$

$$\omega_L = 2\pi \nu_L$$

- (1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ν_E ?
- (2.) What is that frequency ν_E ?

Doppler Jeopardy

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- (2.) What is that frequency ν_E ?

Query (1.) has a Jeopardy-style answer-by-question: What is beam group velocity?

$$u_E = V_{group} = \frac{\nu_{group}}{\kappa_{group}} = \frac{\nu_R - \nu_L}{\kappa_R - \kappa_L} = c \frac{\nu_R - \nu_L}{\nu_R + \nu_L}$$

Doppler Jeopardy

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Query (2.) similarly: What ν_E is blue-shift $b\nu_L$ of ν_L and red-shift ν_R/b of ν_R ?

$$\nu_E = b\nu_L = \nu_R/b \quad \Rightarrow \quad b = \sqrt{\nu_R/\nu_L} \quad \Rightarrow \quad \nu_E = \sqrt{\nu_R \cdot \nu_L}$$

Doppler Jeopardy

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$$\nu_E = b\nu_L = \nu_R/b \quad \Rightarrow \quad b = \sqrt{\nu_R/\nu_L} \quad \Rightarrow \quad \nu_E = \sqrt{\nu_R \cdot \nu_L}$$

Geometric mean

Doppler Jeopardy

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$$\nu_L = 300 \text{ THz} \quad \omega_L = 2\pi \nu_L$$

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Query (2.) similarly: What ν_E is blue-shift $b\nu_L$ of ν_L and red-shift ν_R/b of ν_R ?

$$\nu_E = b\nu_L = \nu_R/b \quad \Rightarrow \quad b = \sqrt{\nu_R/\nu_L} \quad \Rightarrow \quad \nu_E = \sqrt{\nu_R \cdot \nu_L}$$

↑
Geometric mean

Doppler Jeopardy

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$$\nu_E = b\nu_L = \nu_R/b \quad \Rightarrow \quad b = \sqrt{\nu_R/\nu_L} \quad \Rightarrow \quad \nu_E = \sqrt{\nu_R \cdot \nu_L}$$

$$\nu_E = \sqrt{\nu_R \cdot \nu_L} = \sqrt{180000} = 424$$

Geometric mean

Doppler Jeopardy

$$\omega_R = 2\pi \nu_R \quad \nu_R = 600 \text{THz}$$



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Query (2.) similarly: What ν_E is blue-shift $b\nu_L$ of ν_L and red-shift ν_R/b of ν_R ?

$$\nu_E = b\nu_L = \nu_R/b \Rightarrow b = \sqrt{\nu_R/\nu_L} \Rightarrow \nu_E = \sqrt{\nu_R \cdot \nu_L} \quad \begin{aligned} \nu_E &= \sqrt{\nu_R \cdot \nu_L} \\ &= \sqrt{180000} \\ &= 424 \end{aligned}$$

V_{group}/c is ratio of difference mean $\nu_{group} = \frac{\nu_R - \nu_L}{2}$ to arithmetic mean $\nu_{phase} = \frac{\nu_R + \nu_L}{2}$. Frequency $\nu_E = B$ is the geometric mean $\sqrt{\nu_R \cdot \nu_L}$ of left and right-moving frequencies defining the geometry

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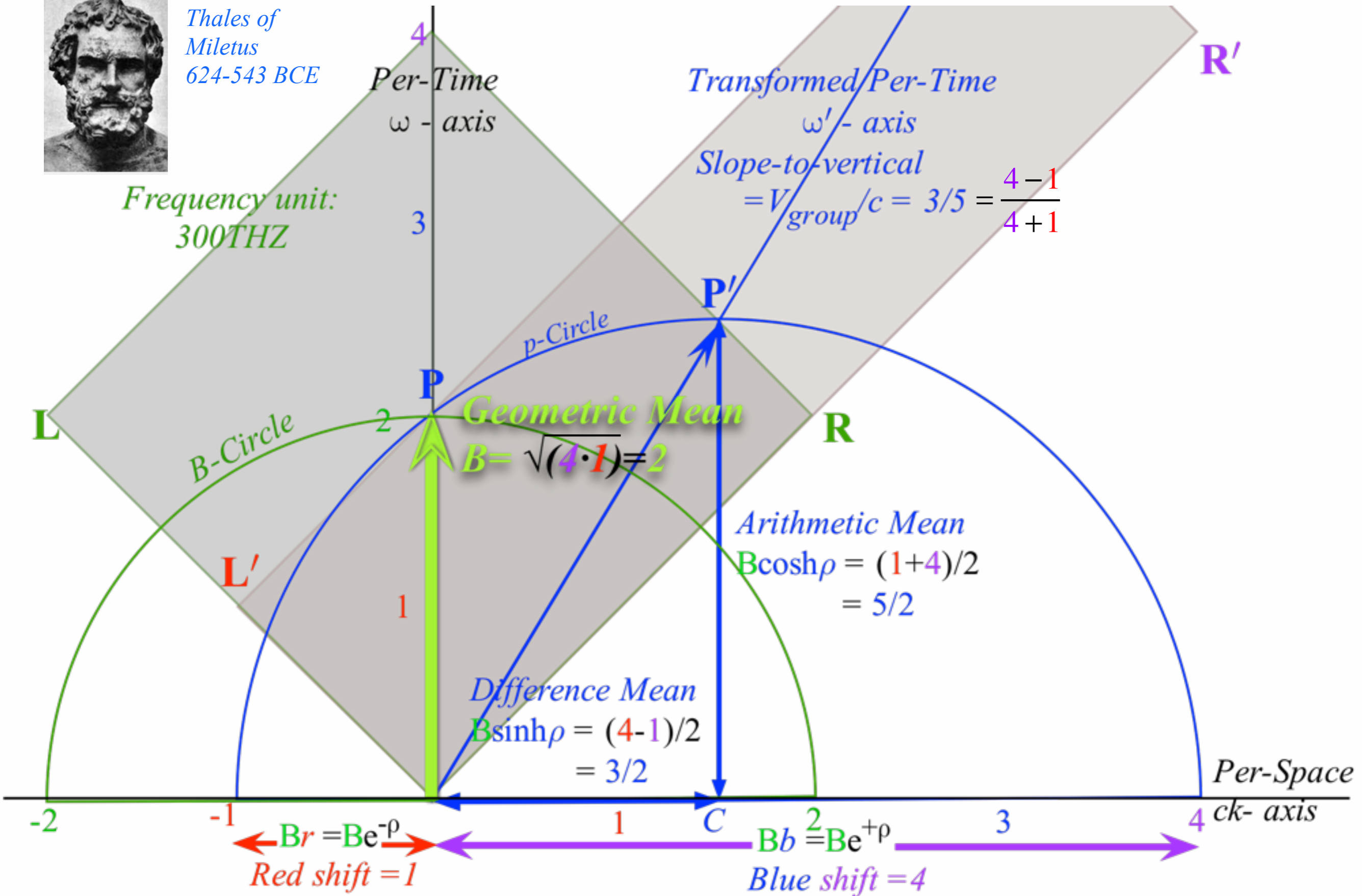
Thales Mean Geometry (600BCE)

helps “Relativity”



Thales of Miletus
624-543 BCE

Frequency unit:
300THZ

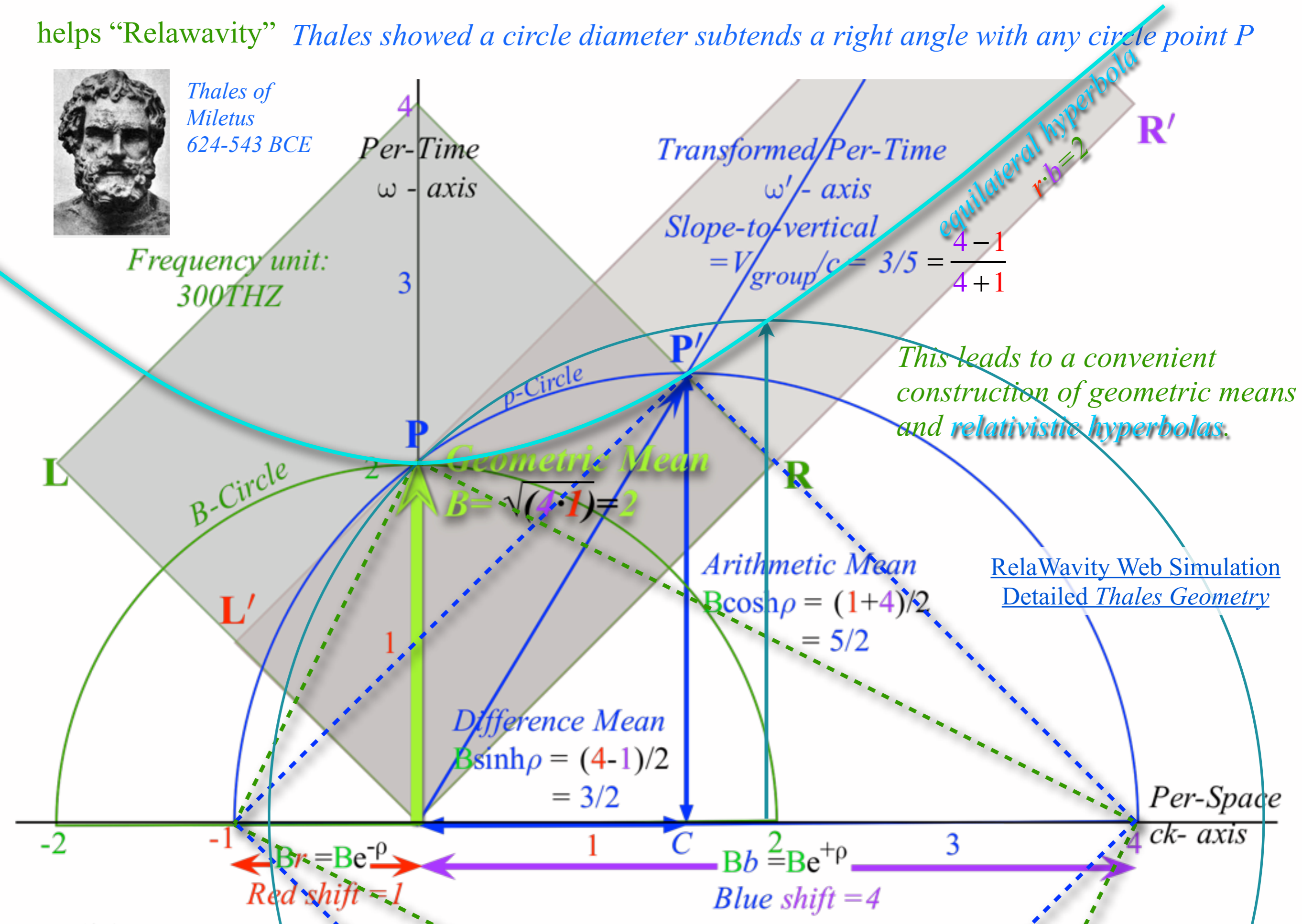


Thales Mean Geometry (600BCE)

helps “Relativity” *Thales showed a circle diameter subtends a right angle with any circle point P*



Thales of Miletus
624-543 BCE

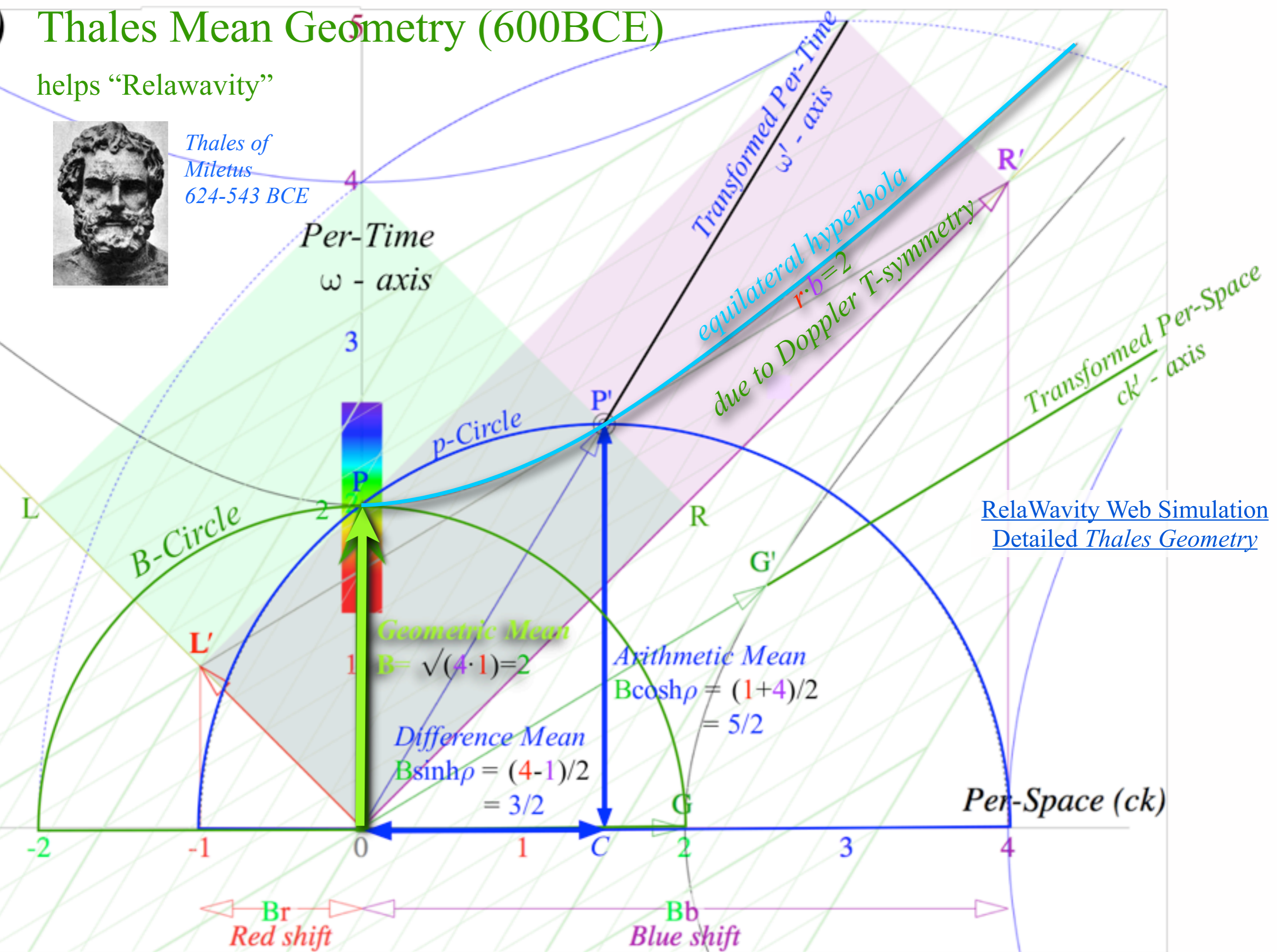


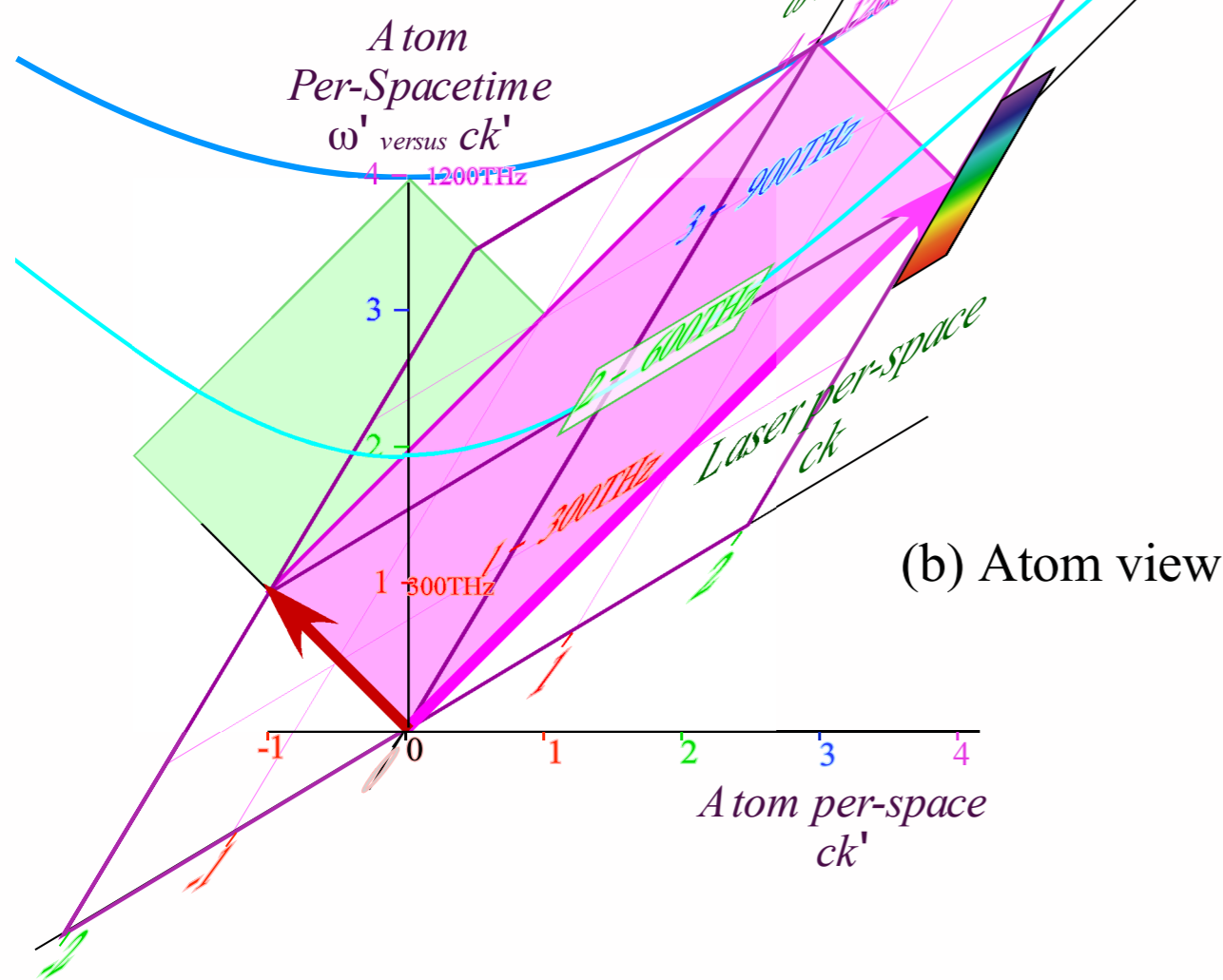
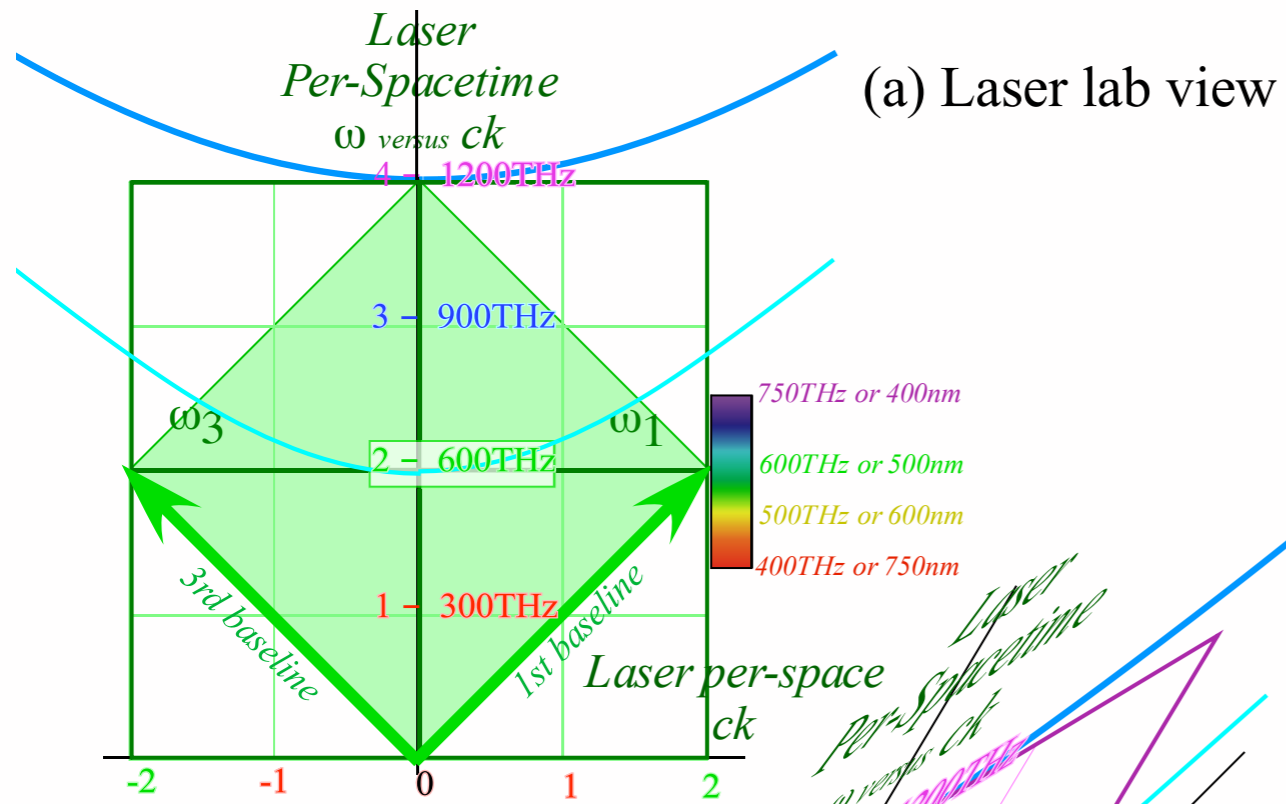
Thales Mean Geometry (600BCE)

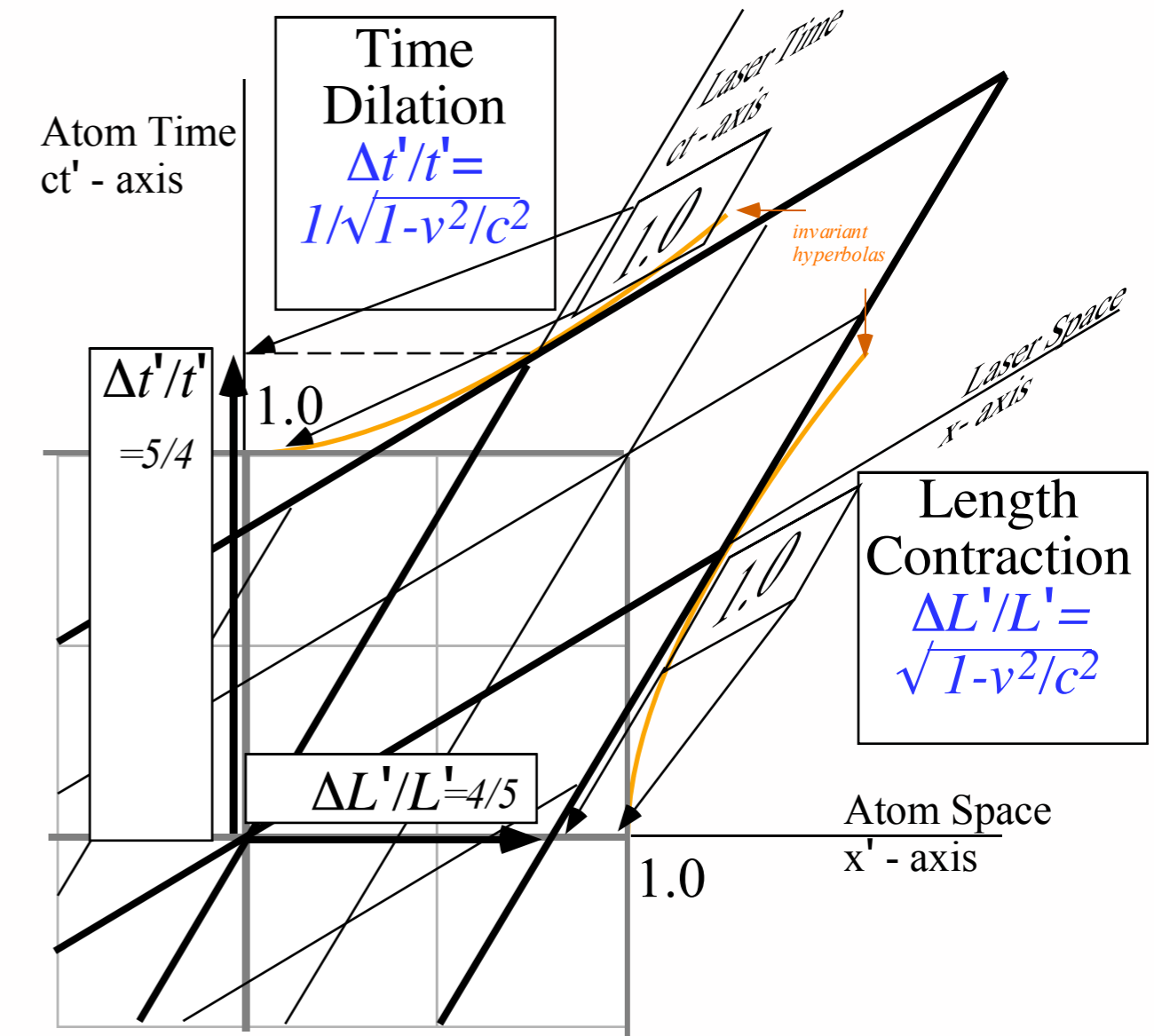
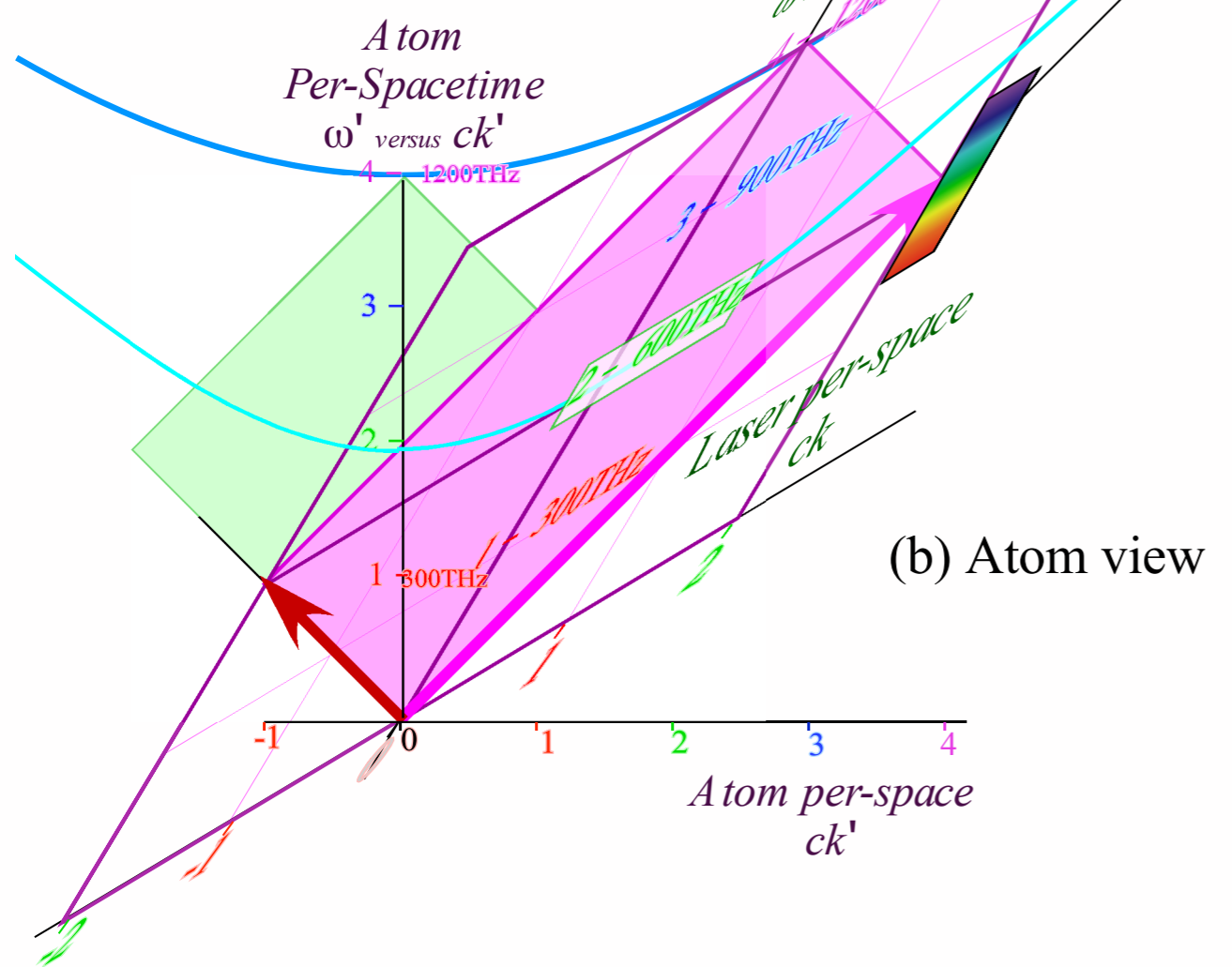
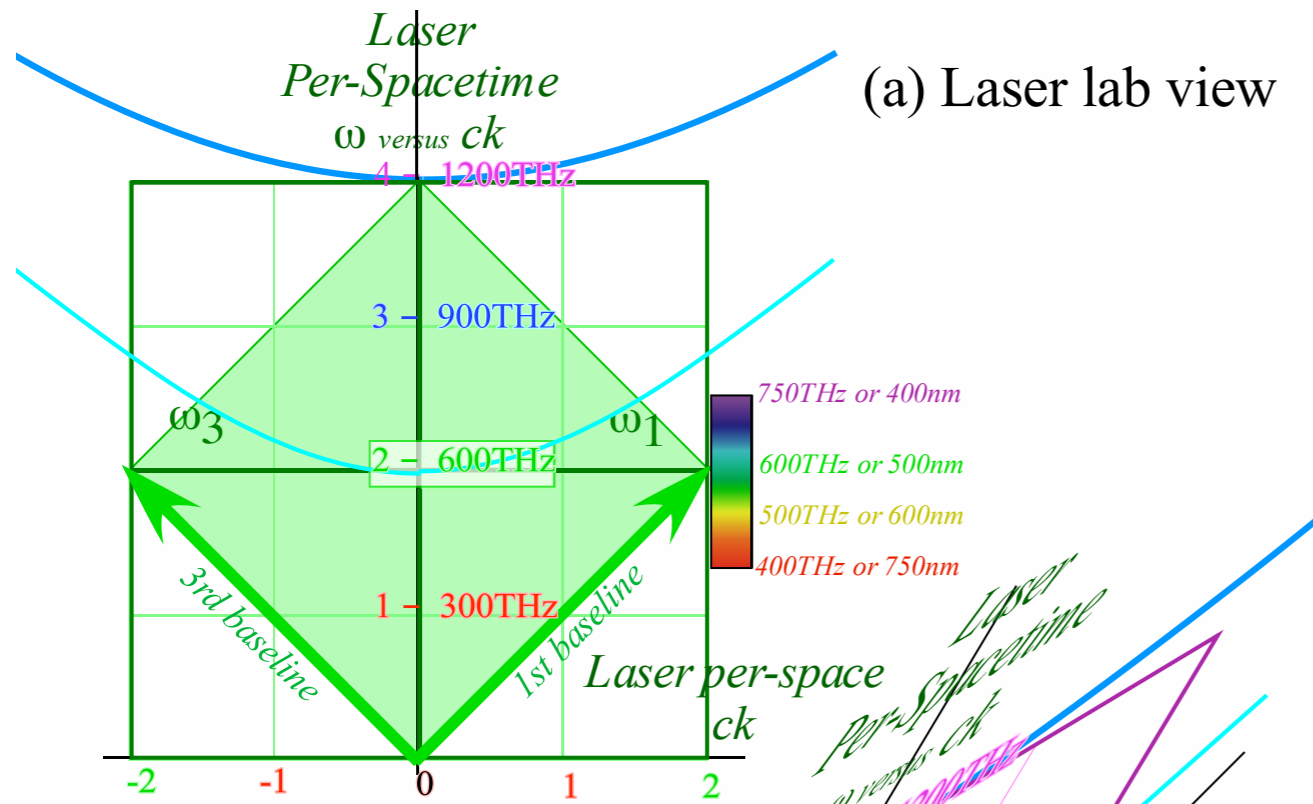
helps "Relativity"

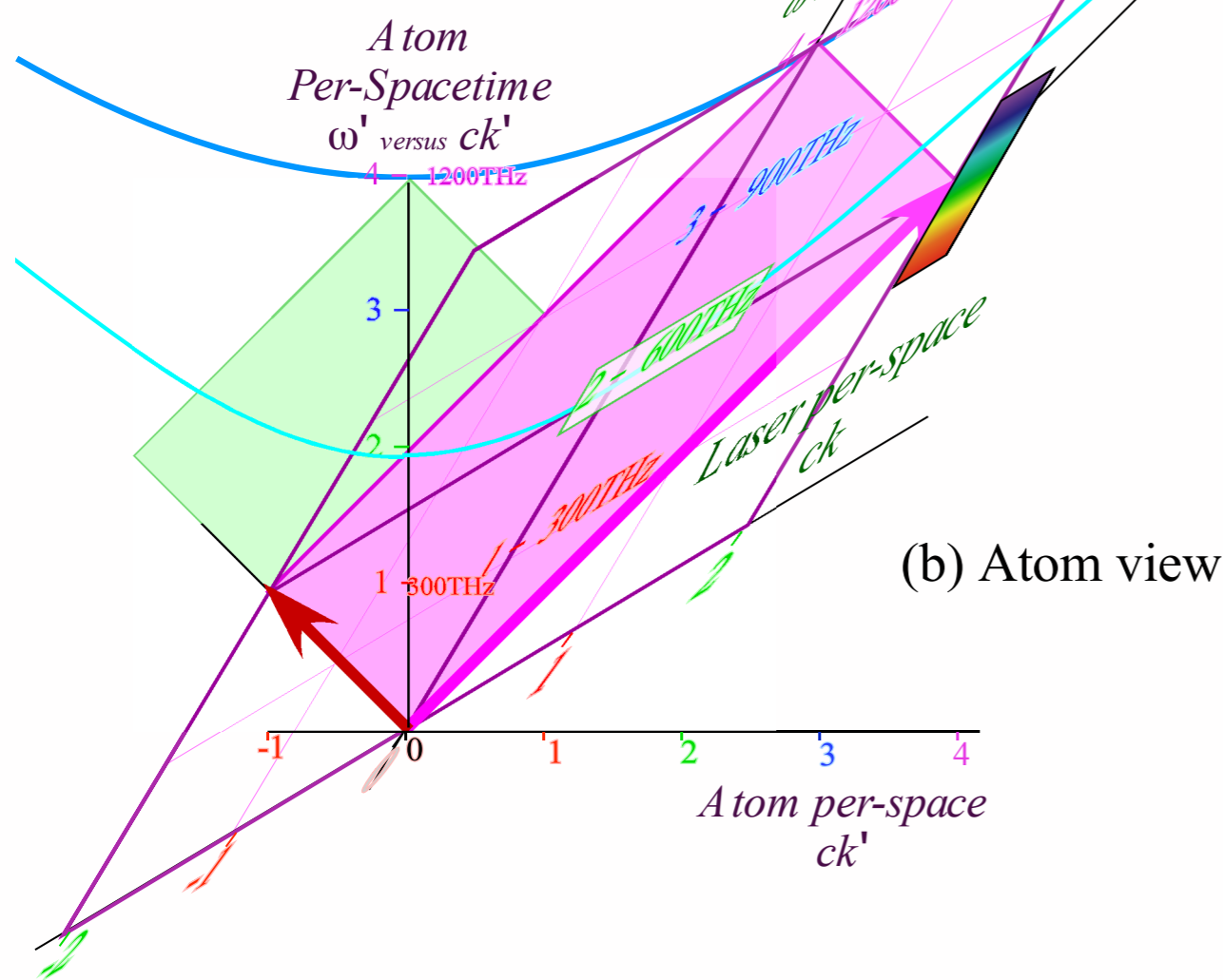
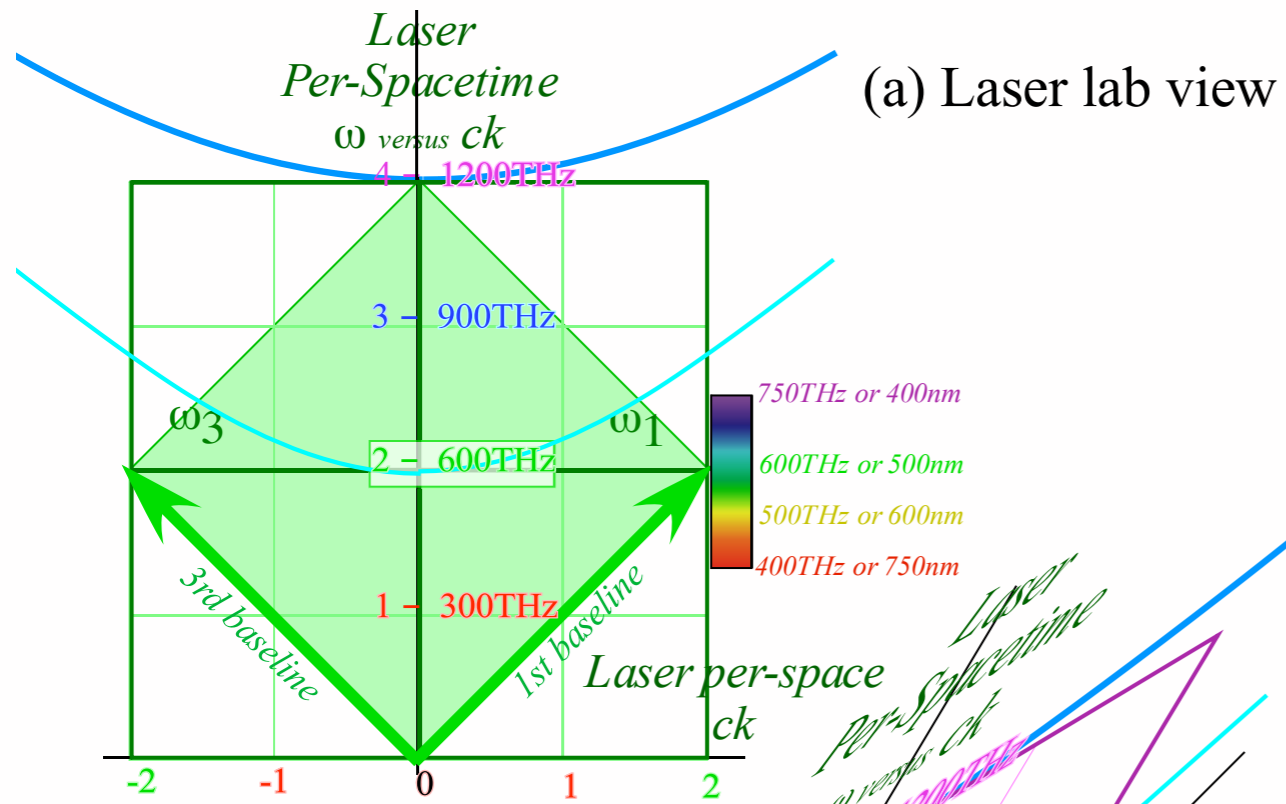


Thales of Miletus
624-543 BCE

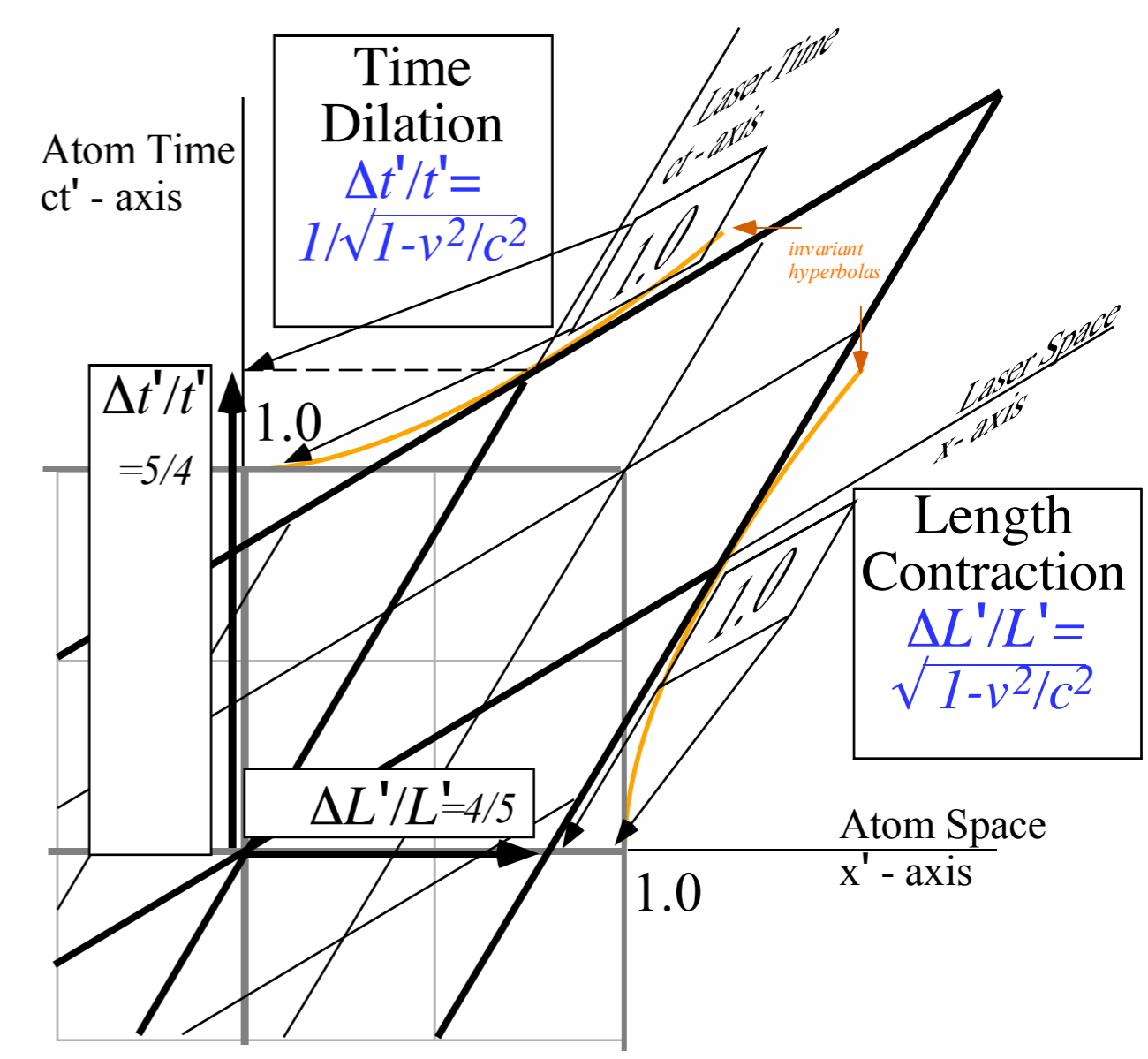








OK! But...
 What about “Time Contraction”?
 or
 “Length dilation”?



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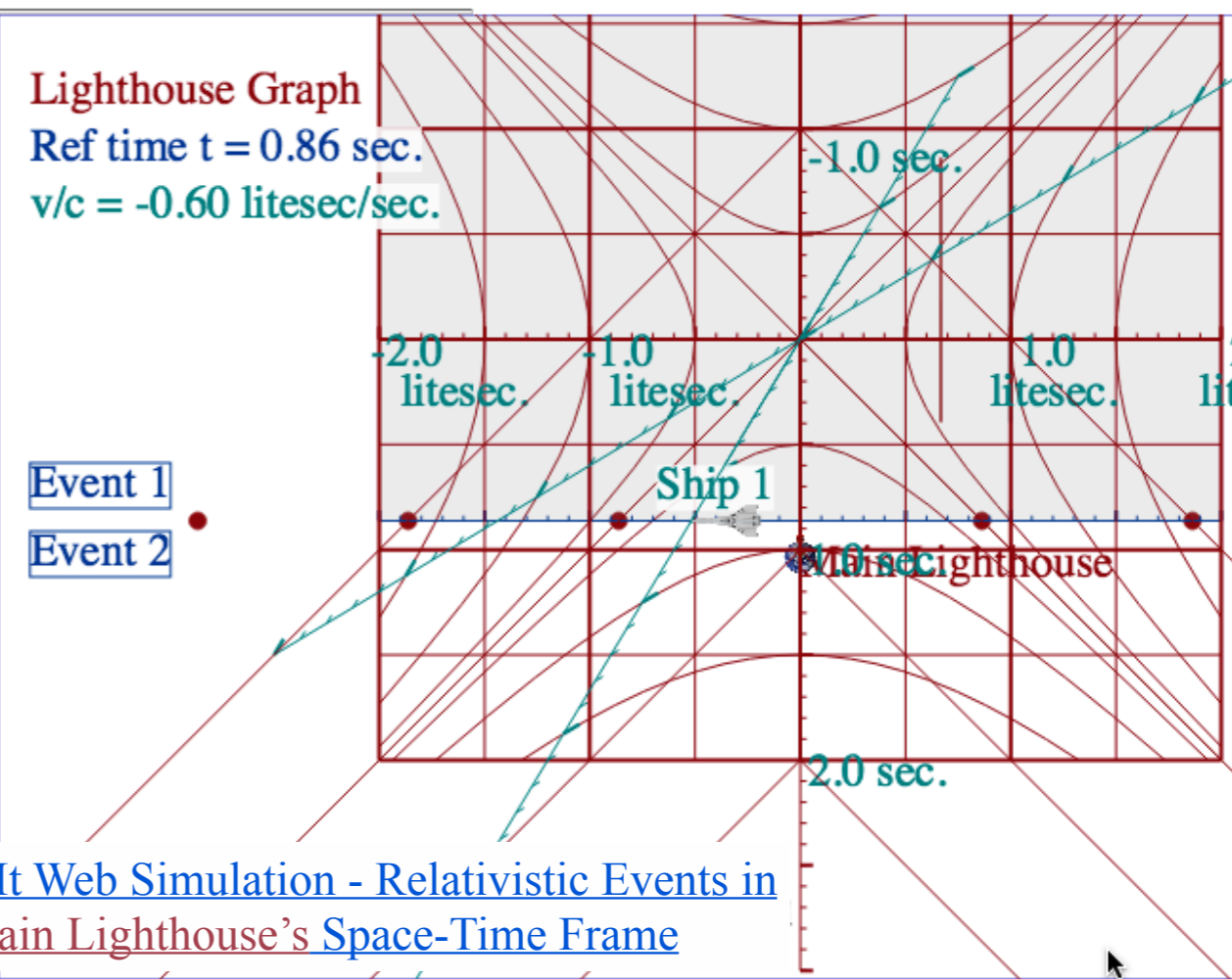
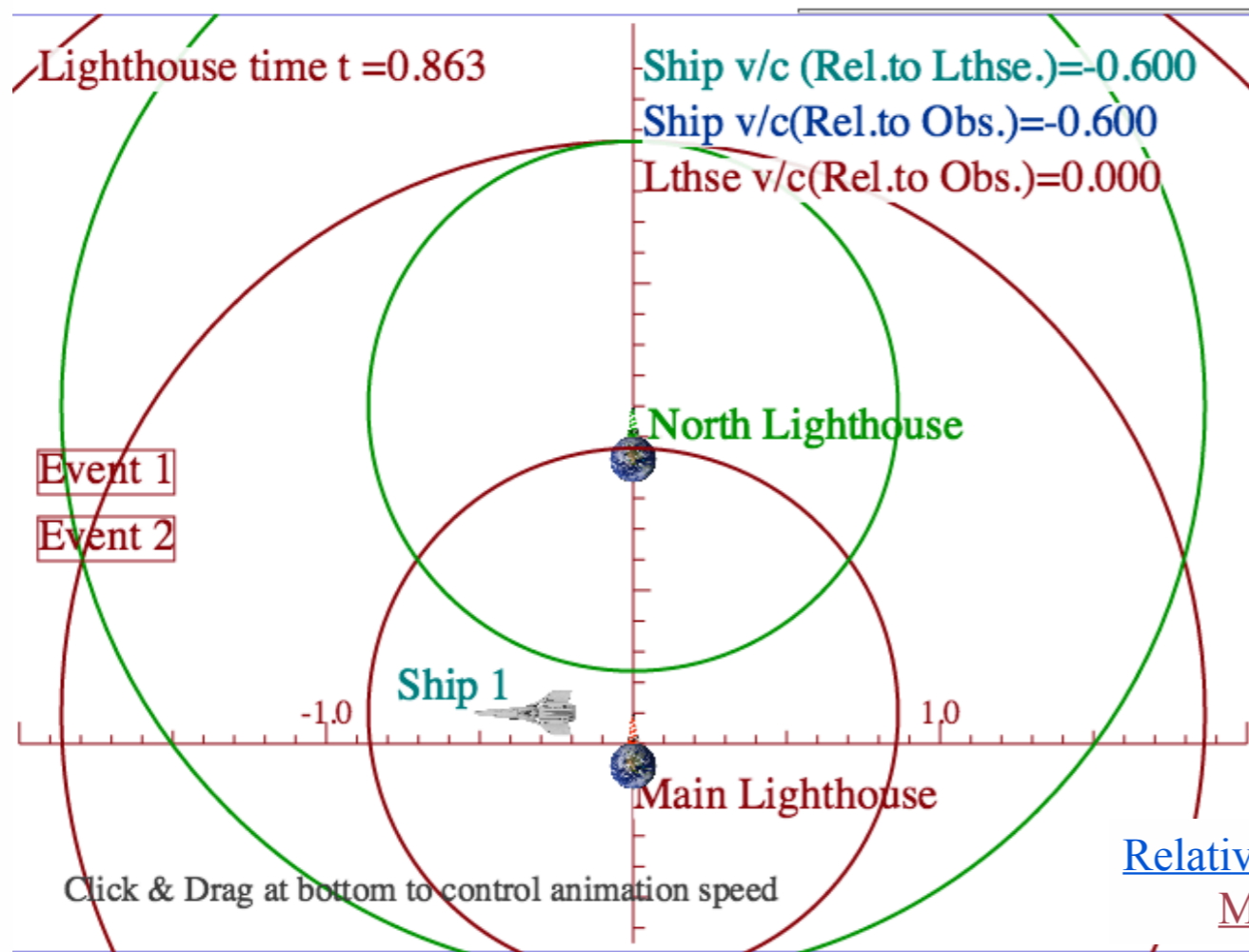
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Relativity in accelerated frames

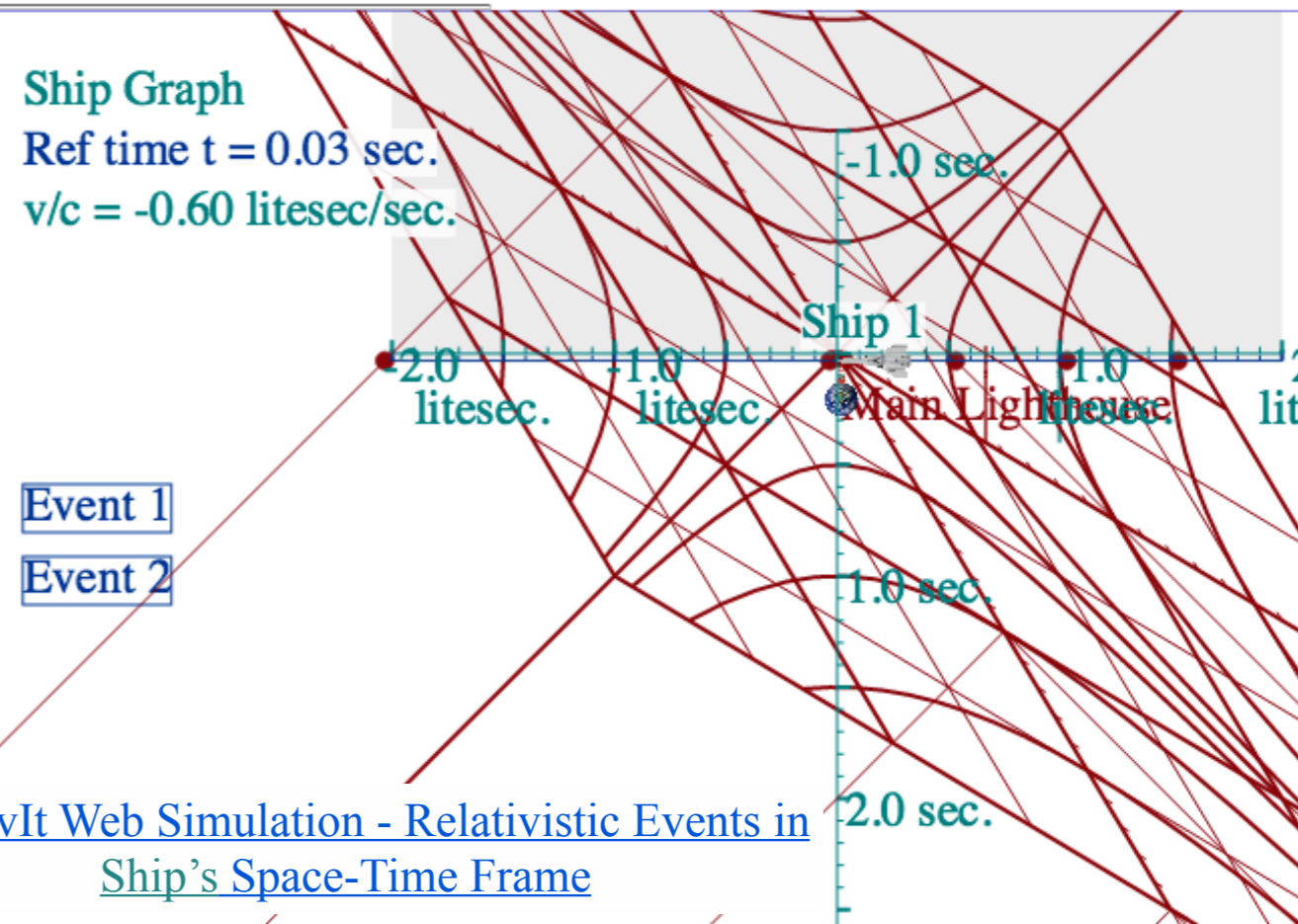
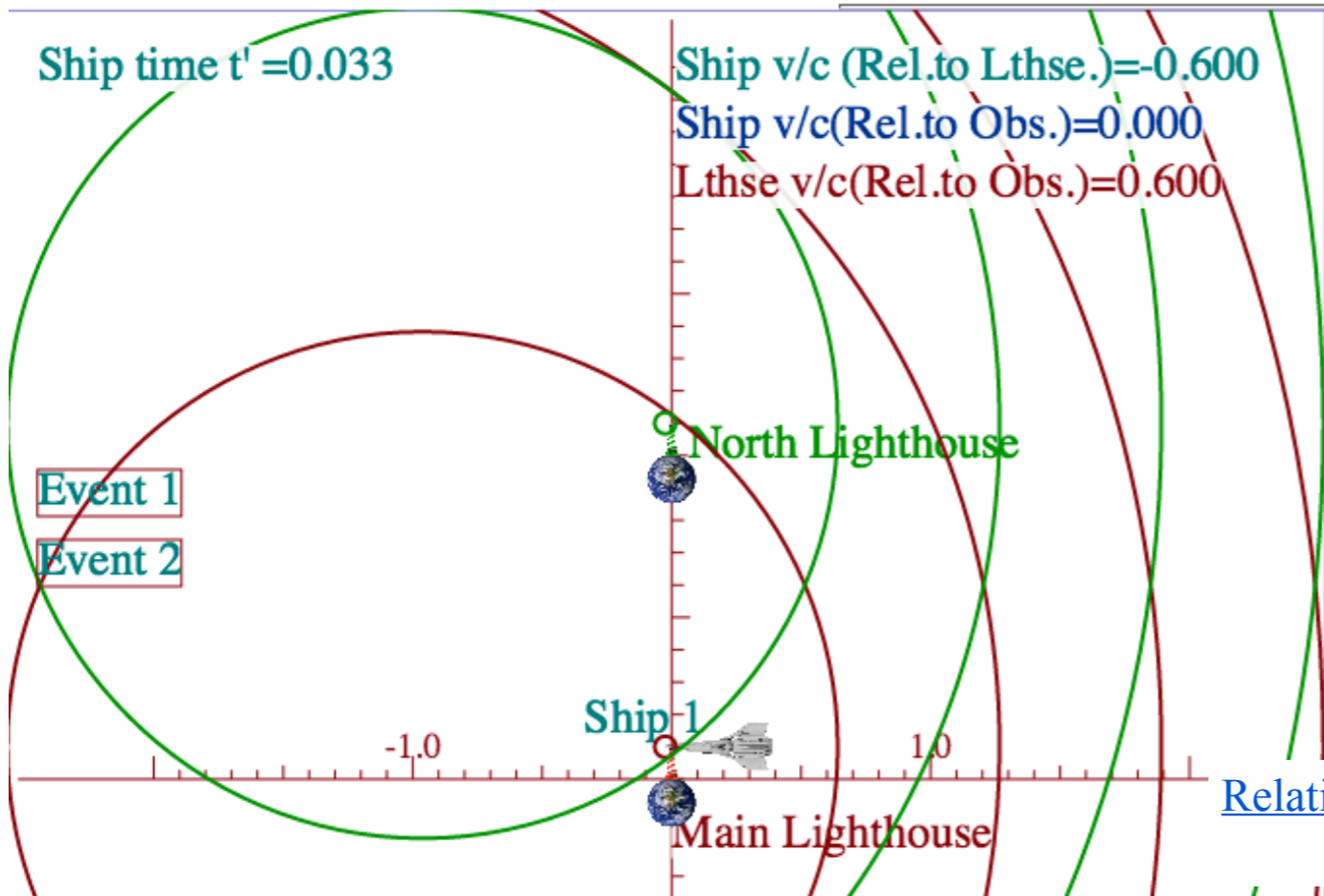
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RelativIt Web Simulation - Relativistic Events in Main Lighthouse's Space-Time Frame



RelativIt Web Simulation - Relativistic Events in Ship's Space-Time Frame

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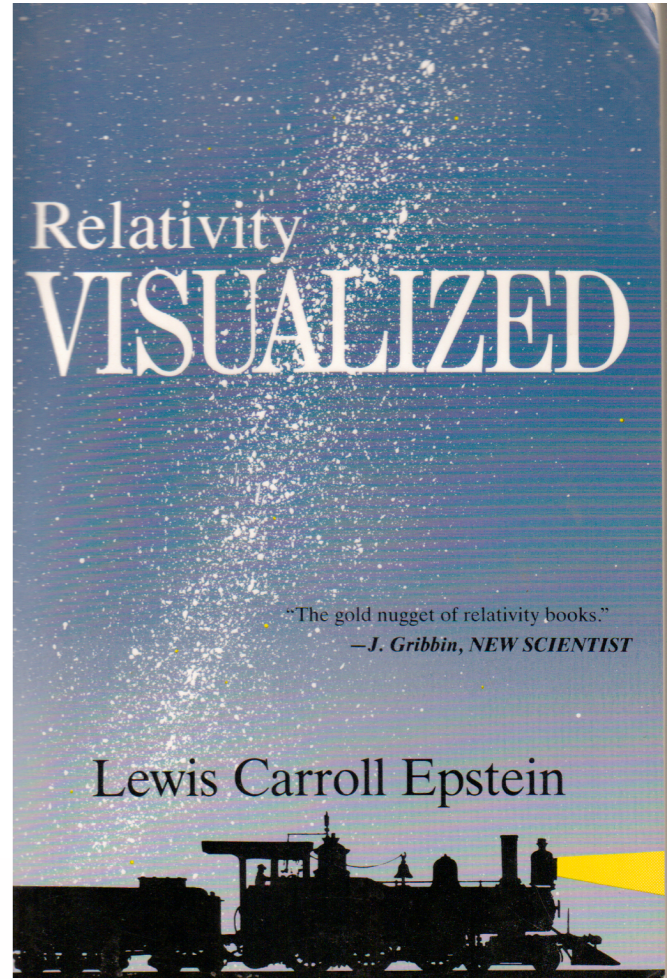
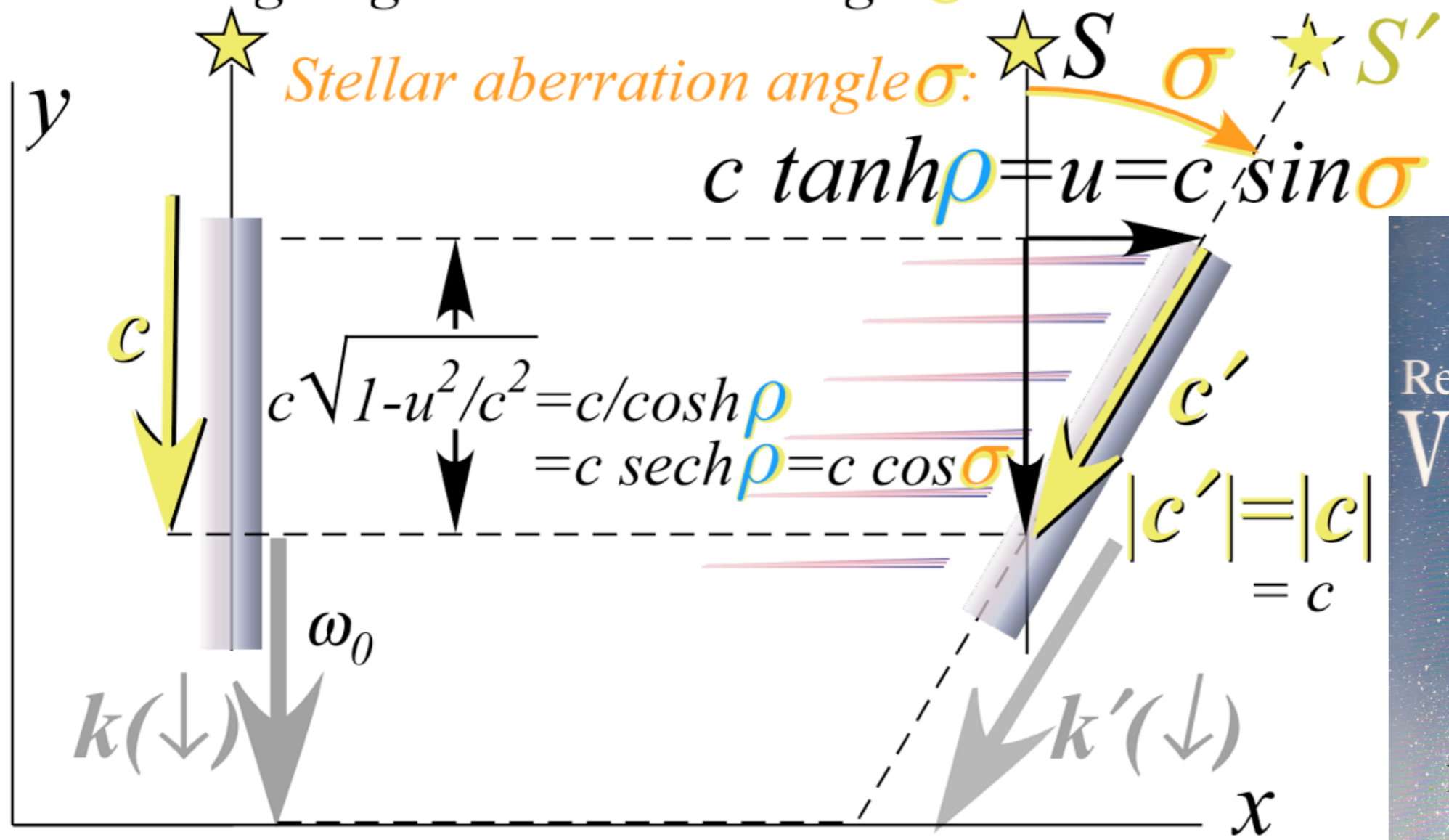
Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

to a Transverse*relativity parameter: Stellar aberration angle σ

*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.
 Observer going u sees star at angle σ in u direction.

We used notion σ for stellar-ab-angle, (a “flipped-out” ρ). Epstein not interested in ρ analysis or in relation of σ and ρ .



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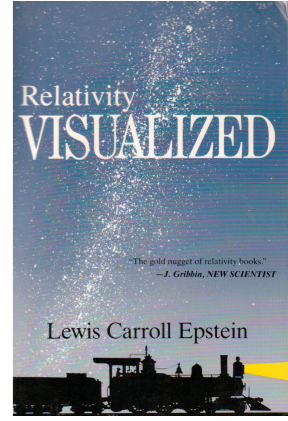
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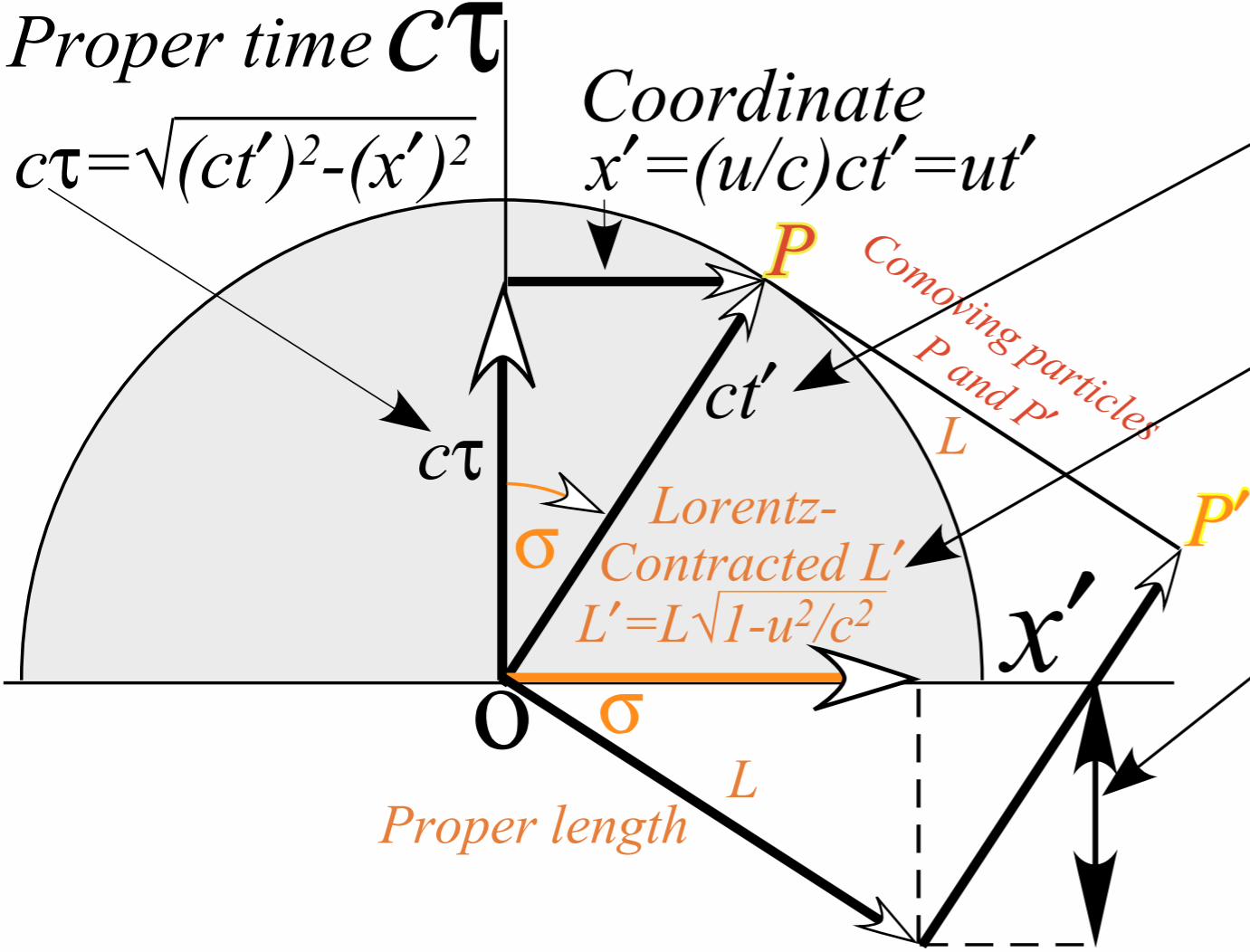
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Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$



Einstein time dilation:
 $ct' = c\tau \sec\sigma = c\tau \cosh\rho = c\tau / \sqrt{1-u^2/c^2}$

Lorentz length contraction:
 $L' = L \operatorname{sech}\rho = L \cos\sigma = L \cdot \sqrt{1-u^2/c^2}$

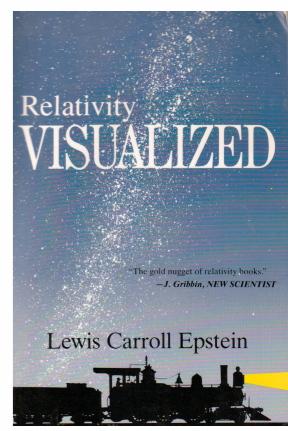
Proper Time asimultaneity:
 $c \Delta\tau = L' \sinh\rho = L \cos\sigma \sinh\rho$
 $= L \cos\sigma \tan\sigma$
 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$

Epstein's trick is to turn a hyperbolic form $c\tau = \sqrt{(ct')^2 - (x')^2}$ into a circular form: $\sqrt{(c\tau)^2 + (x')^2} = (ct')$

Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

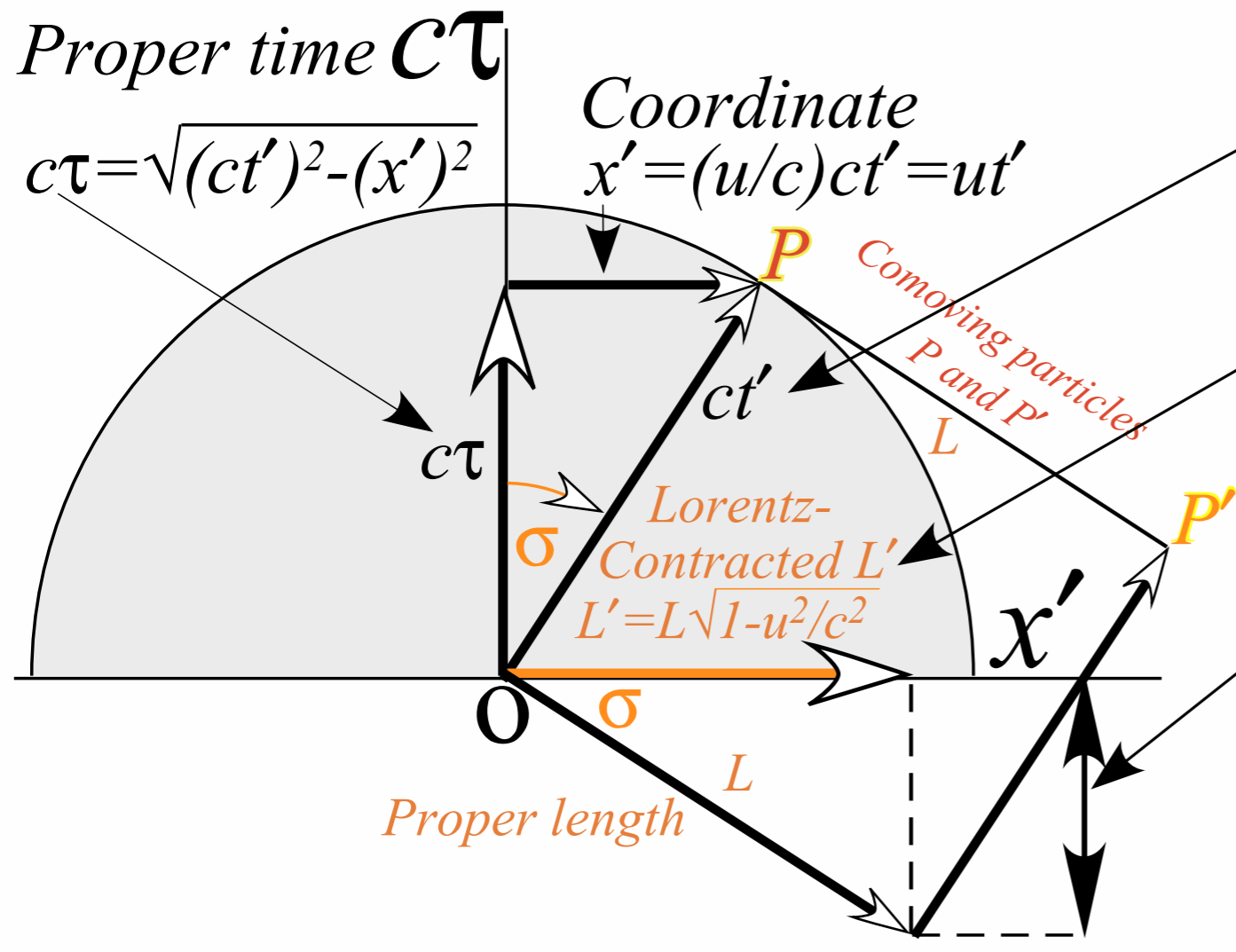
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 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$

Epstein's trick is to turn a hyperbolic form $c\tau = \sqrt{(ct')^2 - (x')^2}$ into a circular form: $\sqrt{(c\tau)^2 + (x')^2} = (ct')$
 Then everything (and everybody) always goes speed c through $(x', c\tau)$ space!

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This map has circle sector arc-area $\sigma = 0.6435$

set to angle $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

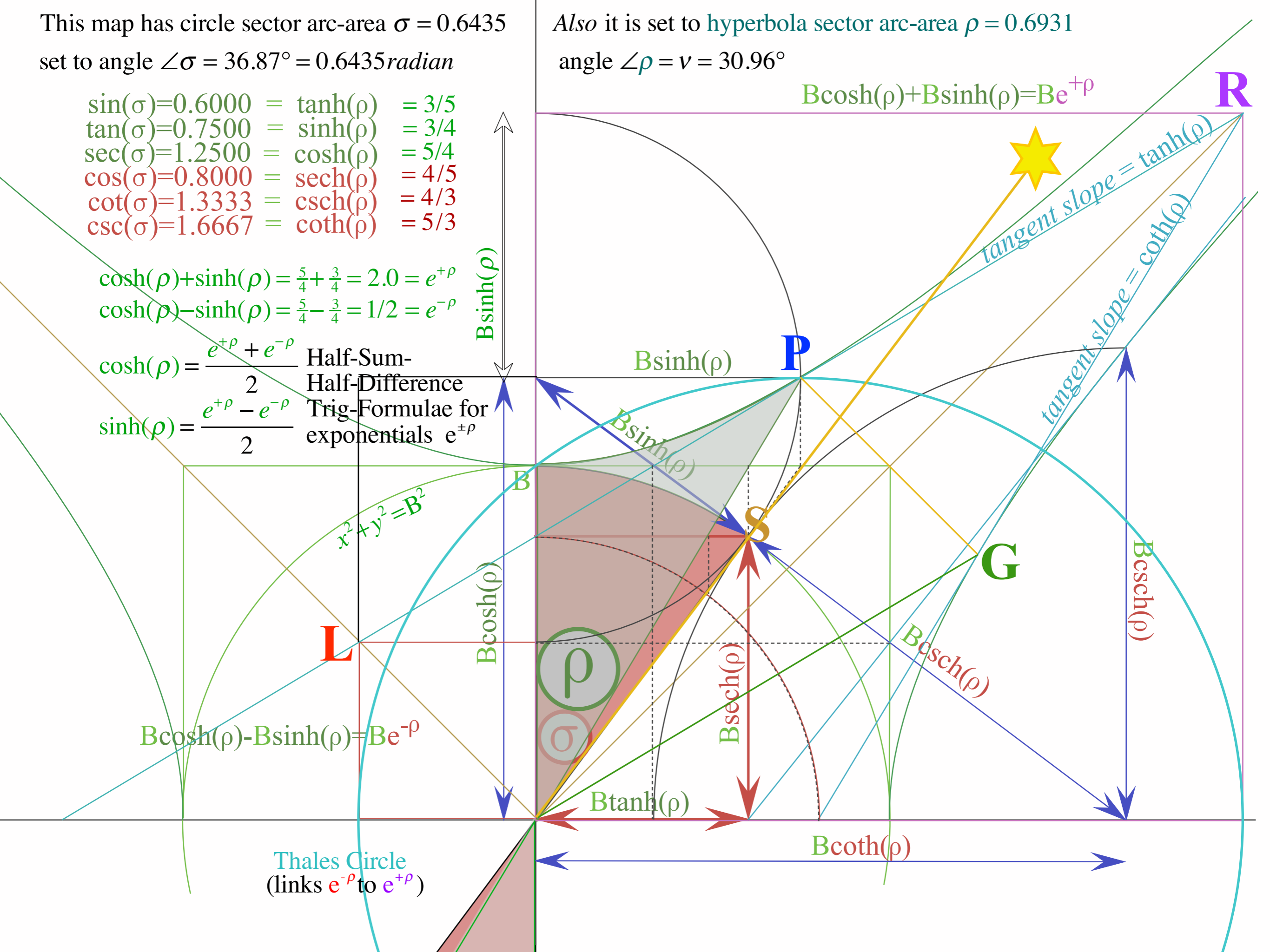
$$\begin{aligned} \sin(\sigma) &= 0.6000 &= \tanh(\rho) &= 3/5 \\ \tan(\sigma) &= 0.7500 &= \sinh(\rho) &= 3/4 \\ \sec(\sigma) &= 1.2500 &= \cosh(\rho) &= 5/4 \\ \cos(\sigma) &= 0.8000 &= \operatorname{sech}(\rho) &= 4/5 \\ \cot(\sigma) &= 1.3333 &= \operatorname{csch}(\rho) &= 4/3 \\ \csc(\sigma) &= 1.6667 &= \operatorname{coth}(\rho) &= 5/3 \end{aligned}$$

$$\begin{aligned} \cosh(\rho) + \sinh(\rho) &= \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho} \\ \cosh(\rho) - \sinh(\rho) &= \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho} \end{aligned}$$

$$\begin{aligned} \cosh(\rho) &= \frac{e^{+\rho} + e^{-\rho}}{2} && \text{Half-Sum-} \\ &&& \text{Half-Difference} \\ \sinh(\rho) &= \frac{e^{+\rho} - e^{-\rho}}{2} && \text{Trig-Formulae for} \\ &&& \text{exponentials } e^{\pm\rho} \end{aligned}$$

Also it is set to hyperbola sector arc-area $\rho = 0.6931$

angle $\angle\rho = \nu = 30.96^\circ$



Thales Circle
(links $e^{-\rho}$ to $e^{+\rho}$)

$$B\cosh(\rho) + B\sinh(\rho) = B e^{+\rho}$$

$$B\cosh(\rho) - B\sinh(\rho) = B e^{-\rho}$$

tangent slope = $\tanh(\rho)$

tangent slope = $\coth(\rho)$

R

P

G

L

ρ

S

$B\sinh(\rho)$

$B\sinh(\rho)$

$B\cosh(\rho)$

$B\operatorname{sech}(\rho)$

$B\tanh(\rho)$

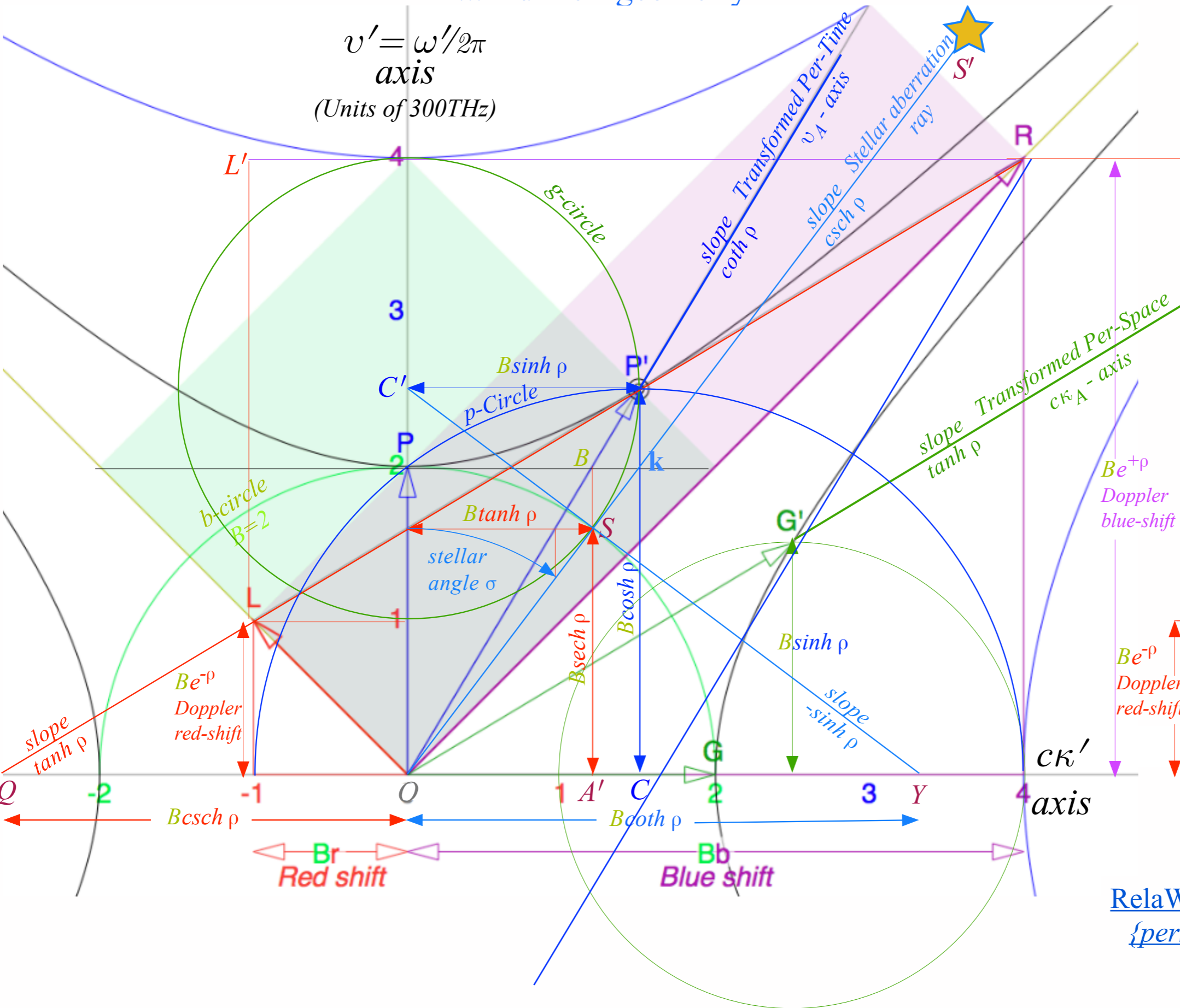
$B\coth(\rho)$

$B\operatorname{csch}(\rho)$

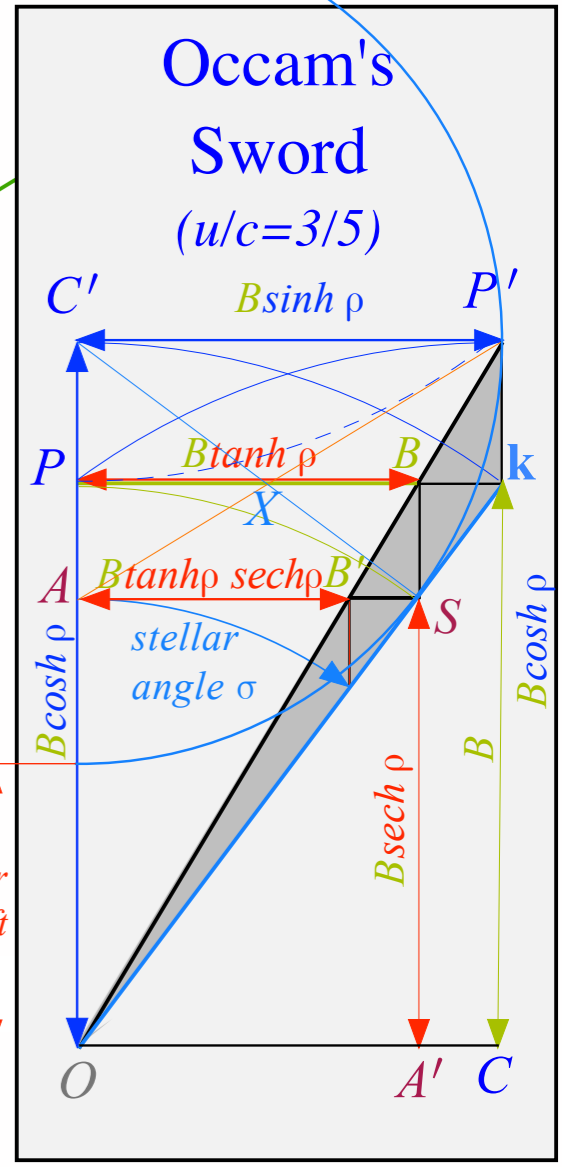
$B\operatorname{csch}(\rho)$

Summary of optical wave parameters for relativity and QM

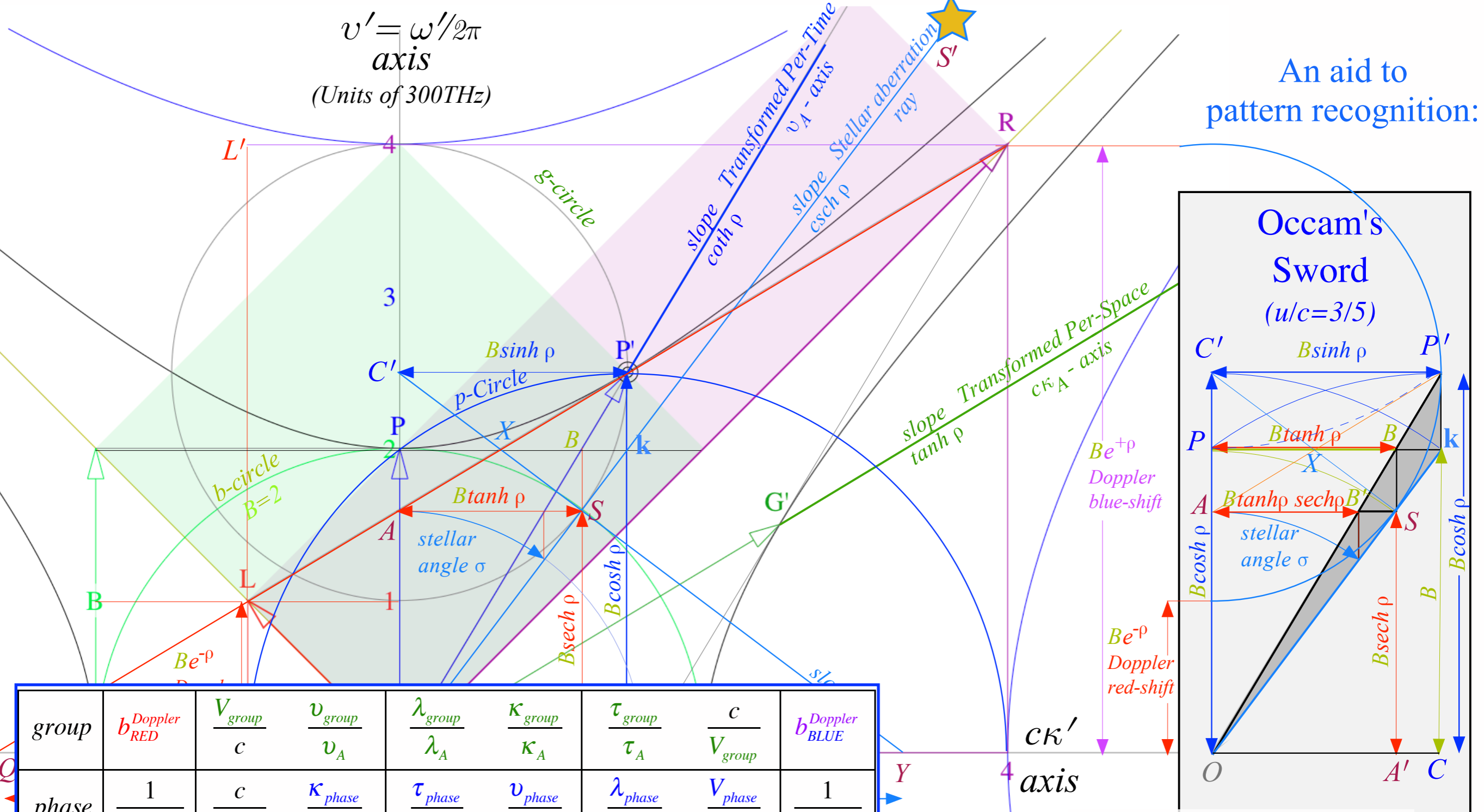
...and their geometry



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RelaWavity Web Simulation
 {perSpace - perTime All}



group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{\nu_{group}}{\nu_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{\nu_{phase}}{\nu_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

RelaWavity Web Simulation
{perSpace - perTime All}

Table of 12 wave parameters (includes inverses) for relativity

...and values for $u/c=3/5$

RelaWavity Web Simulation
Expanded Table of Relativistic Relations

Lecture 31

Thur. 12.10.2015

Review: Relativity ρ functions Two famous ones Extremes and plot vs. ρ
Doppler jeopardy Geometric mean and Relativistic hyperbolas
Animation of $e^{\rho}=2$ spacetime and per-spacetime plots

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity

Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

“Occams Sword” and summary of 16 parameter functions of ρ and σ

➔ Applications to optical waveguide, spherical waves, and accelerator radiation

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Compton recoil related to rocket velocity formula

Comparing 2nd-quantization “photon” number N and 1st-quantization wavenumber κ

Relativity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes g -acceleration grid

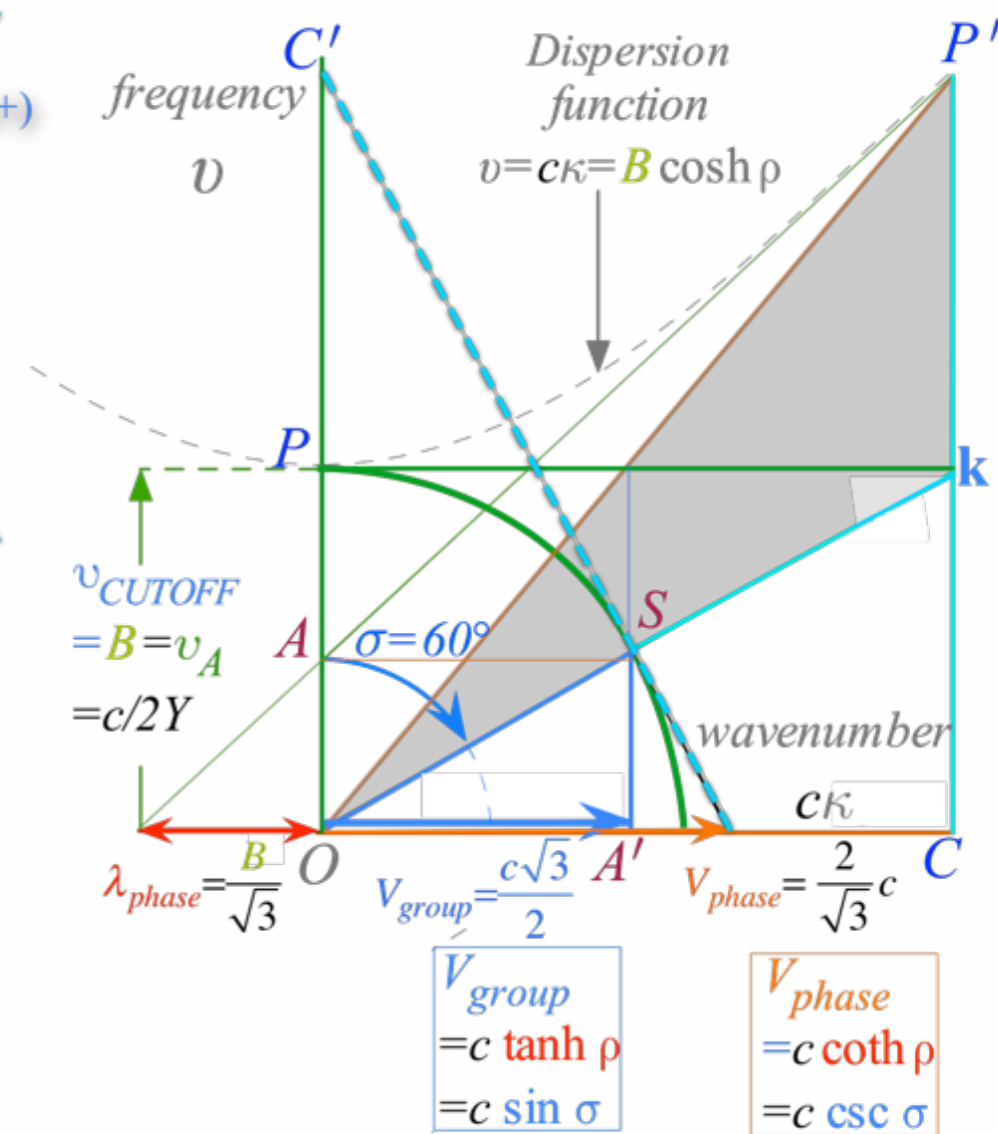
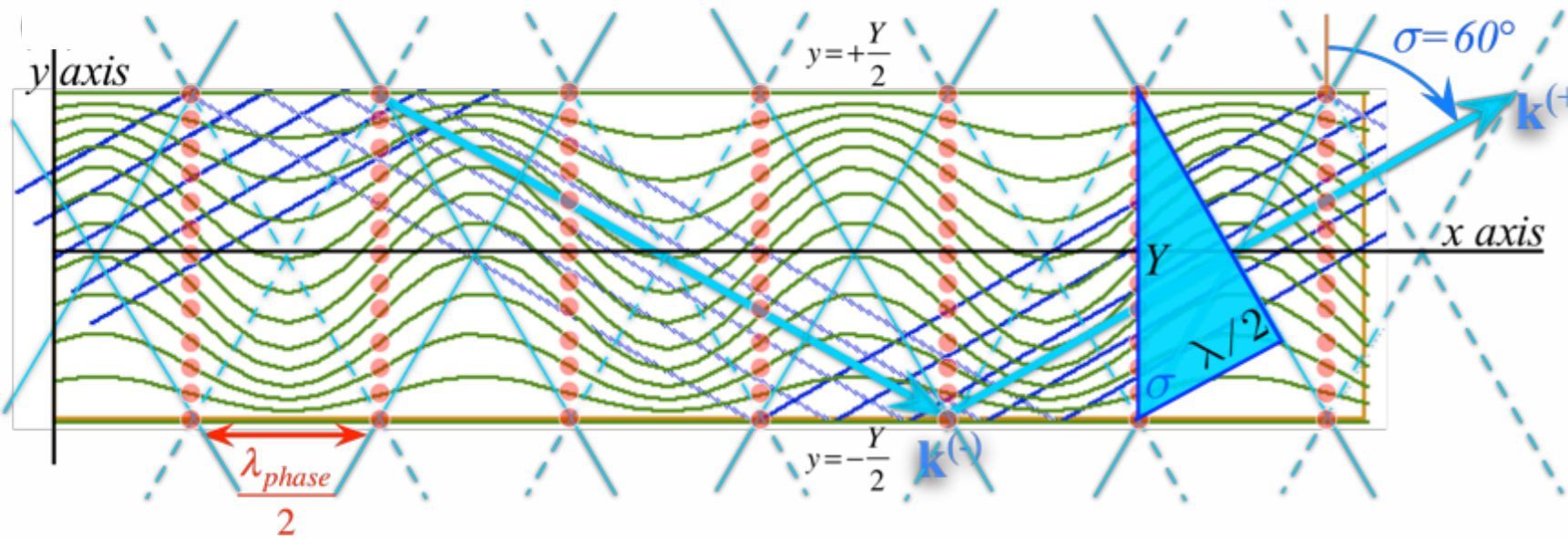
Analysis of constant- g grid compared to zero- g Minkowski grid

Animation of mechanics and metrology of constant- g grid

Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to (x,y) space-space
 to (k_x, k_y) per-space-per-space
 to (x, ct) space-time

Relativistic mode with near-c $V_{group}=c/2$ and $V_{phase}=2c$. (Low dispersion.)

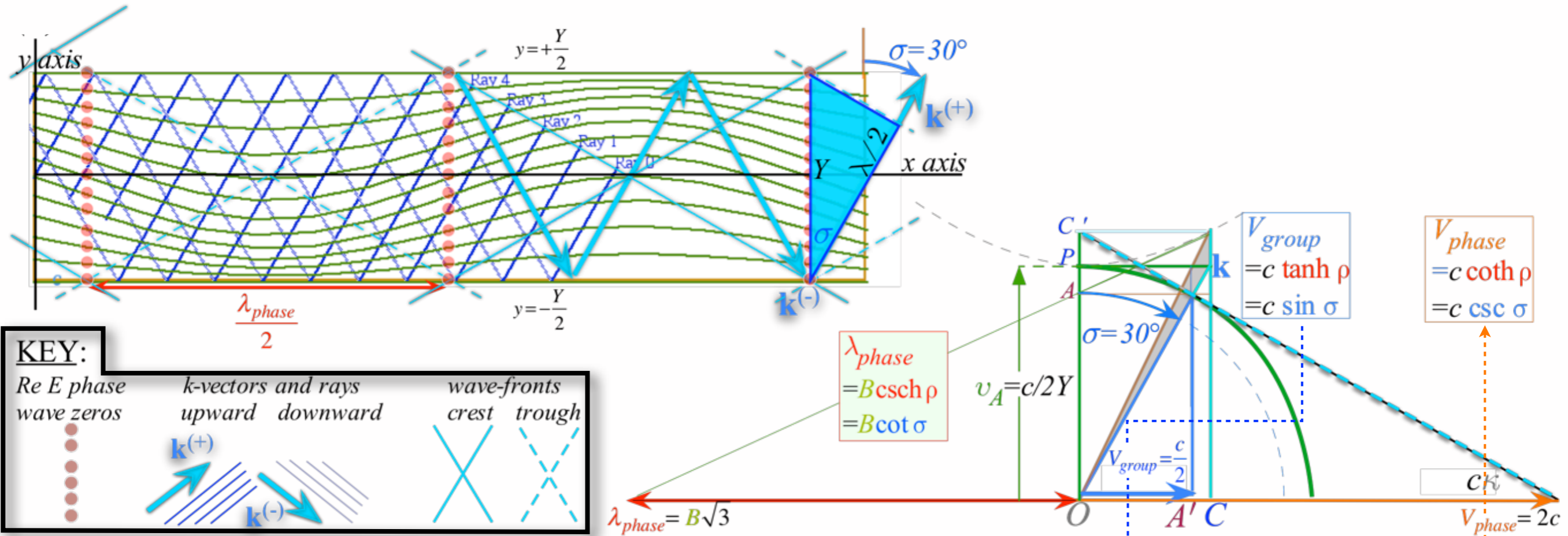


KEY:

Re E phase wave zeros	k-vectors and rays upward downward	wave-fronts crest trough

Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to (x,y) space-space
 to (k_x, k_y) per-space-per-space
 to (x, ct) space-time

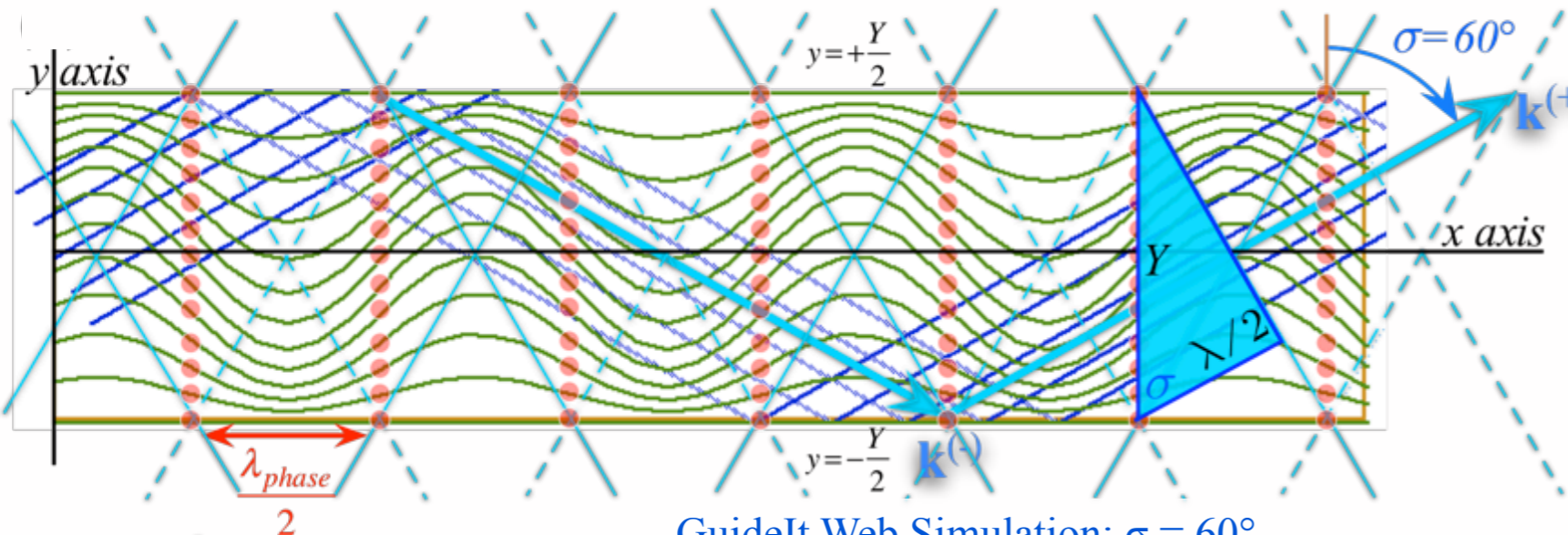


Example of near-cut-off mode with low $V_{group} = c/2$ and high $V_{phase} = 2c$. (High dispersion.)

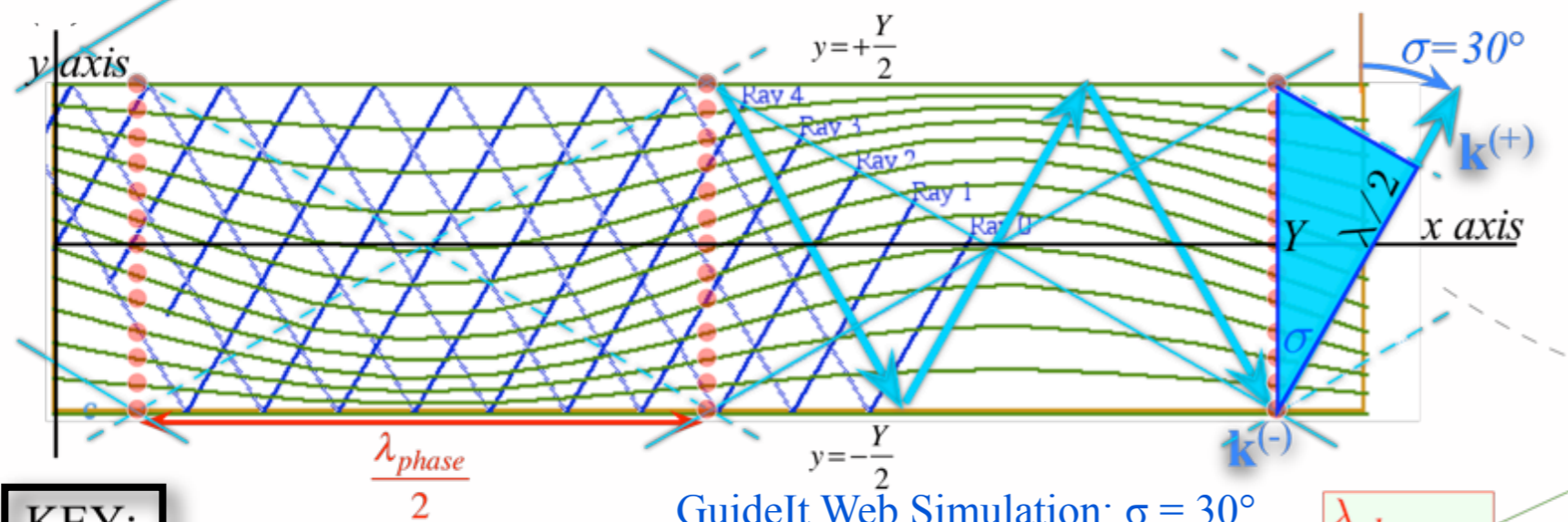
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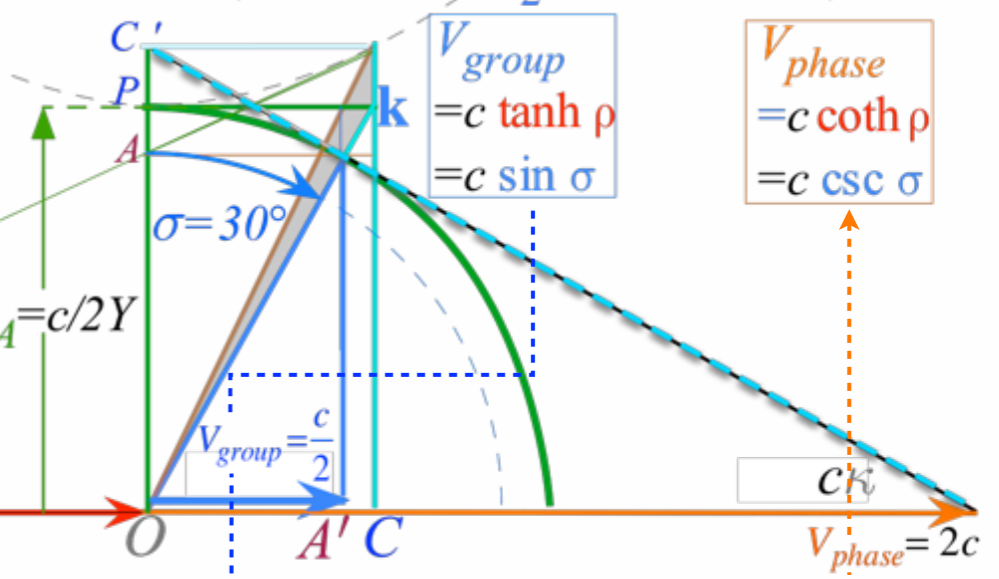
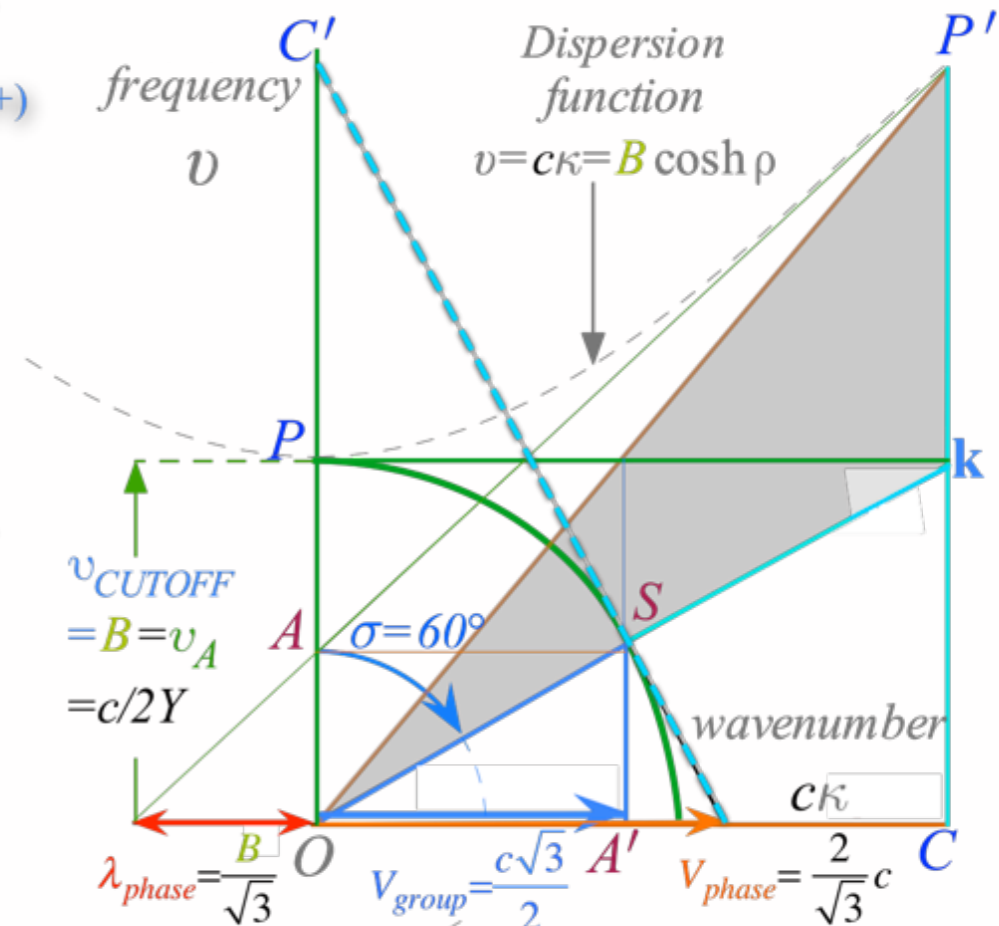
GuideIt Web Simulation: $\sigma = 60^\circ$



GuideIt Web Simulation: $\sigma = 30^\circ$

KEY:

Re E phase wave zeros	k-vectors and rays upward downward	wave-fronts crest trough



$$\lambda_{phase} = B \csc \rho = B \cot \sigma$$

$$V_{group} = c \tanh \rho = c \sin \sigma$$

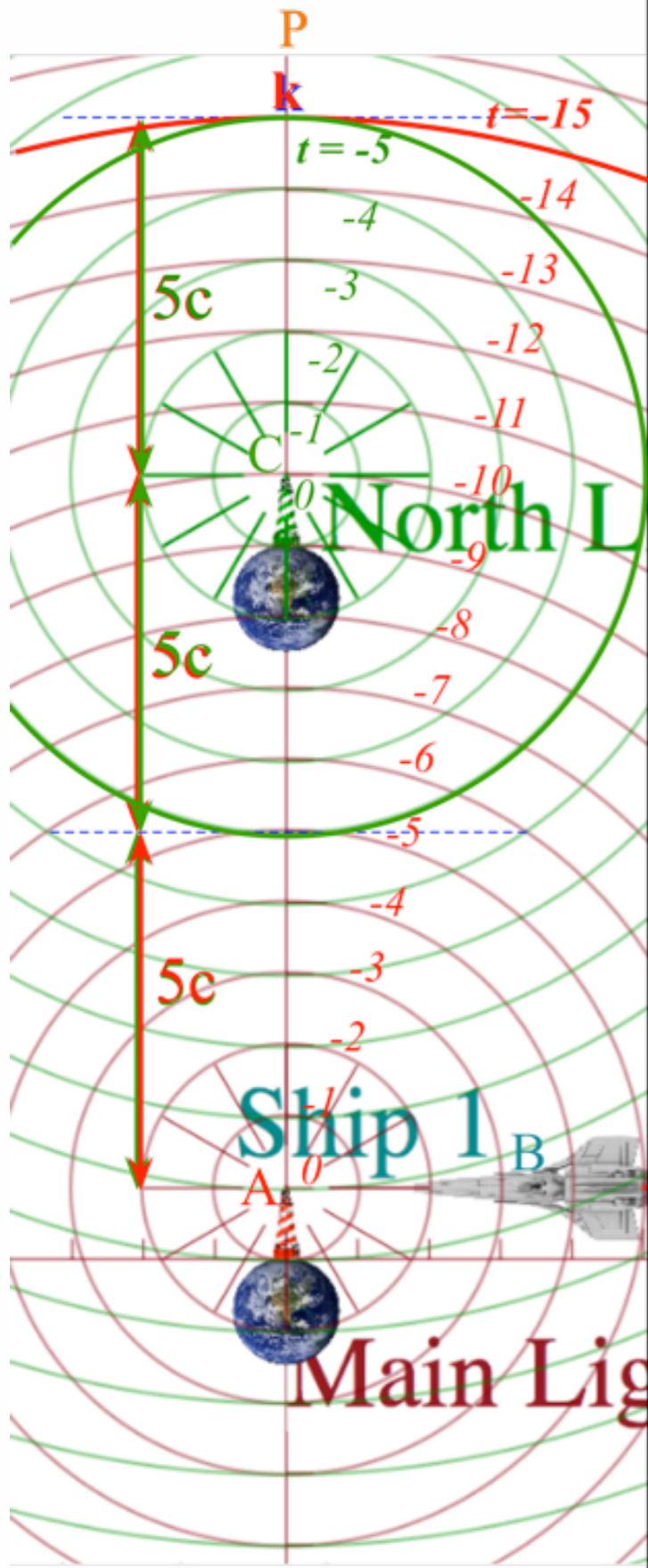
$$V_{phase} = c \coth \rho = c \csc \sigma$$

Example of near-cut-off mode with low $V_{group}=c/2$ and high $V_{phase}=2c$. (High dispersion.)

Spherical wave relativistic geometry

Also, aided by Occam's Sword

(a) Spherical wave pair
In Alice-Carla frame

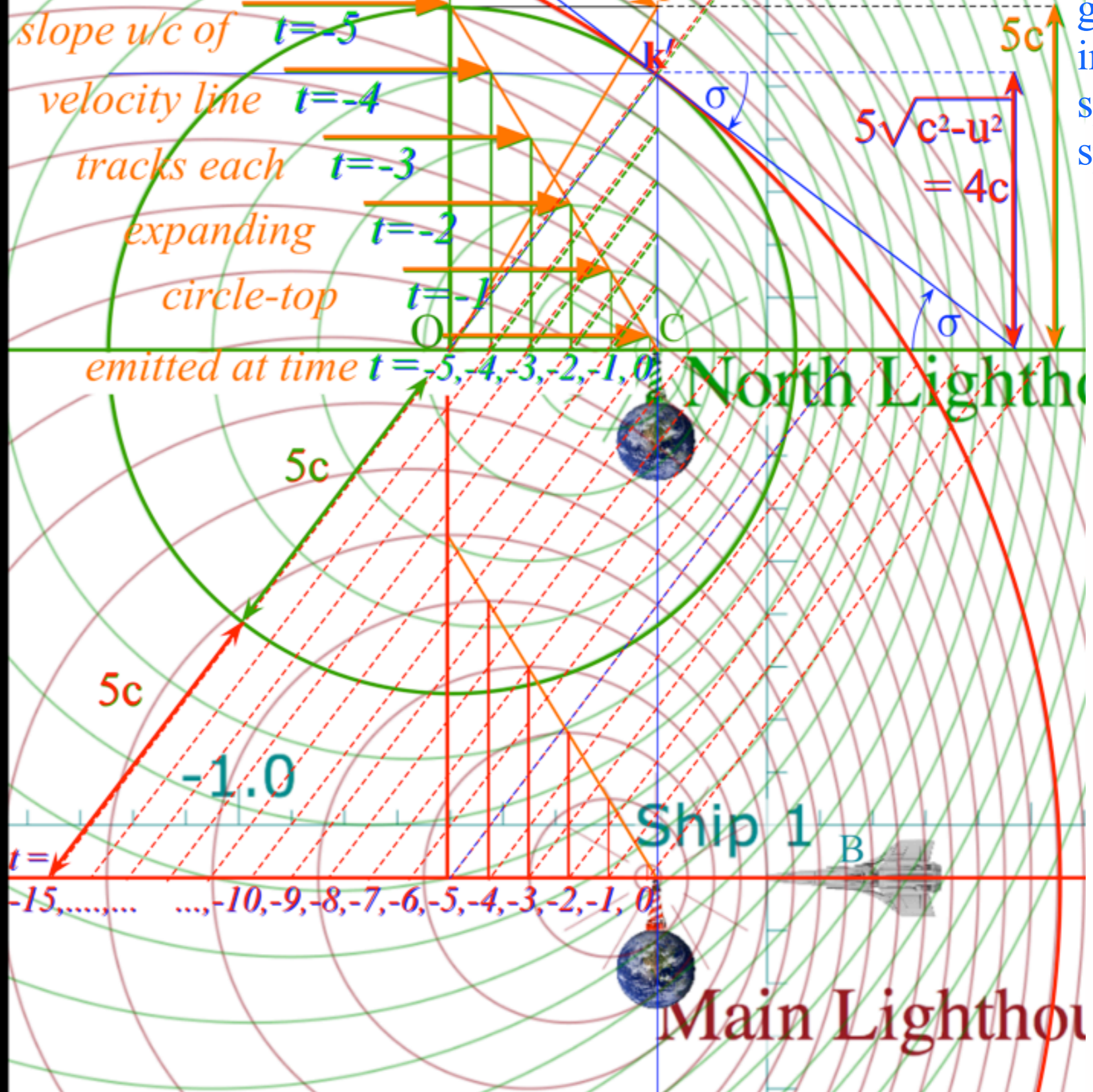
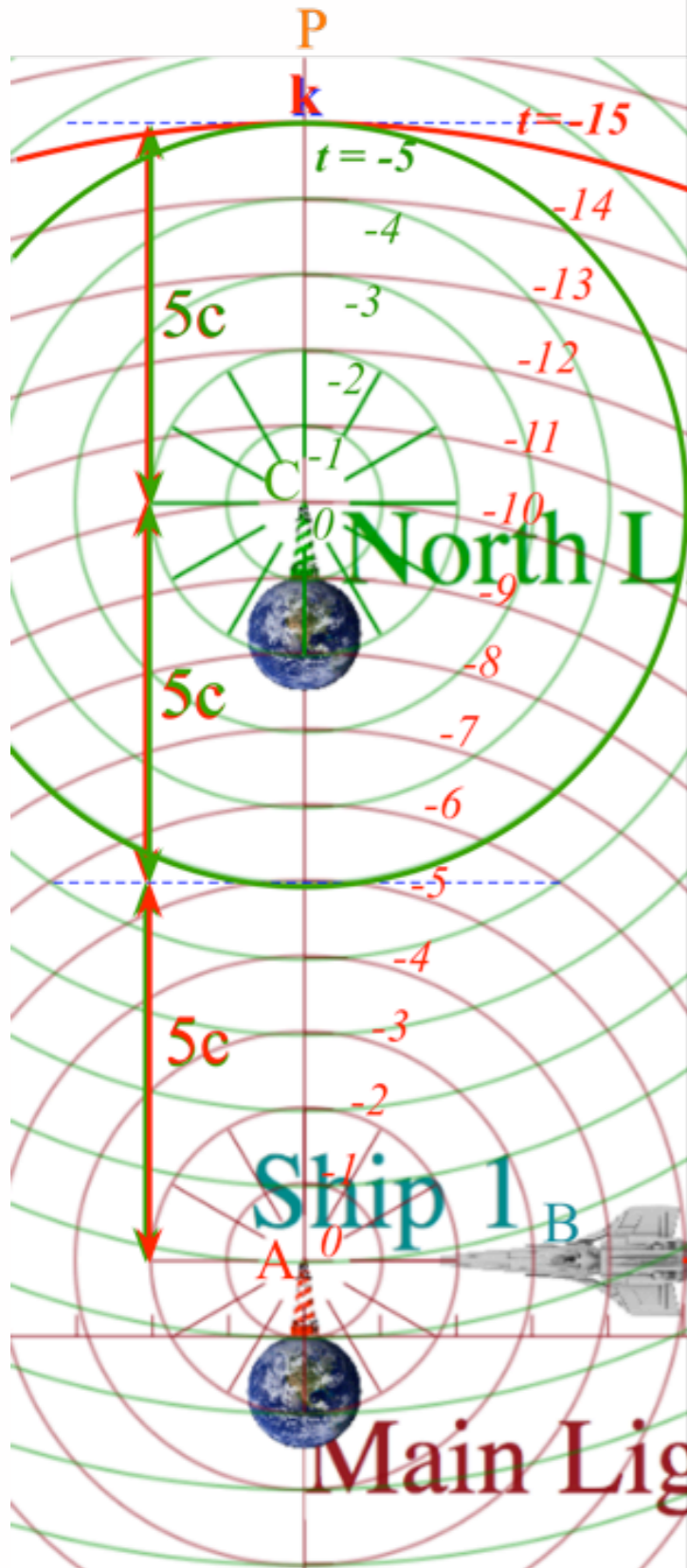


(a) Spherical wave pair
In Alice-Carla frame

stellar angle $\sigma = \sin^{-1}(u/c)$
velocity angle $v = \tan^{-1}(u/c)$
slope u/c of $t=-5$
velocity line $t=-4$
tracks each $t=-3$
expanding $t=-2$
circle-top $t=-1$
emitted at time $t=-5, -4, -3, -2, -1, 0$

(b) Spherical wave pair
In Bob's frame: $u_x/c = -3/5$

Occam
Sword
geometry
in (x,y)
space-
space



Main Lighthouse's Frame

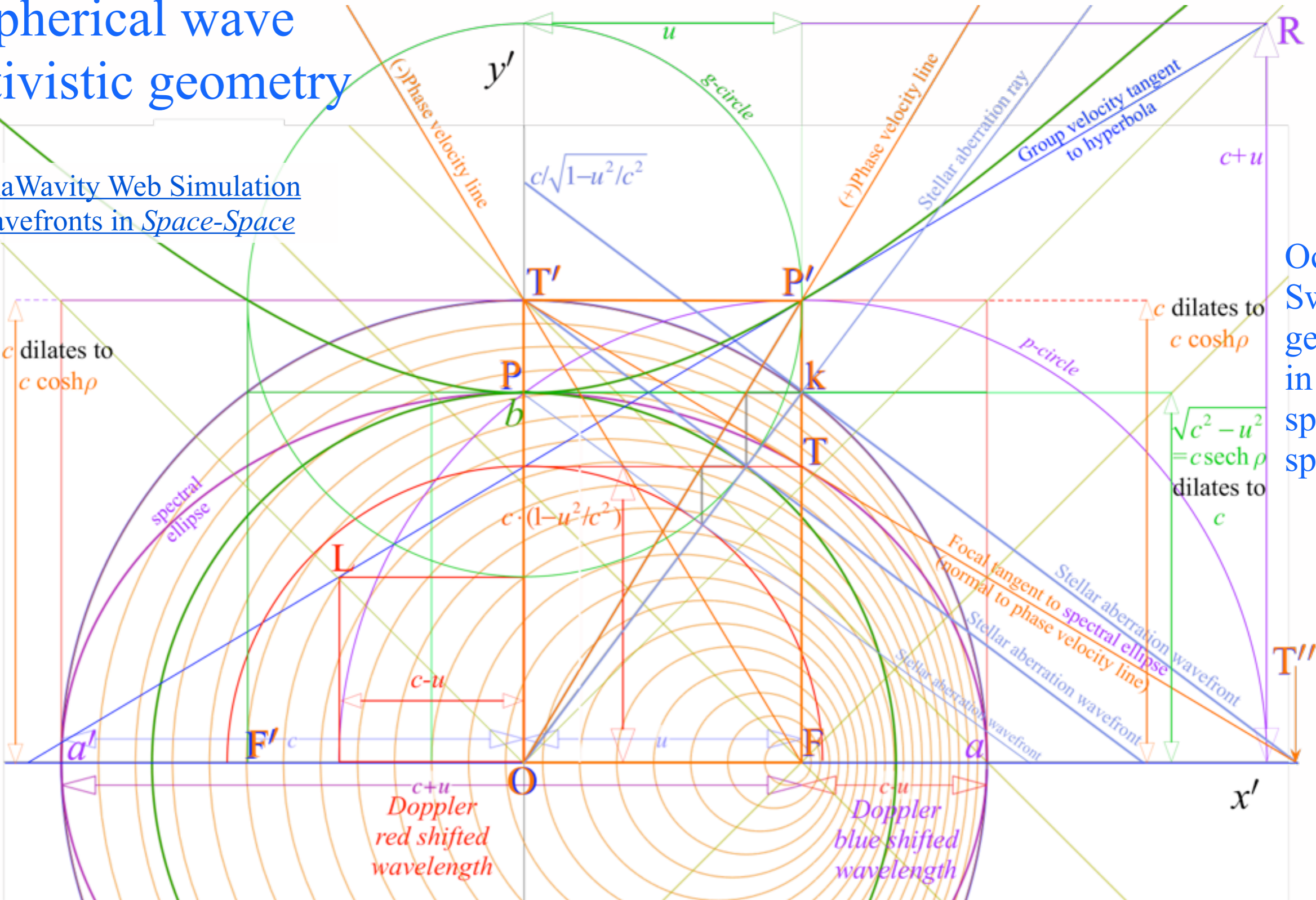
← RelativIt Web Simulation - Space-Time with many blinks →

Ship's Frame

Spherical wave relativistic geometry

RelaWavity Web Simulation
Wavefronts in Space-Space

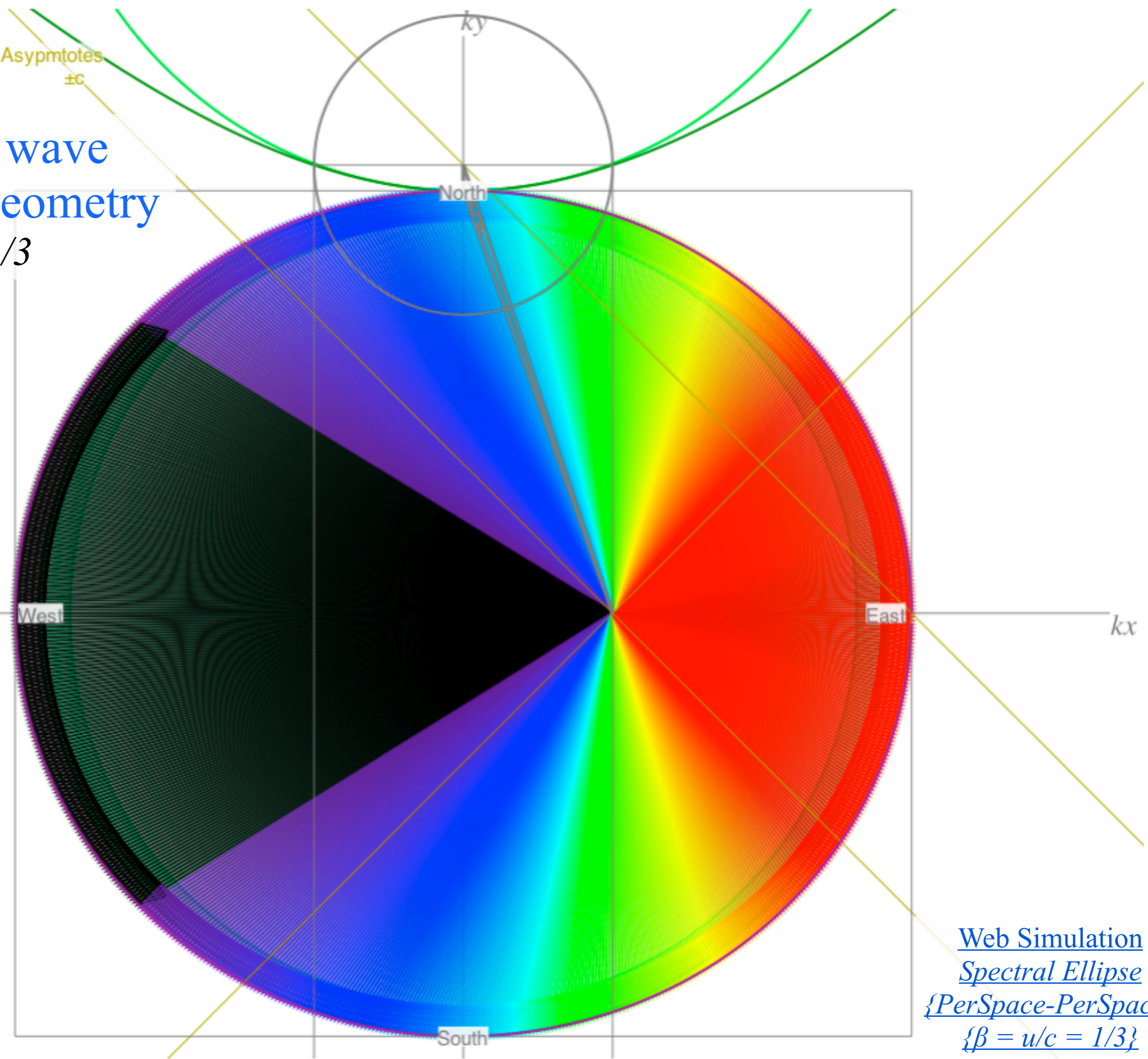
Occam
Sword
geometry
in (x,y)
space-space



<p>Doppler Red $\lambda=c+u$ dilates to: $(c+u) \cosh \rho = c \sqrt{\frac{c+u}{c-u}} = ce^{+\rho}$</p> <p>ellipse major radius $a=OFa=c$ dilates to: $c \cosh \rho = c/\sqrt{1-u^2/c^2}$</p>	<p>Applications of Einstein dilation factor: $\gamma = \cosh \rho = 1/\sqrt{1-u^2/c^2}$</p>	<p>ellipse focal length $FO = u = c \tanh \rho$ dilates to: $u \cosh \rho = c \sinh \rho$</p> <p>ellipse latus radius $FT = c(1-u^2/c^2)$ dilates to: $c(1-u^2/c^2) \cosh \rho = c\sqrt{1-u^2/c^2} = c \operatorname{sech} \rho$</p>	<p>Doppler Blue $\lambda=c-u$ dilates to: $(c-u) \cosh \rho = c \sqrt{\frac{c-u}{c+u}} = ce^{-\rho}$</p> <p>Base height $FTk = \sqrt{c^2 - u^2}$ dilates to: $\sqrt{c^2 - u^2} \cosh \rho = c$ (equal to ellipse minor radius b)</p>
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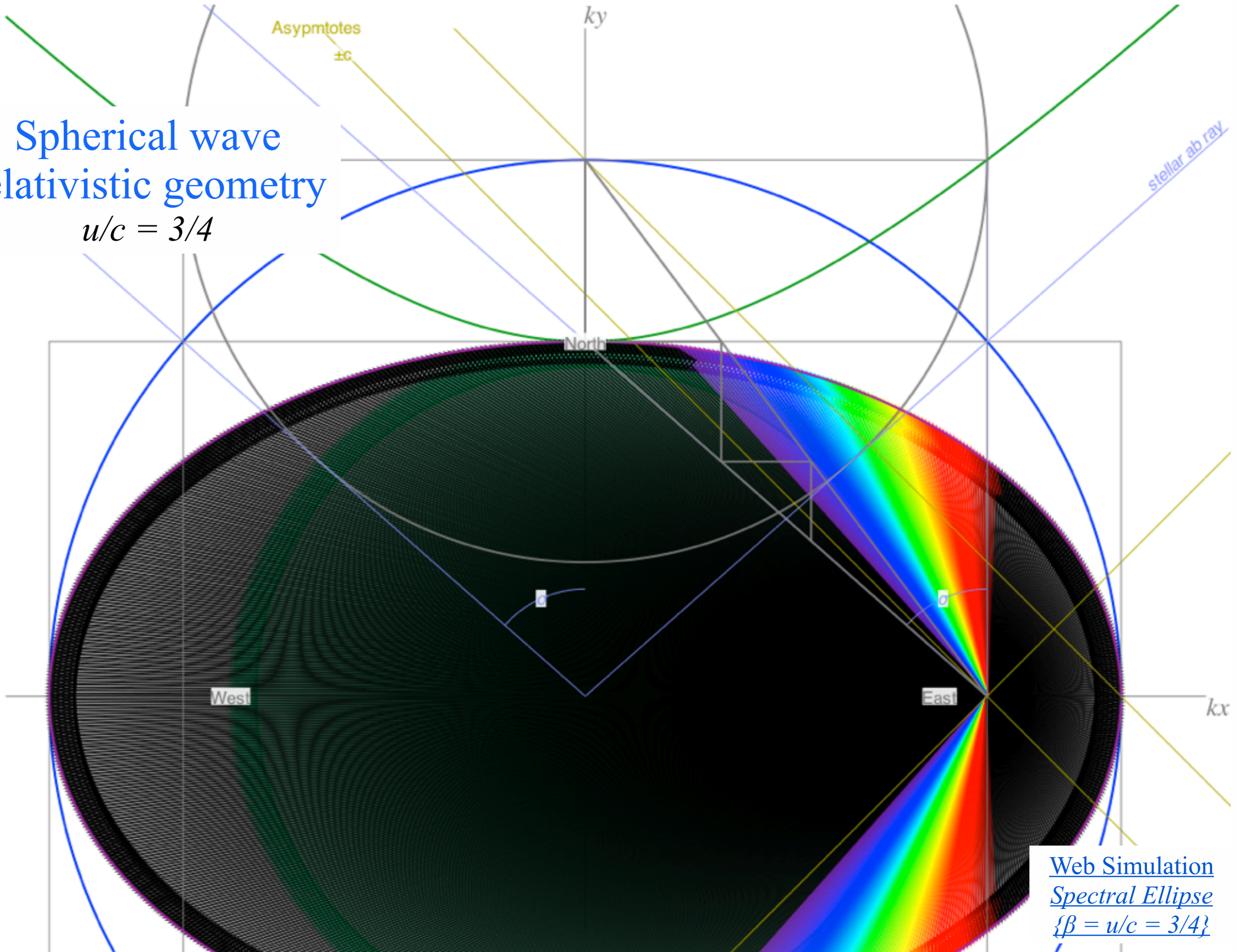
Spherical wave relativistic geometry

$$u/c = 1/3$$



[Web Simulation](#)
[Spectral Ellipse](#)
[{PerSpace-PerSpace}](#)
[{ \$\beta = u/c = 1/3\$ }](#)

Spherical wave
relativistic geometry
 $u/c = 3/4$



Web Simulation
Spectral Ellipse
 $\{\beta = u/c = 3/4\}$

Lecture 31

Thur. 12.10.2015

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Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

“Occams Sword” and summary of 16 parameter functions of ρ and σ

Applications to optical waveguide, spherical waves, and accelerator radiation

Learning about sin! and COS and...

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$$c\kappa_{phase} = B \sinh \rho \approx B \rho \text{ (for } u \ll c \text{)}$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

At low speeds:

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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[RelaWavity Web Simulation - Relativistic Terms](#)
(Expanded Table)

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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← for ($u \ll c$) ⇒

$$K_{phase} \approx \frac{B}{c^2} u$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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$$K_{phase} \approx \frac{B}{c^2} u$$

v_{phase} and K_{phase} resemble formulae for Newton's kinetic energy and momentum

Resembles: $const. + \frac{1}{2} Mu^2$

Resembles: Mu

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
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So attach scale factor h to match units.

Resembles: $const. + \frac{1}{2} Mu^2$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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⇐ for ($u \ll c$) ⇒

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time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
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So attach scale factor h to match units.

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group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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Resembles: Mu

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group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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RelaWavity Web Simulation - Relativistic Terms
(Expanded Table)

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

(old-fashioned notation)

Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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Planck (1900) \uparrow
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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Max Planck
1858-1947

Using (some) wave parameters to develop relativistic quantum theory

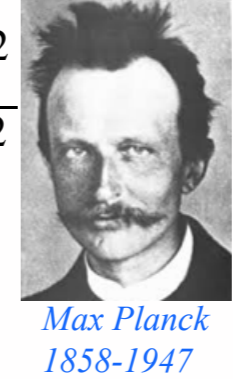
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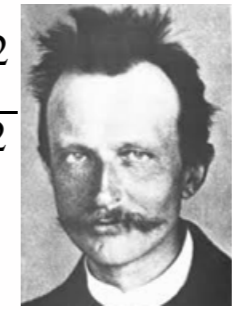
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phase	$b_{BLUE}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
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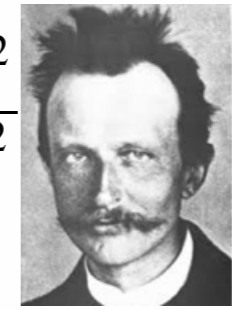
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For more visit the Pirelli Challenge Site
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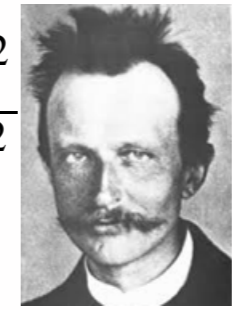
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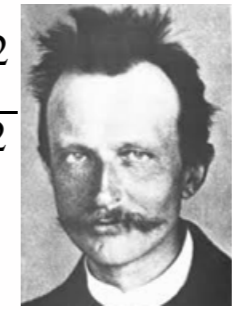
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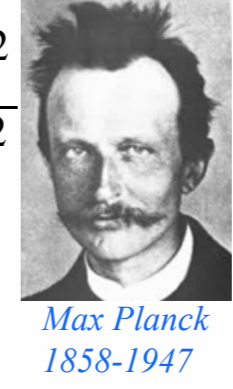
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$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$



$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

⇐ for ($u \ll c$) ⇒

$$K_{phase} \approx \frac{B}{c^2} u$$

v_{phase} and K_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu .

Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

(The famous Mc^2 shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

⇐ for ($u \ll c$) ⇒

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor h (or hN) to match units.

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2$$

⇐ for ($u \ll c$) ⇒

$$hK_{phase} \approx Mu$$

Lucky coincidences?? Cheap trick??
...Try exact v_{phase} and K_{phase} ...

Need to replace h with hN to match e.m. energy density $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hN v_{phase}$

$$hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

↑ Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

↑ Einstein (1905)

$$hK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$

$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}} \quad (\text{old-fashioned notation})$$

$$cp = \frac{Mc u}{\sqrt{1-u^2/c^2}}$$

Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

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for $(u \ll c) \Rightarrow$

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~~Natural wave conspiracy~~
~~Lucky coincidences??~~ ~~Expensive Cheap trick??~~
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group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	
phase	$b_{BLUE}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}} \quad (\text{old-fashioned notation})$$

$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

Momentum: $hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$

DeBroglie (1921)

Using (some) wave parameters to develop relativistic quantum theory

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

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$$hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

↑ Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

↑ Einstein (1905)

This motivates the "particle" normalization $\int \Psi^* \Psi dV = N$ $\Psi = \sqrt{\frac{\epsilon_0}{h\nu}} E$

$$h c K_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
phase	$b_{BLUE}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$		
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$		
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$

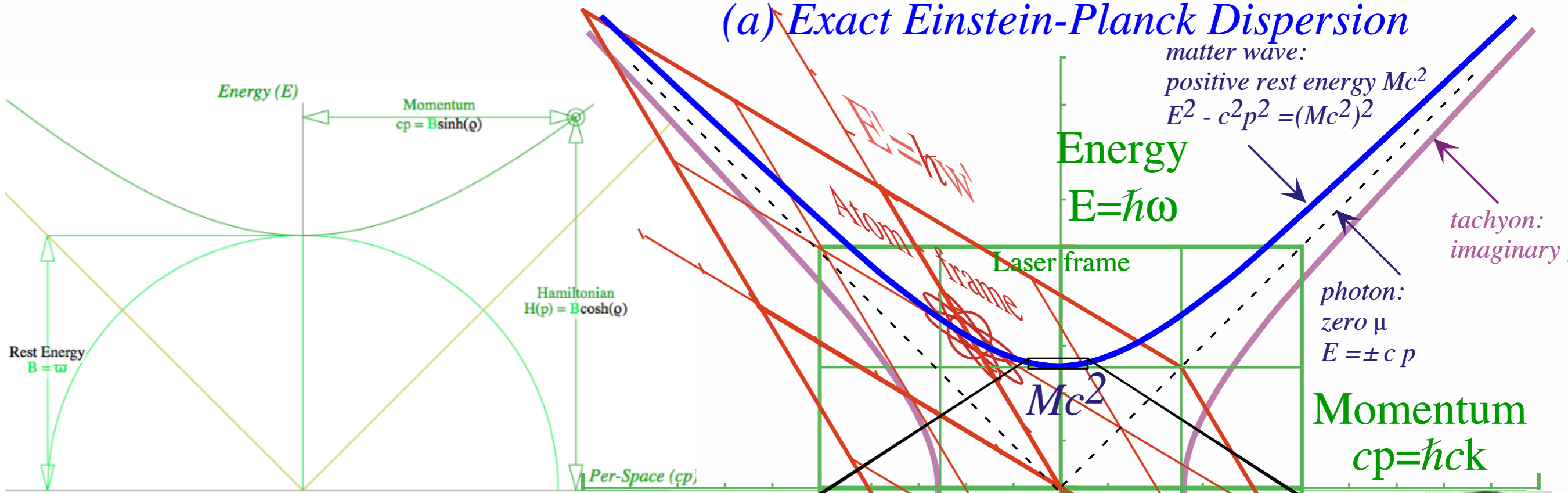
$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}} \quad (\text{old-fashioned notation})$$

$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

Momentum: $h c K_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$

DeBroglie (1921)

Using (some) wave coordinates for relativistic quantum theory



Mass (resting)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Energy

$$h\nu_{\text{phase}} = E = h\nu_A \cosh \rho$$

Momentum

$$hc\kappa_{\text{phase}} = cp = hc\kappa_A \sinh \rho = h\nu_A \sinh \rho$$

Energy versus Momentum

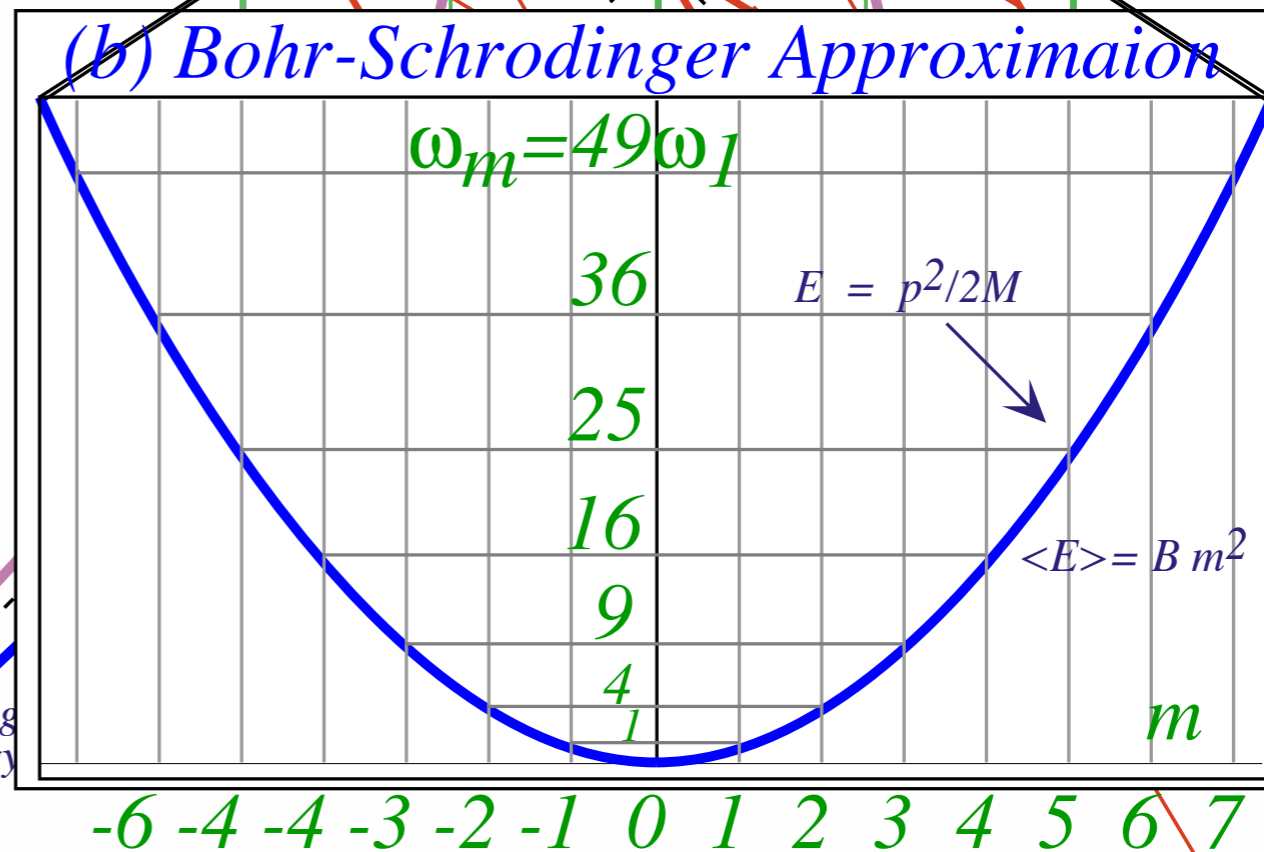
$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$



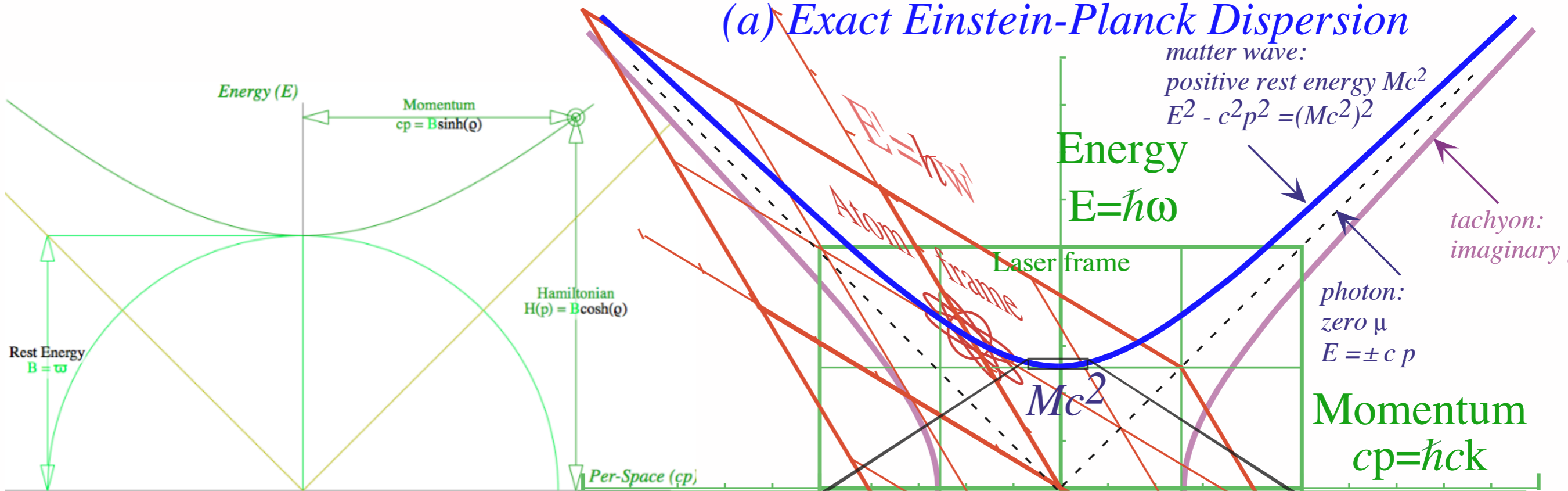
Niels Bohr
1885-1962

(b) Bohr-Schrodinger Approximaion



Erwin Schrodinger
1887-1961

Using (some) wave coordinates for relativistic quantum theory



Mass (resting)

$$hB = \hbar\omega_A = Mc^2 = \hbar ck_A$$

Energy

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

Momentum

$$\hbar ck_{phase} = cp = \hbar ck_A \sinh \rho = \hbar\omega_A \sinh \rho$$

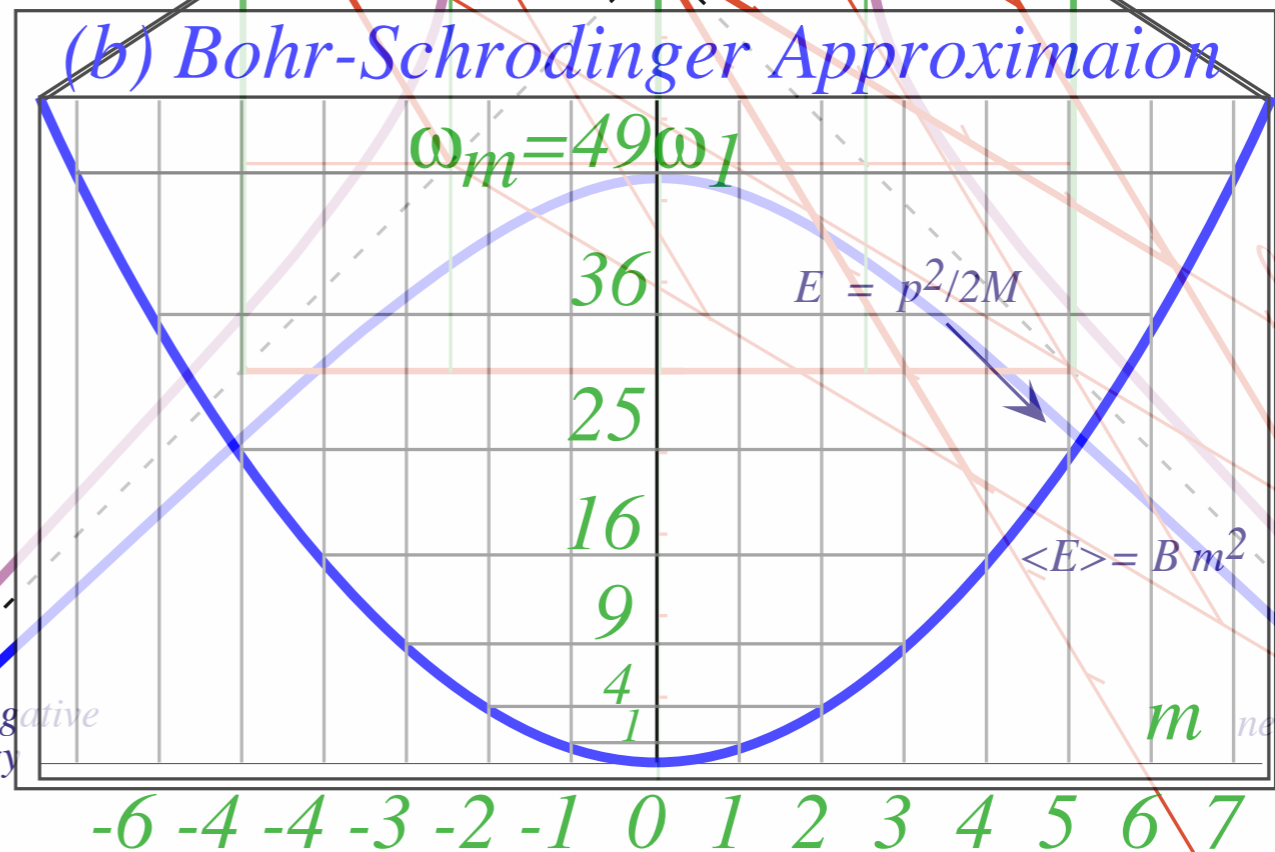
Energy versus Momentum

$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$

low speed approximation

(b) Bohr-Schrodinger Approximaion



Relativity variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\operatorname{csc} \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>effects</i>	$b_{RED}^{Doppler}$	V_{group}	<i>past-future asymmetry</i> (off-diagonal Lorentz-transform)	<i>x-contraction</i> ^(Lorentz) τ_{phase} -contraction	<i>t-dilation</i> ^(Einstein) v_{phase} -dilation (on-diagonal Lorentz-transform)	<i>inverse asymmetry</i>	V_{phase}	$b_{BLUE}^{Doppler}$

Relativistic quantum mechanics variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\operatorname{csc} \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{\beta}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{\beta}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>functions</i>		$V_{group} = c \tanh \rho$	<i>momentum</i> $cp = Mc^2 \sinh \rho$	<i>-Lagrangian</i> $L = -Mc^2 \operatorname{sech} \rho$	<i>Hamiltonian</i> $H = Mc^2 \cosh \rho$	<i>DeBroglie</i> $\lambda = \alpha \operatorname{csch} \rho$	$V_{phase} = c \coth \rho$	

Lecture 31

Thur. 12.10.2015

Review: Relativity ρ functions Two famous ones Extremes and plot vs. ρ
Doppler jeopardy Geometric mean and Relativistic hyperbolas
Animation of $e^{\rho}=2$ spacetime and per-spacetime plots

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity

Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

“Occams Sword” and summary of 16 parameter functions of ρ and σ

Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics

➔ What's the matter with mass? Shining some light on the Elephant in the room
Relativistic action and Lagrangian-Hamiltonian relations
Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization “photon” number N and 1st-quantization wavenumber κ

Relativity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes g -acceleration grid

Analysis of constant- g grid compared to zero- g Minkowski grid

Animation of mechanics and metrology of constant- g grid

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$$= h\nu_{\text{phase}}$$

momentum: $cp = Mc^2 \sinh \rho$

$$= hc\kappa_{\text{phase}}$$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

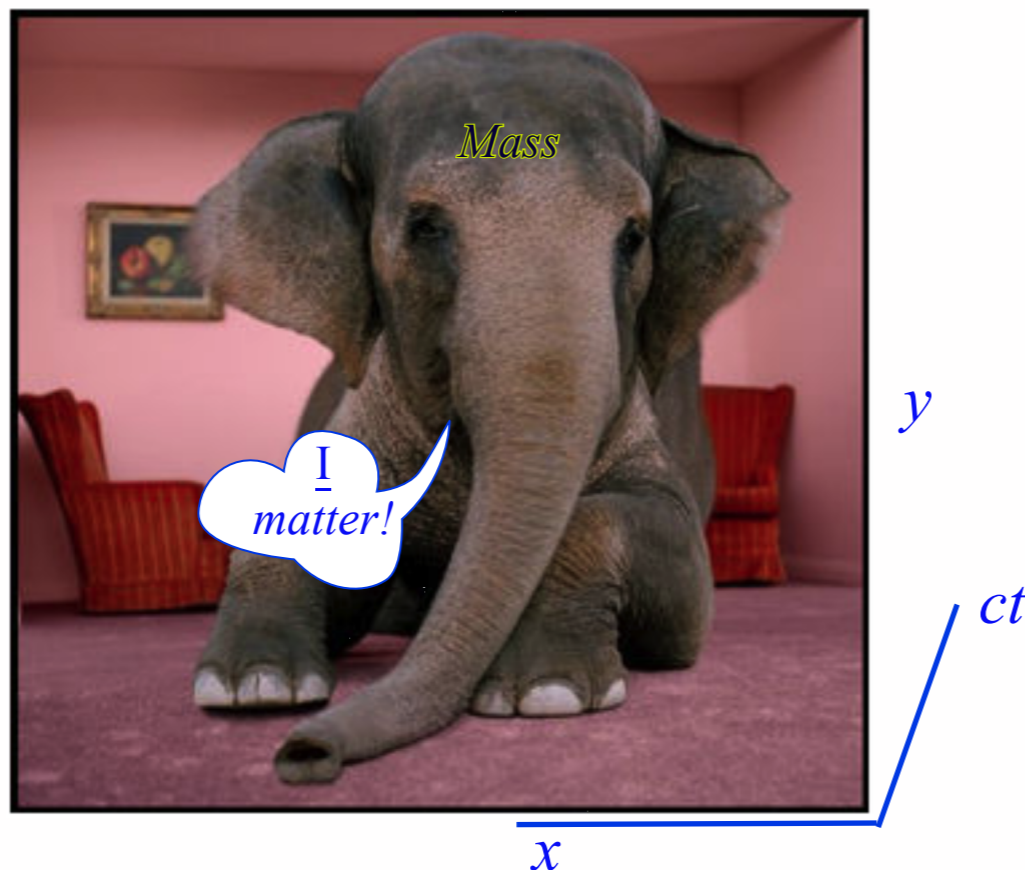
Rest Mass M_{rest} (Einstein's mass)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

- *What's the matter with Mass?*



Shining some light on the elephant in the spacetime room

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$$= h\nu_{phase}$$

momentum: $cp = Mc^2 \sinh \rho$

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$$\frac{h\nu_{phase}}{c^2} = M_{rest}$$

Rest
Mass

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$$= h\nu_{phase}$$

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Defines invariant hyperbola(s)

momentum: $cp = Mc^2 \sinh \rho$

$$h\mathbf{B} = h\mathbf{v}_A = Mc^2 = hc\mathbf{K}_A$$

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

$$= hc\mathbf{K}_{phase}$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} \quad \text{Rest Mass}$$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\mathbf{K}}$

Momentum Mass M_{mom} (Galileo's mass) Defined by ratio p/u of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Definition(s) of mass for relativity/quantum

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$$= hc\kappa_{phase}$$

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$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}} \quad \text{Momentum Mass}$$

Definition(s) of mass for relativity/quantum

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$= h\nu_{phase}$

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$$\frac{h\nu_{phase}}{c^2} = M_{rest} \quad \begin{matrix} \text{Rest} \\ \text{Mass} \end{matrix}$$

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$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Limiting cases:

$$M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho} / 2$$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}} \quad \begin{matrix} \text{Momentum} \\ \text{Mass} \end{matrix}$$

Definition(s) of mass for relativity/quantum

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$$\frac{h\nu_{phase}}{c^2} = M_{rest} \quad \text{Rest Mass}$$

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$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}} \quad \text{Momentum Mass}$$

Effective Mass M_{eff} (Newton's mass) Defined by ratio $F/a = dp/du$ of relativistic force to acceleration.

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$= h\nu_{phase}$

momentum: $cp = Mc^2 \sinh \rho$

$= hc\kappa_{phase}$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

Rest Mass M_{rest} (Einstein's mass)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} \quad \text{Rest Mass}$$

Momentum Mass M_{mom} (Galileo's mass) Defined by ratio p/u of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Limiting cases:

$$M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho/2}$$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \quad \text{Momentum Mass}$$

Effective Mass M_{eff} (Newton's mass) Defined by ratio $F/a = dp/du$ of relativistic force to acceleration.

That is ratio of change $dp = Mc \cosh \rho d\rho$ in momentum to change $du = c \operatorname{sech}^2 \rho d\rho$ in velocity

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$= h\nu_{phase}$

momentum: $cp = Mc^2 \sinh \rho$

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velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

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Limiting cases: $M_{eff} \xrightarrow{u \rightarrow c} M_{rest} e^{3\rho/2}$

$M_{eff} \xrightarrow{u \ll c} M_{rest}$

More common derivation using group velocity: $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\nu}{d\kappa}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}}$$

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

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general wave formula

to accompany $V_{group} = \frac{d\omega}{dk}$

Definition(s) of mass for relativity/quantum

Rest Mass M_{rest} (Einstein's mass)

$$h\mathbf{B} = h\mathbf{v}_A = Mc^2 = h\mathbf{c}\mathbf{K}_A$$

$$\frac{h\omega_{phase}}{c^2} = M_{rest} = \frac{h\mathbf{c}\mathbf{K}_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

Momentum Mass M_{mom} (Galileo's mass) Defined by p/u

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \quad \frac{\text{Momentum}}{\text{Mass}}$$

Effective Mass M_{eff} (Newton's mass) Defined by $F/a = dp/du$

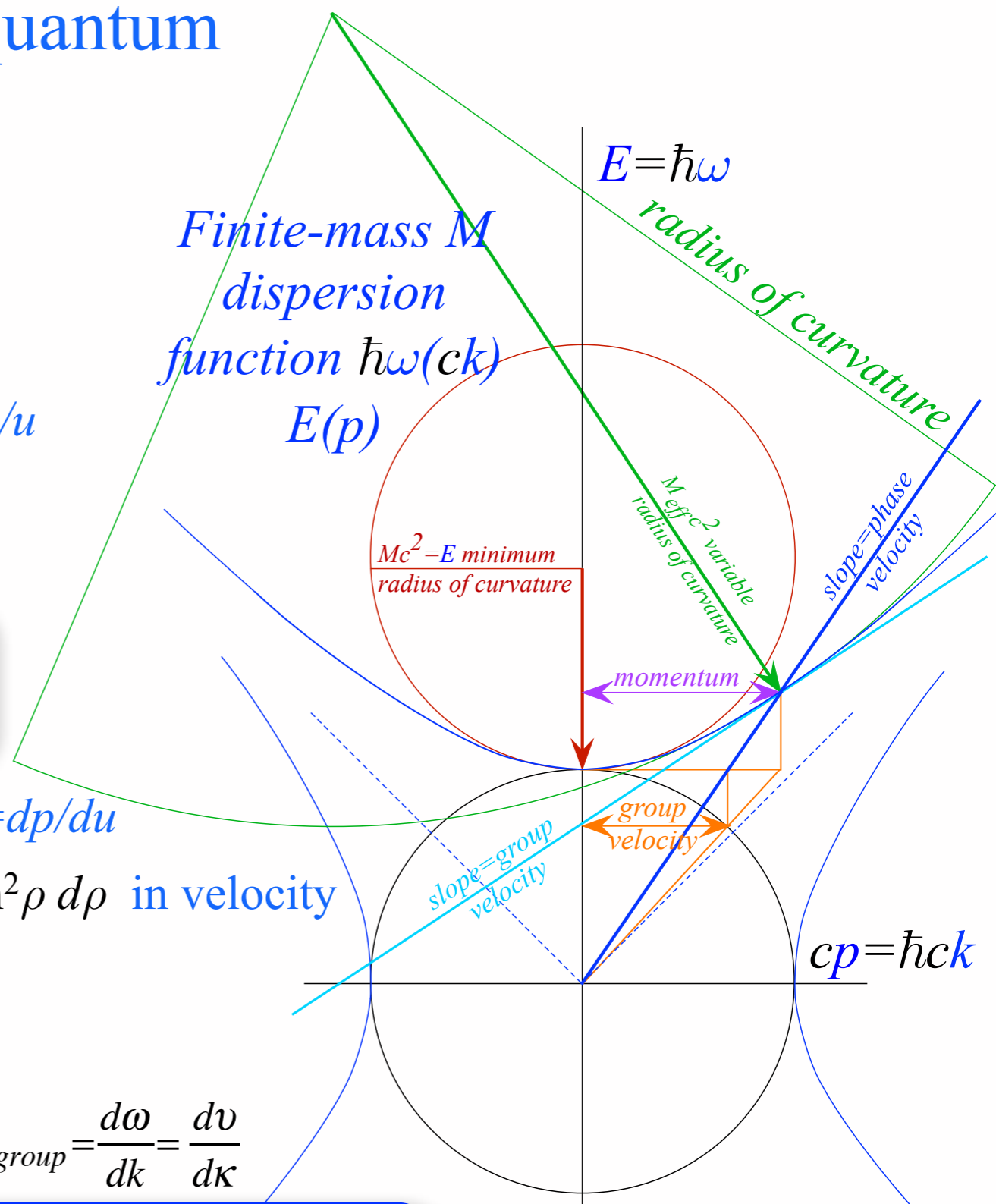
That is ratio of $dp = Mc \cosh \rho d\rho$ to change $du = c \operatorname{sech}^2 \rho d\rho$ in velocity

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho \quad \frac{\text{Effective Mass}}$$

More common derivation using group velocity: $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{dv}{d\mathbf{K}}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2\omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}} = M_{rest} \cosh^3 \rho \quad \frac{\text{Effective Mass}}$$

general wave formula to accompany $V_{group} = \frac{d\omega}{dk}$



Effective mass is proportional to the radius of curvature of $\omega(k)$ dispersion.

Definition(s) of mass for relativity/quantum

How much mass does a γ -photon have?

Rest Mass (a) γ -rest mass: $M_{rest}^{\gamma} = 0$,

Momentum Mass (b) γ -momentum mass: $M_{mom}^{\gamma} = \frac{p}{c} = \frac{h\kappa}{c} = \frac{h\nu}{c^2}$,

Effective Mass (c) γ -effective mass: $M_{eff}^{\gamma} = \infty$.

Newton complained about his “corpuscles” of light having “fits” (going crazy).

(All this would be evidence of *triple Schizophrenia*.)

$$M_{mom}^{\gamma} = \frac{h\nu}{c^2} = \nu(1.2 \cdot 10^{-51}) \text{kg} \cdot \text{s} = 4.5 \cdot 10^{-36} \text{kg} \quad (\text{for: } \nu=600\text{THz})$$

Lecture 31

Thur. 12.10.2015

Review: Relativity ρ functions Two famous ones Extremes and plot vs. ρ
Doppler jeopardy Geometric mean and Relativistic hyperbolas
Animation of $e^{\rho}=2$ spacetime and per-spacetime plots

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity

Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

“Occams Sword” and summary of 16 parameter functions of ρ and σ

Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

➔ **Relativistic action and Lagrangian-Hamiltonian relations**

Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization “photon” number N and 1st-quantization wavenumber κ

Relativity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes g -acceleration grid

Analysis of constant- g grid compared to zero- g Minkowski grid

Animation of mechanics and metrology of constant- g grid

Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$$

$$\hbar \equiv \frac{h}{2\pi}$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz
format

angular phasor
format →

$$\begin{aligned} \hbar\omega_A &= Mc^2 = \hbar ck_A \\ \hbar\omega_{phase} &= E = \hbar\omega_A \cosh \rho \\ \hbar ck_{phase} &= cp = \hbar\omega_A \sinh \rho \end{aligned}$$

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Prior wave relations

← linear Hz format angular phasor →
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$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

$$p = \hbar k = Mc \sinh \rho$$

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Legendre transformation

Use *Group velocity* : $u = \frac{dx}{dt} = c \tanh \rho$

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$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

$$L \text{ is : } Mc^2 \frac{-1}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

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$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

Note: $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Compare *Lagrangian* L

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

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Prior wave relations

← linear Hz format angular phasor →
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Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p = \hbar k$ relation and Planck-energy $E = \hbar \omega$ relation to define *Hamiltonian H = E*

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity: $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

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Note: $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Compare *Lagrangian L*

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

with *Hamiltonian H = E*

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c k_A$$

Prior wave relations

← linear Hz
format

angular phasor
format →

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c k_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c k_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c k_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \equiv \frac{h}{2\pi}$$

Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi=kx-\omega t=k'x'-\omega't'$ for wave of $k=k_{phase}$ and $\omega=\omega_{phase}$.

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$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar\omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity : $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar\omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

Note: $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Also: $cp = Mc^2 \sinh \rho$

Compare Lagrangian L

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

with Hamiltonian $H=E$

$$H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

$$\hbar\omega_A = Mc^2 = \hbar c k_A$$

Prior wave relations $\hbar = h/2\pi$

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

← linear Hz format

angular phasor format →

$$\hbar c k_{phase} = cp = \hbar\omega_A \sinh \rho$$

$$\hbar\omega_A = Mc^2 = \hbar c k_A$$

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$$\hbar \equiv \frac{h}{2\pi}$$

Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase $\Phi=kx-\omega t=k'x'-\omega't'$ for wave of $k=k_{phase}$ and $\omega=\omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define *Hamiltonian H=E*

$$\frac{dS}{dt} = L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar\omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity : $u = \frac{dx}{dt} = c \tanh \rho = c \sin \sigma$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar\omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Note: $Mc u = Mc^2 \tanh \rho = Mc^2 \sin \sigma$

Also: $cp = Mc^2 \sinh \rho$

$= \hbar ck = Mc^2 \tan \sigma$

Compare Lagrangian L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with Hamiltonian H=E

$$H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Including stellar angle σ

Define Action $S = \hbar\Phi$

$$\hbar\omega_A = Mc^2 = \hbar ck_A$$

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

$$\hbar ck_{phase} = cp = \hbar\omega_A \sinh \rho$$

Prior wave relations

← linear Hz format

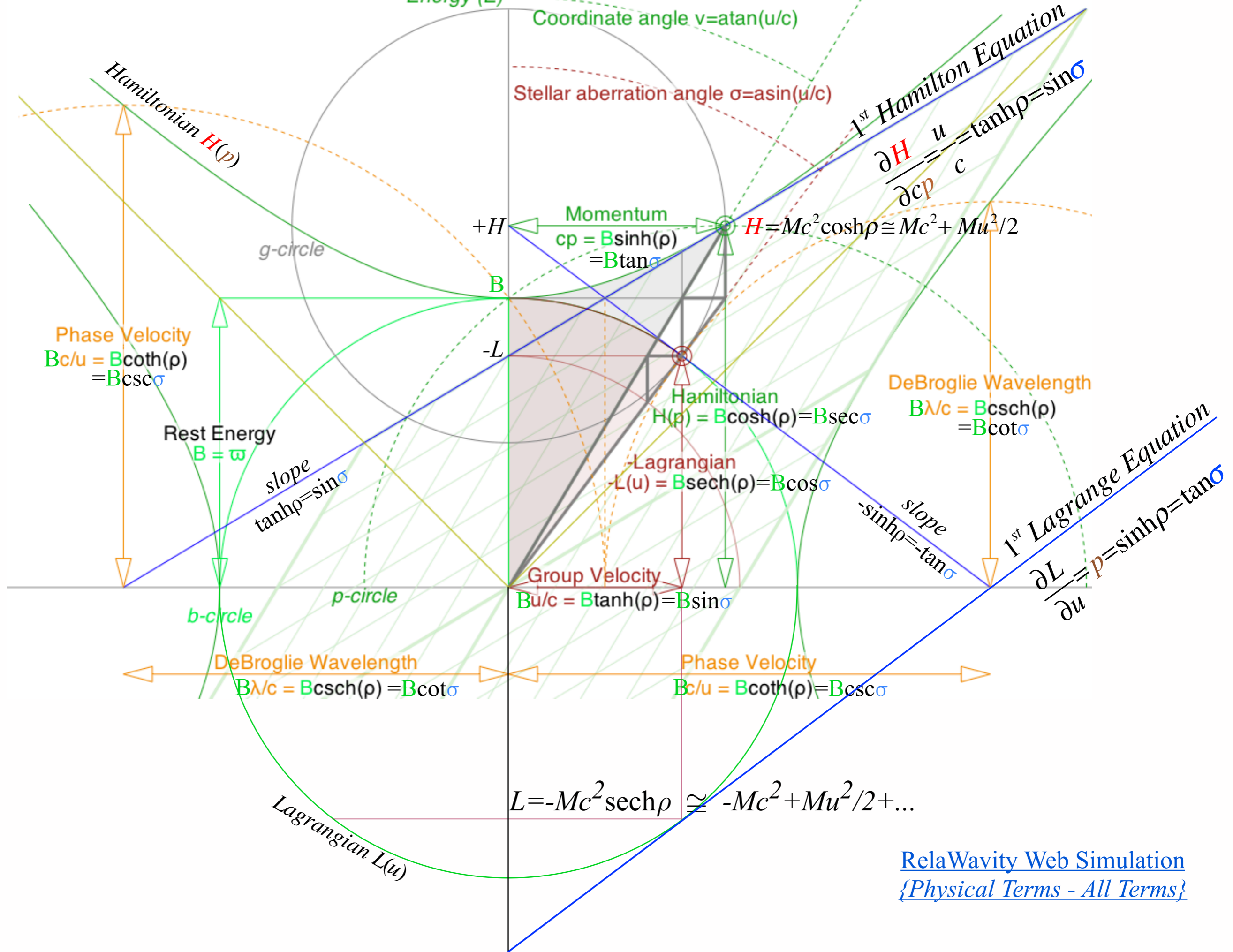
angular phasor → format

$$\hbar\omega_A = Mc^2 = \hbar ck_A$$

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Define *Action S = \hbar \Phi*

$$\begin{aligned} \hbar \nu_A &= Mc^2 = \hbar c \kappa_A \\ \hbar \nu_{phase} &= E = \hbar \nu_A \cosh \rho \\ \hbar c \kappa_{phase} &= cp = \hbar \nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format angular phasor →
format format

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$$dS \equiv Ldt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt \quad \text{Poincare Invariant action differential}$$

Compare *Lagrangian L*

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$$dS \equiv L dt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt$$

Poincare Invariant action differential

$$\frac{\partial S}{\partial x} = p \quad \frac{\partial S}{\partial t} = -H$$

Hamilton-Jacobi equations

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

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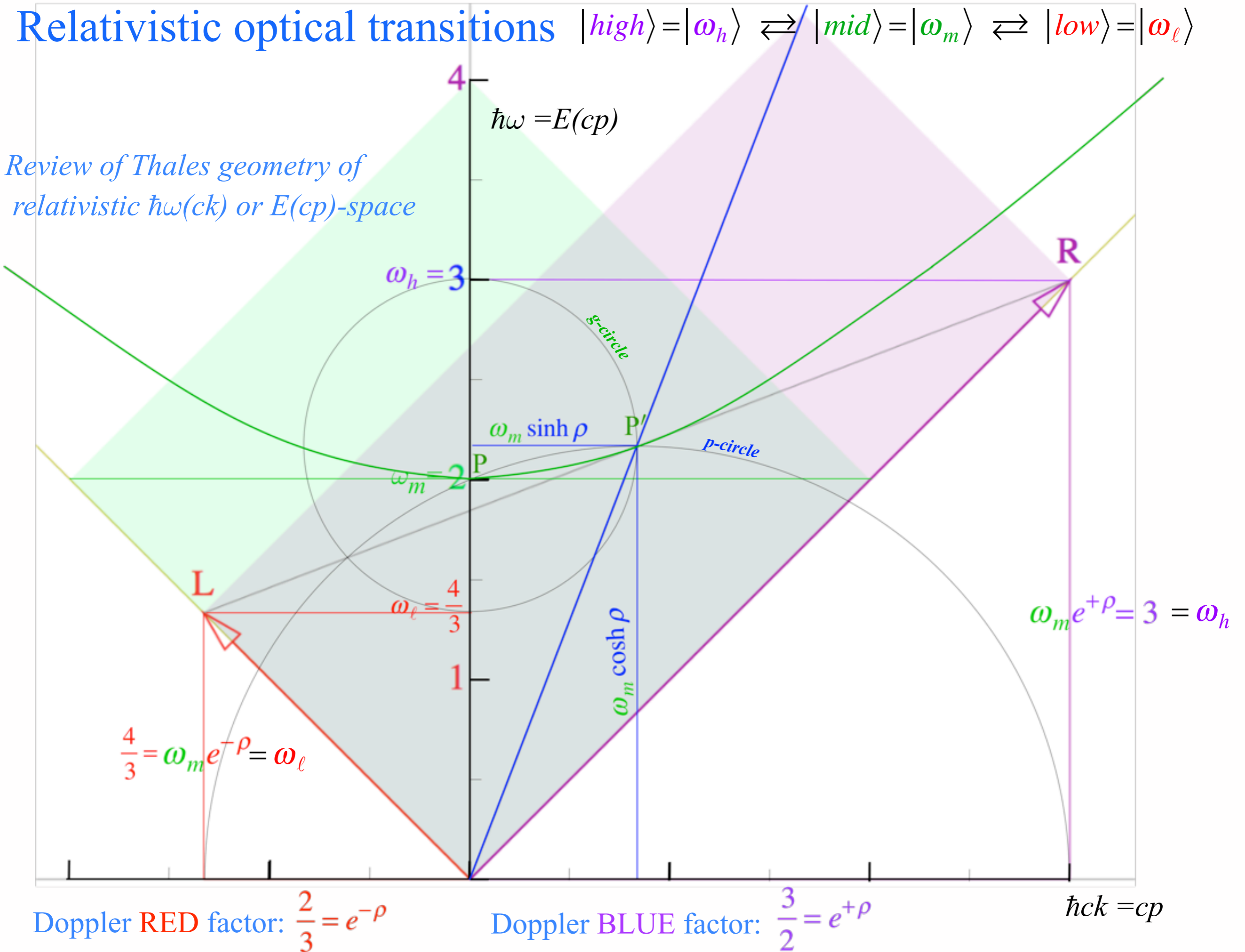
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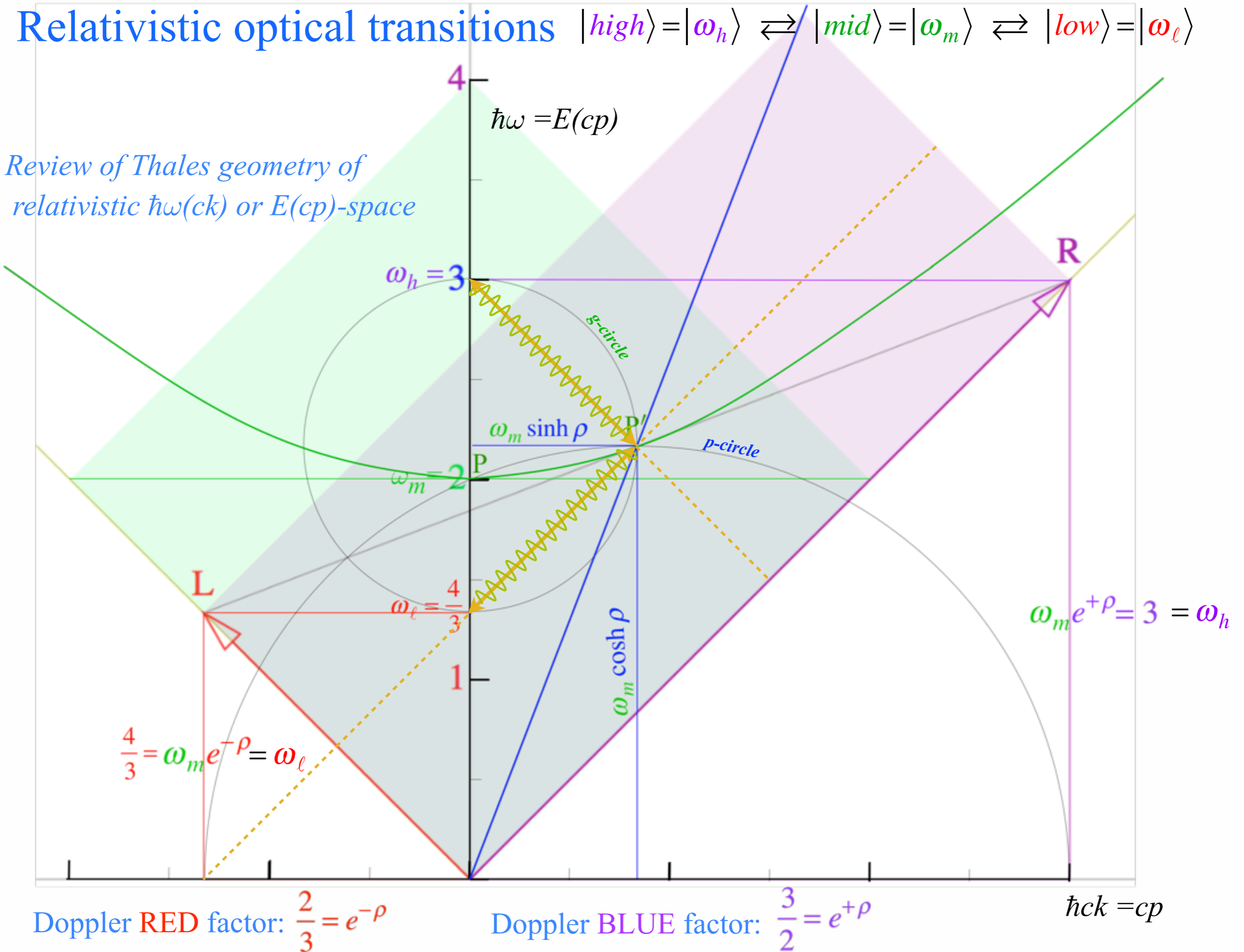
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Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



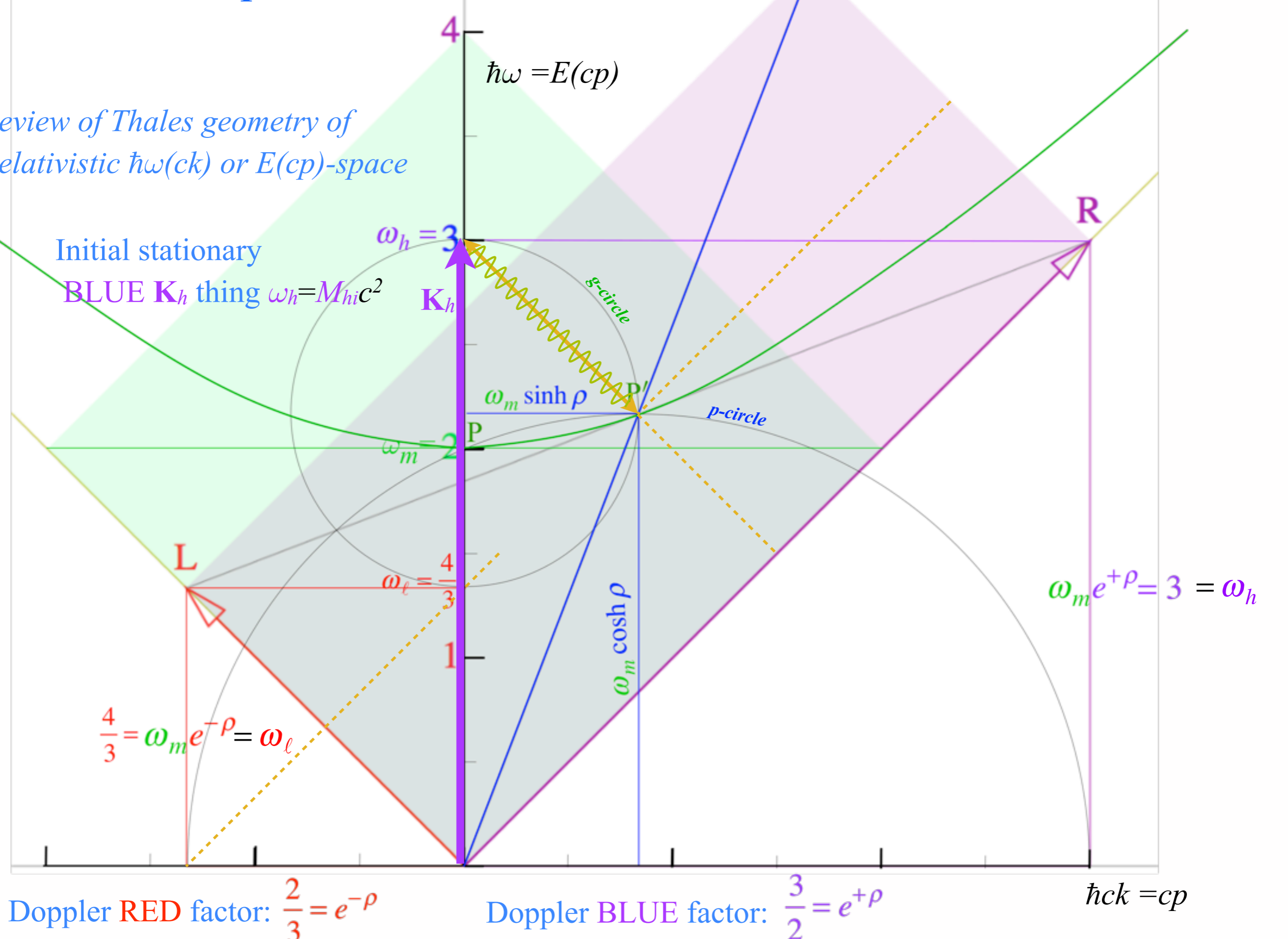
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Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

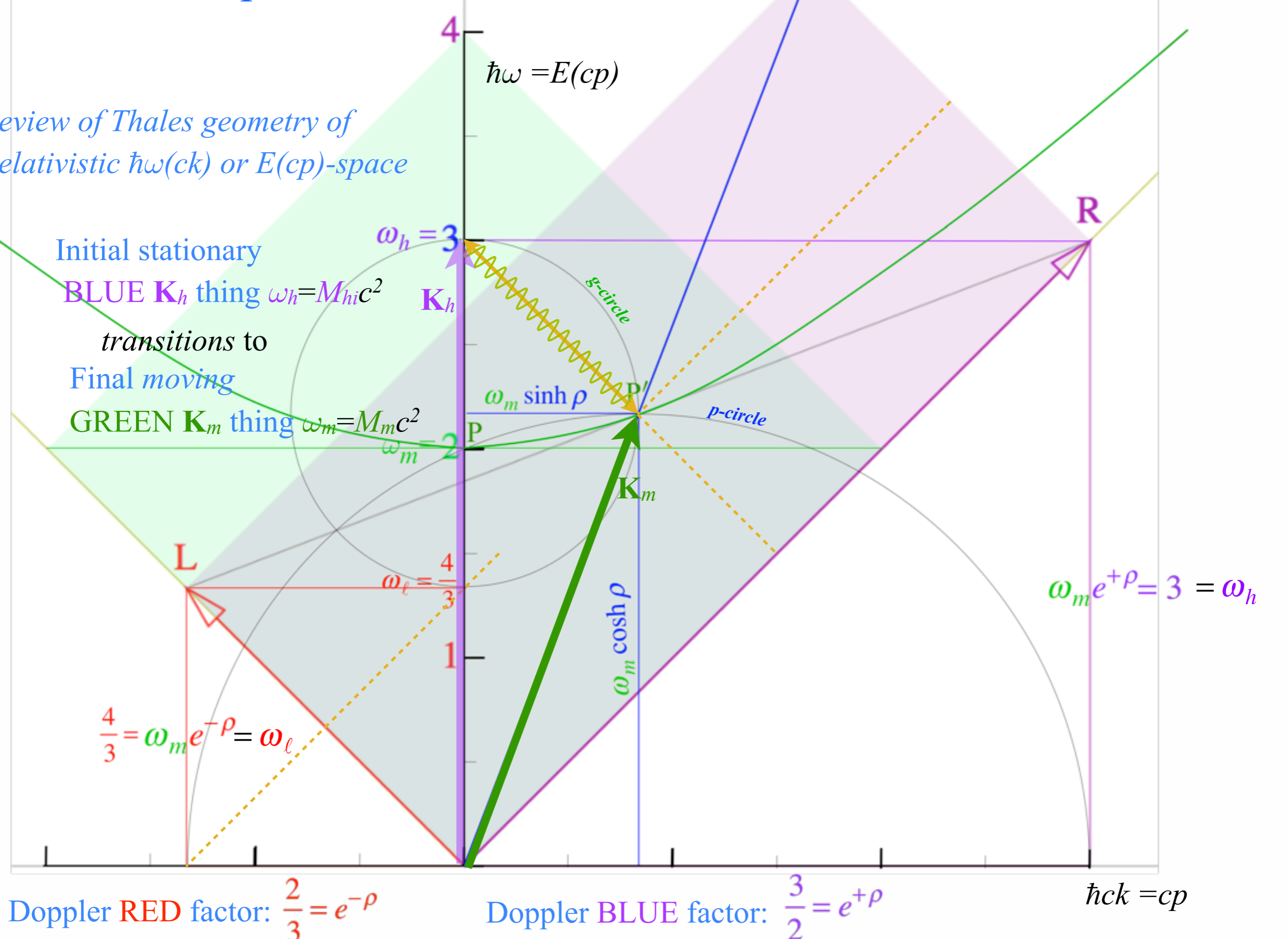
Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

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Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary
BLUE K_h thing $\omega_h = M_h c^2$
 transitions to
 Final moving
GREEN K_m thing $\omega_m = M_m c^2$



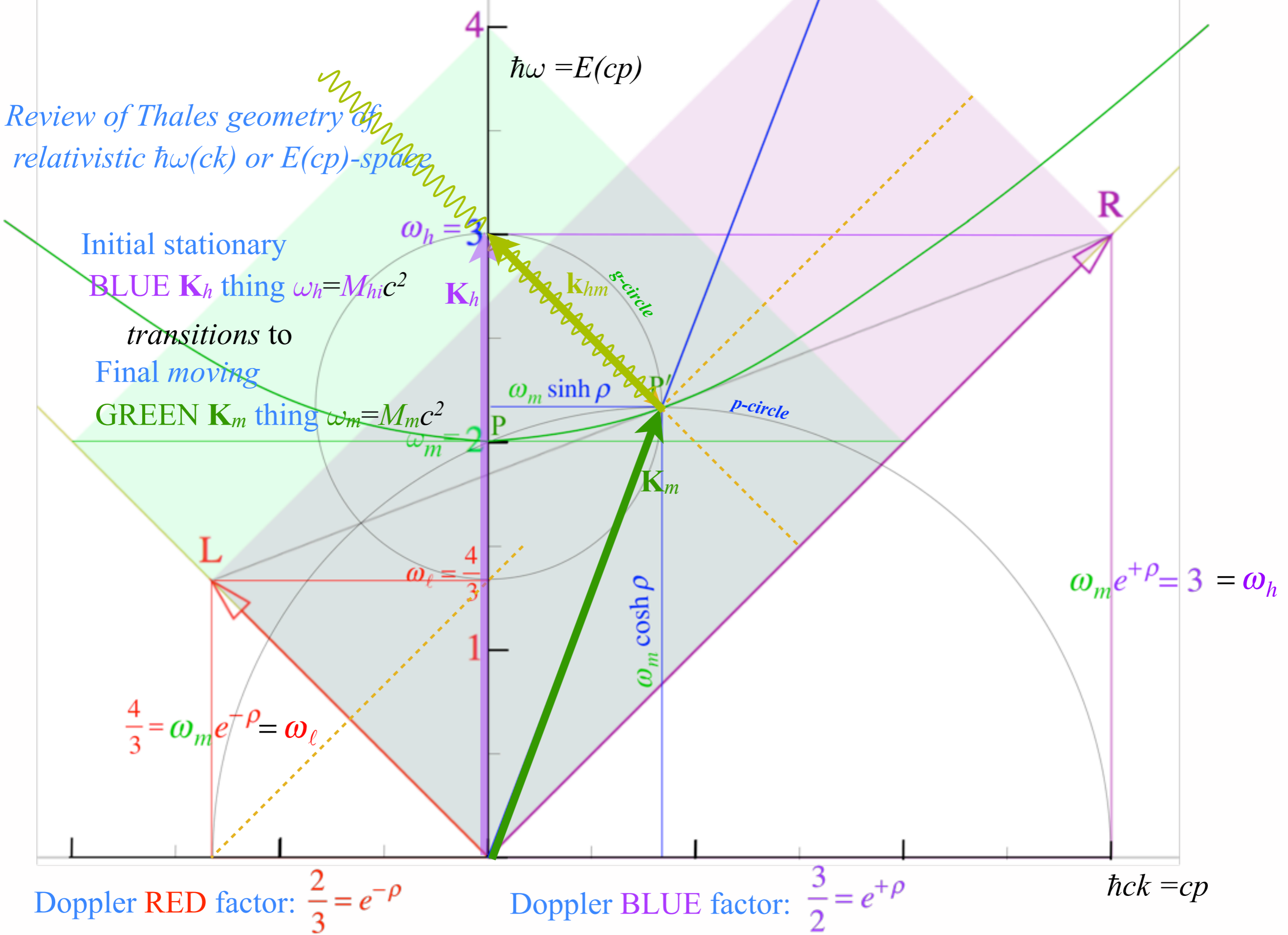
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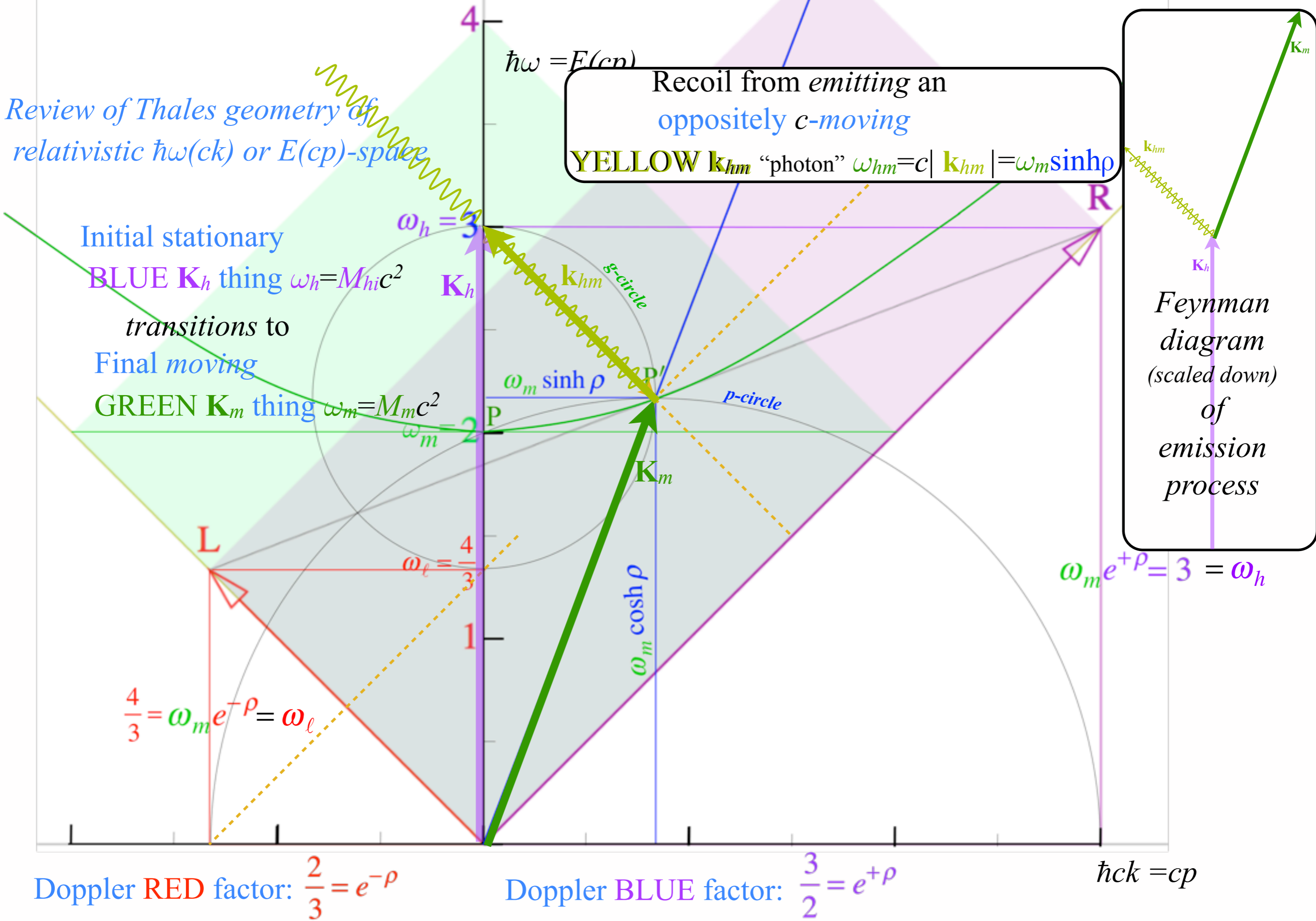
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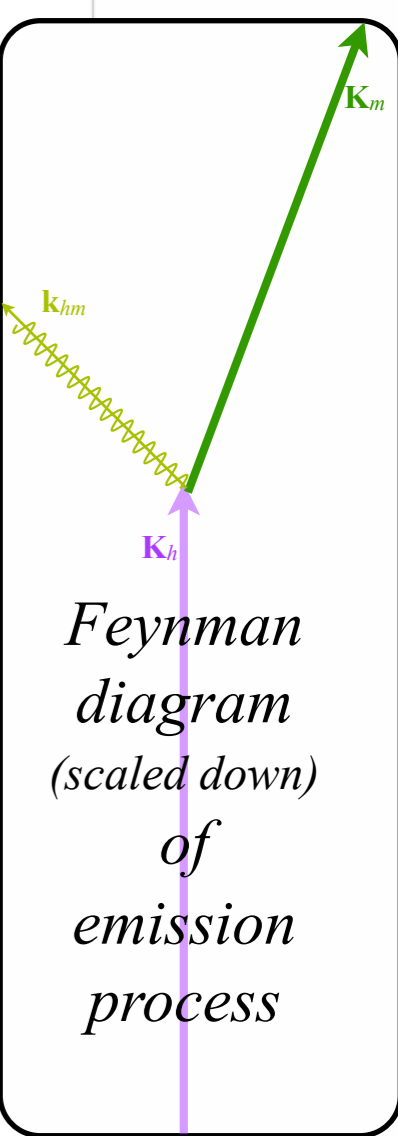


Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary BLUE K_h thing $\omega_h = M_h c^2$
 transitions to Final moving GREEN K_m thing $\omega_m = M_m c^2$

Recoil from emitting an oppositely c -moving YELLOW K_{hm} "photon" $\omega_{hm} = c |k_{hm}| = \omega_m \sinh \rho$



Take-away point 0
 Classical (and spectroscopic) Energy-momentum conservation is due to conservation in quantum-phase space-time "wiggle-count"

Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

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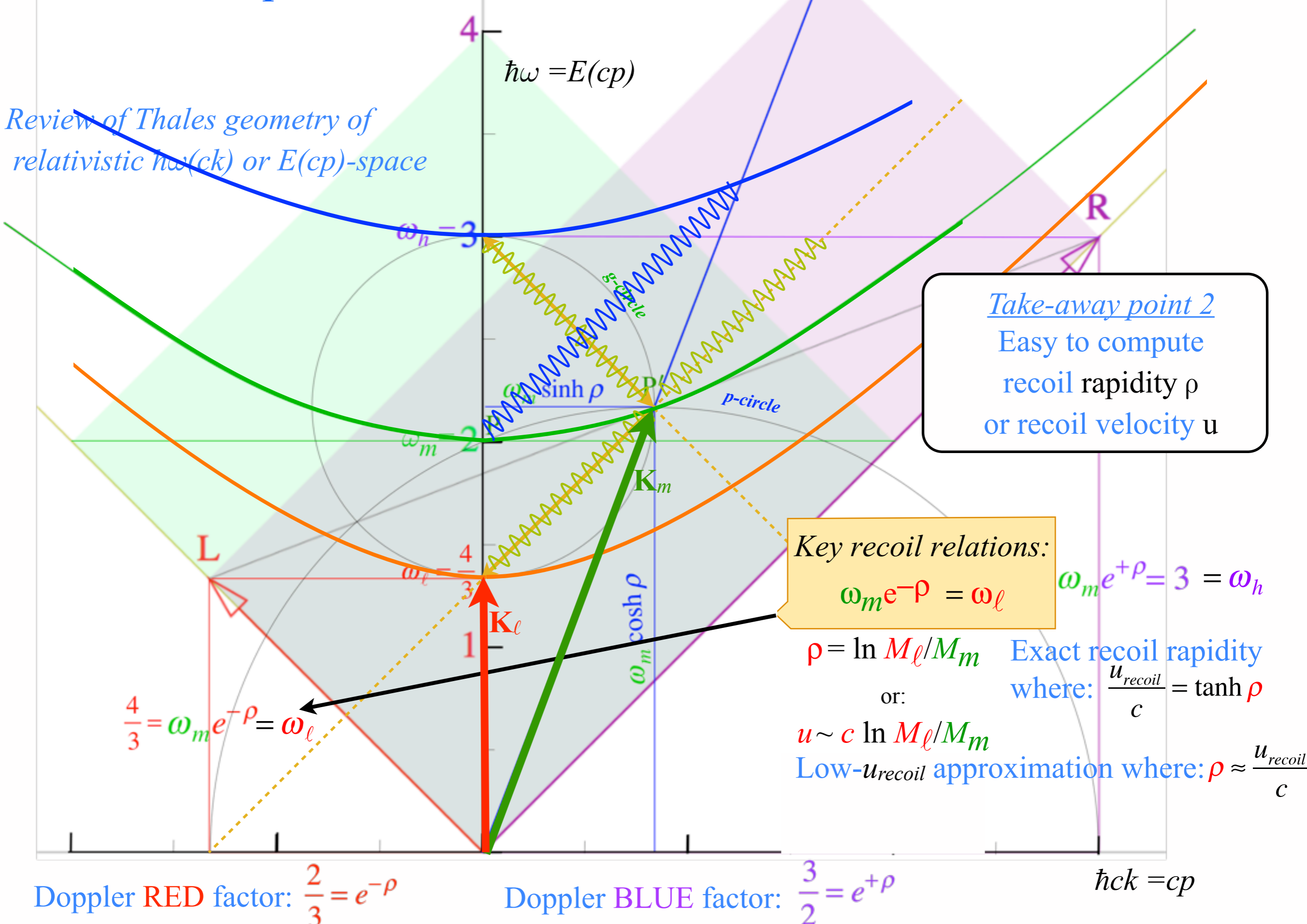
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Take-away point 2
Easy to compute recoil rapidity ρ or recoil velocity u

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_l$
 $\omega_m e^{+\rho} = 3 = \omega_h$

Exact recoil rapidity where: $\frac{u_{recoil}}{c} = \tanh \rho$

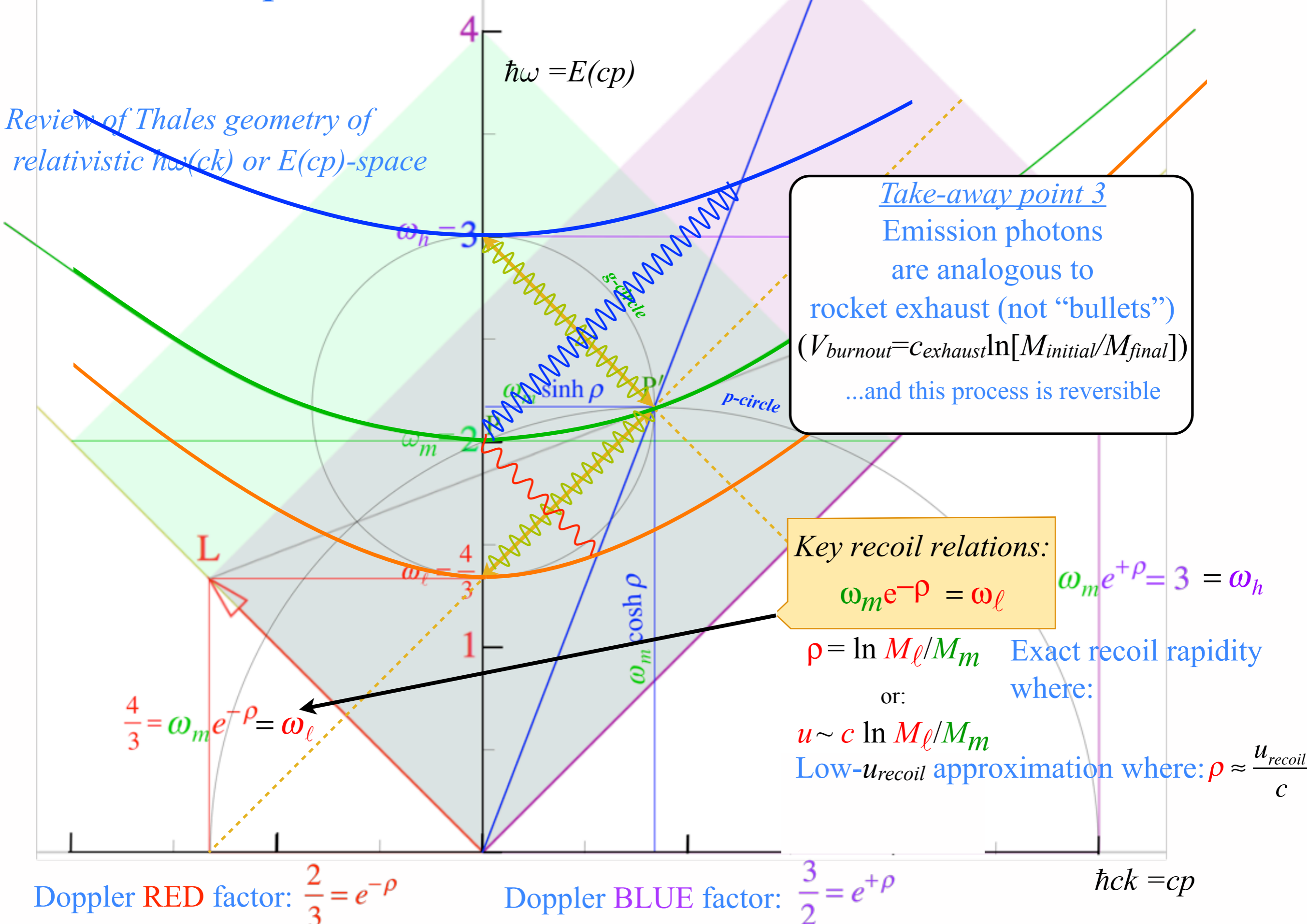
or: $u \sim c \ln M_l/M_m$
Low- u_{recoil} approximation where: $\rho \approx \frac{u_{recoil}}{c}$

Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_\ell\rangle$



Review of Thales geometry of relativistic $\hbar\omega(cp)$ or $E(cp)$ -space

Take-away point 3
 Emission photons are analogous to rocket exhaust (not "bullets")
 ($V_{burnout} = c_{exhaust} \ln[M_{initial}/M_{final}]$)
 ...and this process is reversible

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_\ell$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_\ell / M_m$ Exact recoil rapidity where:

or:
 $u \sim c \ln M_\ell / M_m$

Low- u_{recoil} approximation where: $\rho \approx \frac{u_{recoil}}{c}$

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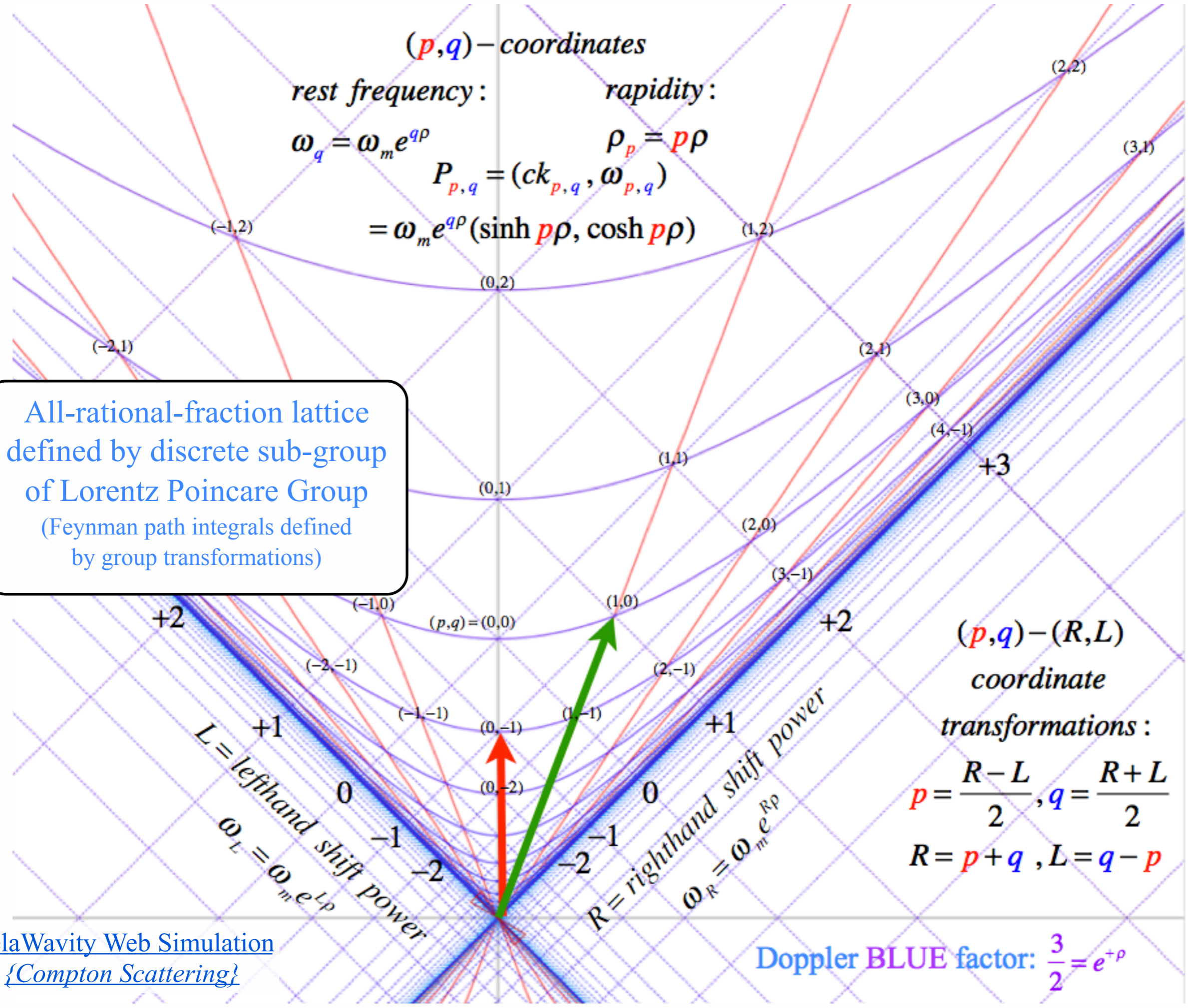
(p,q) -coordinates

rest frequency: rapidity:

$$\omega_q = \omega_m e^{qp} \qquad \rho_p = p\rho$$

$$P_{p,q} = (ck_{p,q}, \omega_{p,q})$$

$$= \omega_m e^{qp} (\sinh p\rho, \cosh p\rho)$$



All-rational-fraction lattice
 defined by discrete sub-group
 of Lorentz Poincare Group
 (Feynman path integrals defined
 by group transformations)

$(p,q) - (R,L)$
coordinate

transformations:

$$p = \frac{R-L}{2}, \quad q = \frac{R+L}{2}$$

$$R = p+q, \quad L = q-p$$

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2nd Quantization:

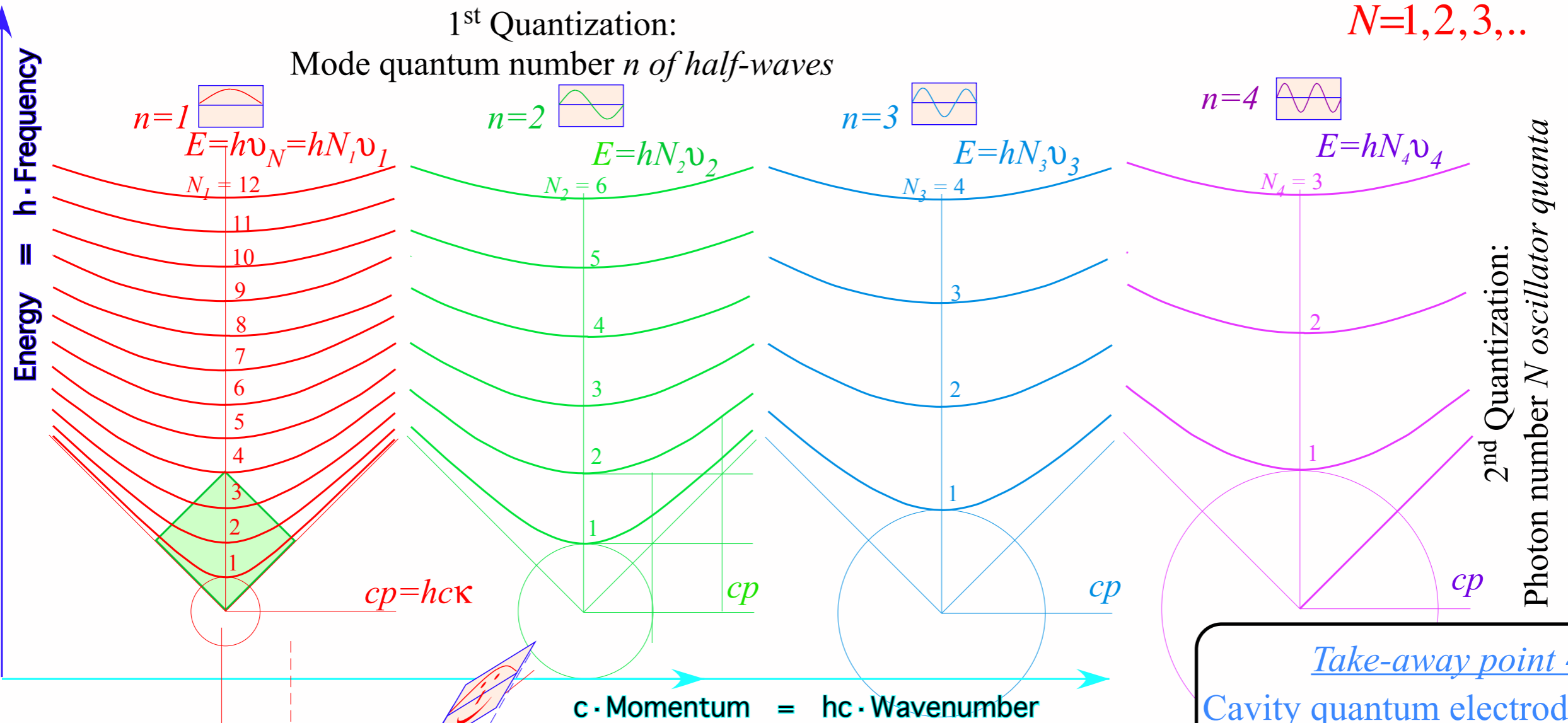
$h\nu$ is actually $hN\nu$

($h\nu_{phase}=E=h\nu_A \cosh \rho$) is actually ($hN\nu_{phase}=E_N=hN\nu_A \cosh \rho$ with quantum numbers)

$N=1,2,3,..$

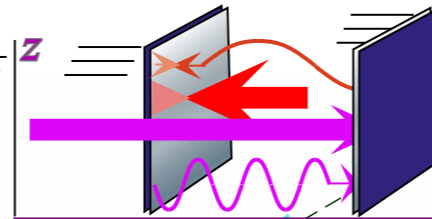
1st Quantization:

Mode quantum number n of half-waves

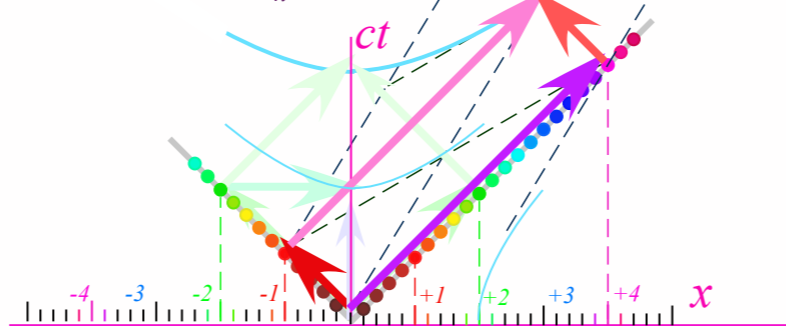


Boosted wave mode

Boosted cavity wave has invariant mode number n photon number N_n



Lorentz contracted cavity length $L=3.2$
Proper length $l=4.0$



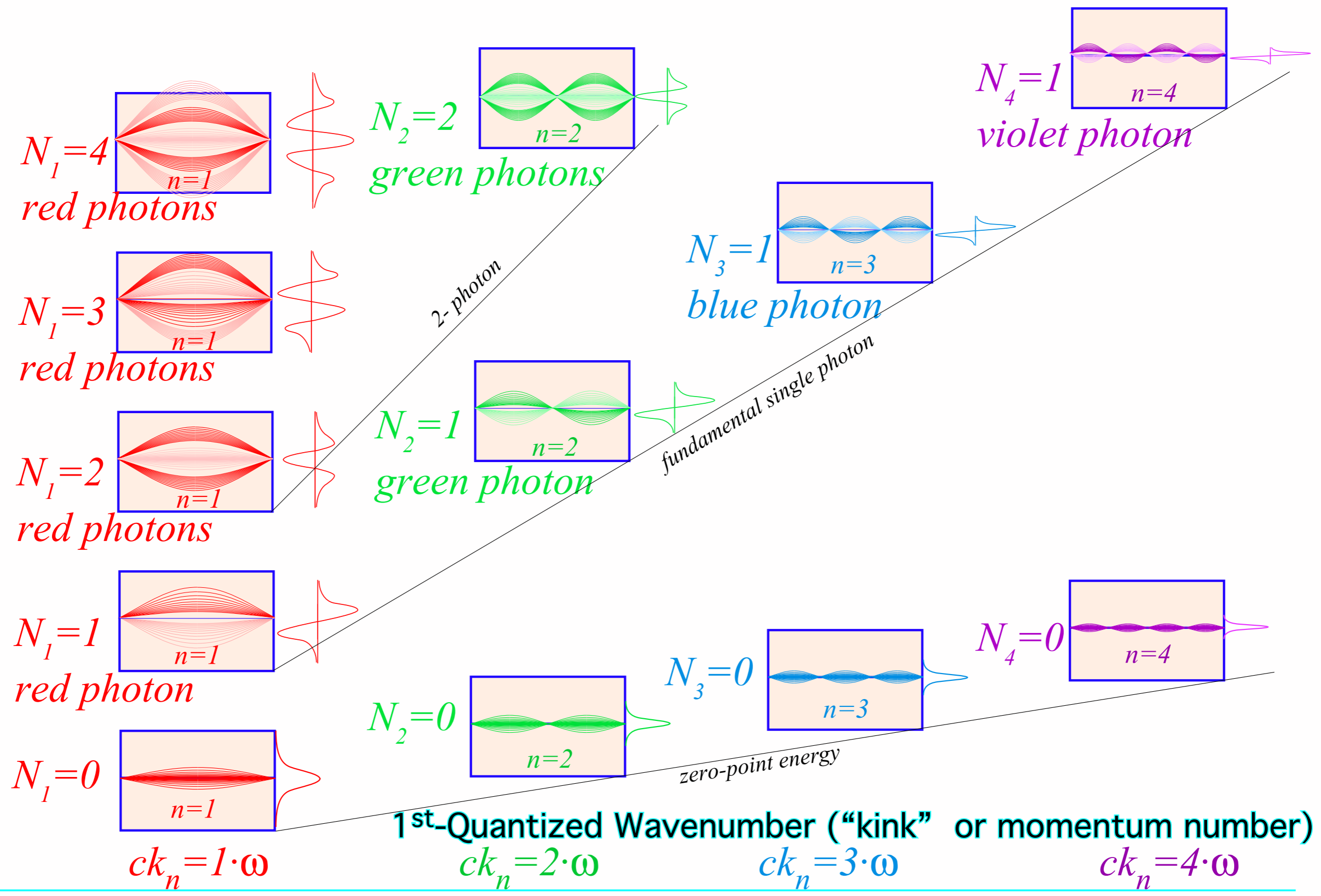
Take-away point 4
Cavity quantum electrodynamics (CQED) and spectra are analogous to molecular rovibronic dynamics with rotation-vibration algebra replaced by Lorentz-Poincare-Dirac algebra (and geometry!)

2nd Quantization:

$h\nu$ is actually $hN\nu$

$(h\nu_{phase}=E=h\nu_A \cosh \rho)$ is actually $(hN\nu_{phase}=E_N=hN\nu_A \cosh \rho \quad (N=1,2,..))$

2nd-Quantized Amplitude ("photon" number)



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Learning about sin! and COS and...

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Acceleration by chirping laser pairs

Varying acceleration (General case)

From Lect. 35
ModPhys (2012)

Only green-light is seen by observers on the green accelerated trajectory

Varying local acceleration $\rho = \rho(\tau)$

$$u = \frac{dx}{dt} = c \tanh(\tau)$$

$$\frac{dt}{d\tau} = \cosh \rho(\tau)$$

$$\frac{dx}{d\tau} = \frac{dx}{dt} \frac{dt}{d\tau} = c \tanh \rho(\tau) \cosh \rho(\tau) = c \sinh \rho(\tau)$$

$$ct = c \int \cosh \rho(\tau) d\tau$$

$$x = c \int \sinh \rho(\tau) d\tau$$

Constant local acceleration $\rho = \frac{g\tau}{c}$ "Einstein Elevator"

$$ct = c \int \cosh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \sinh \frac{g\tau}{c}$$

$$x = c \int \sinh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \cosh \frac{g\tau}{c}$$

Previous examples involved constant velocity

Constant velocity $\rho = \rho_0 = \text{const.}$ "Lorentz transformation"

$$ct = c \int \cosh \rho_0 d\tau = c\tau \cosh \rho_0$$

$$x = c \int \sinh \rho_0 d\tau = c\tau \sinh \rho_0$$

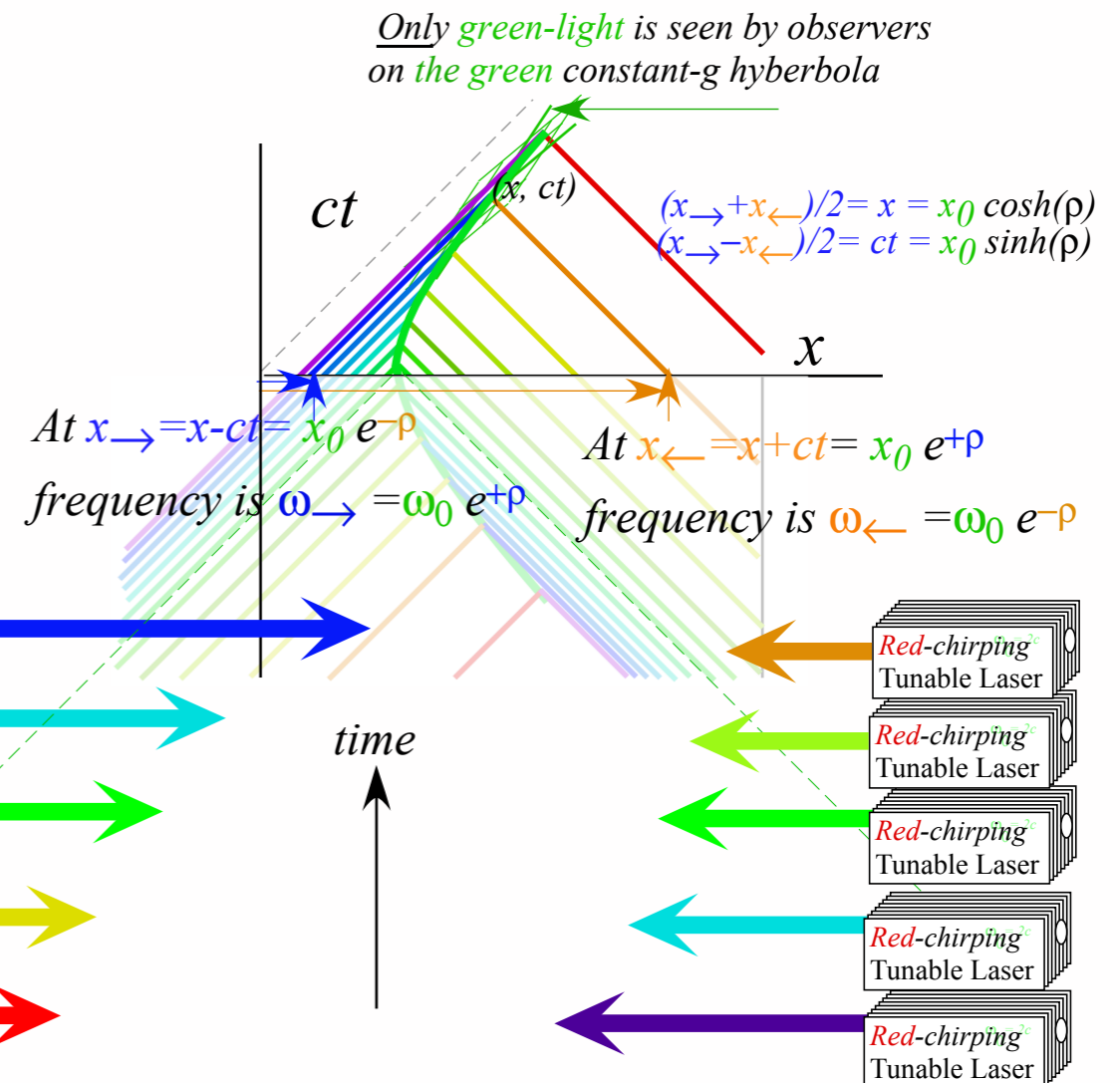
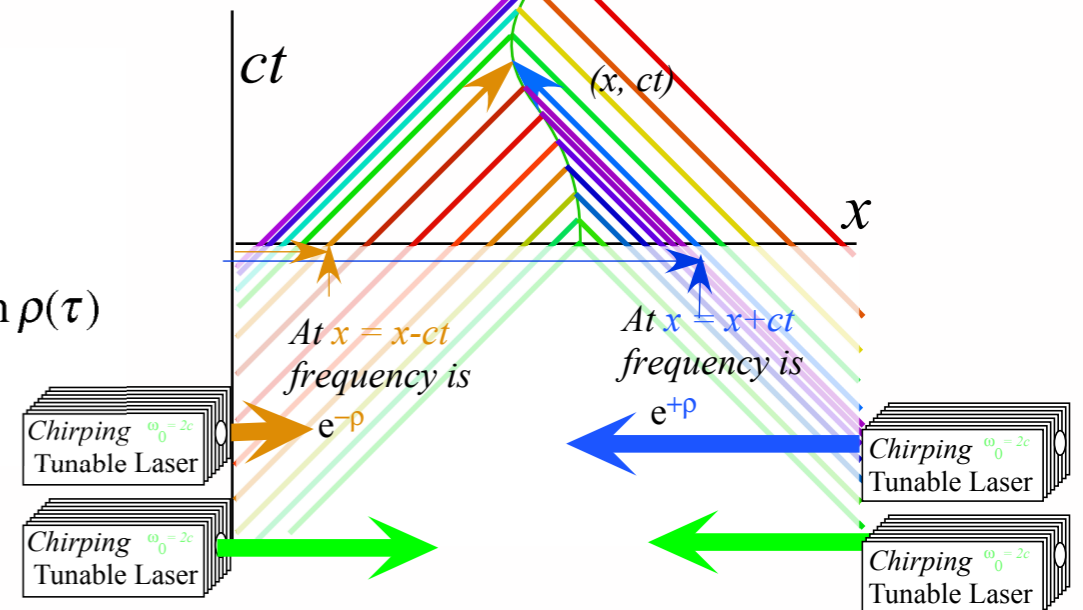
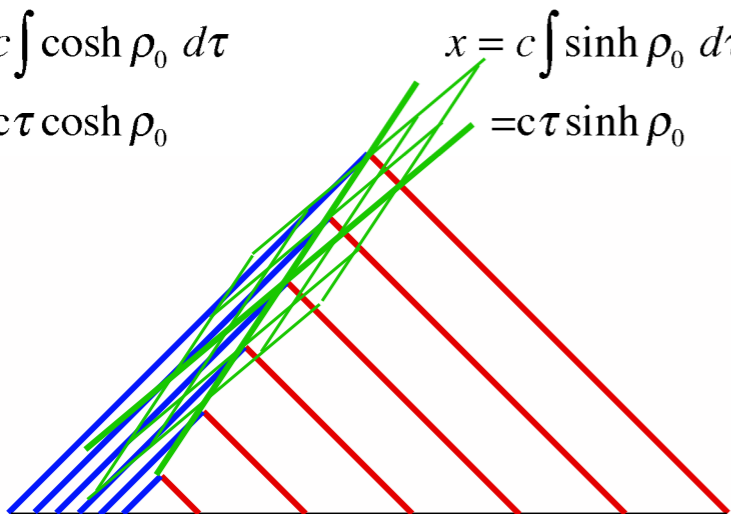
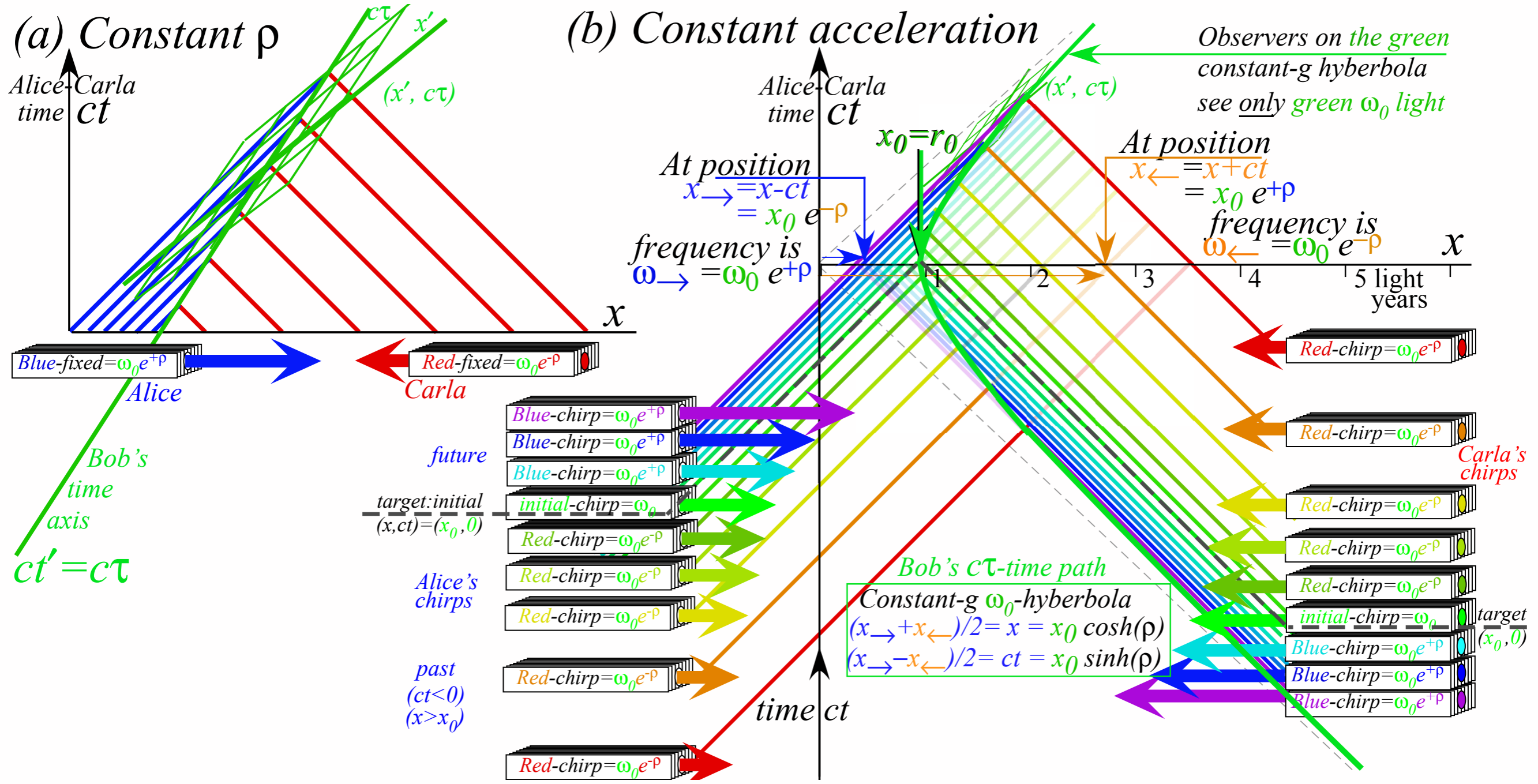
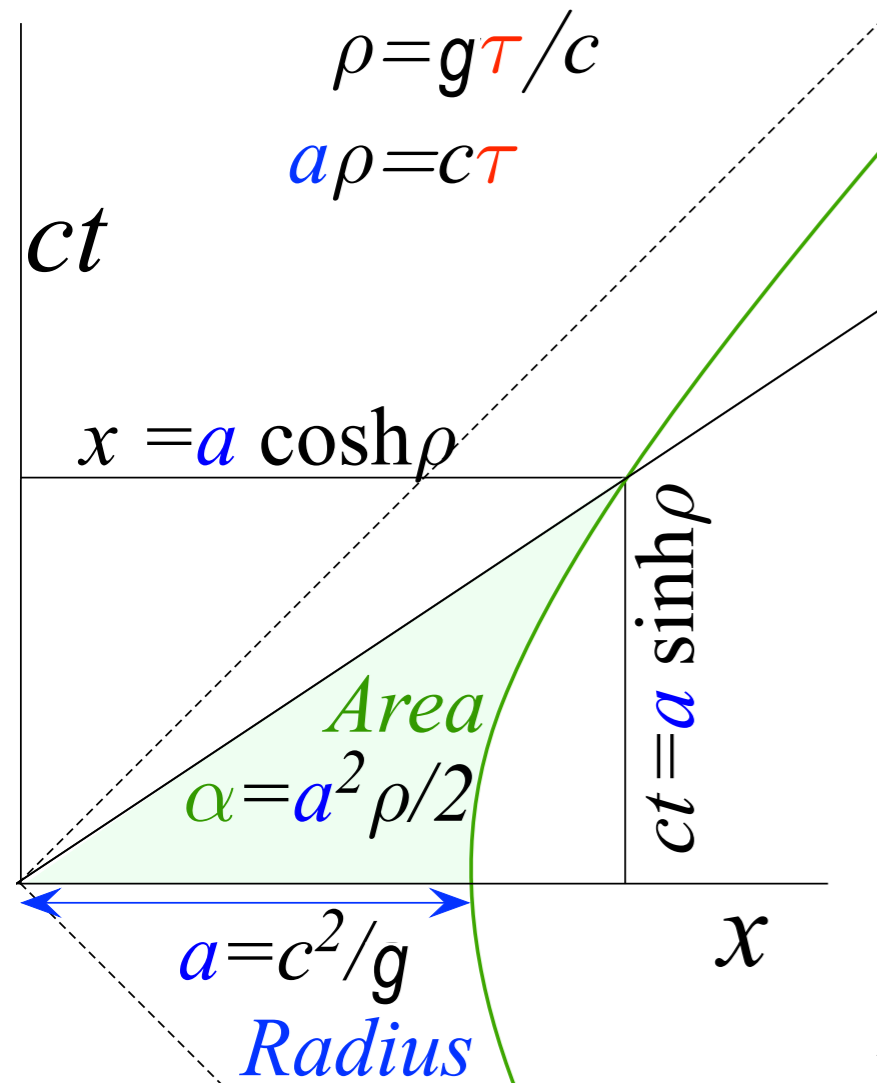


Fig. 8.1 Optical wave frames by red-and-blue-chirped lasers (a)Varying acceleration (b)Constant g



(a) Constant acceleration g
 Rapidity ρ vs proper time τ



(b) Traveler paths of acceleration g_q

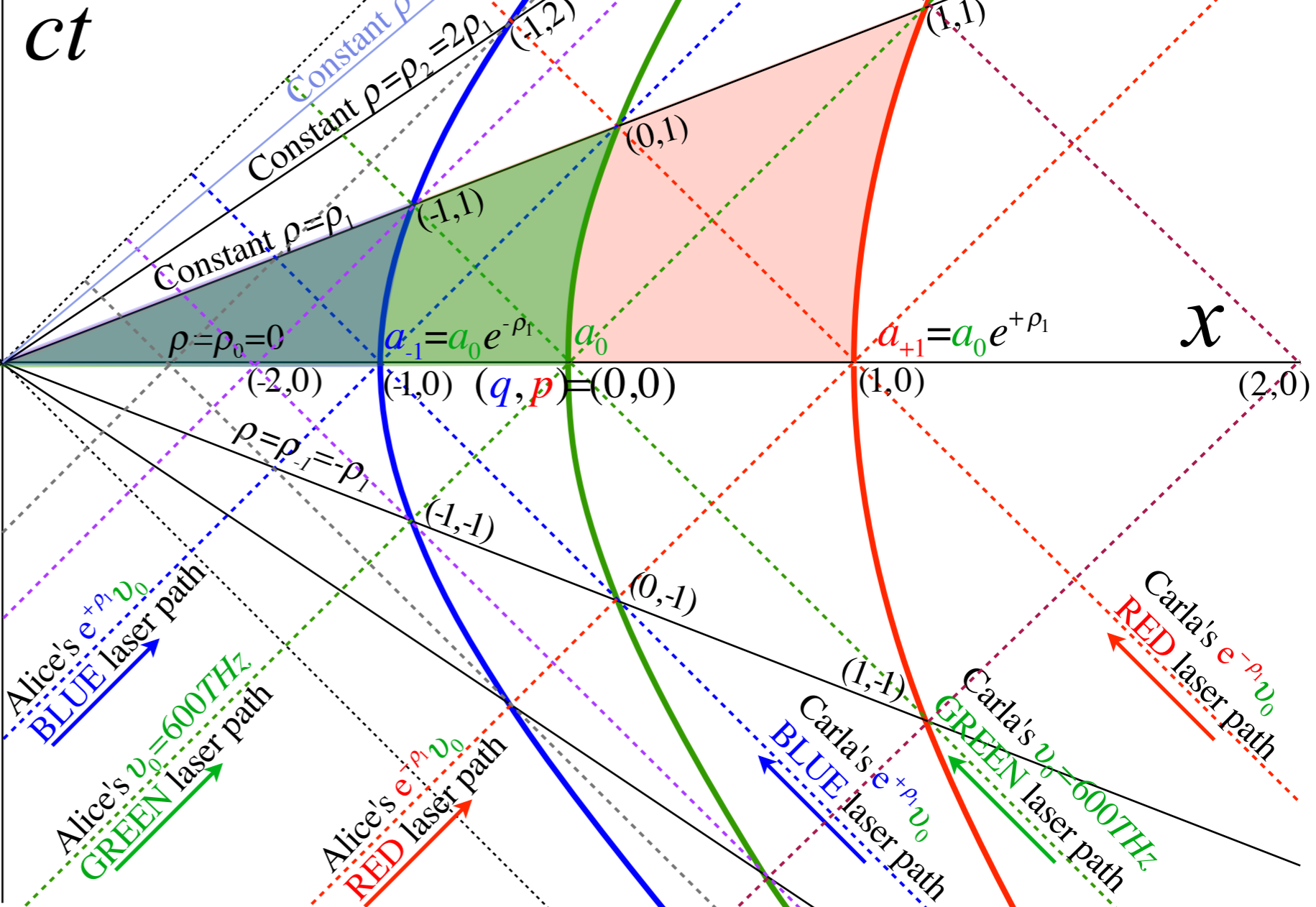
Al: $g_{-1} = g_0 e^{+\rho_1}$ Bob: $g_0 = \frac{c^2}{a_0}$ Carl: $g_{+1} = g_0 e^{-\rho_1}$

Inertial frame coordinates

$(x_{q,p}, ct_{q,p}) = a_0 e^{q\rho_1} (\cosh p\rho_1, \sinh p\rho_1)$

Geometric scale:

$e^{q\rho_1} = \left(\frac{3}{2}\right)^q$



Lecture 31

Thur. 12.10.2015

Review: Relativity ρ functions Two famous ones Extremes and plot vs. ρ
Doppler jeopardy Geometric mean and Relativistic hyperbolas
Animation of $e^{\rho}=2$ spacetime and per-spacetime plots

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity

Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

“Occams Sword” and summary of 16 parameter functions of ρ and σ

Applications to optical waveguide, spherical waves, and accelerator radiation

Learning about sin! and COS and...

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

Relativistic action and Lagrangian-Hamiltonian relations

Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

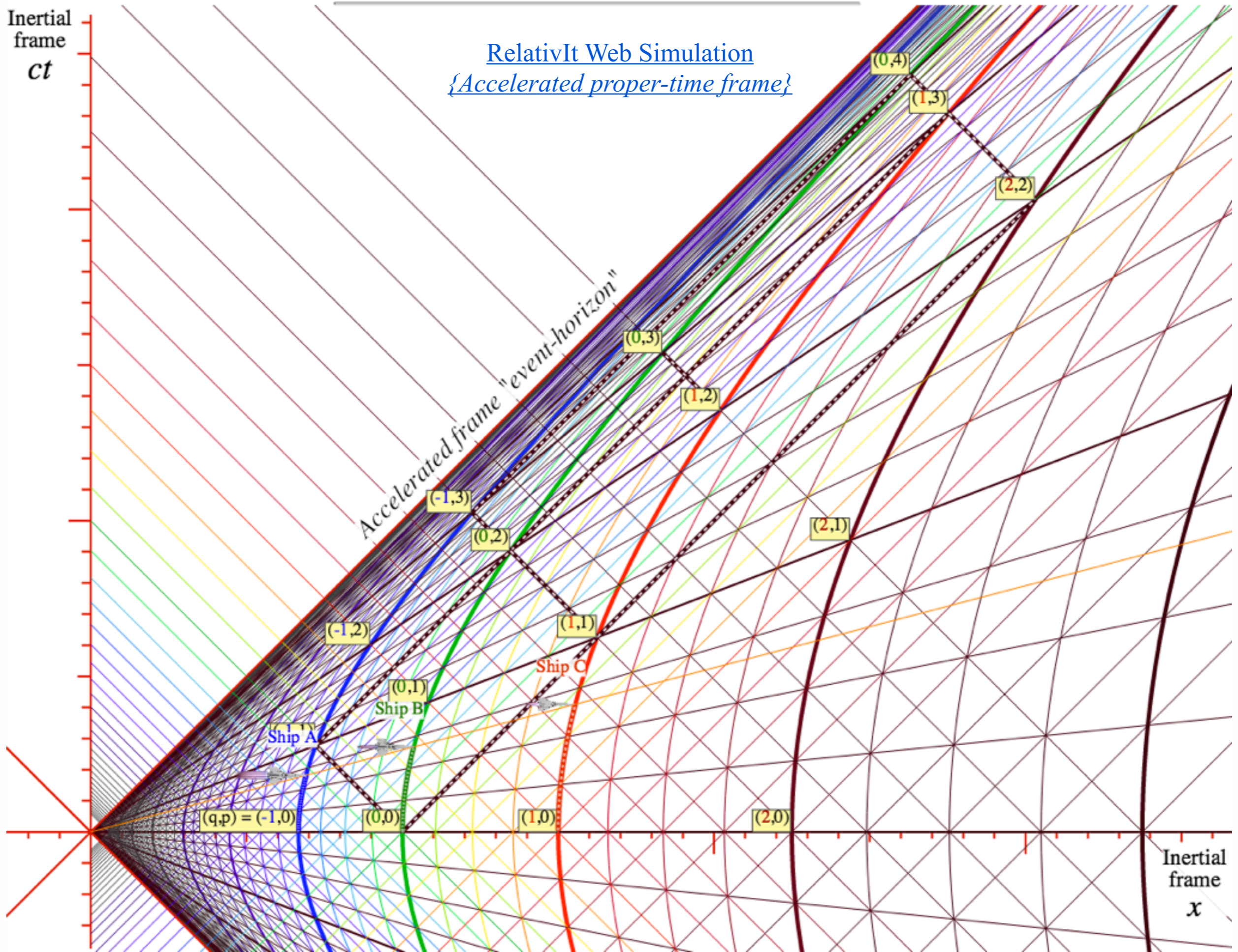
Comparing 2nd-quantization “photon” number N and 1st-quantization wavenumber κ

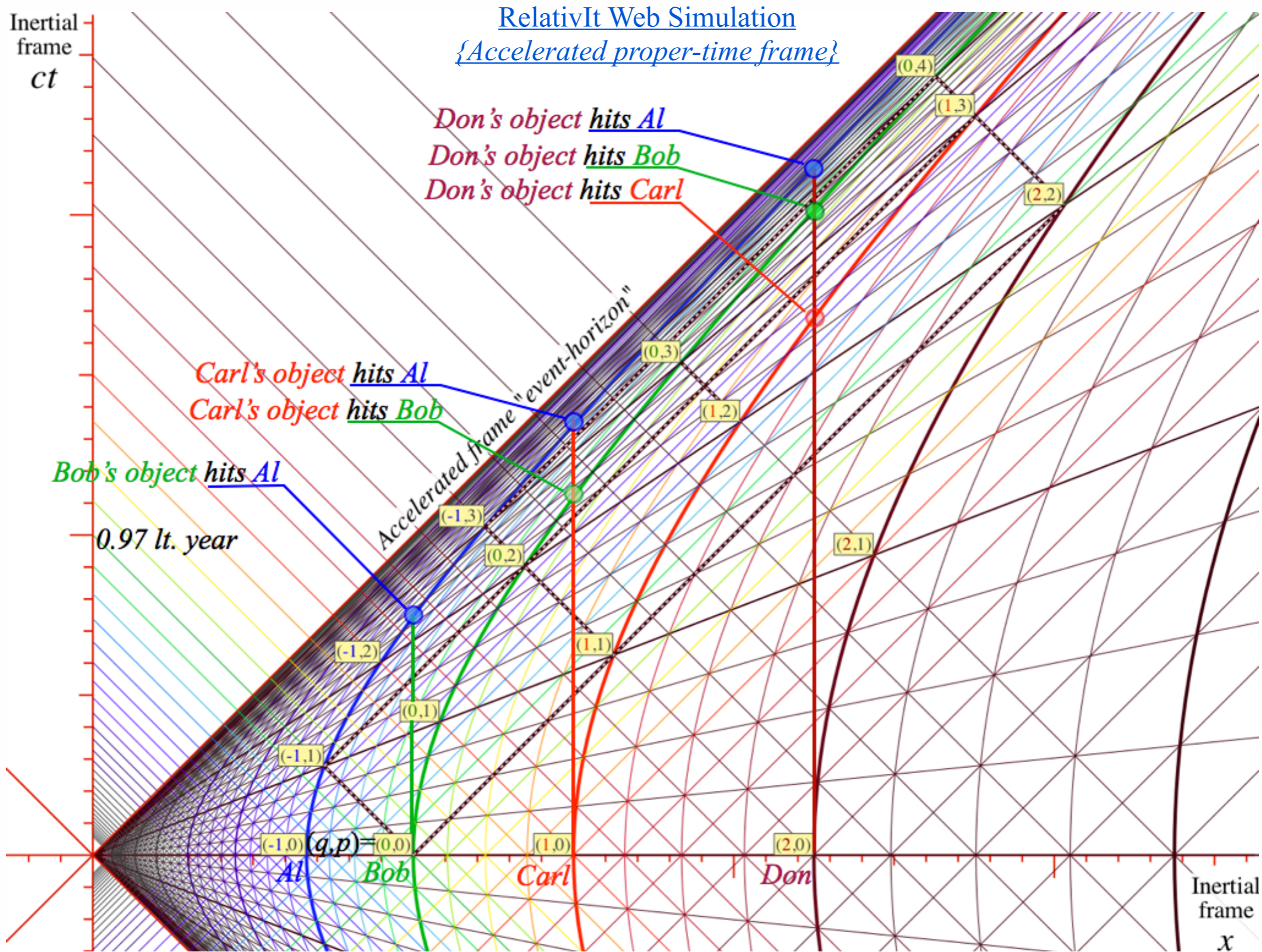
Relativity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes g -acceleration grid

Analysis of constant- g grid compared to zero- g Minkowski grid

➔ Animation of mechanics and metrology of constant- g grid





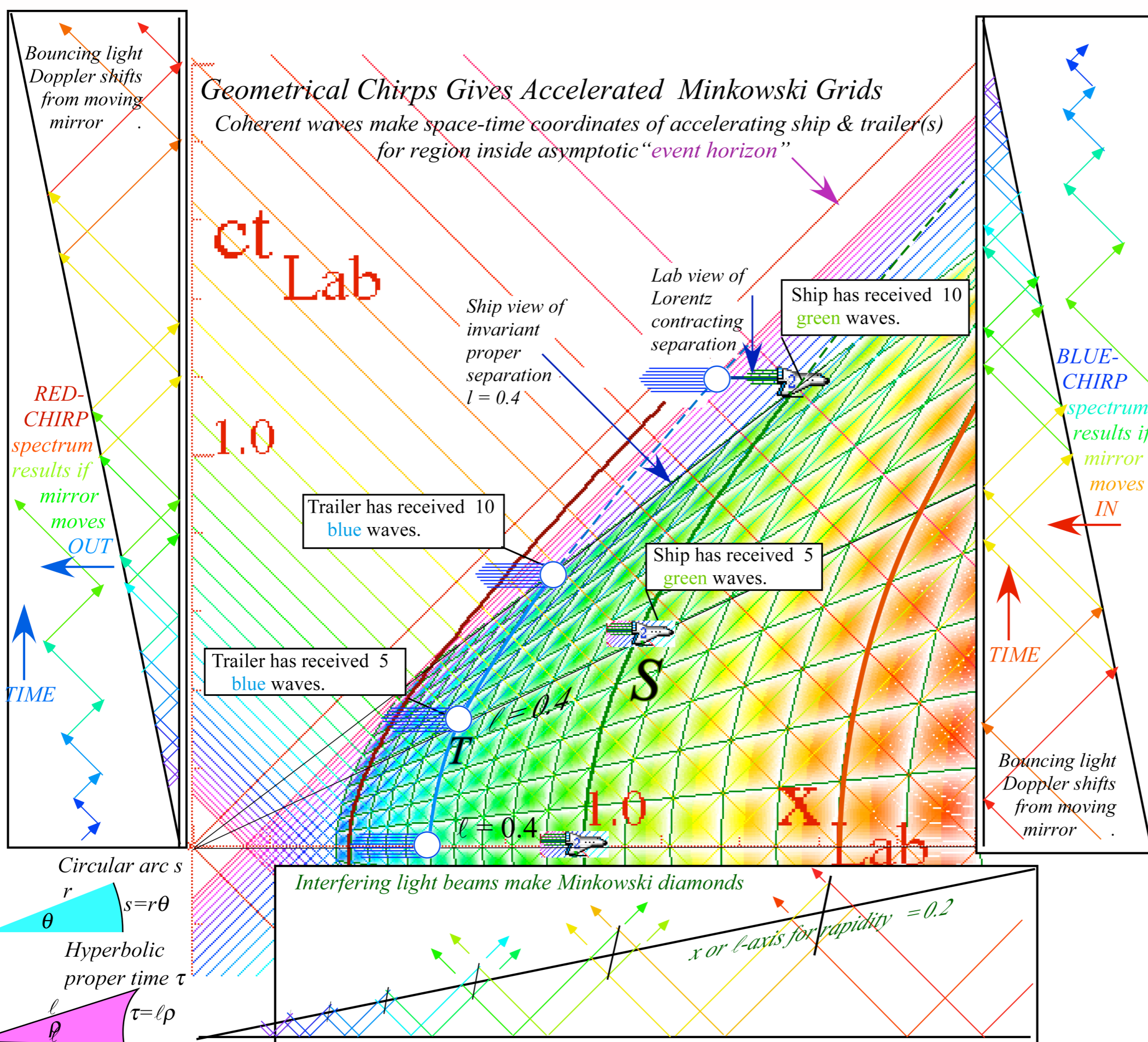
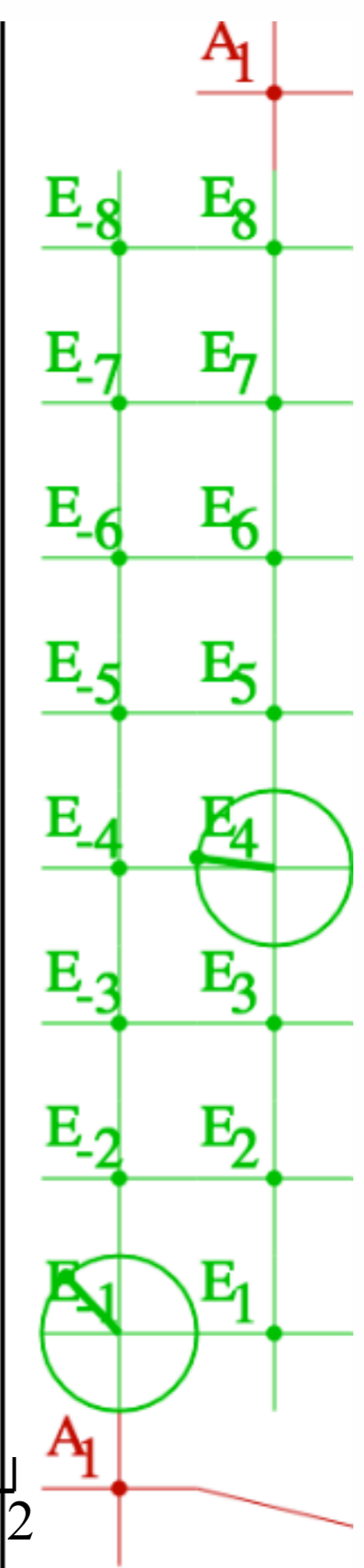
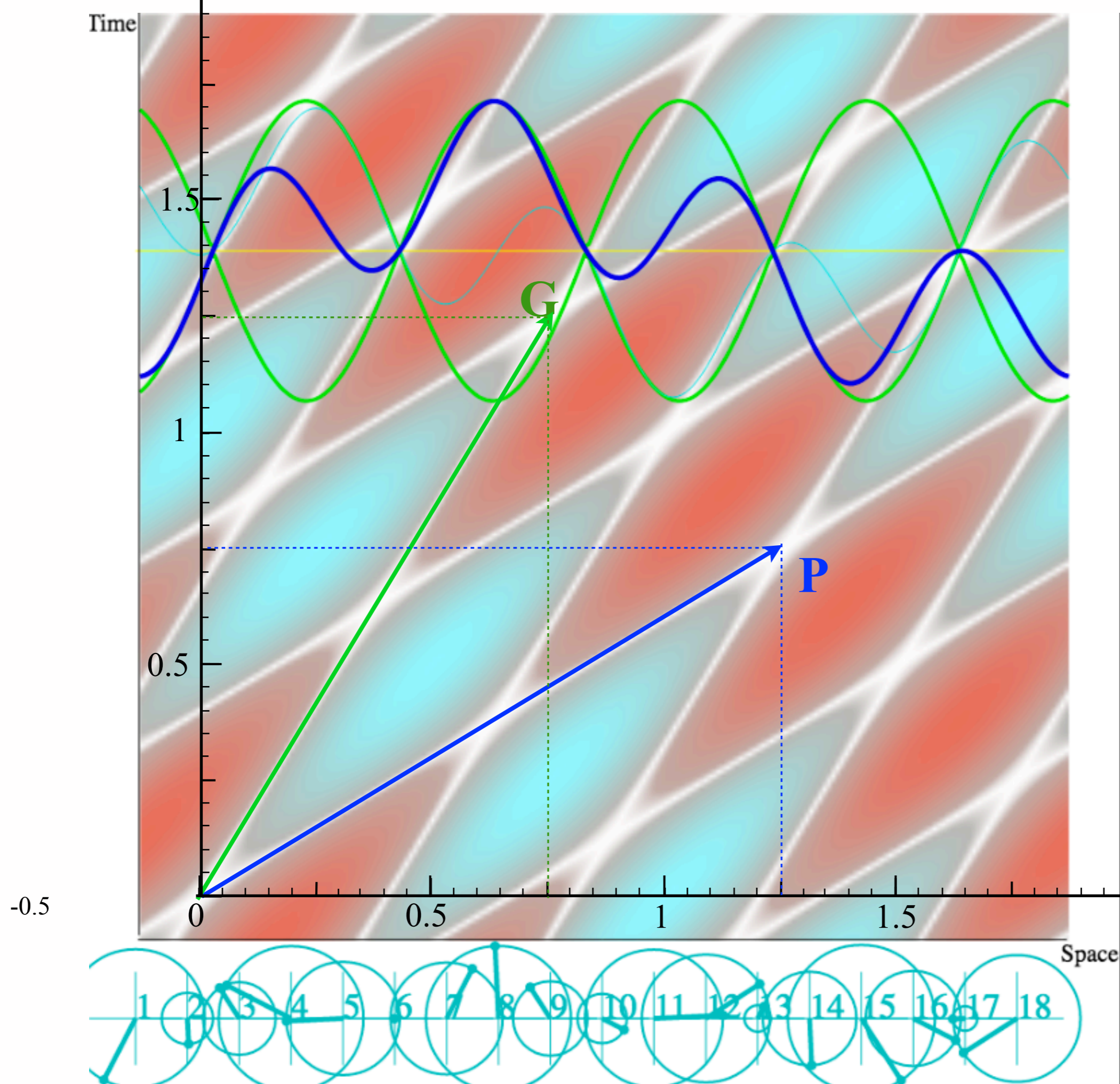
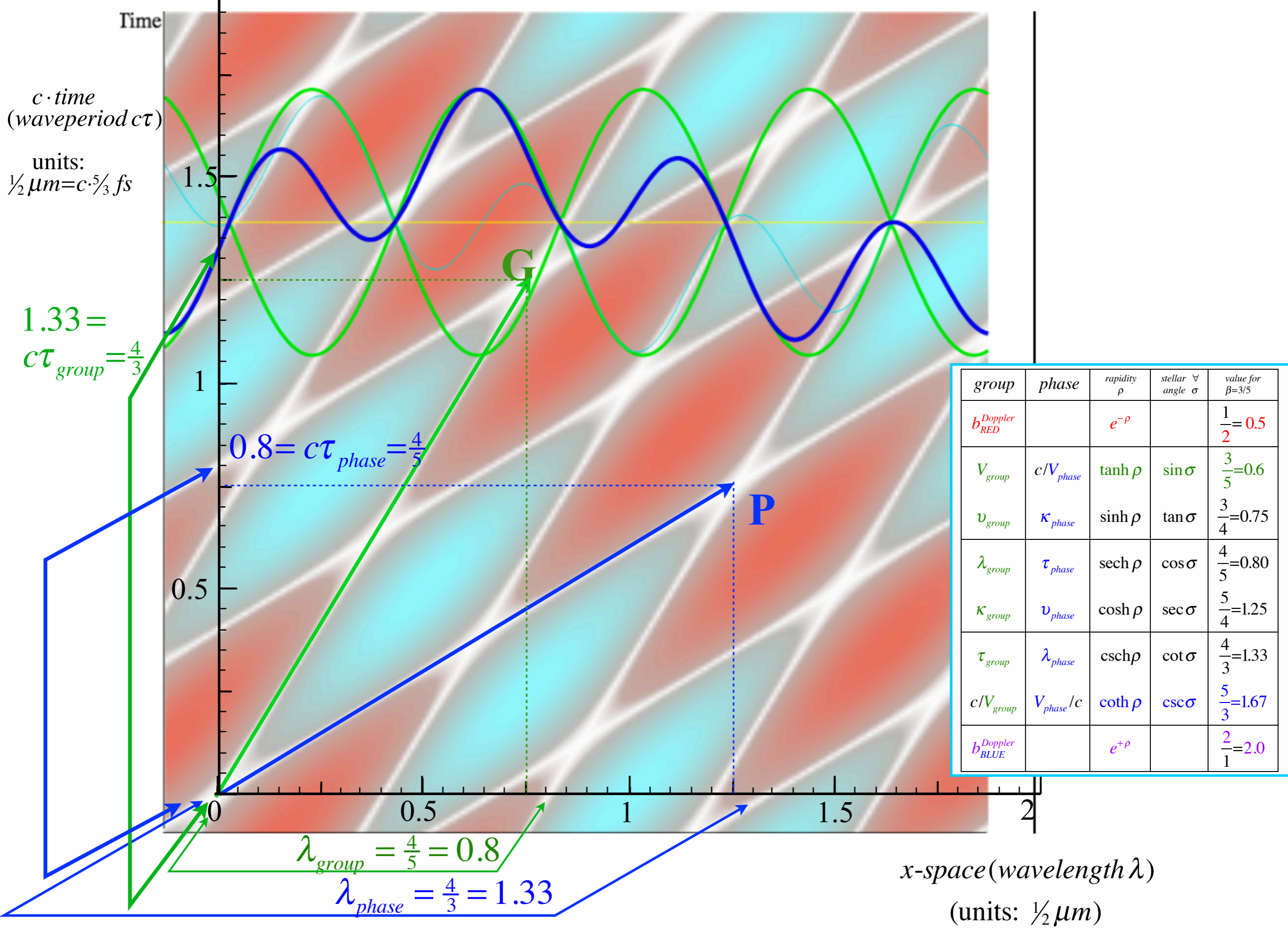
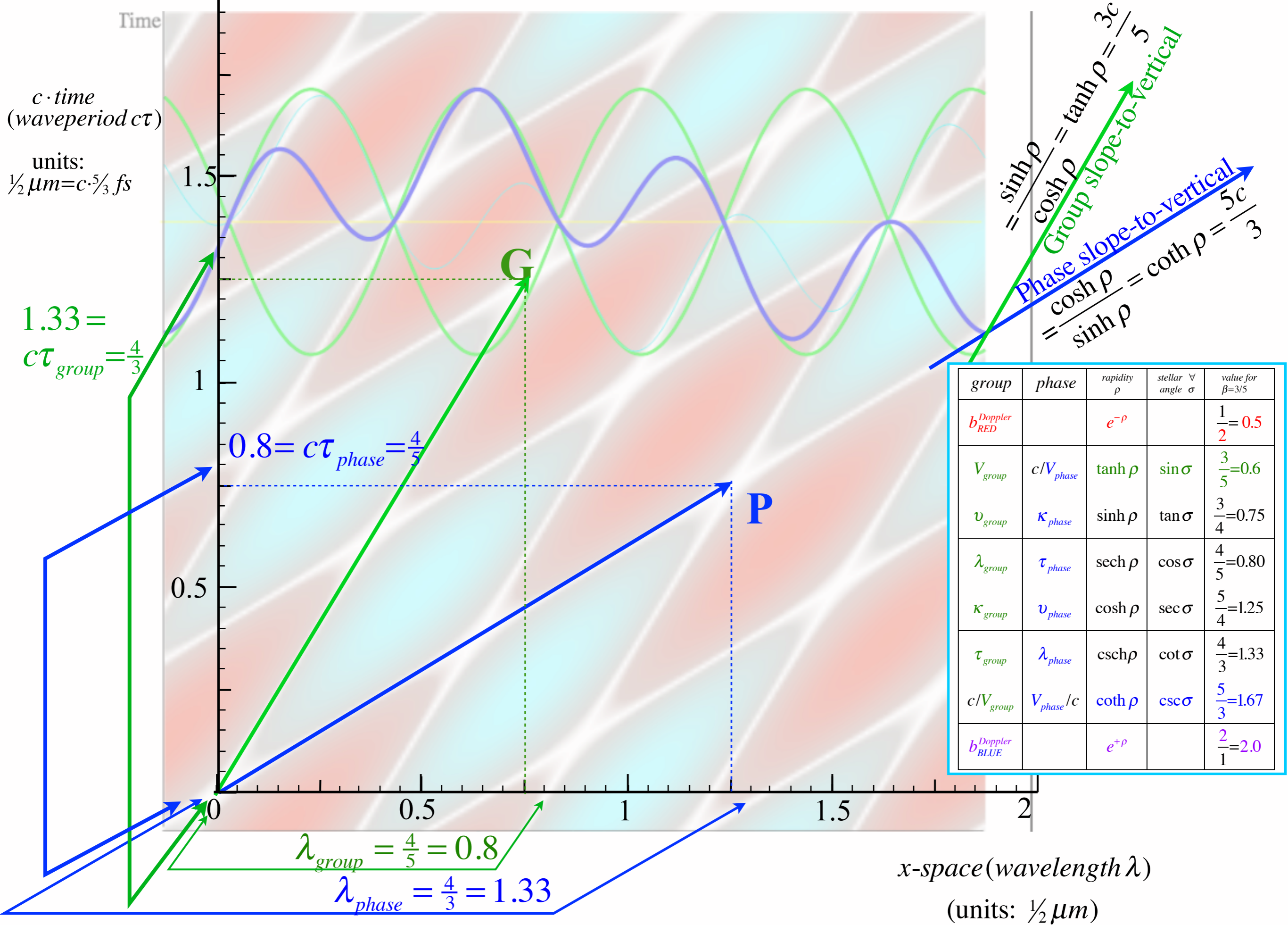
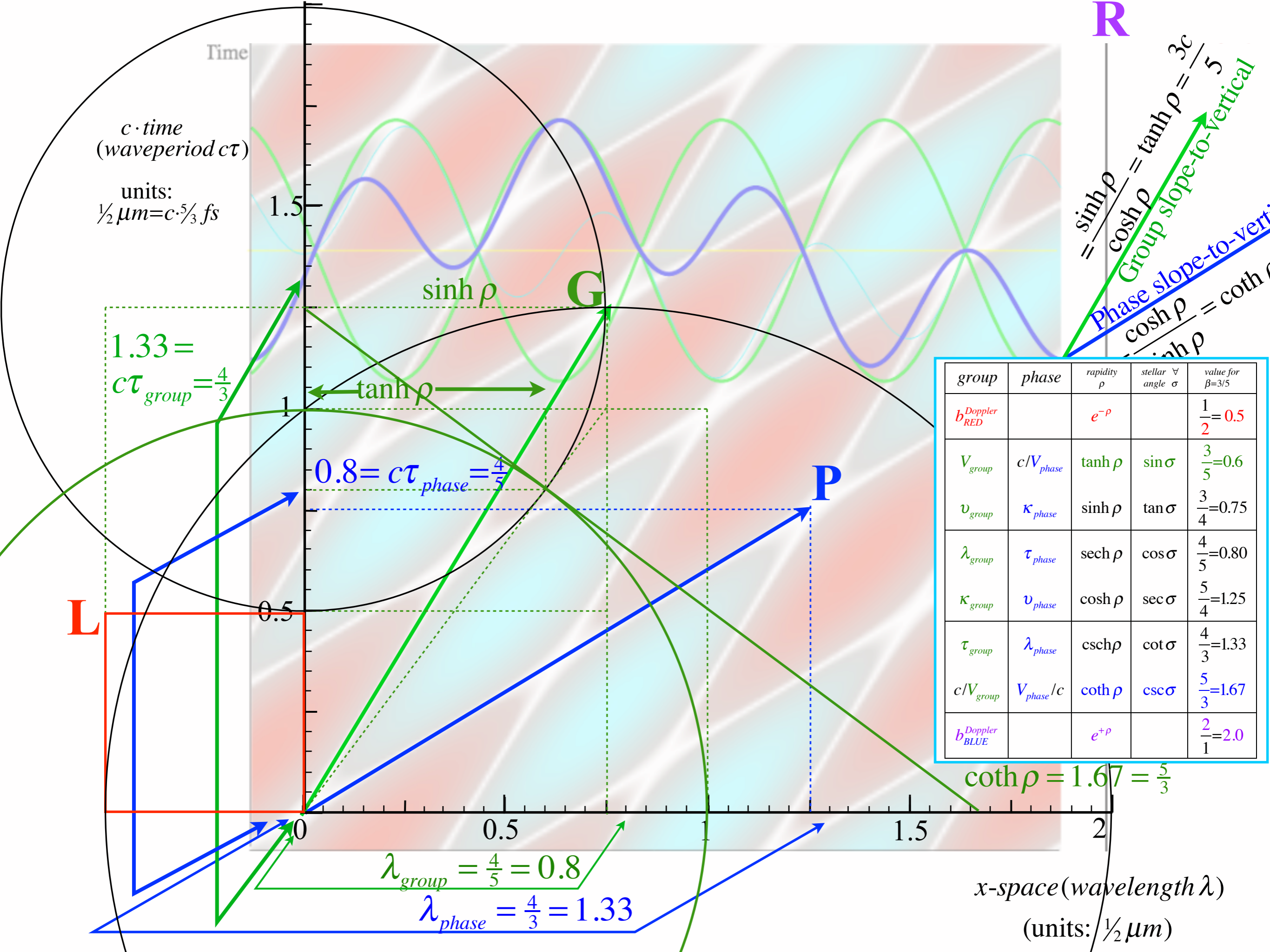


Fig. 8.2 Accelerated reference frames and their trajectories painted by chirped coherent light

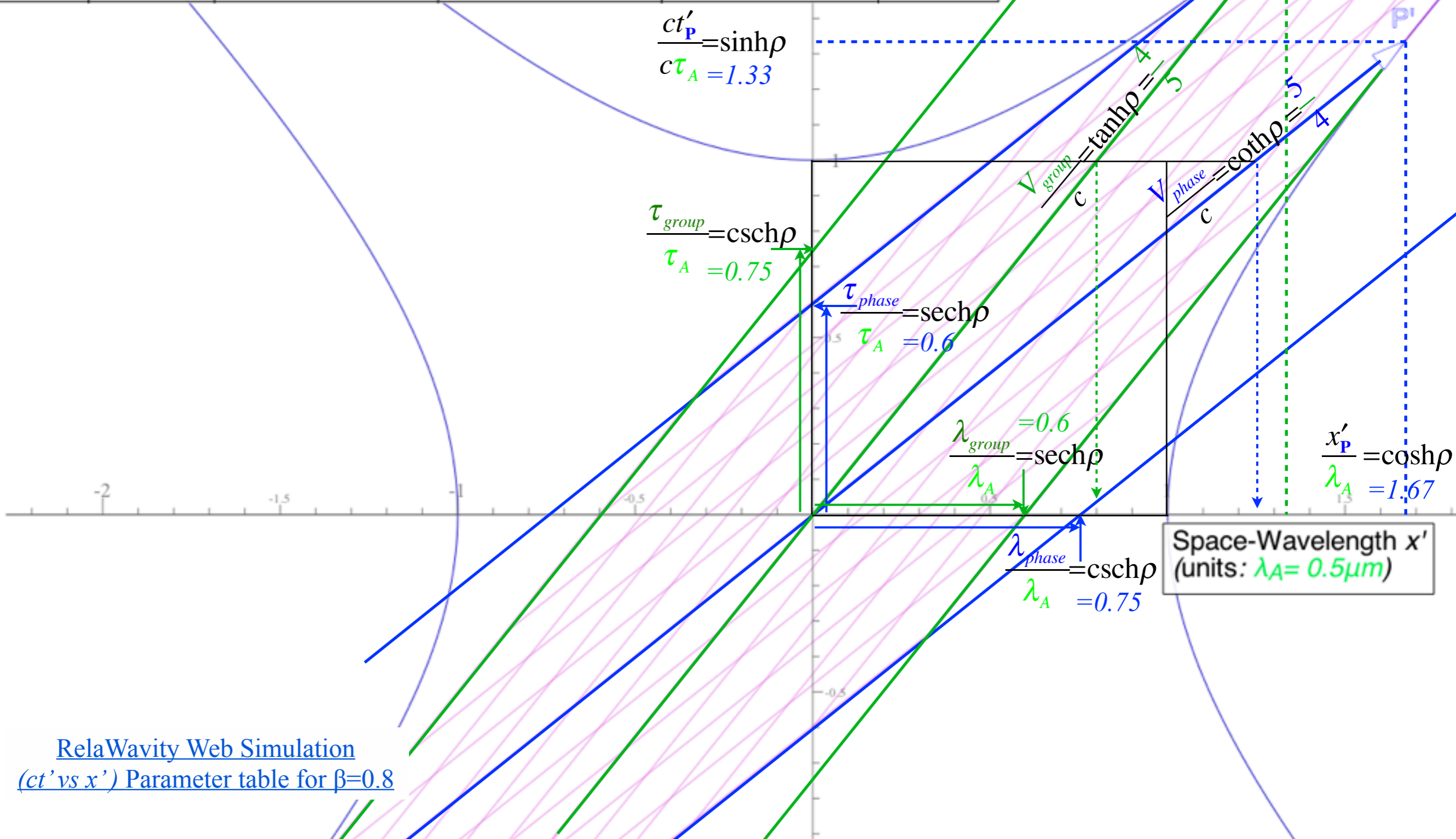








time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
value for $\beta=0.80$	0.33	0.80	1.34	0.60	1.67	0.75	1.25	3.01



Relativity Web Simulation
(ct' vs x') Parameter table for $\beta=0.8$

Bob's coordinates for Alice's G-point

$$x'_G = \lambda_A \sinh \rho$$

$$= c\tau_A \sinh \rho$$

$$ct'_G = \lambda_A \cosh \rho$$

$$= c\tau_A \cosh \rho$$

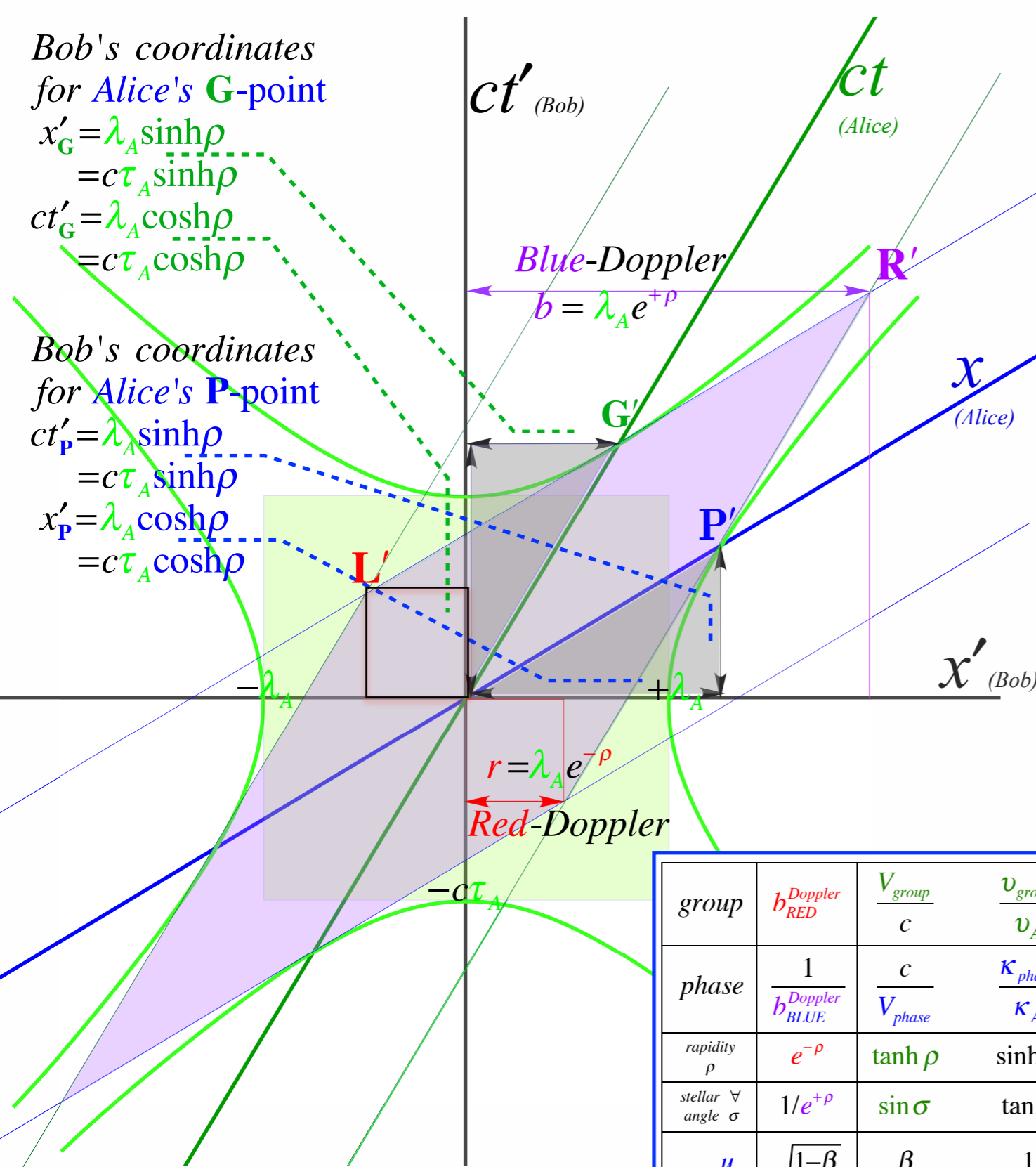
Bob's coordinates for Alice's P-point

$$ct'_P = \lambda_A \sinh \rho$$

$$= c\tau_A \sinh \rho$$

$$x'_P = \lambda_A \cosh \rho$$

$$= c\tau_A \cosh \rho$$



Space-time parameters

$$\lambda_{phase} = \lambda_A \operatorname{csch} \rho$$

$$\lambda_{group} = \lambda_A \operatorname{sech} \rho$$

$$c\tau_{phase} = c\tau_A \operatorname{sech} \rho$$

$$c\tau_{group} = c\tau_A \operatorname{csch} \rho$$

Per-space-time parameters

$$cK_{phase} = cK_A \sinh \rho$$

$$cK_{group} = cK_A \cosh \rho$$

$$v_{phase} = v_A \cosh \rho$$

$$v_{group} = v_A \sinh \rho$$

[RelaWavity Web Simulation](#)
[Comprehensive dual plots](#)
[with parameter table](#)

[RelaWavity Web Simulation](#)
[\(ct' vs x'\) with parameter table](#)

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$
effects	$b_{RED}^{Doppler}$	V_{group}	past-future asymmetry (off-diagonal Lorentz-transform)	x -contraction ^(Lorentz) τ_{phase} -contraction	t -dilation ^(Einstein) v_{phase} -dilation (on-diagonal Lorentz-transform)	inverse asymmetry	V_{phase}	$b_{BLUE}^{Doppler}$

