

# Lecture 29

Wed. 11.28.2018

## *Relativity : a novel introduction to relativistic mechanics I.*

(CMwBang! Unit 8 , AMOP Ch.0, )

*Why Men in Black shot little Suzie...Learning about sin!, cos and...Trigonometric road maps*

Hyper-Trigonometric *Relativity* geometry and Euler exponential algebra

1CW wavefunctions and phasors

Per-space-per-time vs Space-time (How to understand wave parameters)

Wave velocity formulas

Introducing Doppler shifting

Why is  $c$  so constant?!

Introducing Doppler Arithmetic and *Rapidity*  $\rho$

Optical interference “baseball-diamond” displays *phase* and *group* velocity

Details of 2CW wavefunctions in rest frame

Pulse waves (PW) versus Continuous Waves (CW)

Doppler shifted “baseball-diamond” displays Lorentz frame transformation

Analyzing wave velocity by *per-space-per-time* and *space-time* graphs

16 coefficients of relativistic 2CW interference

Two “famous-name” coefficients and the Lorentz transformation

Thales geometry of Lorentz transformation

*Rapidity*  $\rho$  related to *stellar aberration angle*  $\sigma$  and L. C. Epstein’s approach to relativity

Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry

“Occams Sword” and geometry of 16 parameter functions of  $\rho$  and  $\sigma$

Application to TE-Waveguide modes and synchrotron beam relativity

# *A running collection of links to course-relevant sites and articles*

## *Physics Web Resources*

[Comprehensive Harter-Soft Resource Listing](#)

[UAF Physics YouTube channel](#)

[LearnIt Physics Web Applications](#)

Neat external material to start the class:

[AIP publications](#)

[AJP article on superball dynamics](#)

[AAPT summer reading](#)

These are hot off the presses:

[Sorting ultracold atoms in a 3D optical lattice in a realization of Maxwell's demon - Kumar-Nature-Letters-2018](#)

[Synthetic three-dimensional atomic structures assembled atom by atom - Berredo-Nature-Letters-2018](#)

Slightly Older ones:

[Wave-particle duality of C60 molecules](#)

[Optical vortex knots – One Photon at a Time](#)

## *Older Links from Lectures 14-20*

<http://thearmchaircritic.blogspot.com/2011/11/punkin-chunkin.html>

<http://www.sussexcountyonline.com/news/photos/punkinchunkin.html>

[Shooting-range-for-medieval-siege-weapons-Anybody-knows](#)

<https://modphys.hosted.uark.edu/markup/TrebuchetWeb.html>

<https://modphys.hosted.uark.edu/markup/TrebuchetWeb.html?scenario=MontezumasRevenge>

<https://modphys.hosted.uark.edu/markup/TrebuchetWeb.html?scenario=SeigeOfKenilworth>

[The trebuchet, Chevedden, Sci Am 1995](#)

'Simple' Pendulum Sim: <https://modphys.hosted.uark.edu/markup/PendulumWeb.html>

'Cycloid' Pendulum: <https://modphys.hosted.uark.edu/markup/CycloidulumWeb.html>

Google search on: "[Satelite view of Patricia](#)" (Images)

[Physics Girl Channel - Fun with Vortex Rings in the Pool](#)

[iBall demo - Quasi-periodicity: https://youtu.be/\\_jntDtULxDe](#)

<https://modphys.hosted.uark.edu/markup/CoulltWeb.html?scenario=SynchrotronMotion>

<https://modphys.hosted.uark.edu/markup/CoulltWeb.html?scenario=SynchrotronMotion2>

[Mechanical Analog to EM Motion \(YouTube video\) - https://youtu.be/hTd5FTJ-vRk](#)

[Coullt Web Simulation: Bound-state motion in parabolic coordinates](#)

[Coullt Web Simulation: Bound-state motion in hyperbolic coordinates](#)

[Oscillt Web App: Simulations of various types of resonance: \[18\]\(#\), \[27\]\(#\), \[31\]\(#\), \[35\]\(#\), \[38\]\(#\), \[39\]\(#\)](#)

[Smith Chart](#)

<http://nobelprize.org/>

*AnalyIt Web Application*, posted 10/22/2018 in our *testing area*:

<https://modphys.hosted.uark.edu/testing/markup/AnalyItBJS.html>

## *"Texts"*

[Classical Mechanics with a Bang!](#)

[Quantum Theory for the Computer Age](#)

[Principles of Symmetry, Dynamics, and Spectroscopy](#)

[Modern Physics and its Classical Foundations](#)

[AMOP Detailed Development of Relativity](#)

[AMOP Ch 0 Space-Time Symmetry - 2019](#)

[Seminar at Rochester Institute of Optics, Auxiliary slides, June 19, 2018](#)

[Springer AMO Handbook - Ch 32 - Harter-Reimer-2019](#)

"Relativity" and quantum basis of *Lagrangian & Hamiltonian* mechanics:

[2-CW laser wave - BohrIt Web App](#)

[Lagrangian vs Hamiltonian - RelaWavity Web App](#)

## *Classes*

[2014 AMOP](#)

[2017 Group Theory for QM](#)

[2018 AMOP](#)

[2018 Adv Mechanics](#)

## *Older Links from Lectures 21-23*

[Advanced Atomic and Molecular Optical Physics 2018 Class #9, pages: \[5\]\(#\), \[61\]\(#\)](#)

[BoxIt Web Simulations](#)

[Pure A-Type w/Cosine](#)

[Pure B-Type w/Cosine](#)

[Pure B-Type w/Freq ratios](#)

[Mixed AB-Type 2:1 Freq ratio](#)

[Pure A-Type A=4.9, B=0, C=0, & D=4.0](#)

[Pure B-Type: A=4.0, B=-0.2, C=0, & D=4.0](#)

[Pure C-Type A,D=4.055, B=0, C=0.1](#)

[Mixed AB-Type w/Cosine](#)

[Mixed AB Type A=4.0, BU2=0.866..., CU2=0, & D=1.0 w/Stokes & Freq rats](#)

[Mixed AB Type A=5.086 B=-0.27 C=0 D=2.024 w/Stokes plot](#)

[Mixed ABC Type A=4.833 B=0.2403 C=0.4162 D=4.277 w/Stokes plot](#)

[Recent mixed ABC Type A=0.325 B=0.375 C=0.825 D=0.05 w/Stokes plot](#)

[Classical Mechanics with a Bang! 2018](#)

[Lectures \[8\]\(#\), \[9\]\(#\), \[23\]\(#\) page \[93\]\(#\)](#)

[Text Unit \[6\]\(#\), page=\[27\]\(#\)](#)

[ColorU2 for the Web - in development](#)

[Group Theory for Quantum Mechanics - 2017 Lectures: \[6\]\(#\), \[7\]\(#\), \[8\]\(#\),](#)

[and the combined \[9-10\]\(#\)](#)

[Quantum Theory for the Computer Age Unit 3 Ch.7-10, page=90](#)

[Web based 3D & XR \( \$x \in \{A, M, V\}\$ , R=Reality\) <https://www.babylonjs.com/>](#)

[Web based 3D graphics WebGL API \(Graphics Layer modeled after OpenGL\)](#)

[Wiki on Pafnuty Chebyshev](#)

*continued* ↘



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[Principles of Symmetry, Dynamics, and Spectroscopy](#)

[Modern Physics and its Classical Foundations](#)

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[2014 AMOP](#)

[2017 Group Theory for QM](#)

[2018 AMOP](#)

[2018 Adv Mechanics](#)

 Repeated from previous page

## *Older Links from Lectures 24-27*

[JerkIt Web App: 2-, 2+, Amp50Omega147-, Amp50Omega296, Amp50Omega602, Gap\(1\)](#)

[MolVibes Web App: C3vN3](#)

[Wavelt Web App:](#)

[Dim = 3 w/Wave Components;](#)

[Static Char Table: 6, 12, 12\(b\), 16, 36, 256](#)

[Quantum Carpet with N=20: Gaussian, Boxcar](#)

[Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-CPL-2015](#)

[QTCA Unit 5 Ch14 2013](#)

[Lester. R. Ford, Am. Math. Monthly 45,586\(1938\)](#)

[John Farey, Phil. Mag.\(1816\) Wolfram](#)

[Harter, J. Mol. Spec. 210, 166-182 \(2001\)](#)

[Harter, Li IMSS \(2013\)](#)

[Li, Harter, Chem.Phys.Letters \(2015\)](#)

[OscillatorPE Web App: IHO Scenario 2, Coulomb Scenario 3](#)

[RelaWavity Web App/Simulator/Calculator: Elliptical - IHO orbits](#)

[Coultt Web App Simulations: p19, p32, p72, p73, p92, R=-0.375, R=+0.5, Rutherford](#)

## *Links to supplement Lecture 29*

[Ruler & Compass - Relawavity Exercise](#)

[2018 RelaWavity Portal Page](#)

[AMOP Chapter 0: Space-Time Symmetry](#)

[AMOP Detailed Development of RelaWavity](#)

[2018 Rochester Talk \(Auxiliary Slides\)](#)

[Special Relativity and Quantum Theory by Ruler and Compass - Earlier, expanded draft](#)

[Pirelli Relativity Challenge Web Site:](#)

[Title Page, Clocks\\_12\\_hr, Clocks\\_24\\_hr\\_QT, Phasors Addition](#)

[Bohrlt Web App/Simulations: -130022, -30001, -30104, 30004, 30022](#)

[Guidelt Web App/Scenarios: 230, 260](#)

[RelativIt Web App/Scenarios: 22, 24](#)

[RelaWavity Web App/Scenarios: 0,9, 3,6, 3,6 NoMink, 4,8, 6,1, 6,3a, 6,3b, 7,1, 7,2,1, 7,2,2, 7,2,3, 7,2,7, 8,3, 8,5, 8,7, 8,8](#)

## *Older Links from Lectures 28*

[CMwBang Text 2012 Unit 6 page=5](#)

[Bouncelt Web App/Scenarios:](#)

[5002, 5003](#)

[Coultt Web App/Scenarios:](#)

[TwoParticleCollision LToR, TwoParticleCollision LToR CM, TwoParticleOrbit Coulomb,](#)

[TwoParticleOrbit Coulomb CM, TwoParticleOrbit Hooke, TwoParticleOrbit Hooke CM](#)

[Singular Motion of Asymmetric Rotators AJP 44, 11 p1080 Harter-Kim-1976](#)

[Molecular Eigensolution Symmetry Analysis and Fine Structure - Int.J.MolSci1.4.13 Harter-Mitchell-IJMS-2013](#)

[Lenz Vector and Orbital Analog Computers - AJP 44 p348 1976](#)

[Some Geometric Aspects of Classical Coulomb Scattering AJP 40 4 p1852 1972](#)

[How Molecules do Self-NMR - Harter-Mitchell-Columbus-2009](#)

[Classical Mechanics with a Bang! - Asymmetric Top Demo](#)

[Allbookstores.com - Compare for Heller's SemiClassical Way - 0691163731](#)

["My Bomerang Won't Come Back" \(YouTube: Playlist\)](#)

[Rotating Solid Bodies in Microgravity \(YouTube\)](#)

[Dancing T-handle in zero-g \(YouTube\)](#)

➔ *Why Men in Black shot little Suzie... Learning about **sin!**, **cos** and... Trigonometric road maps*

Hyper-Trigonometric *Relativity* geometry and Euler exponential algebra

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*For an introductory, web based development of this and other concepts in special relativity see our entrant in the 2005 Pirelli Challenge:*

*A Colorful Road to Relativity*  
*Using Occam's Razors and*  
*Evenson's Lasers*





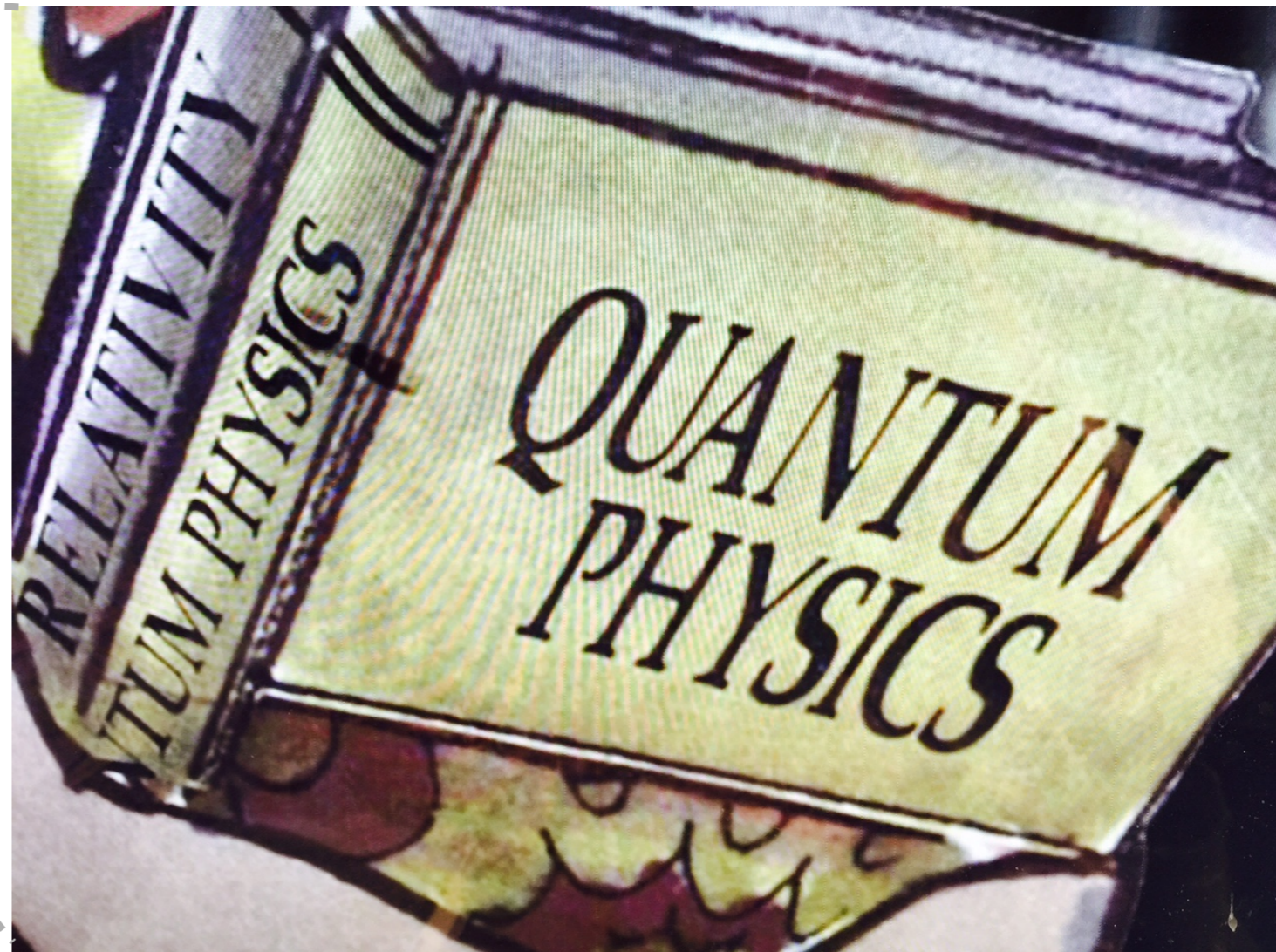
From AMOP Ch.0 article.

Why did a *Men In Black* candidate shoot little Suzy?

*Bad Suzy!*

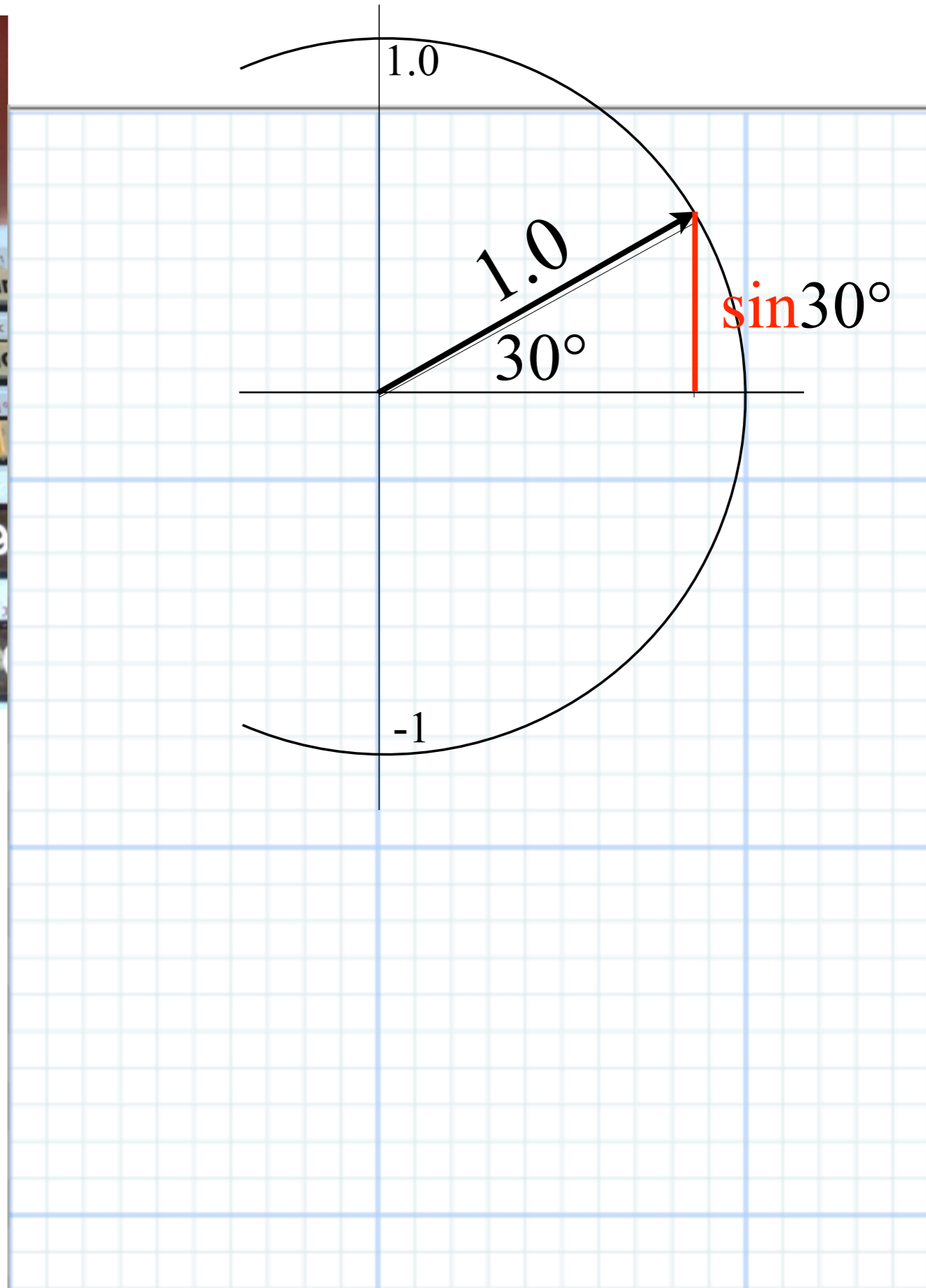
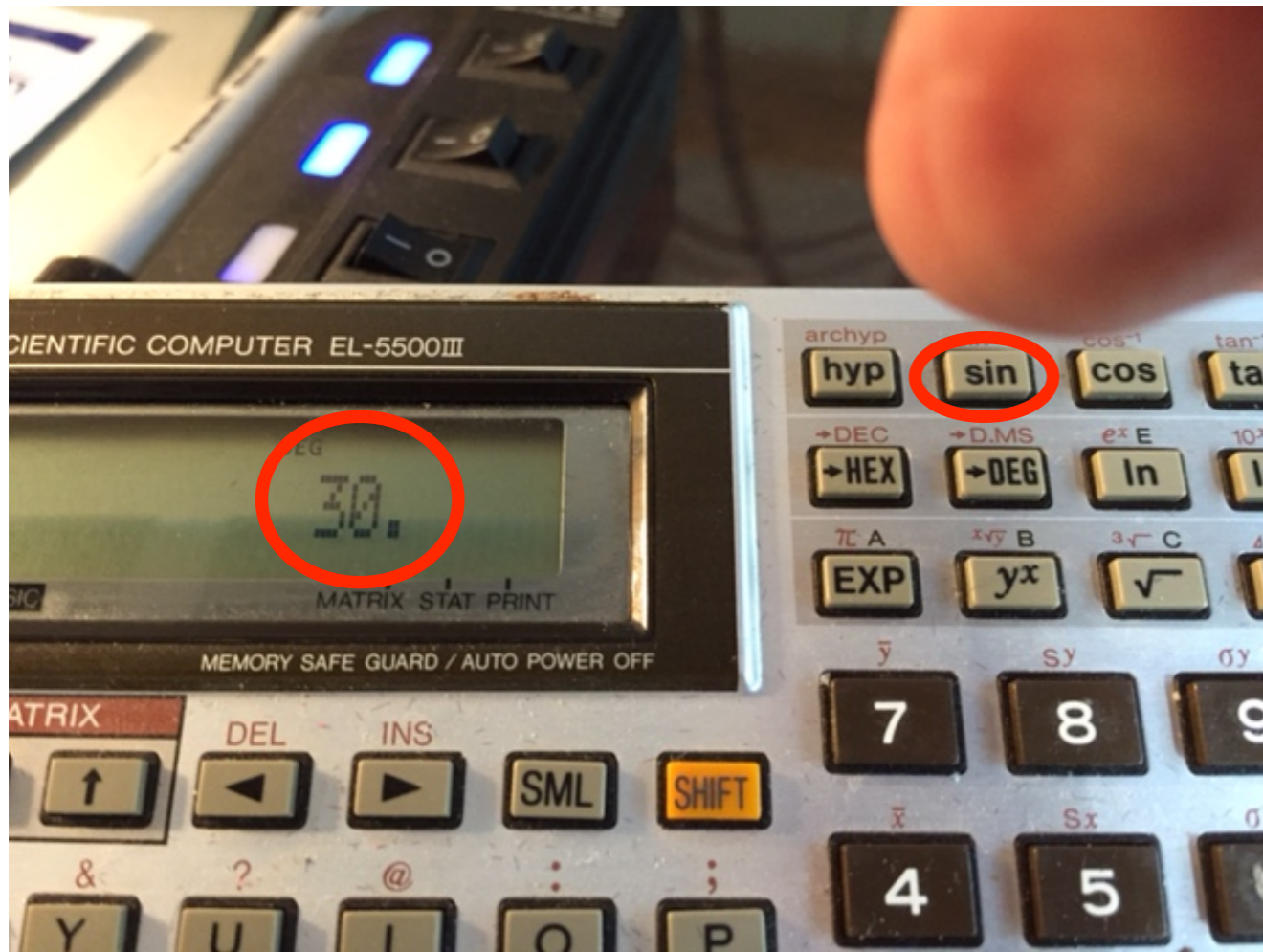
Relativity and Quantum Theory  
need to be unified in *one* book  
*half* the size of those old tomes!

We call that a *Relawavity* book.  
(It's a *lot* **lighter**!)

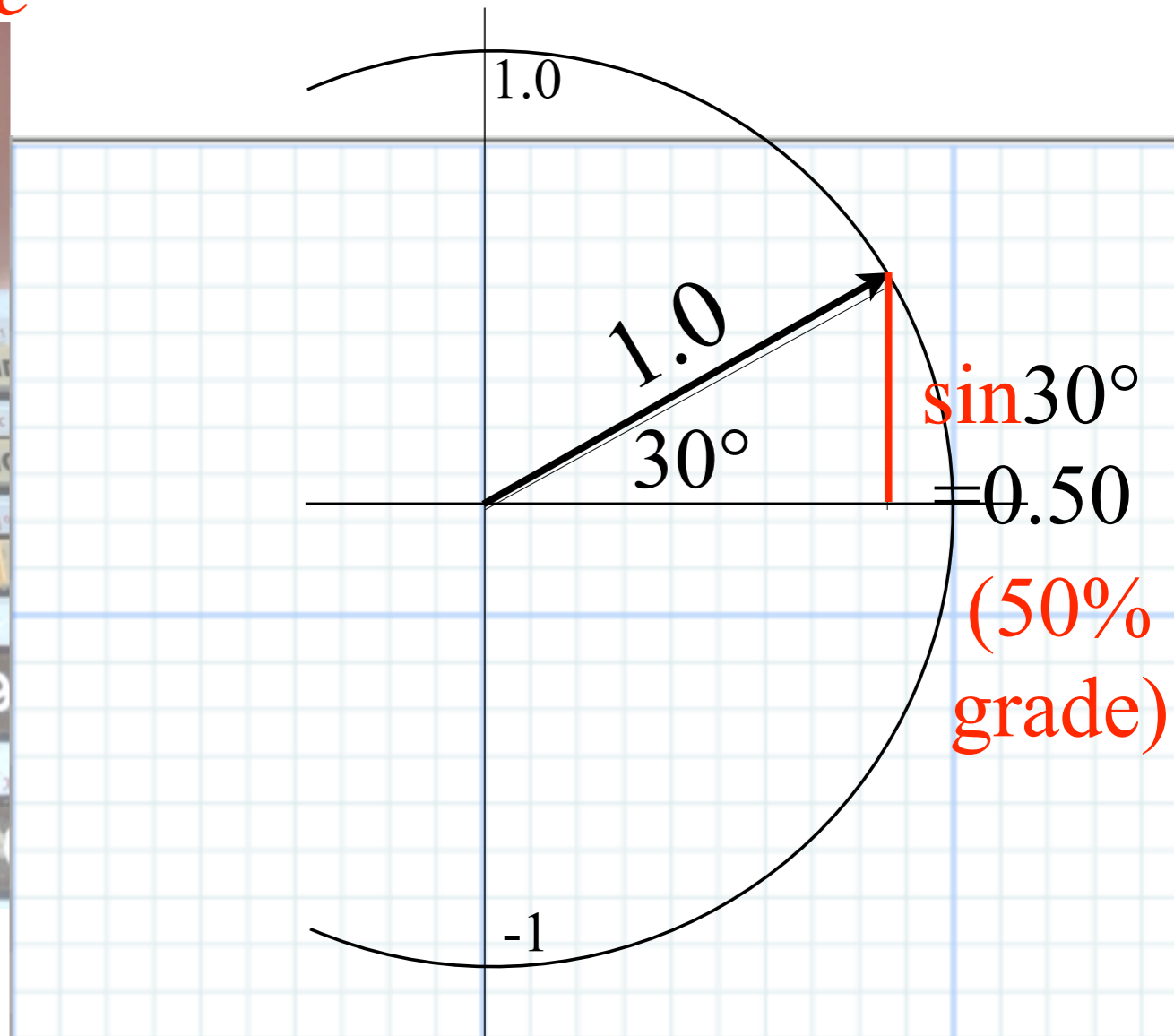
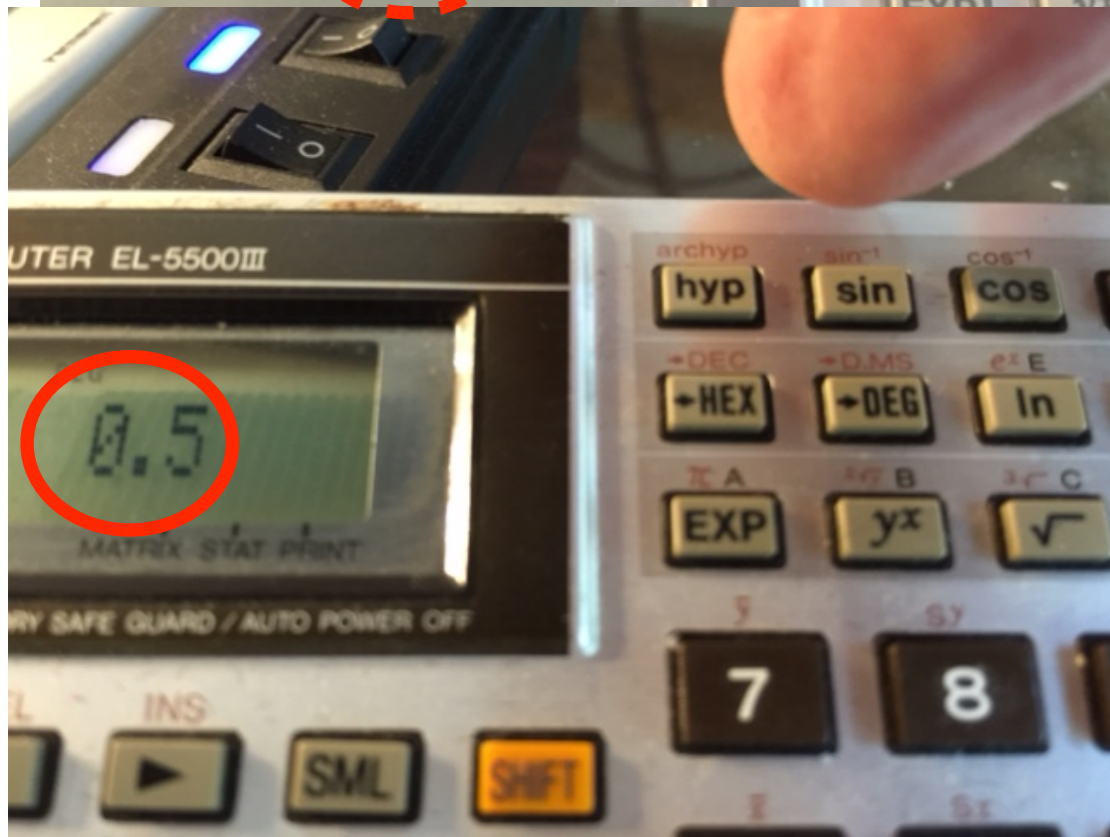
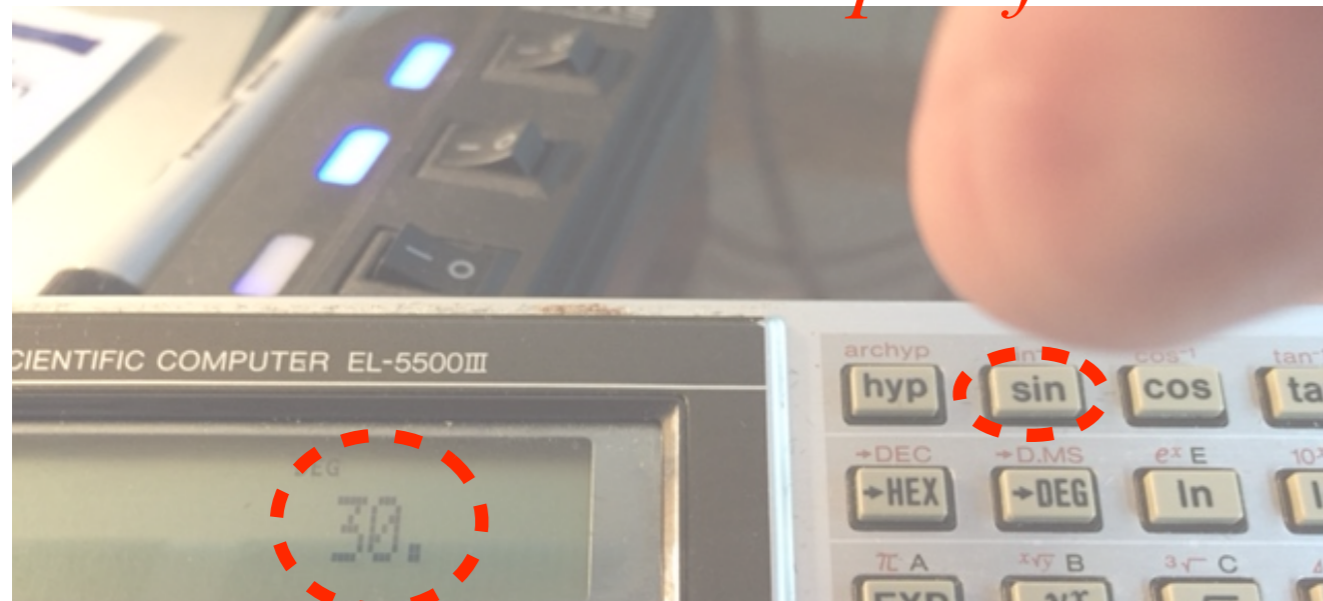




# Learning about SIN



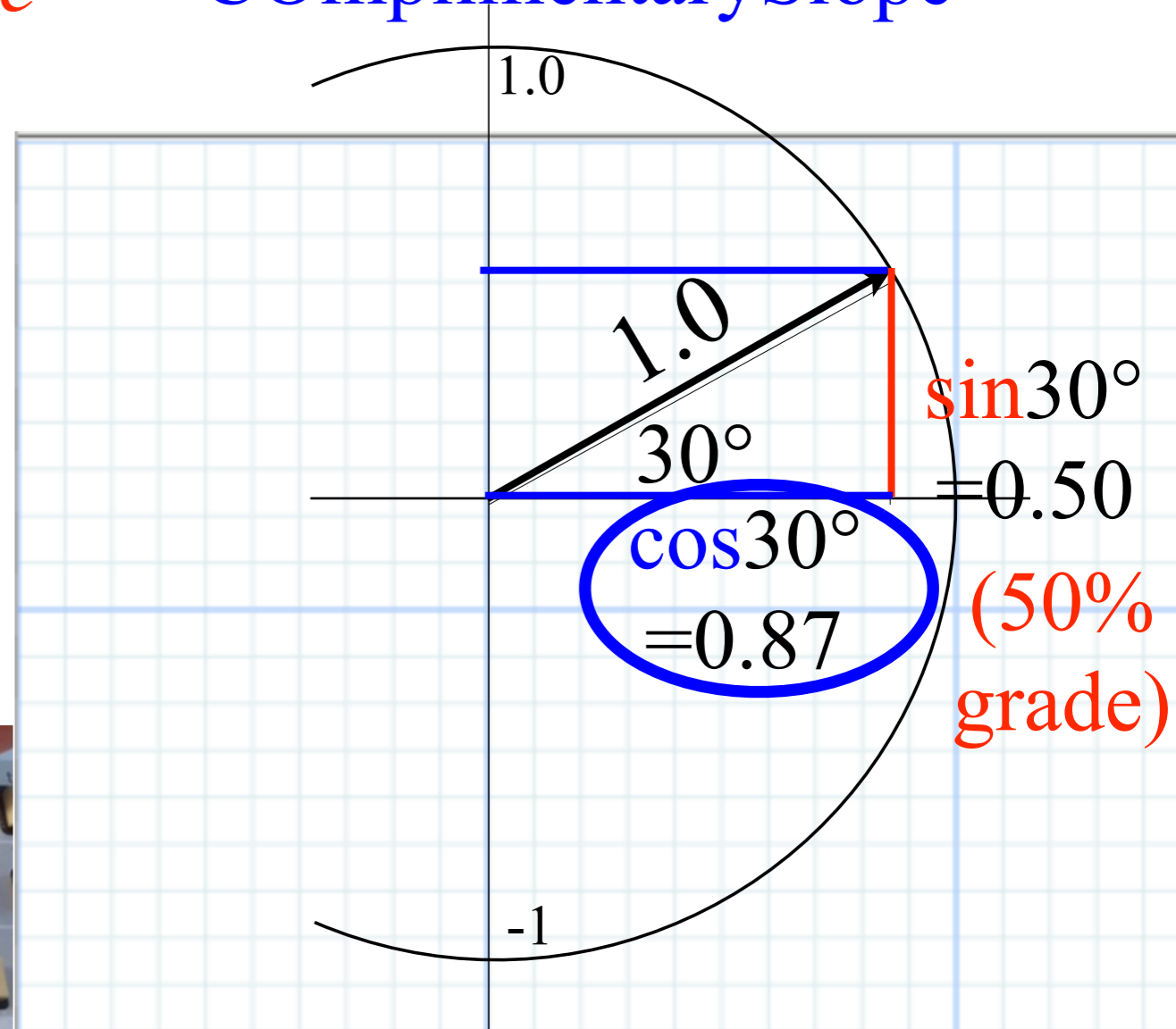
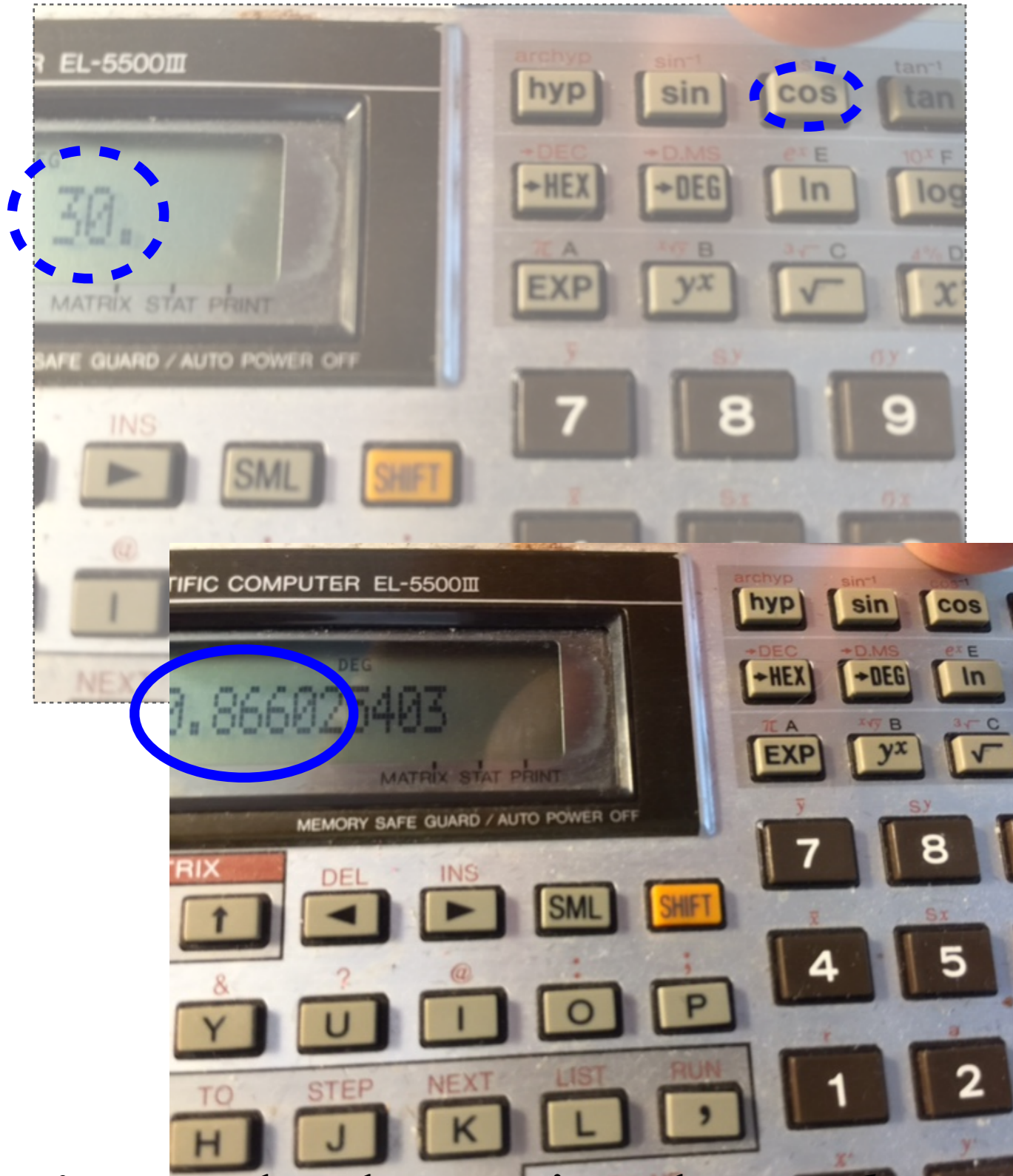
# Learning about SIN “Slope of INcline”



It's mostly about triangles *and sine-waves*



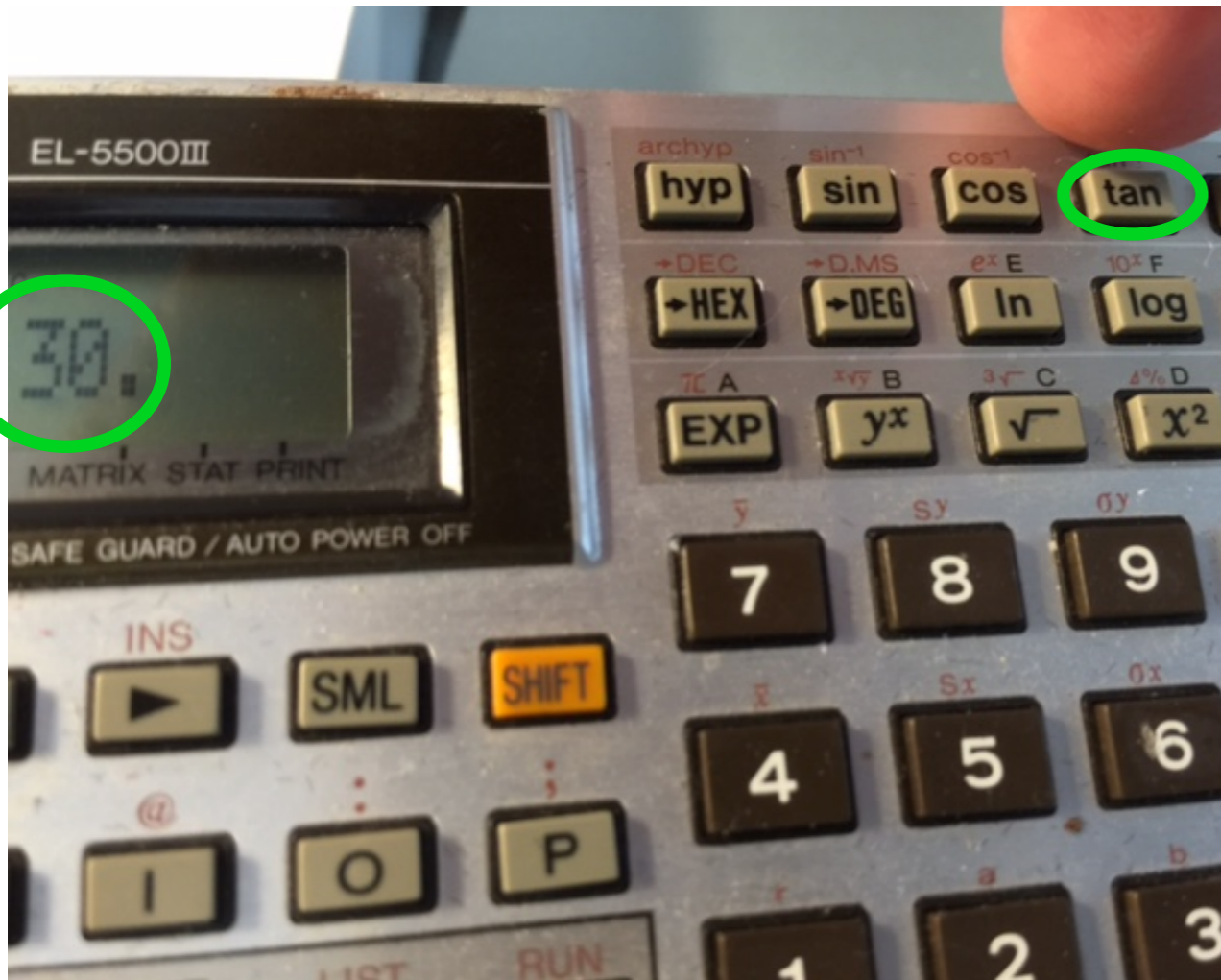
# Learning about **SIN** and the **COS**in “*Slope of INcline*” “**C**Omplimentary**S**lope”



It's mostly about triangles *and sine-waves and cosine-waves*



# Learning about **SIN** and the **COS**in and **TAN**gent and **CO**Tangent *“Slope of INcline”*

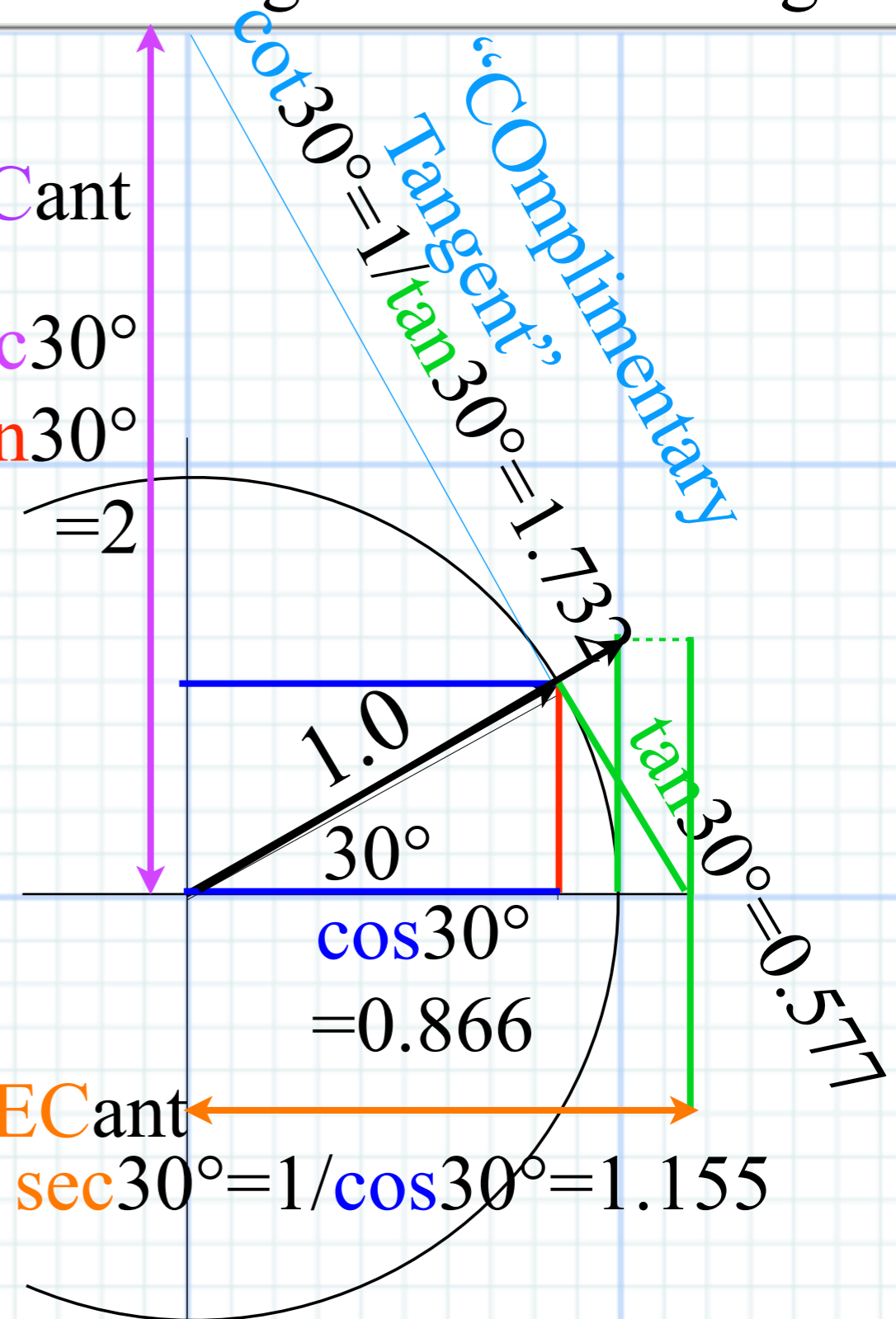


...and  
**CoSeCant**

$$\text{csc}30^\circ = 1/\text{sin}30^\circ = 2$$

...and **SECant**

$$\text{sec}30^\circ = 1/\text{cos}30^\circ = 1.155$$



Fundamental relativity and quantum wave mechanics  
 is mostly about triangles *and sine-waves and cosine-waves*

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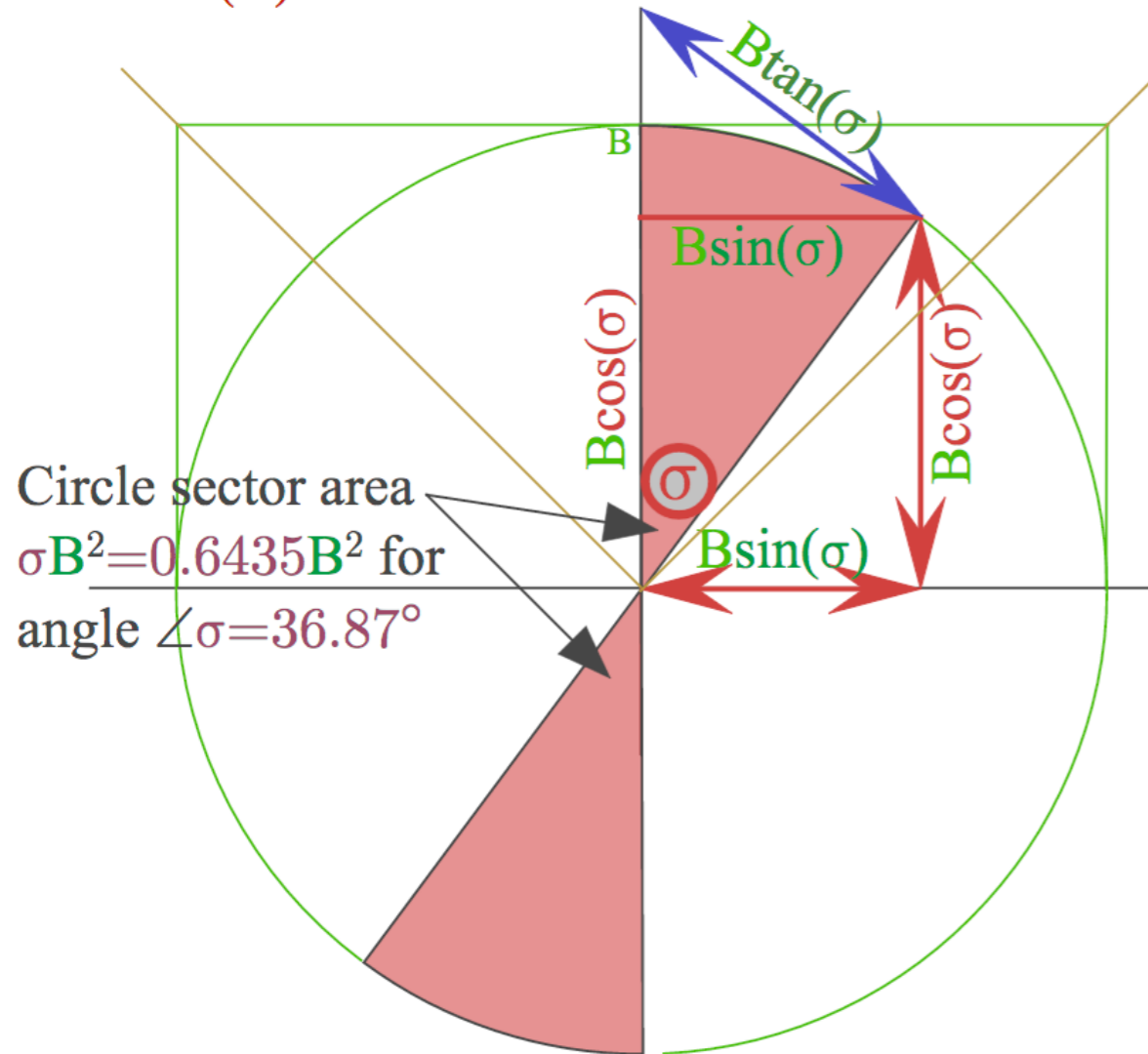
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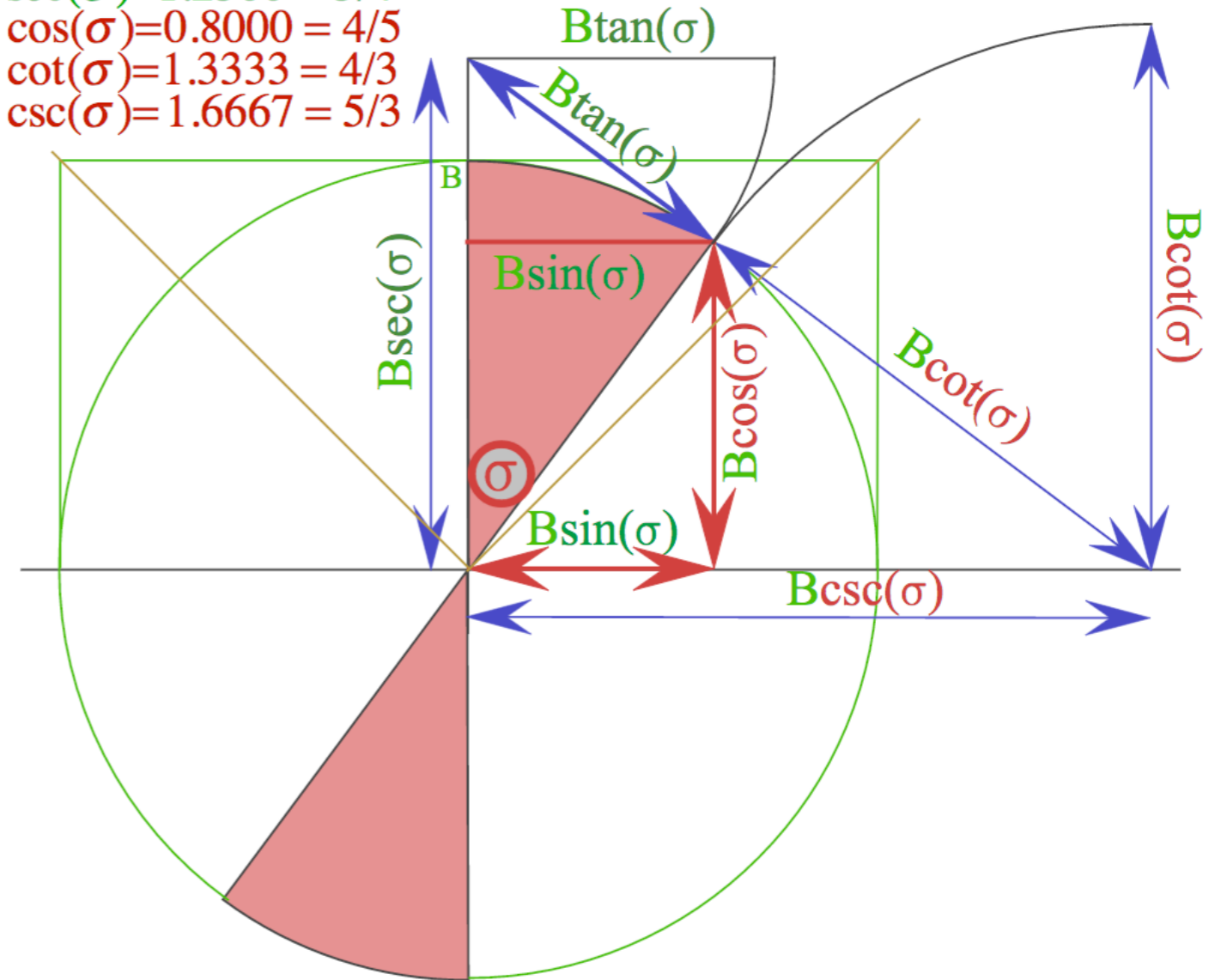
Application to TE-Waveguide modes and synchrotron beam relativity

# Trigonometric road maps

(a)  $\sin(\sigma) = 0.6000 = 3/5$   
 $\tan(\sigma) = 0.7500 = 3/4$   
 $\cos(\sigma) = 0.8000 = 4/5$



(b)  $\sin(\sigma) = 0.6000 = 3/5$   
 $\tan(\sigma) = 0.7500 = 3/4$   
 $\sec(\sigma) = 1.2500 = 5/4$   
 $\cos(\sigma) = 0.8000 = 4/5$   
 $\cot(\sigma) = 1.3333 = 4/3$   
 $\csc(\sigma) = 1.6667 = 5/3$



*All this physics of relativity  
 is mostly simple trigonometry  
 of optical wave interference!*

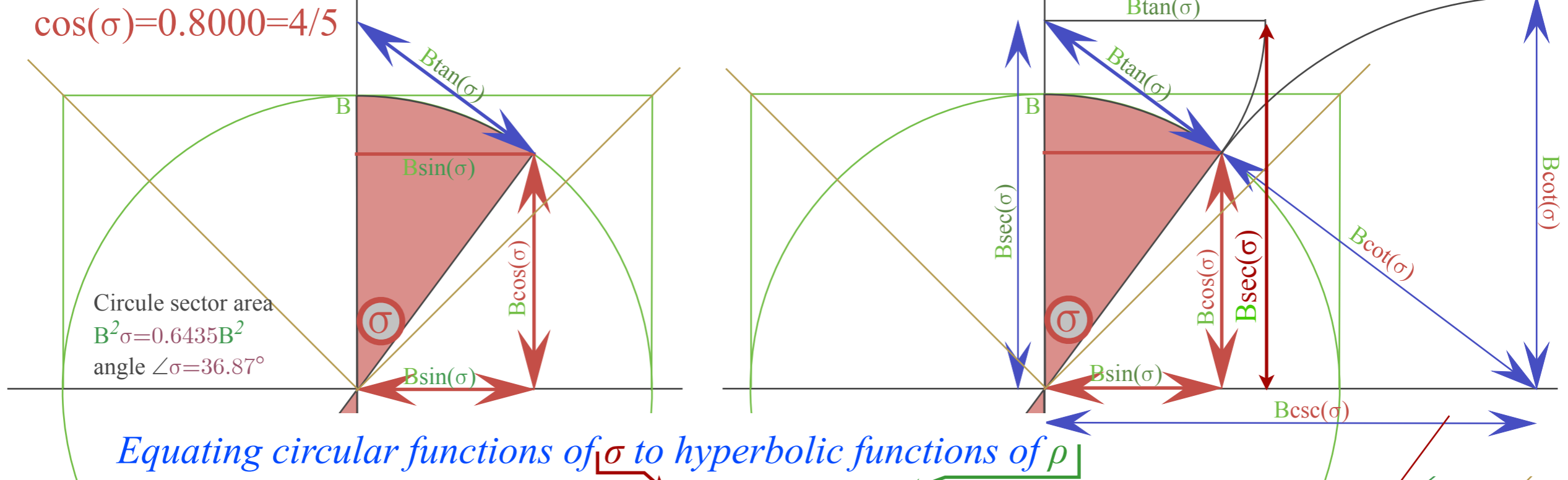
*And, it derives fundamentals  
 of quantum theory, too!*



# Trigonometric road maps become hyperbolic trig maps...

(a)  $\sin(\sigma) = 0.6000 = 3/5$

(b)



Equating circular functions of  $\sigma$  to hyperbolic functions of  $\rho$

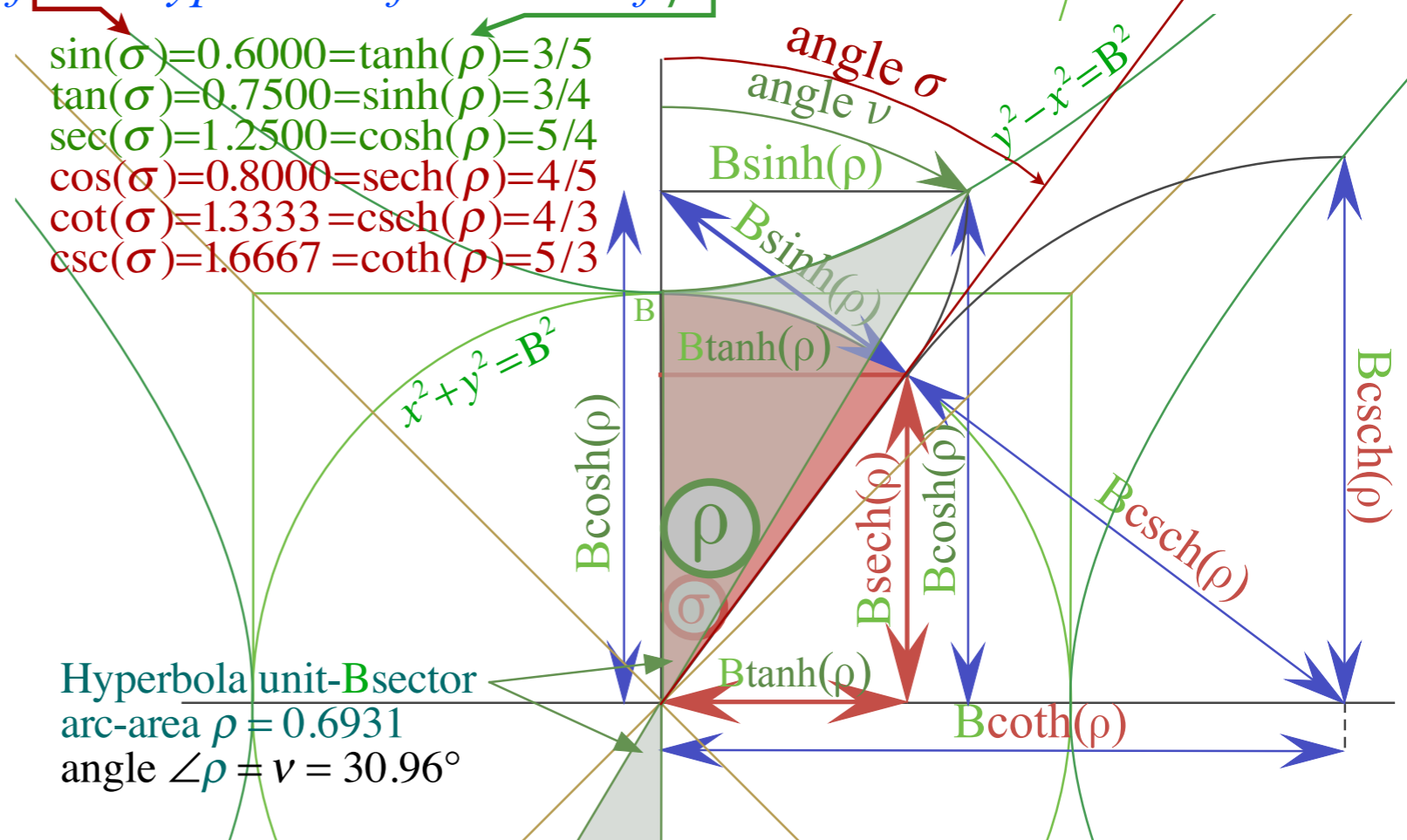
$$\begin{aligned} \sin(\sigma) &= 0.6000 = \tanh(\rho) = 3/5 \\ \tan(\sigma) &= 0.7500 = \sinh(\rho) = 3/4 \\ \sec(\sigma) &= 1.2500 = \cosh(\rho) = 5/4 \\ \cos(\sigma) &= 0.8000 = \operatorname{sech}(\rho) = 4/5 \\ \cot(\sigma) &= 1.3333 = \operatorname{csch}(\rho) = 4/3 \\ \csc(\sigma) &= 1.6667 = \operatorname{coth}(\rho) = 5/3 \end{aligned}$$

[AMOP Ch.0 article p.9.](#)

All this physics of relativity  
 is mostly simple trigonometry  
 of optical wave interference!

And, it derives fundamentals  
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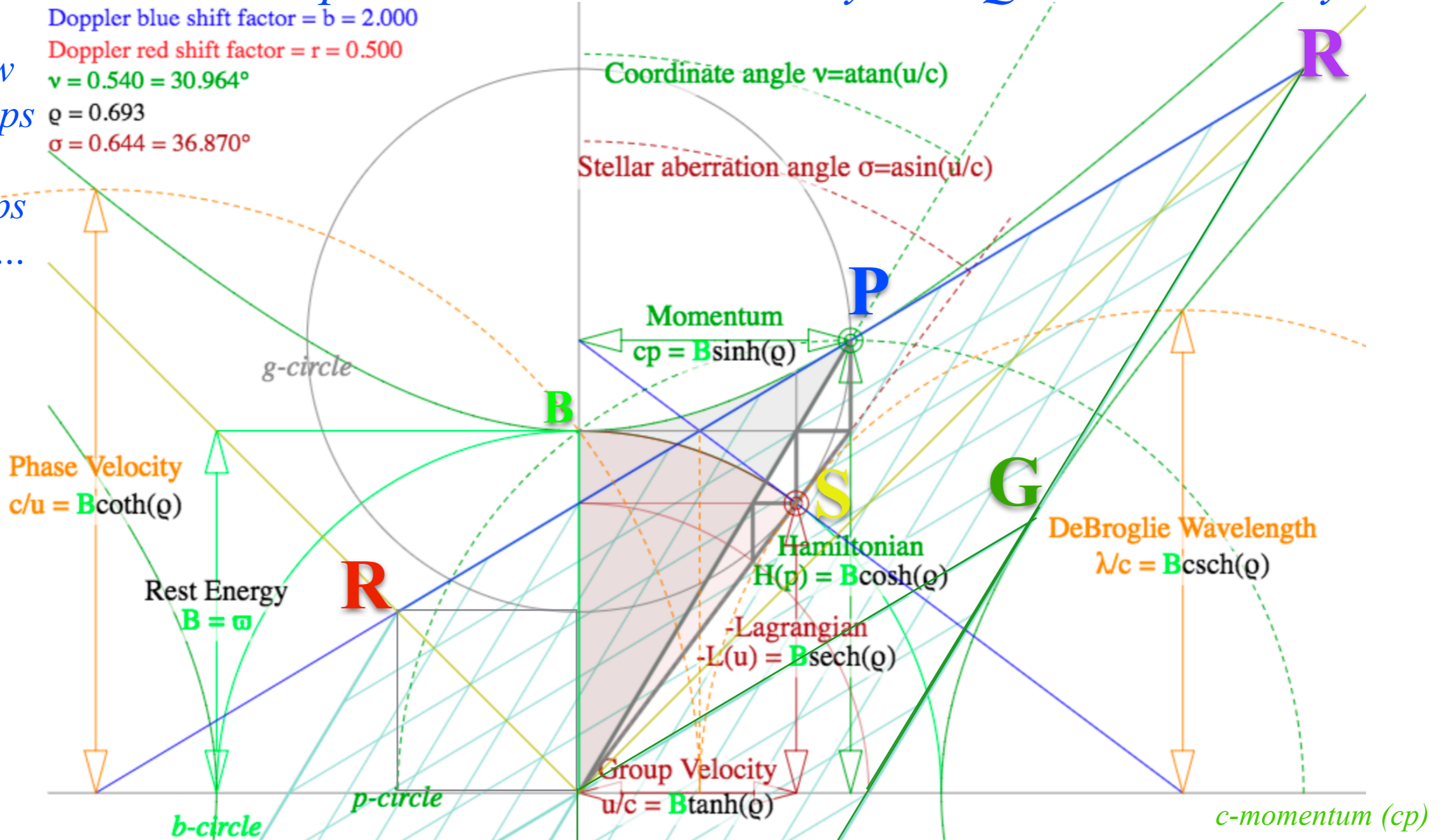
Hyperbola unit-Bsector  
 arc-area  $\rho = 0.6931$   
 angle  $\angle\rho = \nu = 30.96^\circ$





# Trigonometric road maps.... Energy (E) to Relativity and Quantum Theory\*

Need to show trig road maps can match physical maps like this one...



All this physics of relativity is mostly simple trigonometry of optical wave interference.

And, it derives fundamentals of quantum theory, too!

**\*Relativity**

[AMOP Ch.0 article p.20.](#)

[Relativity Web Simulation](#)  
[{Physical Terms - All Terms}](#)



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# Hyper-Trigonometric algebra easily derives Circular-Trigonometric-algebra

Exponential derived by infinite- $n$ -compounding limit of the interest rate- $r$  formula.

$$e^{rt} = \lim_{n \rightarrow \infty} \left( 1 + \frac{rt}{n} \right)^n$$

Infinite- $n$  limit of binomial series is an exponential power- $p$  series of  $(rt)^p$  with  $1/p!$  coefficients.

$$e^{rt} = 1 + rt + \frac{(rt)^2}{2} + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} + \dots$$
$$e^{-rt} = 1 - rt + \frac{(rt)^2}{2} - \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots$$

Half-sum and half difference of  $e^{\pm rt}$  series define the hyperbolic cosine ( $\cosh(rt)$ ) and sine ( $\sinh(rt)$ ).

$$\frac{e^{+rt} + e^{-rt}}{2} = 1 + \frac{(rt)^2}{2} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cosh(rt)$$
$$\frac{e^{+rt} - e^{-rt}}{2} = rt + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \dots = \sinh(rt)$$

*Hyper-Trig*  
 $\cosh \rho$  and  $\sinh \rho$



# Hyper-Trigonometric algebra easily derives Circular-Trigonometric-algebra

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$$e^{-rt} = 1 - rt + \frac{(rt)^2}{2} - \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cosh(rt) - \sinh(rt)$$

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$$\frac{e^{+rt} + e^{-rt}}{2} = 1 + \frac{(rt)^2}{2} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cosh(rt)$$

$$\frac{e^{+rt} - e^{-rt}}{2} = rt + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \dots = \sinh(rt)$$

*Hyper-Trig*  
 $\cosh \rho$  and  $\sinh \rho$

Replace rate  $r$  with imaginary rate  $ir$  and  $i = \sqrt{-1}$  powers  $i^0=1, i^1=i, i^2=-1, i^3=-i, i^4=1, i^5=i, i^6=-1, i^7=-i, \dots$

Then *hyper*-sine-cosine becomes the *circular*-sine-cosine.

$$\frac{e^{+i rt} + e^{-i rt}}{2} = 1 - \frac{(rt)^2}{2} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cos rt$$

$$\frac{e^{+i rt} - e^{-i rt}}{2} = i rt - i \frac{(rt)^3}{2 \cdot 3} + i \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} - \dots = i \sin rt$$

*Circular-Trig*  
 $\cos \sigma$  and  $\sin \sigma$

Sum and difference of this pair gives the Euler-DeMoivre relations of exponentials vs trig-functions.

$$e^{+i\sigma} = \cos \sigma + i \sin \sigma ,$$

$$e^{+\rho} = \cosh \rho + \sinh \rho ,$$

$$e^{-i\sigma} = \cos \sigma - i \sin \sigma .$$

$$e^{-\sigma} = \cosh \rho - \sinh \rho .$$

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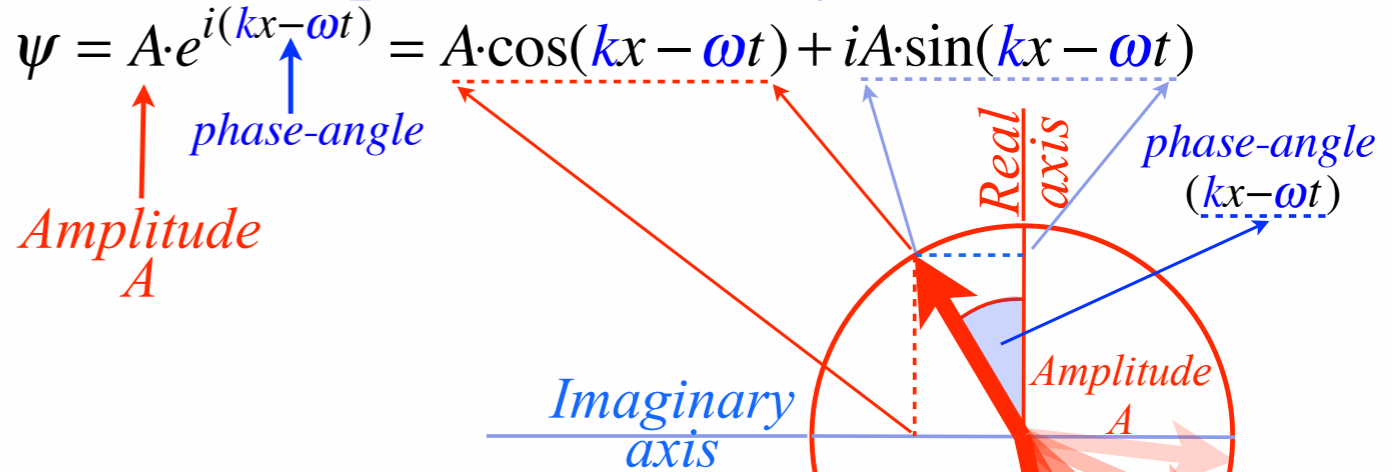
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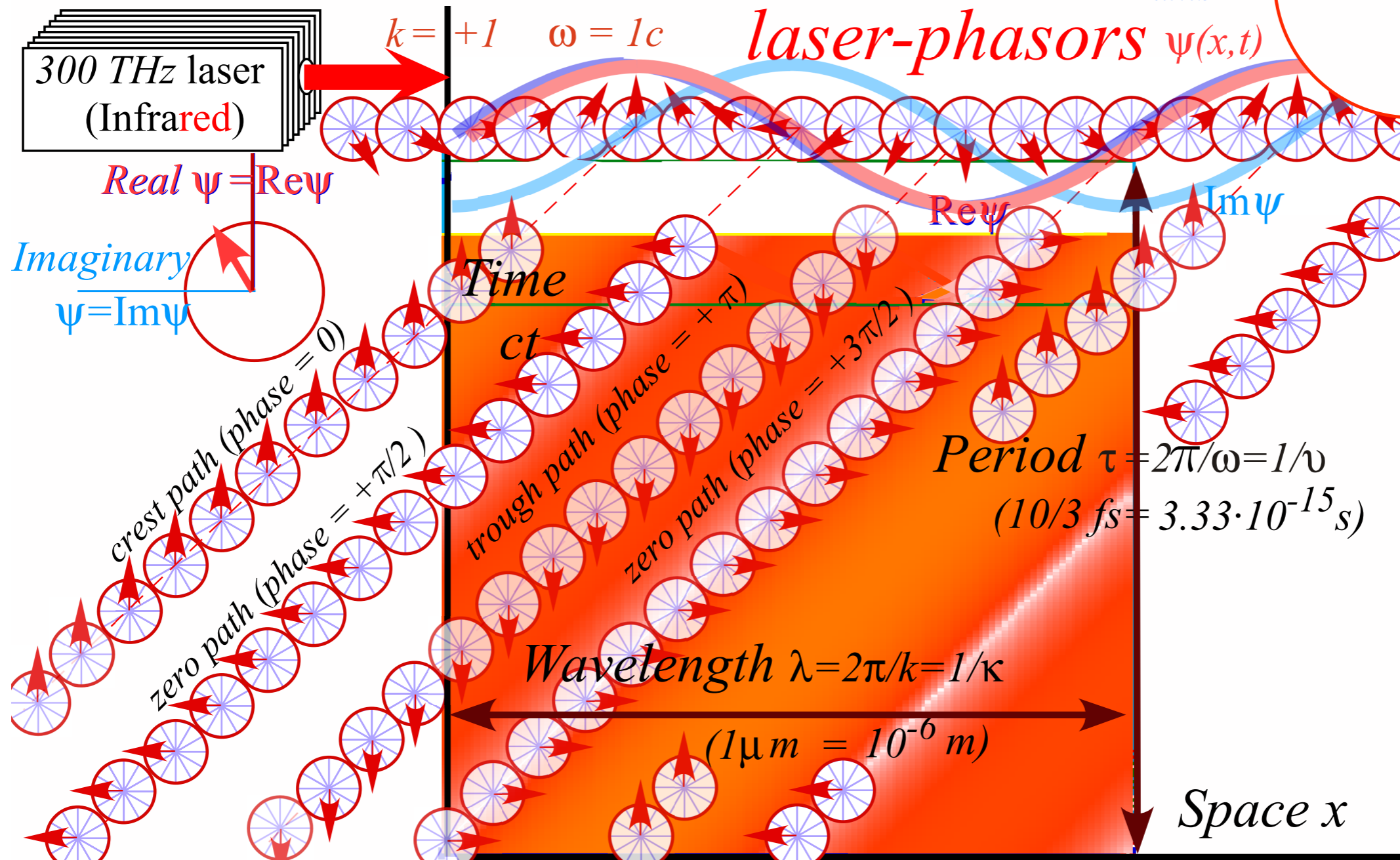
Application to TE-Waveguide modes and synchrotron beam relativity

# 1CW Laser-phasor wave function

$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$



Hyper-Trigonometric phasors in space-time



(a) Single-phasor plot of wave-function at  $(x, ct)$ . (b) Array of phasors at many  $(x, ct)$ -points.



# 1CW Laser-phasor wave function

Dimensionless Light wave-velocity  $c/c=1$

$$\frac{v_{\text{light}}}{c} = \frac{\lambda}{c\tau} = \frac{v}{c\kappa} = 1 = \frac{\omega \text{ angular}}{ck \text{ units}}$$

“winks”  
“n”  
“kinks”

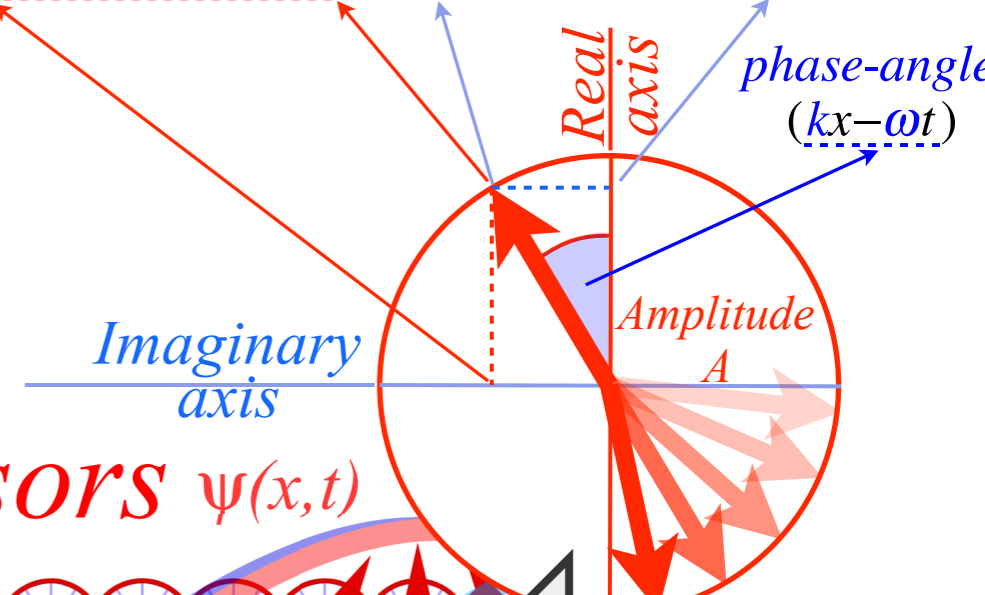
angular frequency:  $\omega = 2\pi\nu$

angular wave number:  $k = 2\pi\kappa$

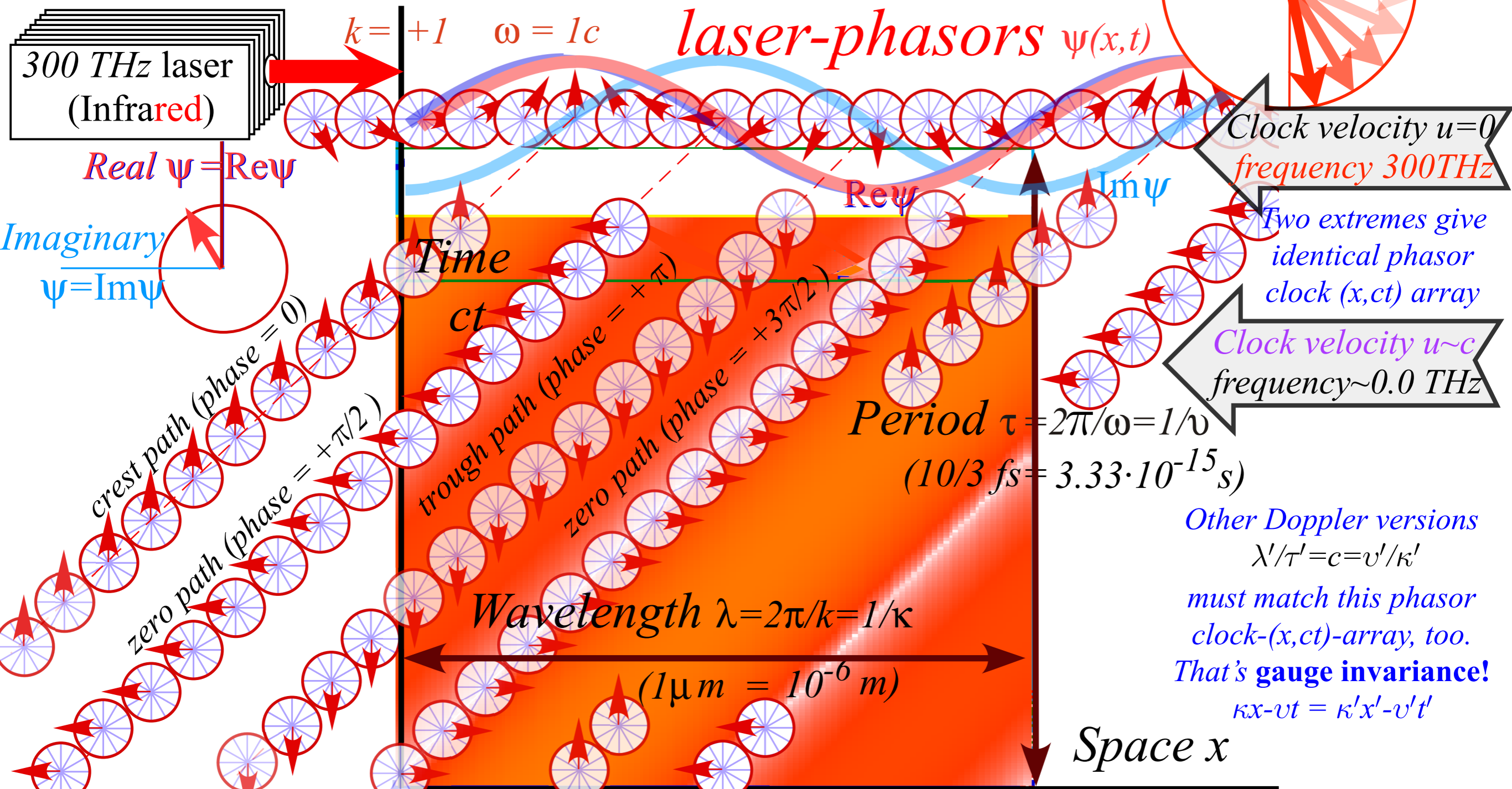
$k = \text{wavevector}$

$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$

Amplitude  $A$   
phase-angle  
 $(kx - \omega t)$



*laser-phasors*  $\psi(x,t)$



Clock velocity  $u=0$   
frequency 300THz

Two extremes give  
identical phasor  
clock  $(x,ct)$  array

Clock velocity  $u \sim c$   
frequency  $\sim 0.0$  THz

Period  $\tau = 2\pi/\omega = 1/\nu$   
(10/3 fs =  $3.33 \cdot 10^{-15}$  s)

Other Doppler versions  
 $\lambda'/\tau' = c = v'/\kappa'$   
must match this phasor  
clock- $(x,ct)$ -array, too.  
**That's gauge invariance!**  
 $\kappa x - \nu t = \kappa' x' - \nu' t'$

Wavelength  $\lambda = 2\pi/k = 1/\kappa$   
(1  $\mu\text{m} = 10^{-6}$  m)

*Why Men in Black shot little Suzie...Learning about sin!, cos and...Trigonometric road maps*

Hyper-Trigonometric *Relativity* geometry and Euler exponential algebra

1CW wavefunctions and phasors

➔ Per-space-per-time vs Space-time (How to understand wave parameters)

Wave velocity formulas

Introducing Doppler shifting

Why is  $c$  so constant?!

Introducing Doppler Arithmetic and *Rapidity*  $\rho$

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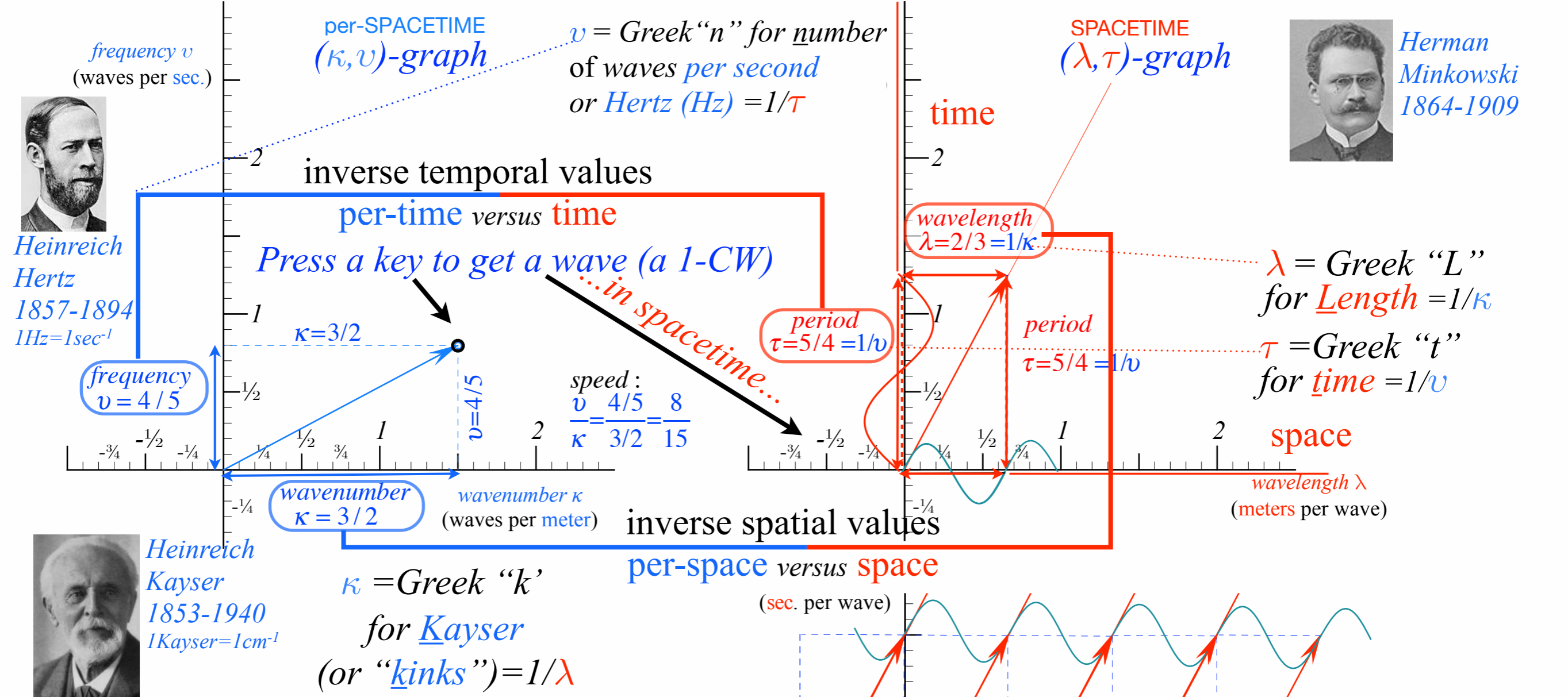
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# The "Keyboard of the gods" : per-space-per-time plot versus space-time Minkowski plot



Per-space-per-time vs Space-time

"Keyboard of the gods" known as "Fourier-space"



**Jean-Baptiste Joseph Fourier**  
 1768-1830

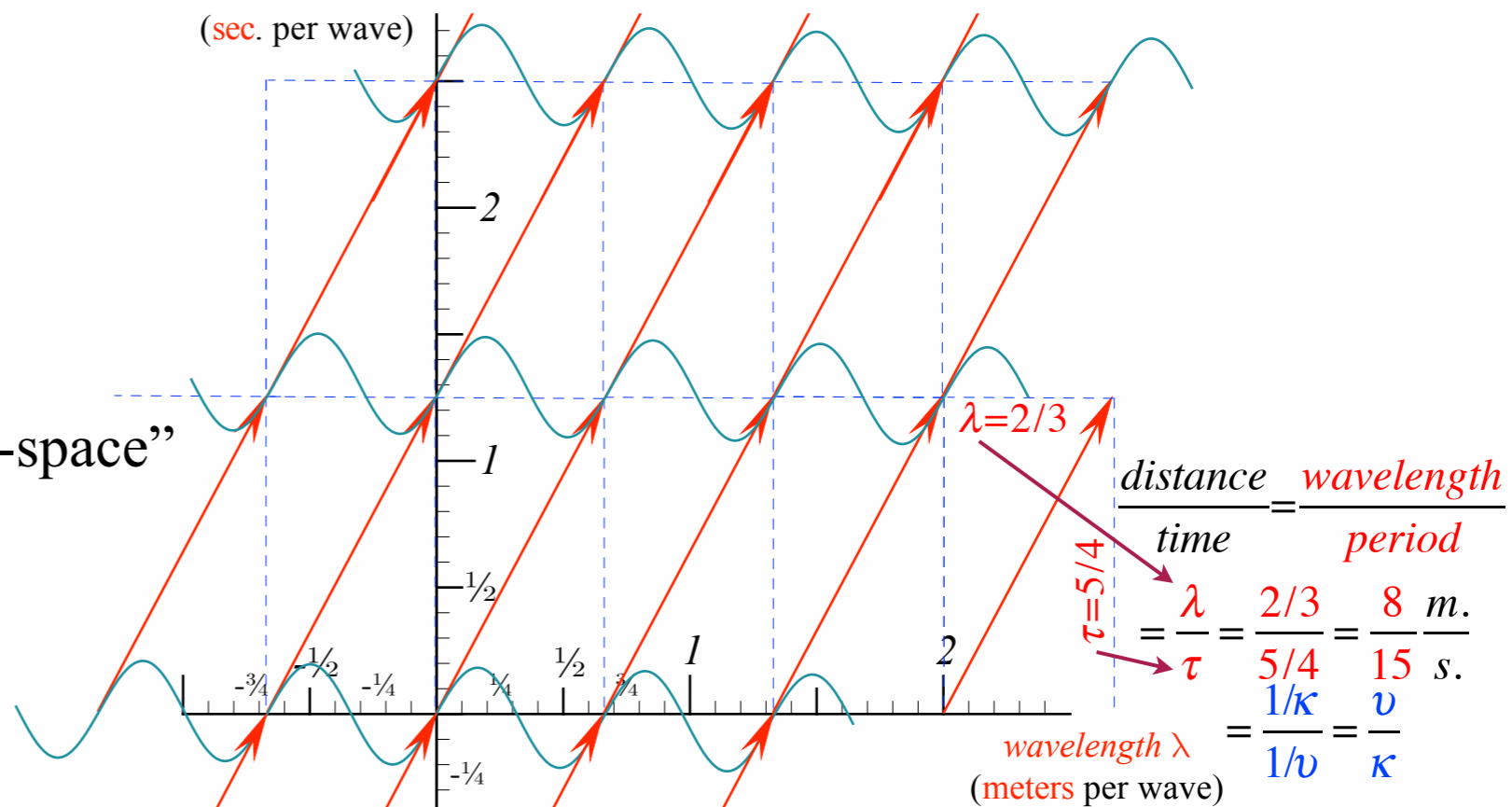
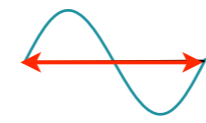
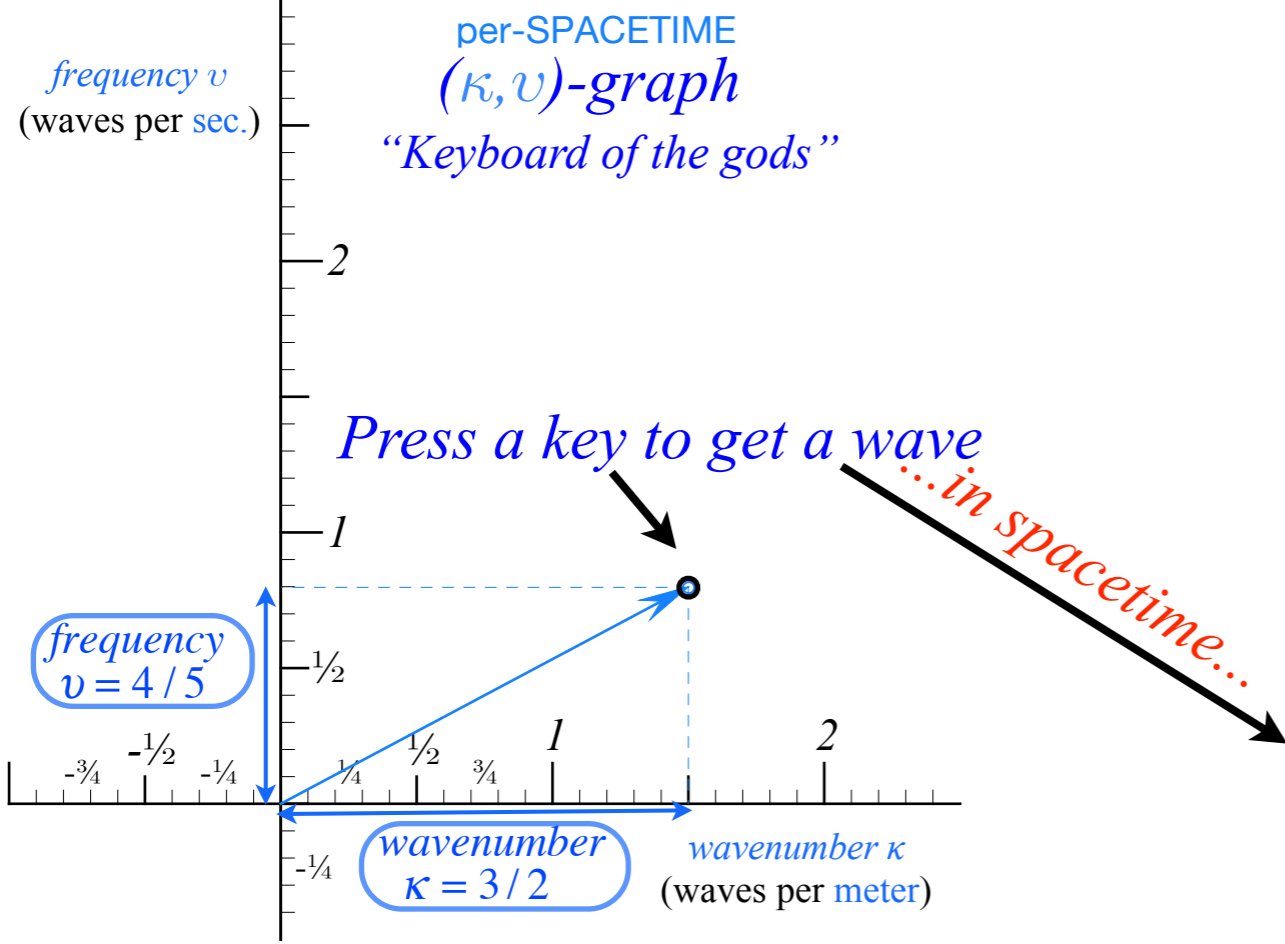
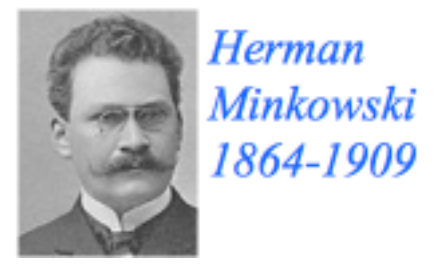


Fig. 5 Comparing a wave point in Kaiser-Hertz per-space-time to its Minkowski space-time view.



Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



"Keyboard of the gods" is known as "Fourier-space"

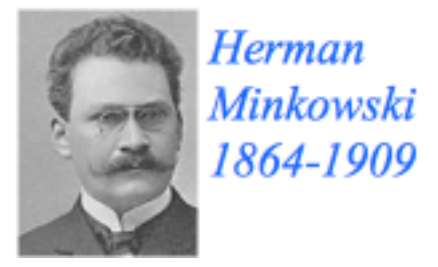


- How to understand waves and wave parameters
- |                      |                      |
|----------------------|----------------------|
| wave frequency $\nu$ | wave period $\tau$   |
| wavenumber $\kappa$  | wavelength $\lambda$ |

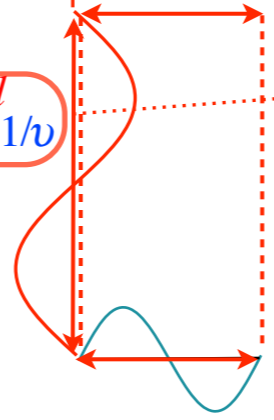
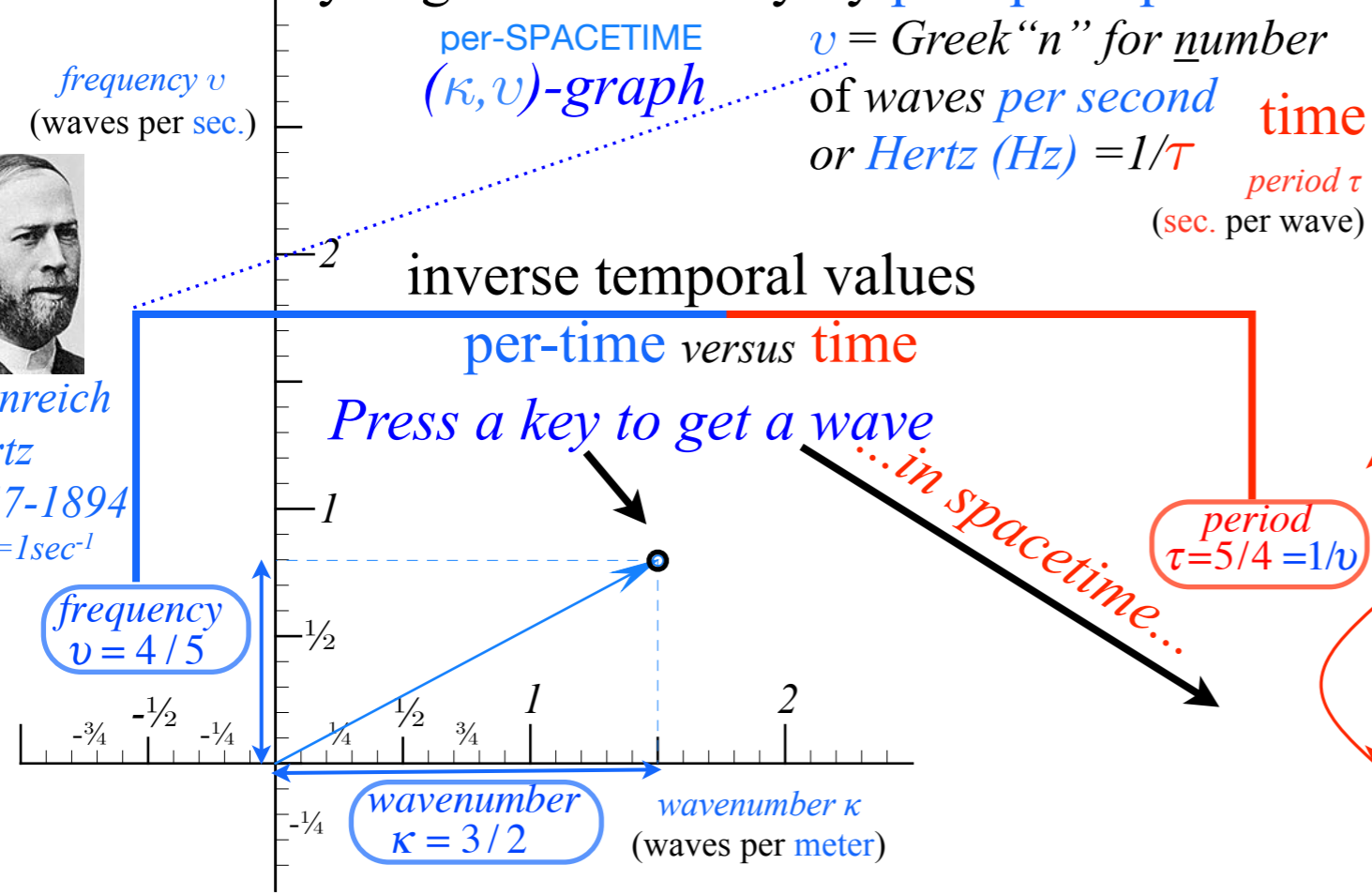
[RelaWavity Web Simulation](#)  
[Keyboard of the Gods \(per-Time vs per-Space\)](#)



# Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



Heinrich Hertz  
1857-1894  
1Hz=1sec<sup>-1</sup>



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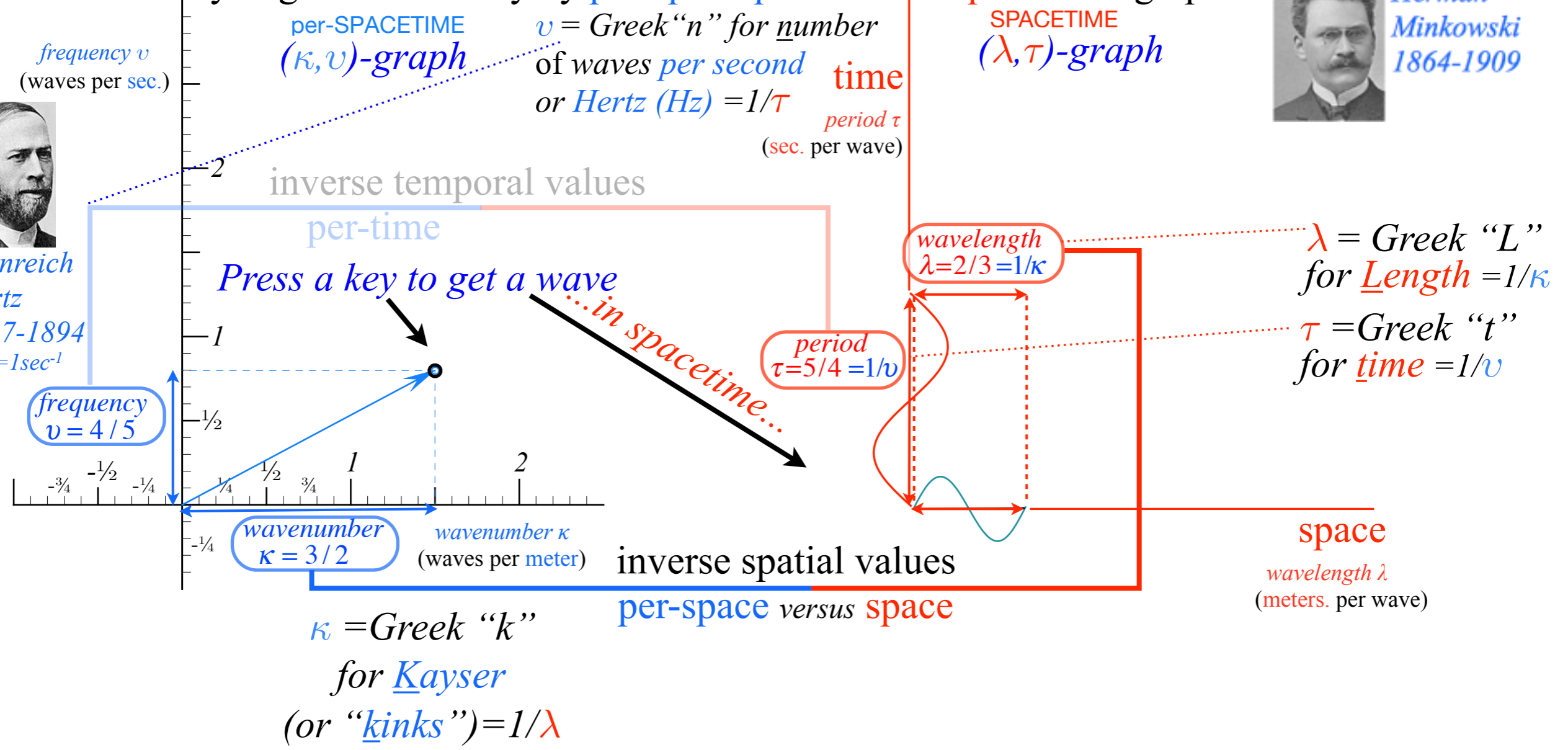
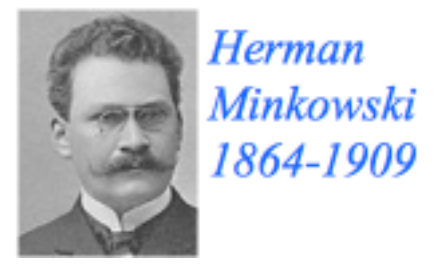


Jean-Baptiste Joseph Fourier  
1768-1830

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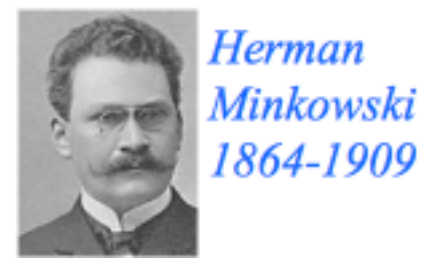


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# Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



frequency  $\nu$   
(waves per sec.)

per-SPACETIME  
 $(\kappa, \nu)$ -graph

$\nu$  = Greek "n" for number  
of waves per second  
or Hertz (Hz) =  $1/\tau$   
time  
period  $\tau$   
(sec. per wave)

SPACETIME  
 $(\lambda, \tau)$ -graph

inverse temporal values

per-time

Press a key to get a wave

...in spacetime...

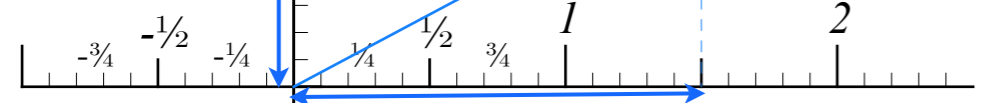
period  
 $\tau = 5/4 = 1/\nu$

wavelength  
 $\lambda = 2/3 = 1/\kappa$

$\lambda$  = Greek "L" for Length =  $1/\kappa$

$\tau$  = Greek "t" for time =  $1/\nu$

frequency  
 $\nu = 4/5$



wavenumber  
 $\kappa = 3/2$

wavenumber  $\kappa$   
(waves per meter)

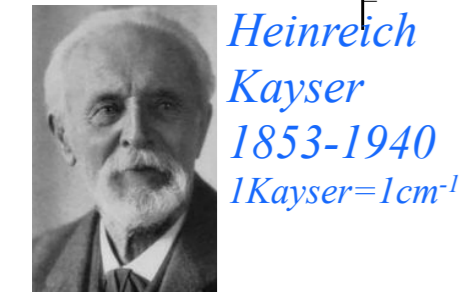
inverse spatial values

per-space versus space

space

wavelength  $\lambda$   
(meters. per wave)

$\kappa$  = Greek "k" for Kayser  
(or "kinks") =  $1/\lambda$



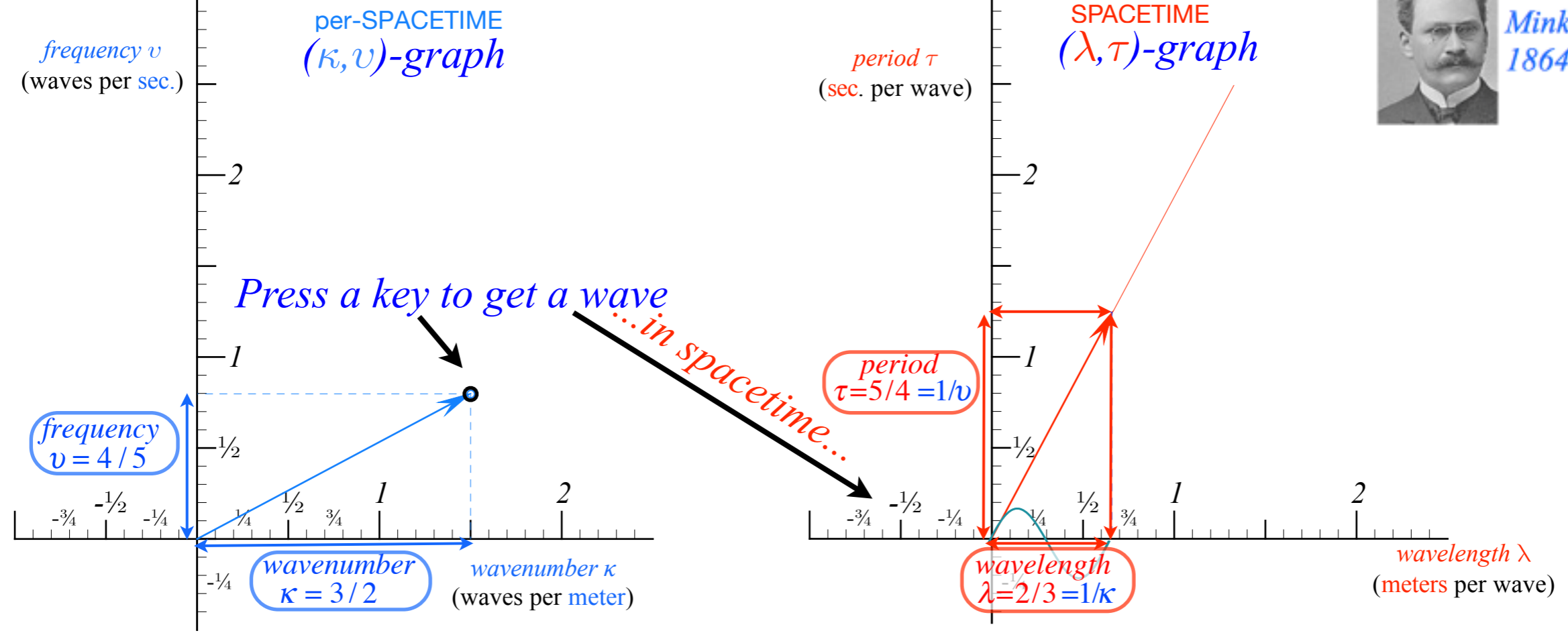
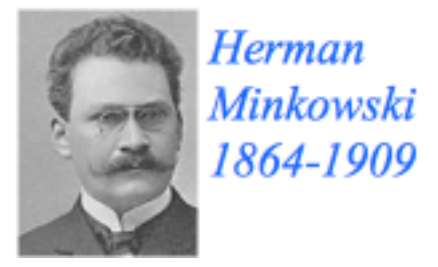
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•How to understand waves and wave parameters

wave frequency  $\nu$       wave period  $\tau$   
wavenumber  $\kappa$         wavelength  $\lambda$

# Analyzing wave velocity by per-space-per-time and space-time graphs



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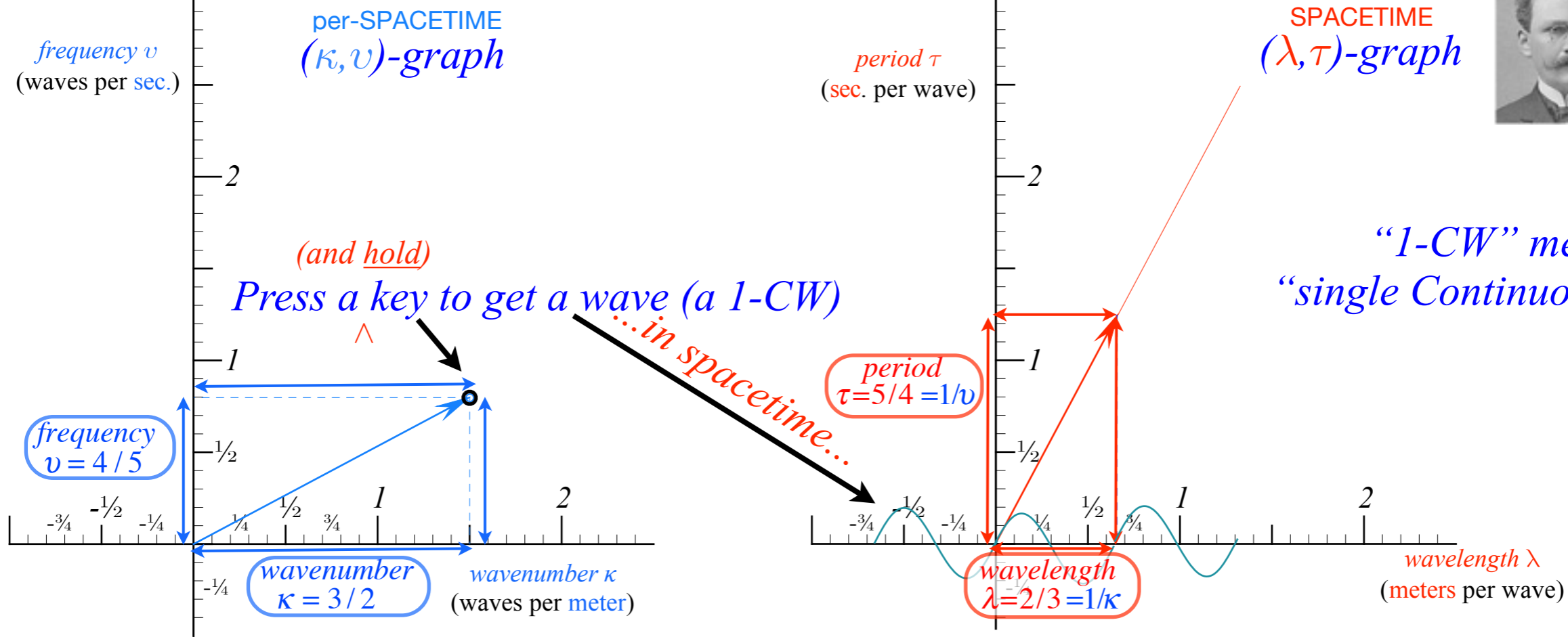
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# Analyzing wave velocity by per-space-per-time and space-time graphs



Herman Minkowski  
1864-1909



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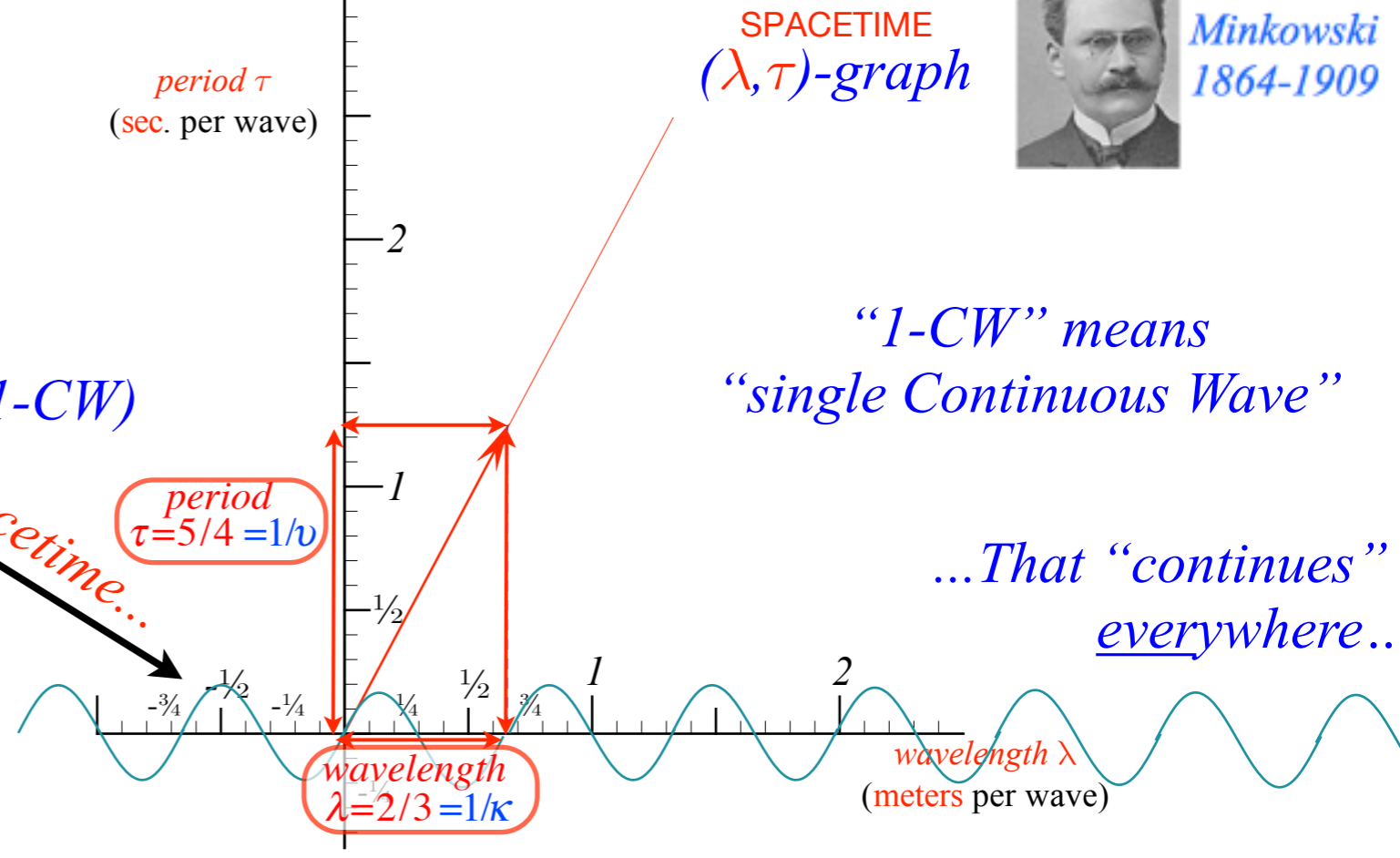
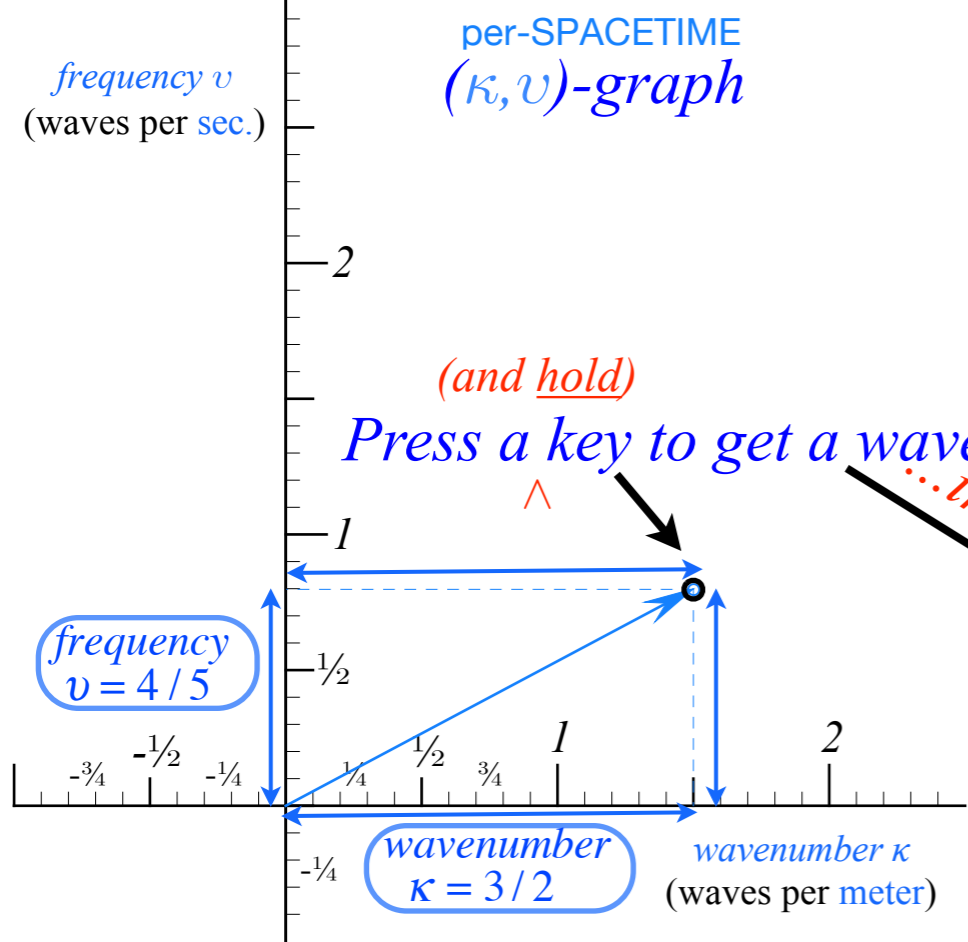
Jean-Baptiste Joseph Fourier  
1768-1830

- How to understand waves and wave parameters
- |                             |                             |
|-----------------------------|-----------------------------|
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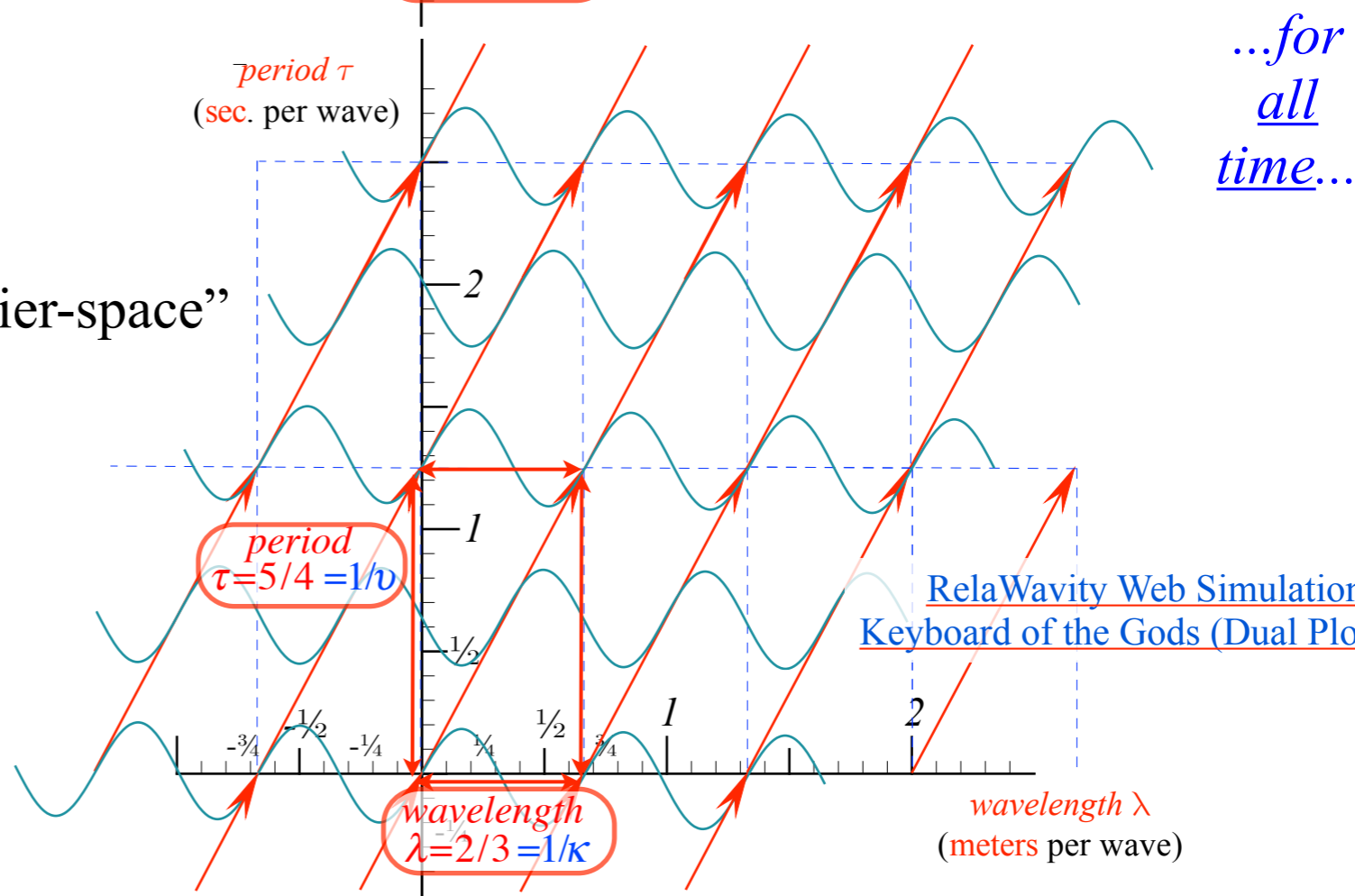
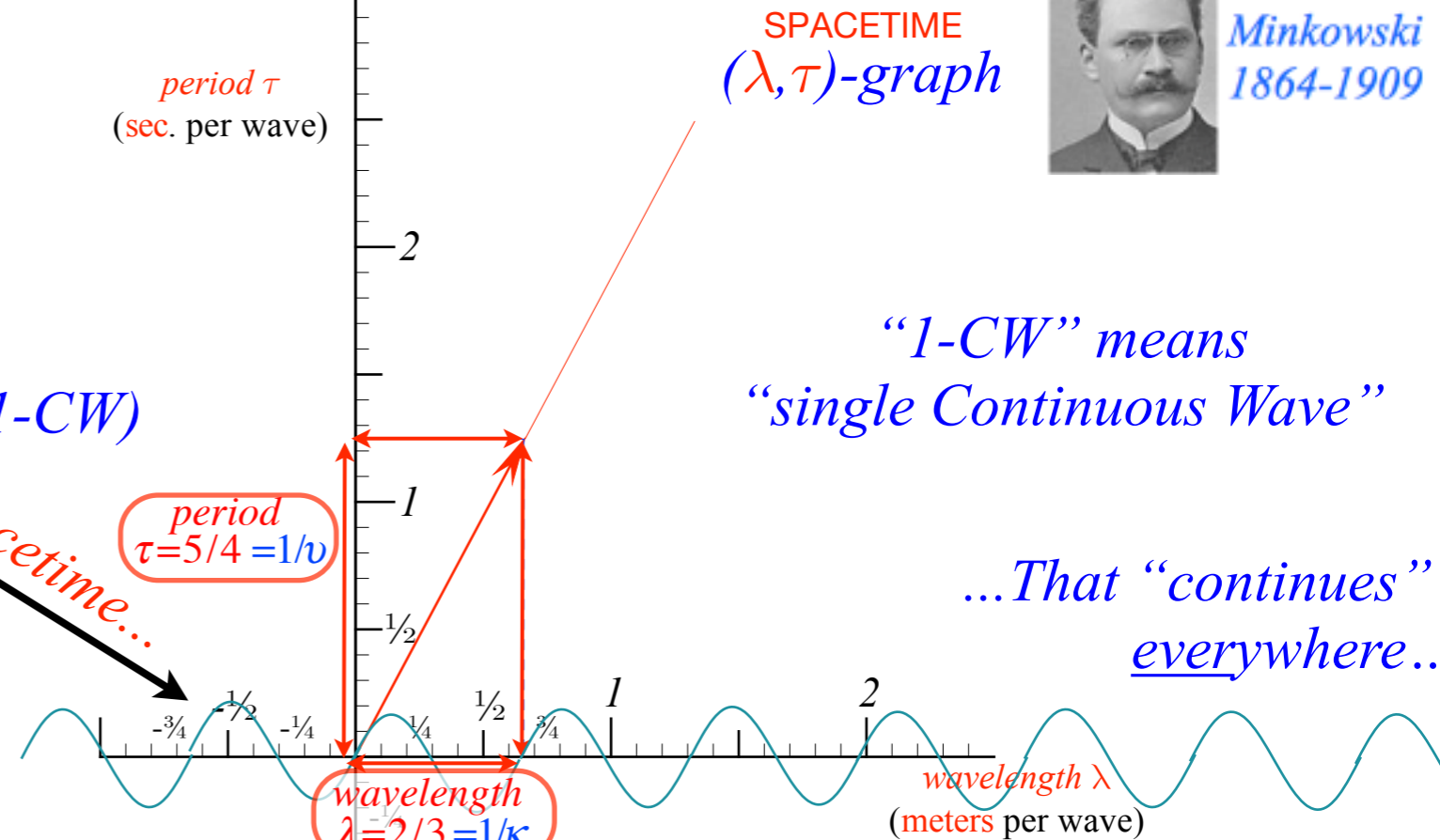
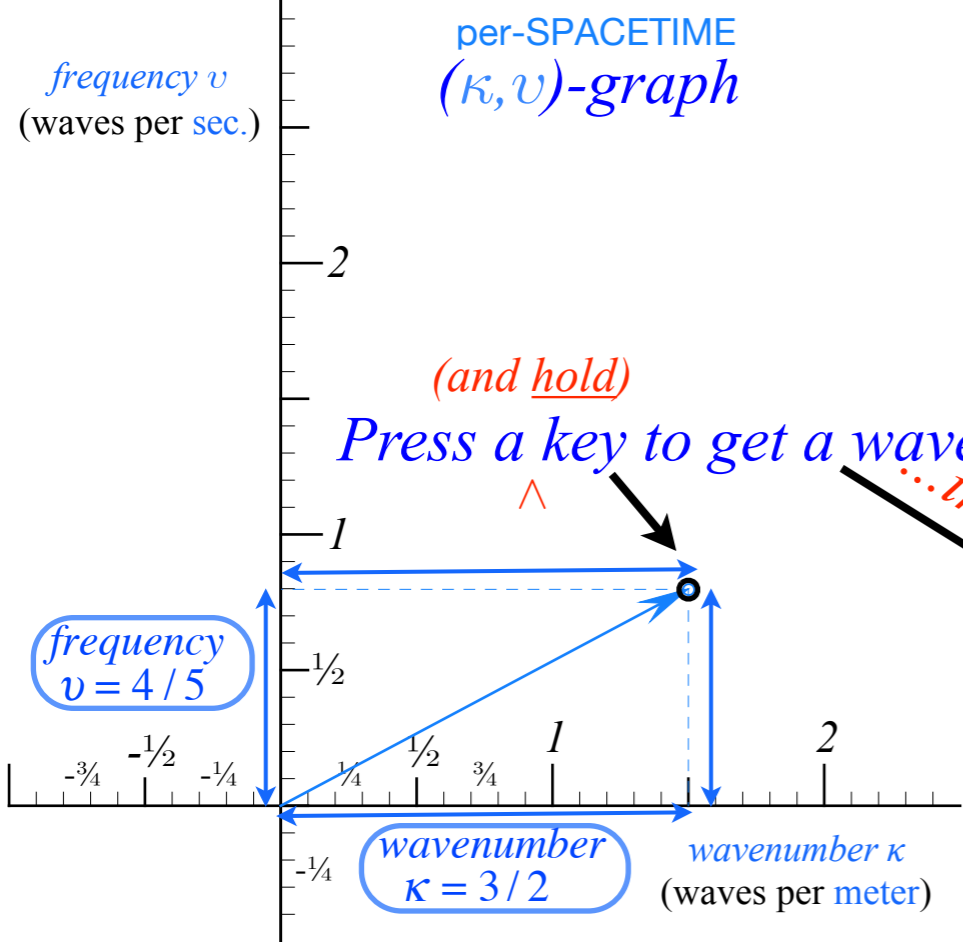
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[RelaWavity Web Simulation Keyboard of the Gods \(Dual Plot #7\)](#)



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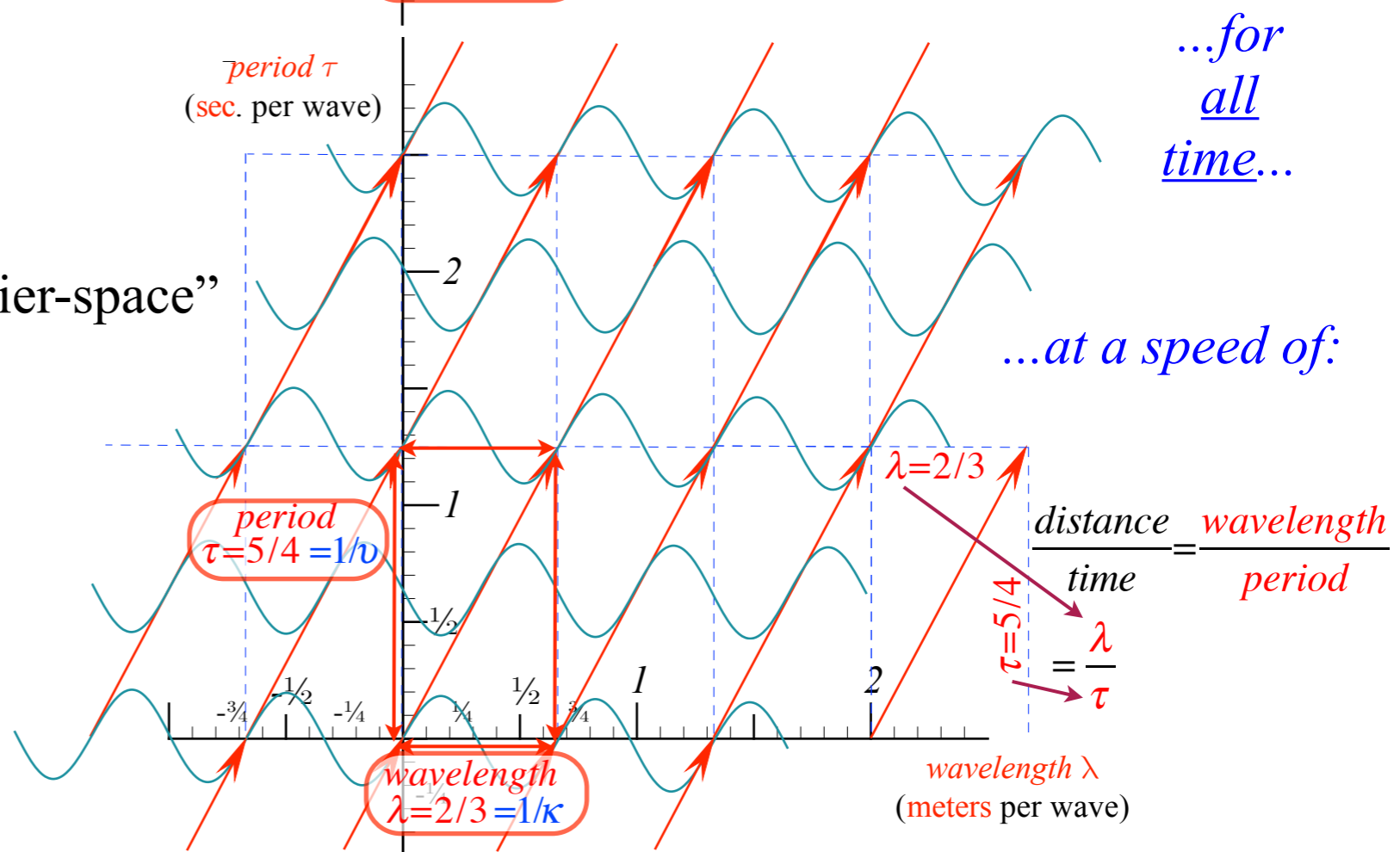
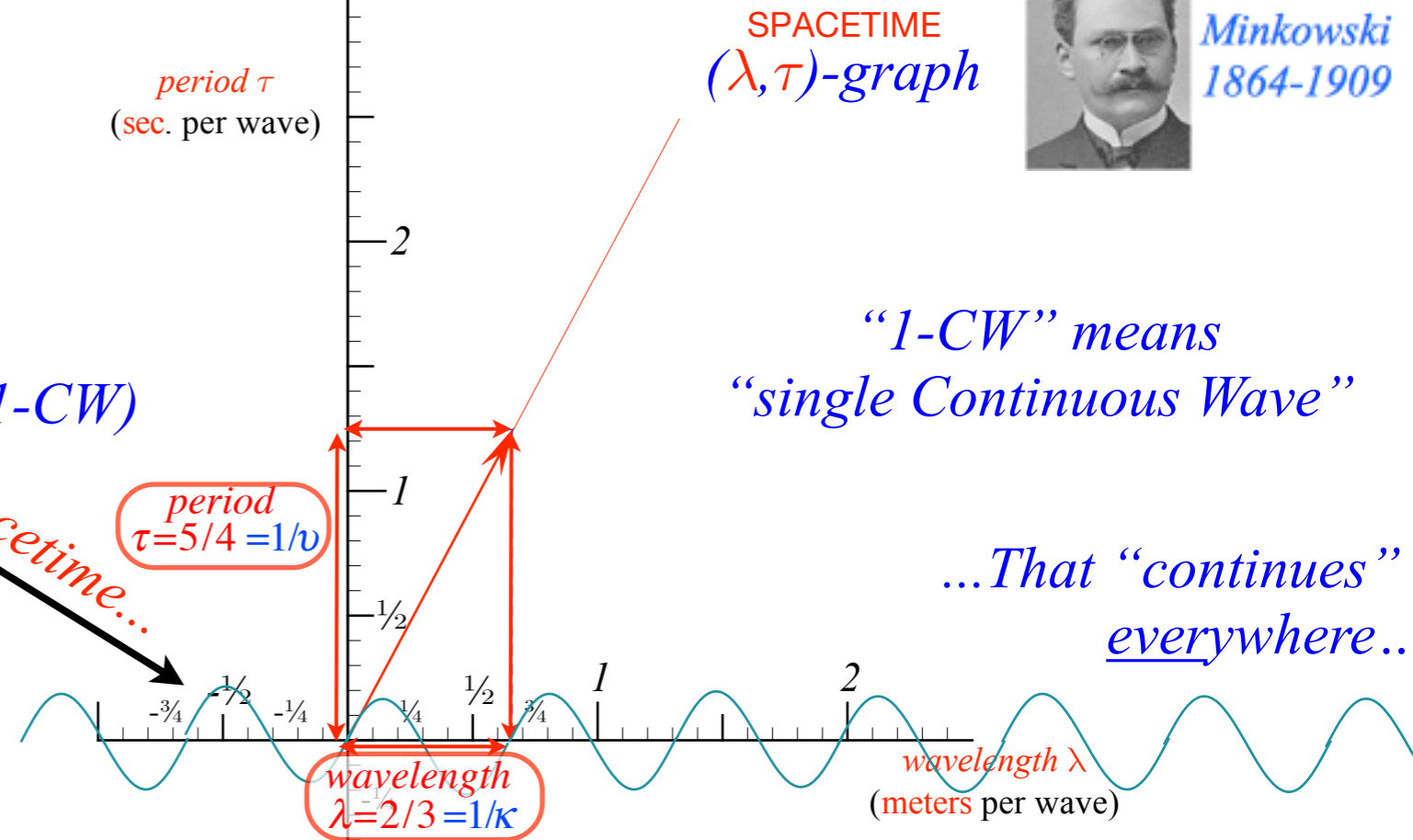
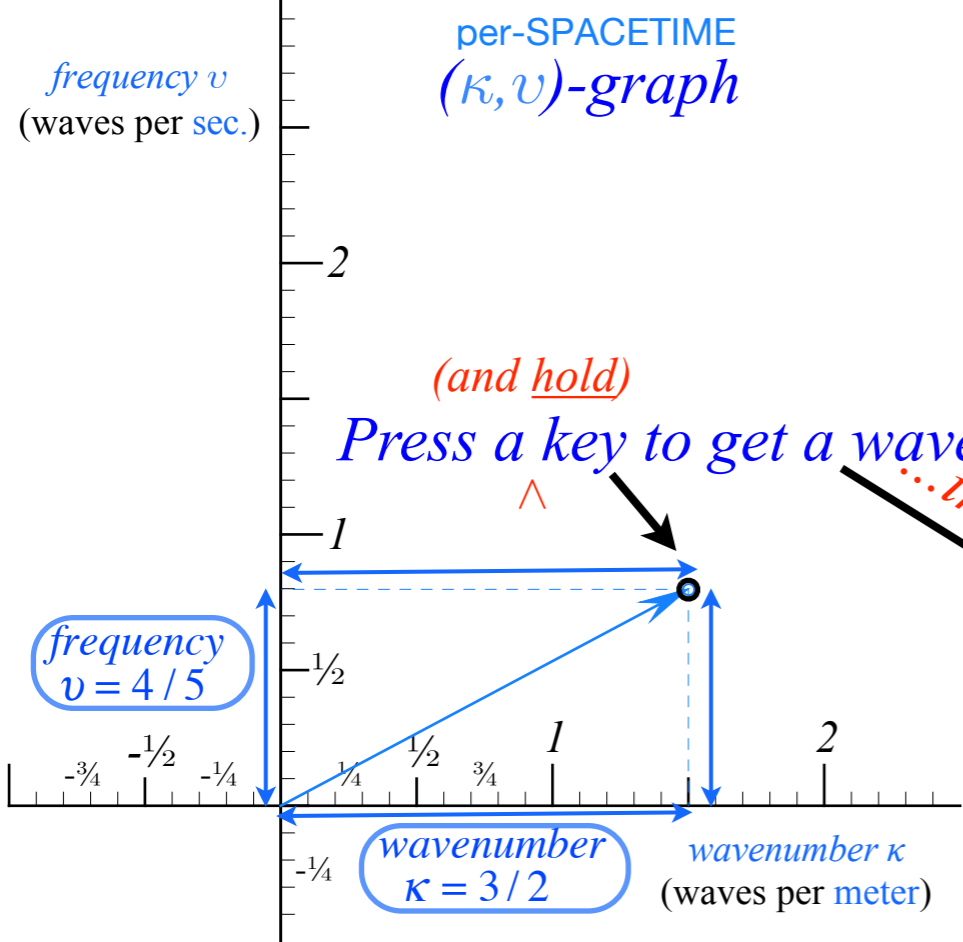
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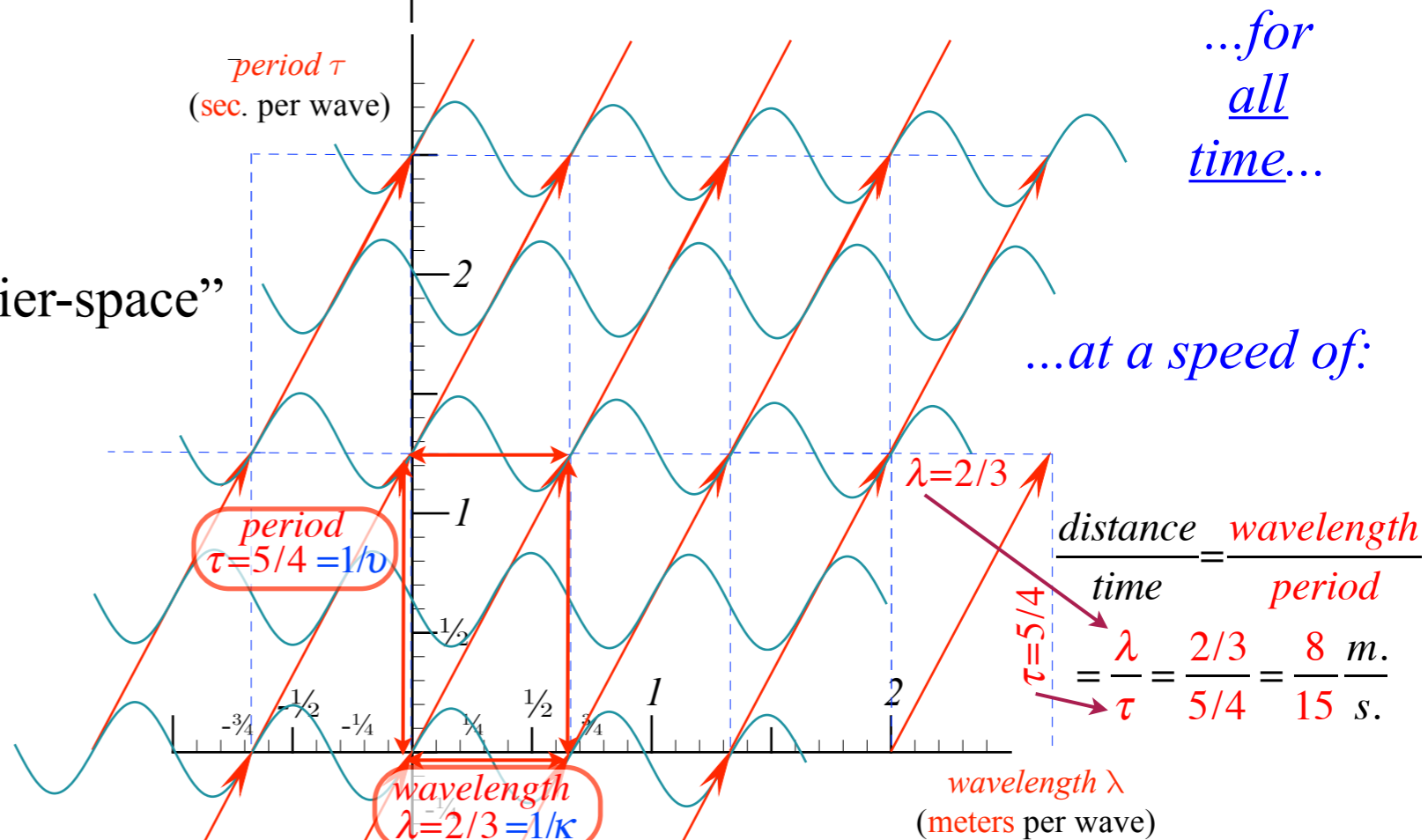
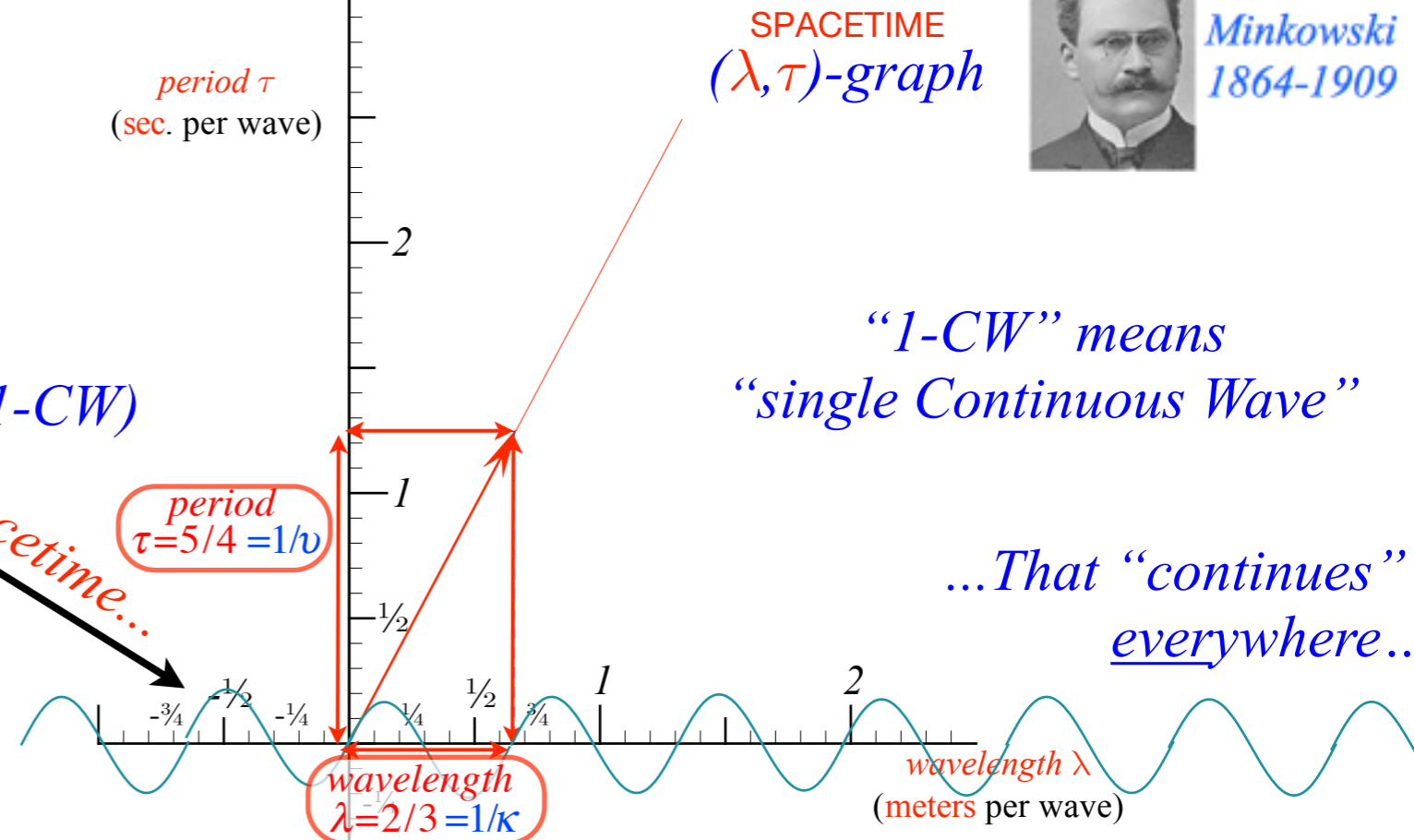
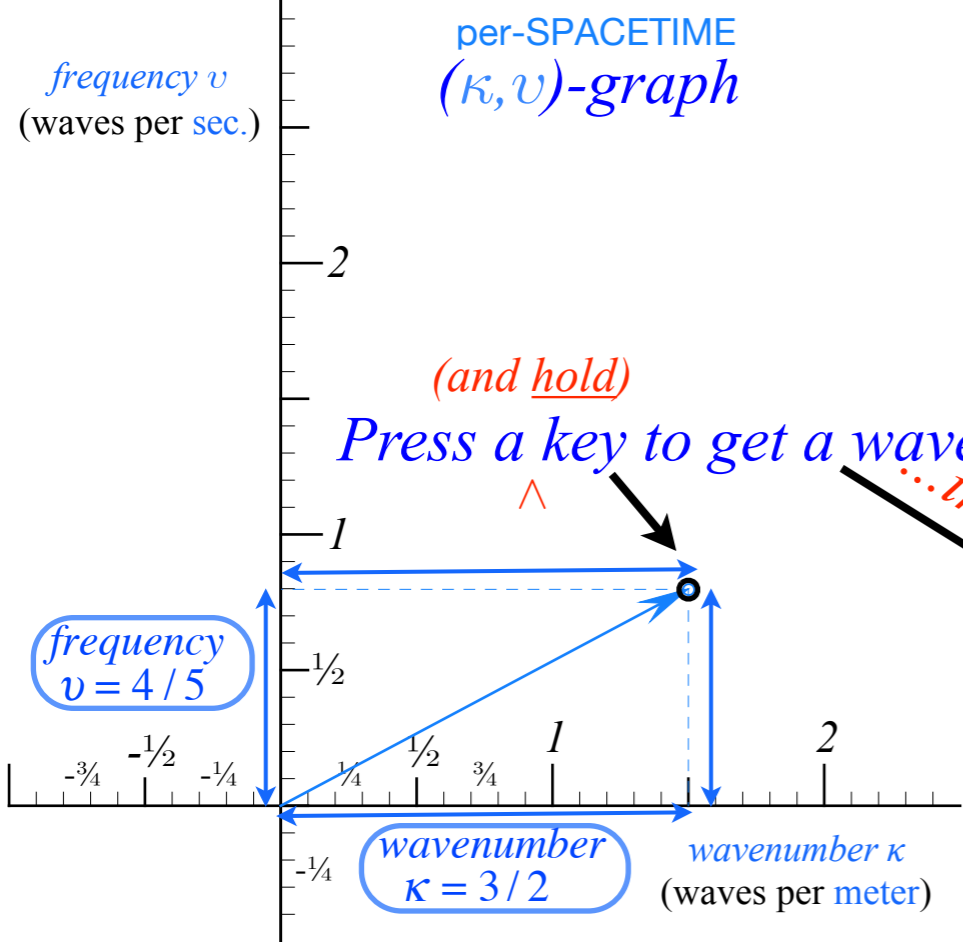
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•How to understand waves  
and  
wave velocity  $V_{\text{wave}}$

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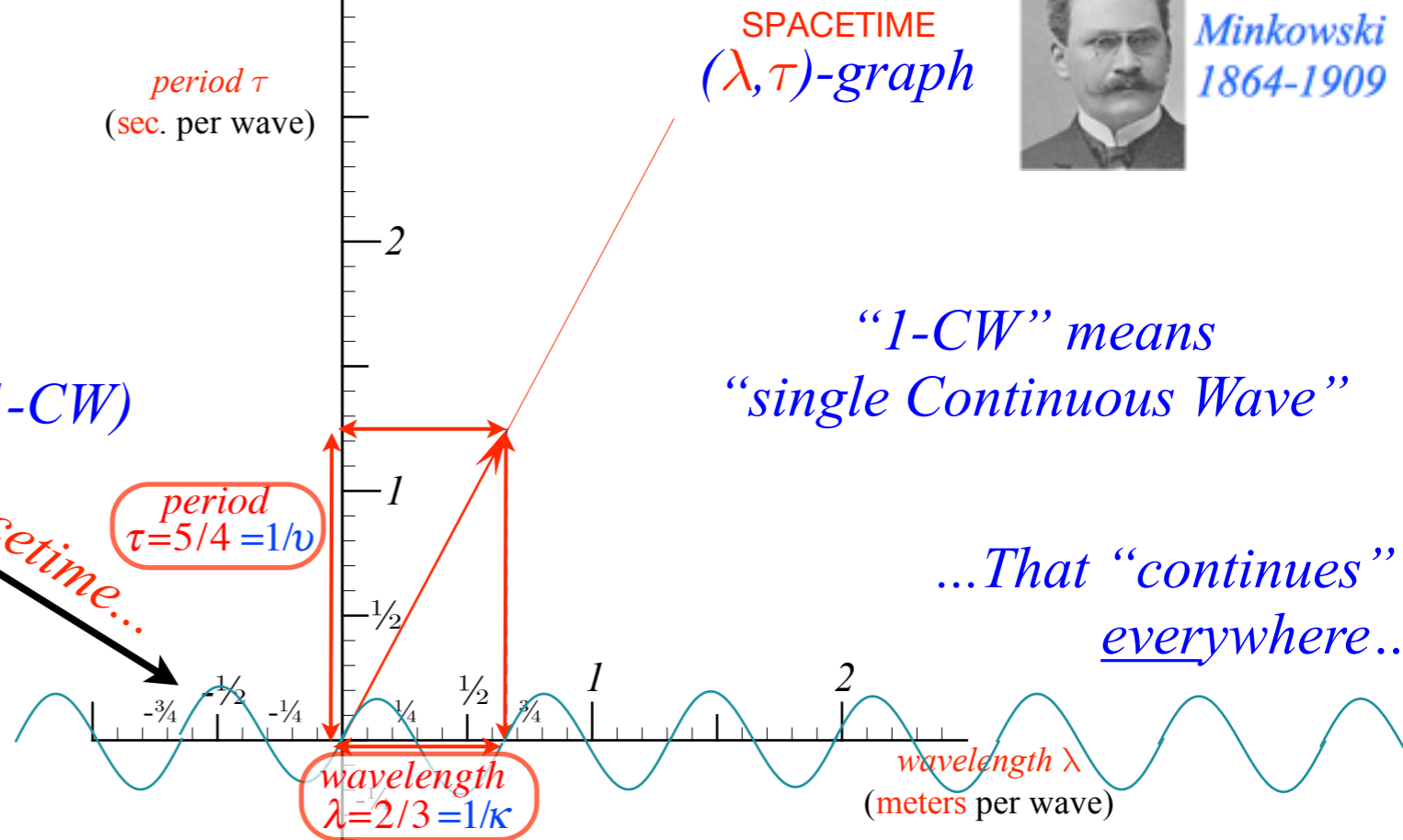
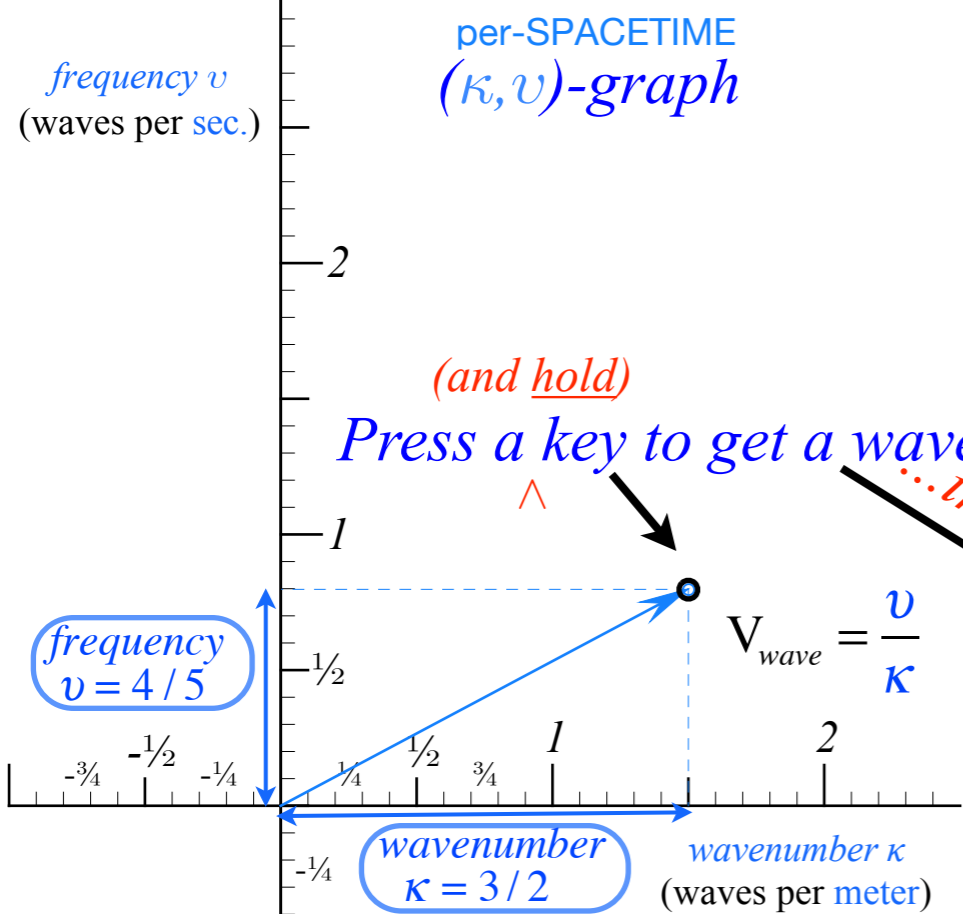
wave-speed equals slope-to-vertical  $\lambda/\tau$  in  $(\lambda, \tau)$ -graph



# Analyzing wave velocity by per-space-per-time and space-time graphs



Herman Minkowski  
1864-1909



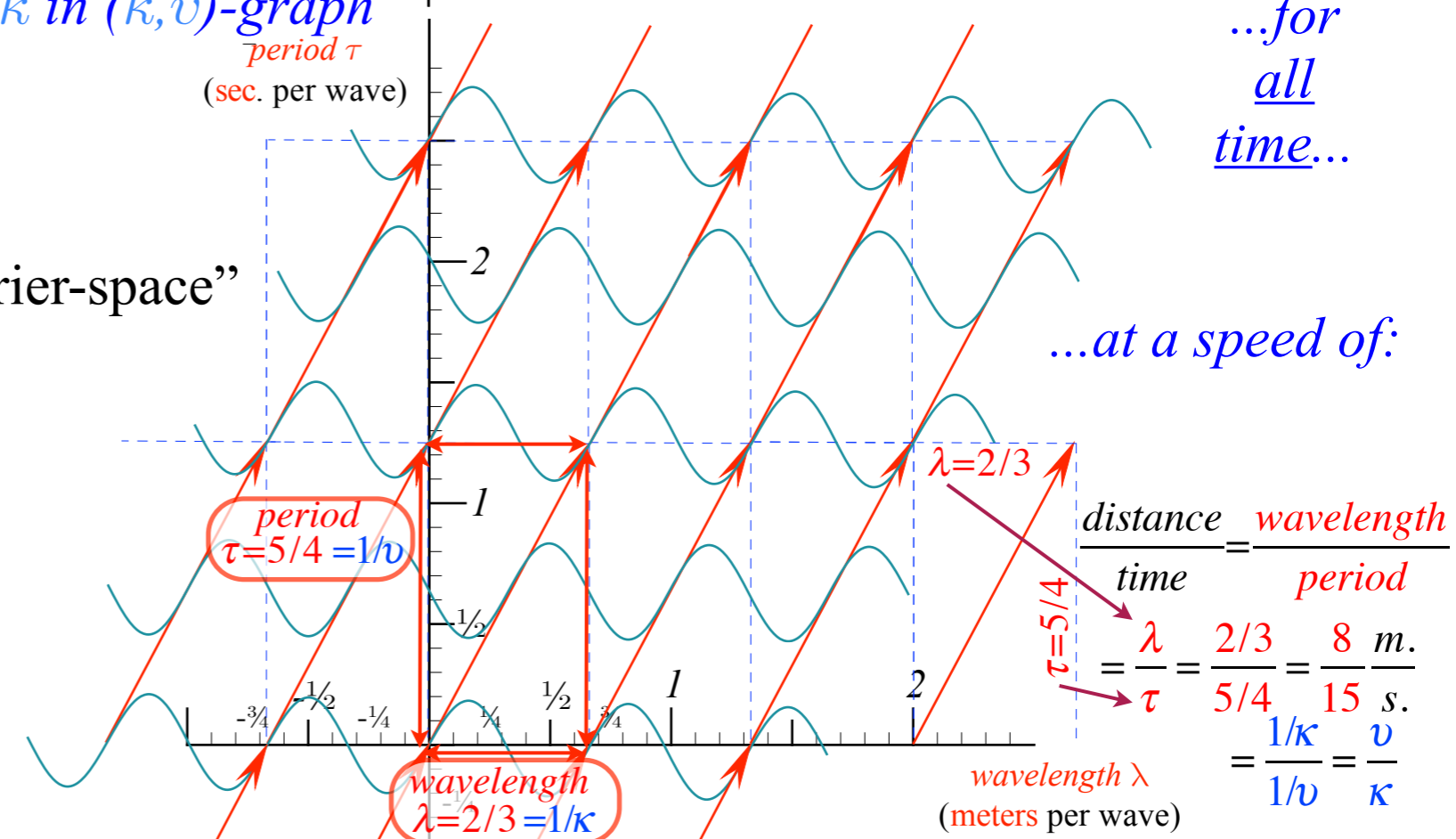
wave-speed equals slope-to-horizontal  $\nu/\kappa$  in  $(\kappa, \nu)$ -graph

...for  
all  
time...

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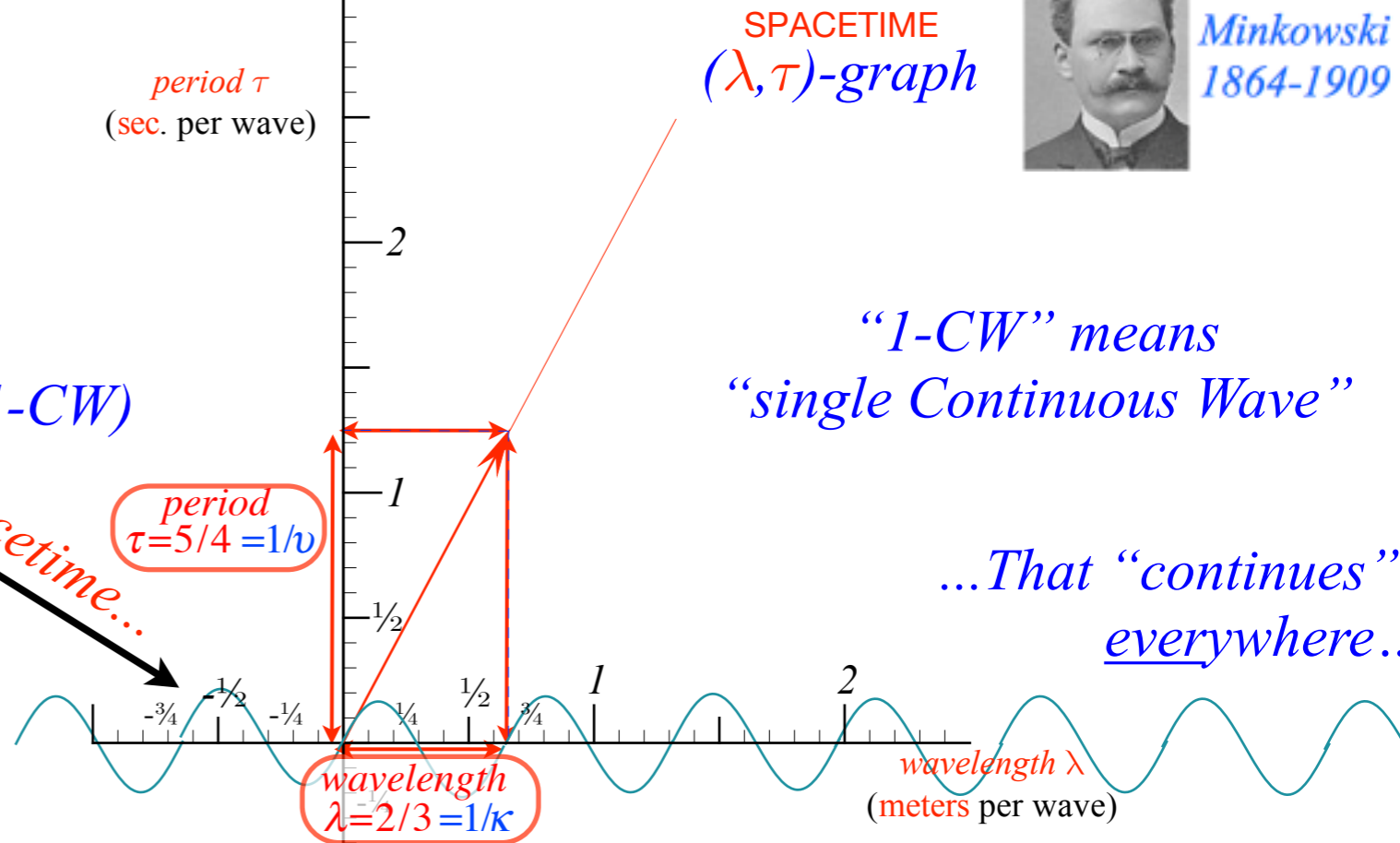
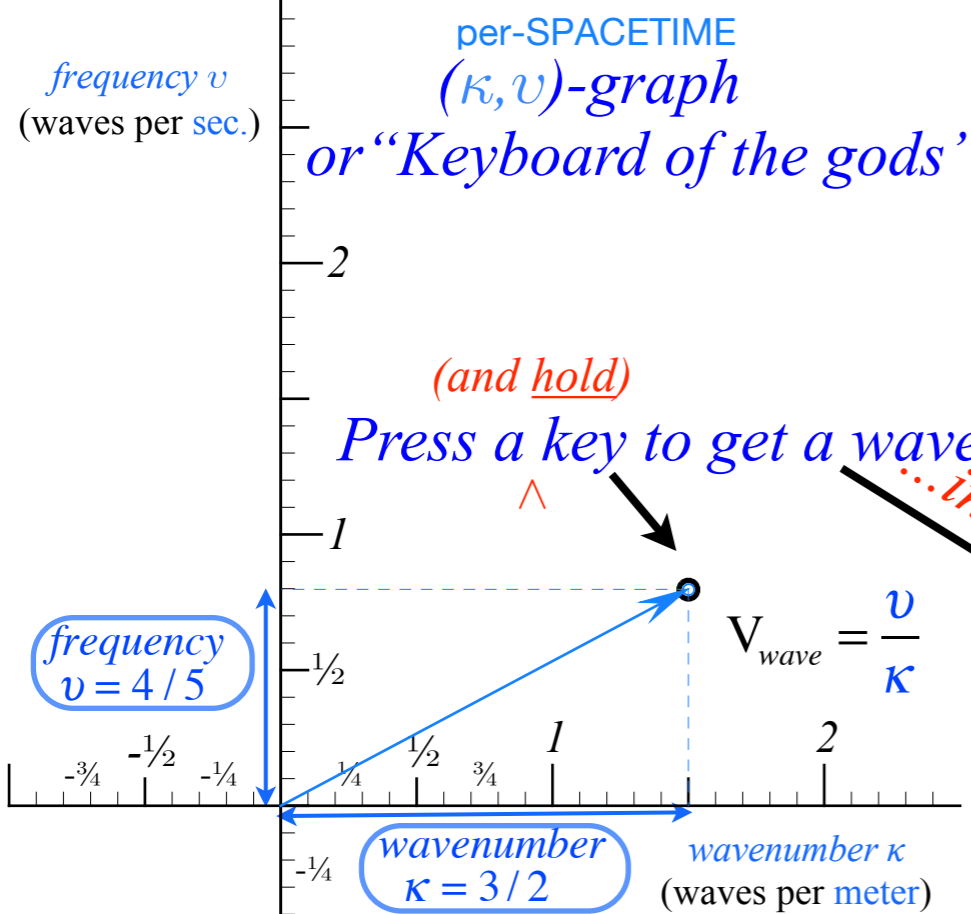
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1864-1909



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wave-velocity formulas

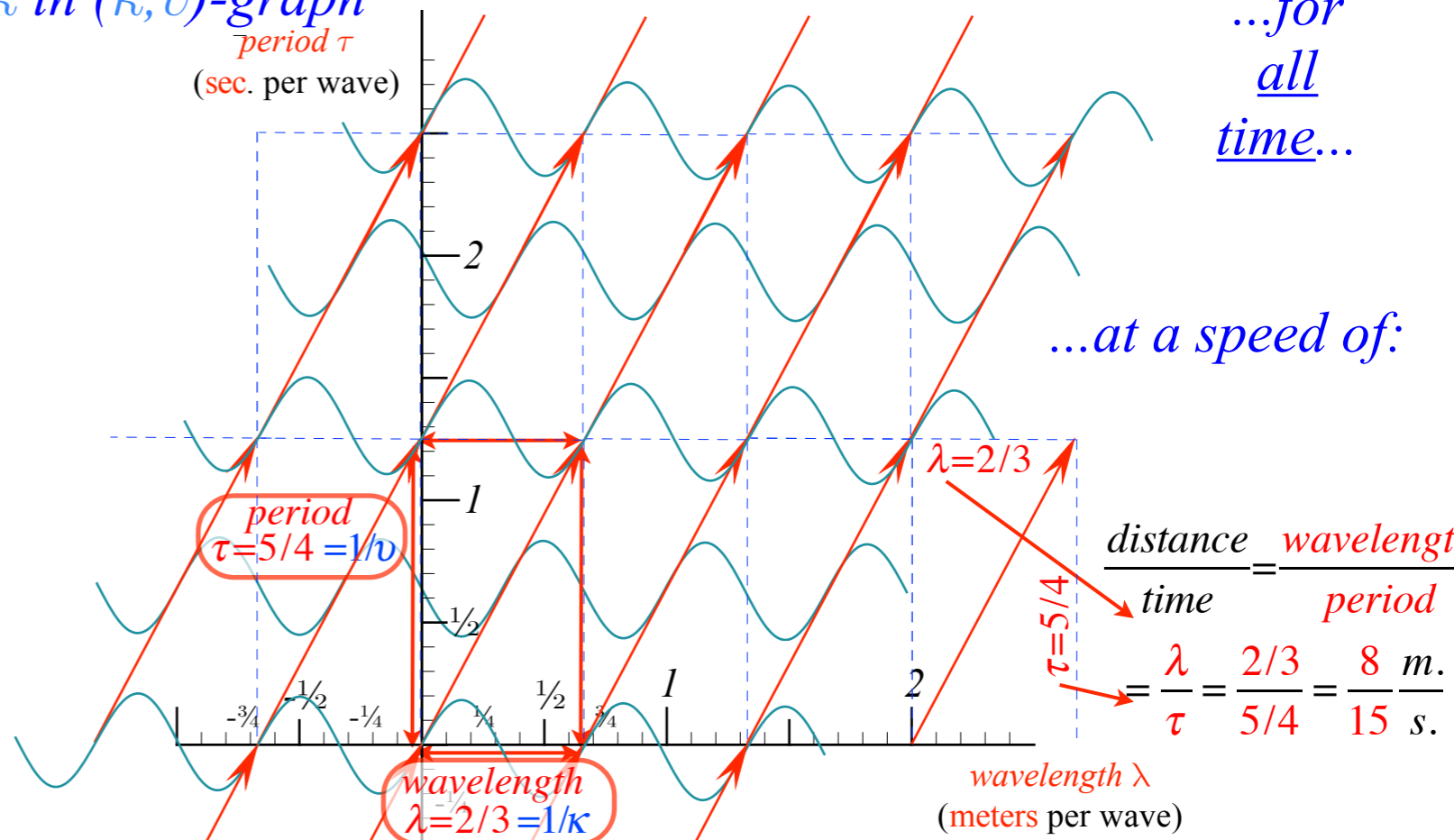
$$\frac{\text{distance}}{\text{time}} = \frac{\text{wavelength}}{\text{period}} = \frac{\text{frequency}}{\text{wavenumber}}$$

$$V_{wave} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/\nu} = \frac{\nu}{\kappa} = \frac{1/\tau}{1/\lambda}$$

$$= \frac{2/3}{5/4} = \frac{4/5}{3/2} = \frac{8 \text{ m.}}{15 \text{ s.}}$$

wave arithmetic is simpler to explain using fractions

•How to understand waves  
and  
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# Introducing Doppler shifting

frequency  $\nu$   
(units: 600THz)  
 $= \nu_A$  1800THz

per-SPACETIME  
 $(c\kappa, \nu)$ -graph

$c \cdot$  time period  $c\tau$   
(units:  $\frac{1}{2}\mu m$ )  
 $c\tau_A = \lambda_A$

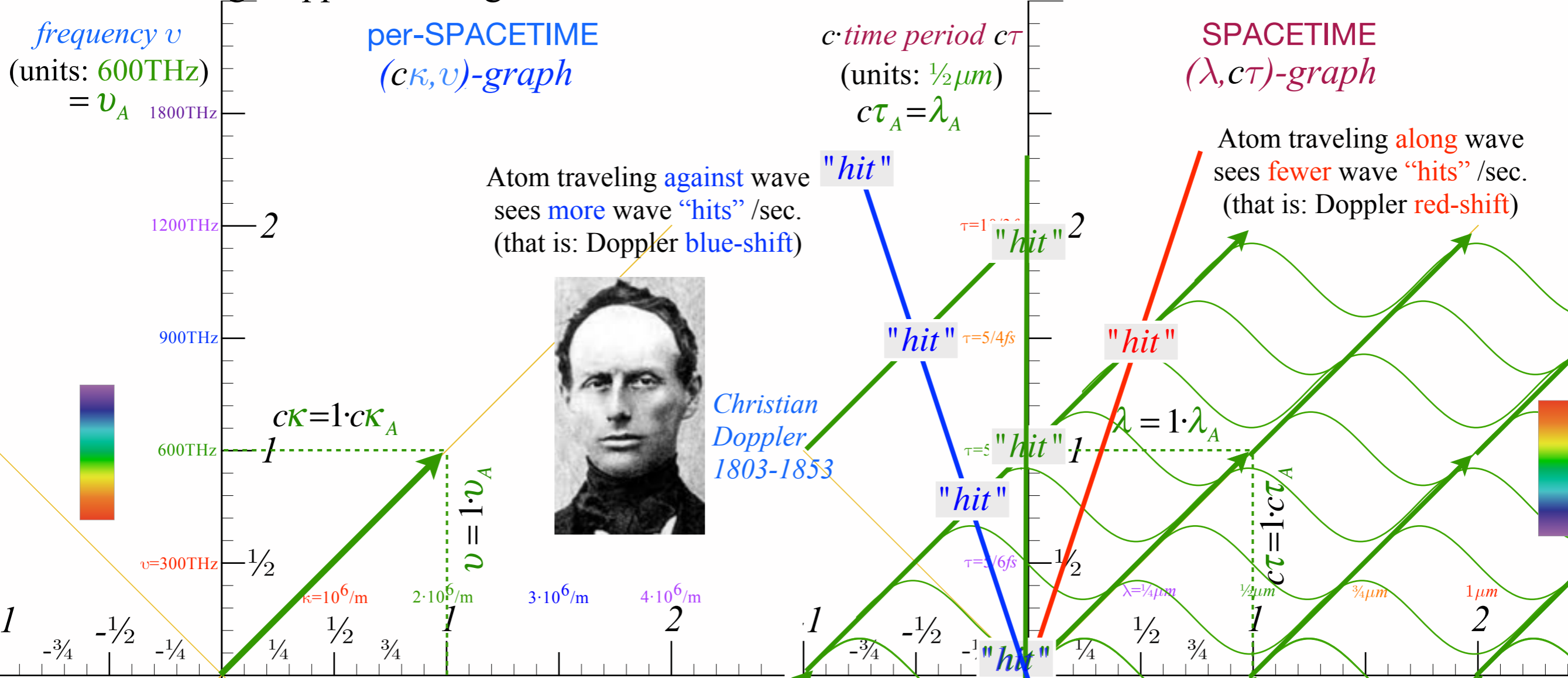
SPACETIME  
 $(\lambda, c\tau)$ -graph

Atom traveling **along** wave  
sees **fewer** wave "hits" /sec.  
(that is: Doppler **red-shift**)

Atom traveling **against** wave  
sees **more** wave "hits" /sec.  
(that is: Doppler **blue-shift**)



Christian Doppler  
1803-1853



$$c = \frac{\lambda}{\tau} = \frac{\nu}{\kappa} = \frac{\omega}{k}$$

rescaled by  $c$  to:

$$1 = \frac{\lambda}{c\tau} = \frac{\nu}{c\kappa} = \frac{\omega}{ck}$$

Move fast enough this way then the "green" wave gets **redder** and **redder** until it dies

Move fast enough this way then the "green" wave gets **bluer** and **bluer** until YOU die

Frequency AND Amplitude decrease exponentially

Frequency AND Amplitude increase exponentially

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# Introducing Doppler shifting and why $c$ is so constant (and so slow)

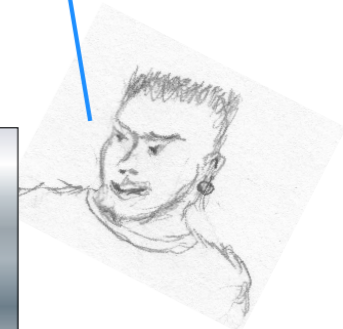
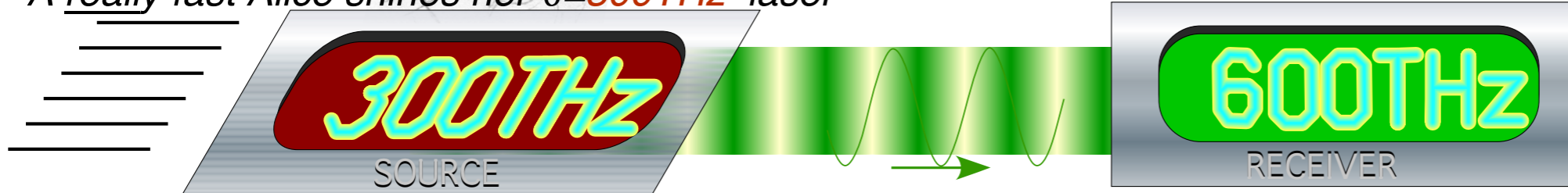
(a)



Bob: "Alice! My frequency meter reads  $\nu=600\text{THz}$  for your laser beam."

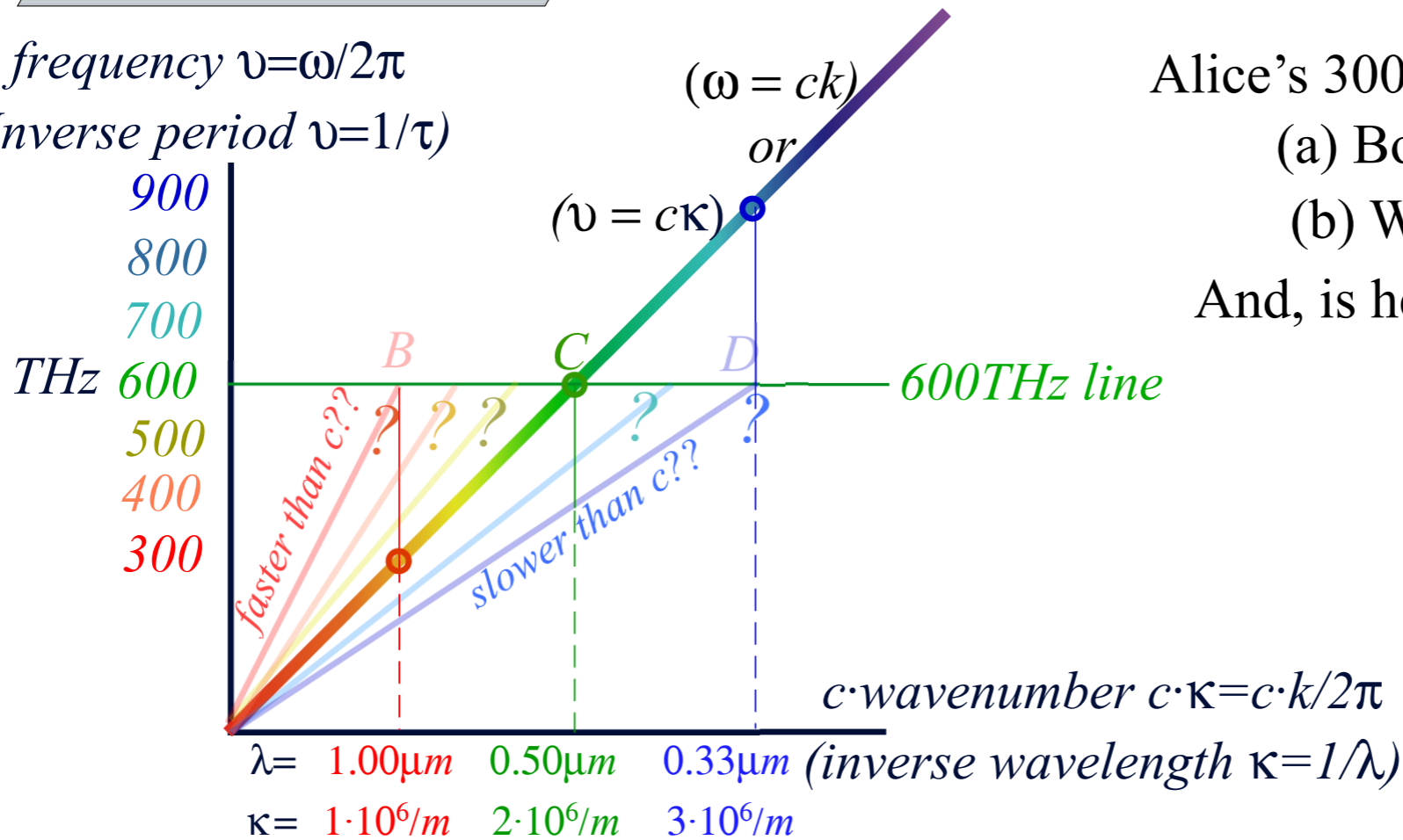
Alice: "Well, what is its wavelength  $\lambda$ , Bob!"

A really fast Alice shines her  $\nu=300\text{THz}$  laser



(b)

frequency  $\nu=\omega/2\pi$   
(Inverse period  $\nu=1/\tau$ )



Alice's 300THz laser approaches Bob.

(a) Bob sees  $\nu=600\text{THz}$ .

(b) What  $\lambda=1/\kappa$  does Bob measure?

And, is he seeing a 'phony' green?



# Introducing Doppler shifting and why $c$ is constant

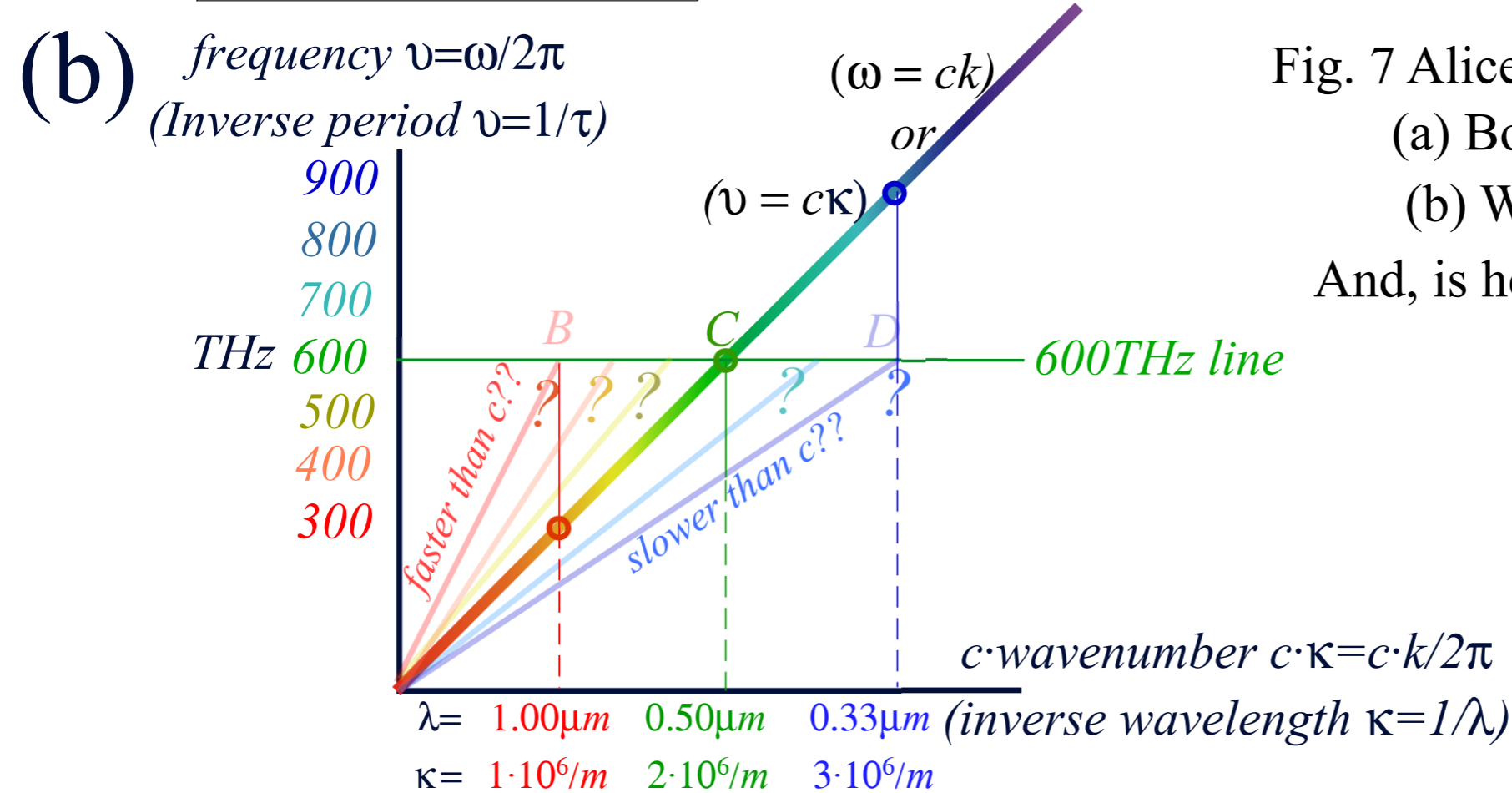
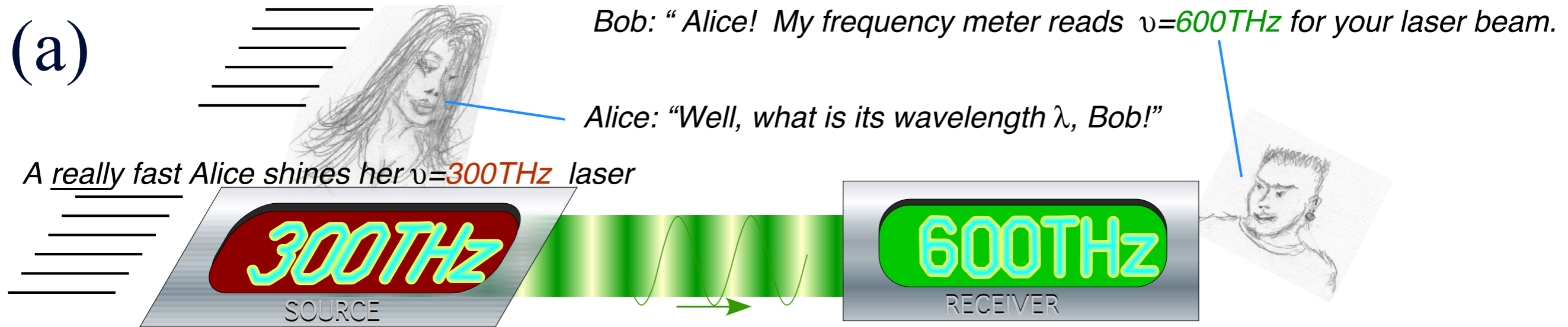


Fig. 7 Alice's 300THz laser approaches Bob.

(a) Bob sees  $\nu=600\text{THz}$ .

(b) What  $\lambda=1/\kappa$  does Bob measure?

And, is he seeing a 'phony' green?

Years of spectroscopy rule out 'phony' 600THz blue-green that do not have wavelength  $\lambda=0.5\text{micron}$ .

The only choice is C.

# Introducing Doppler shifting and why c is constant

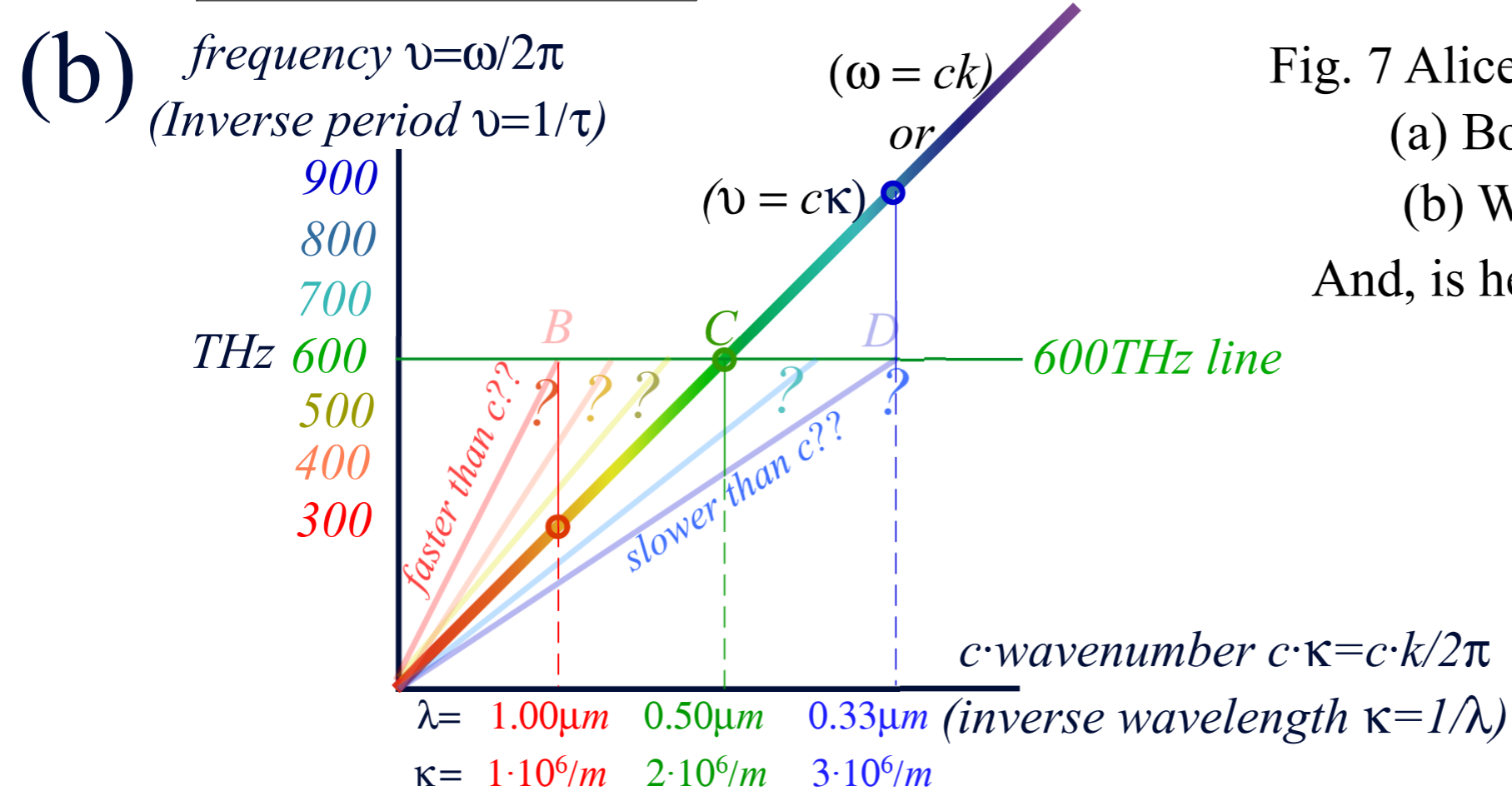
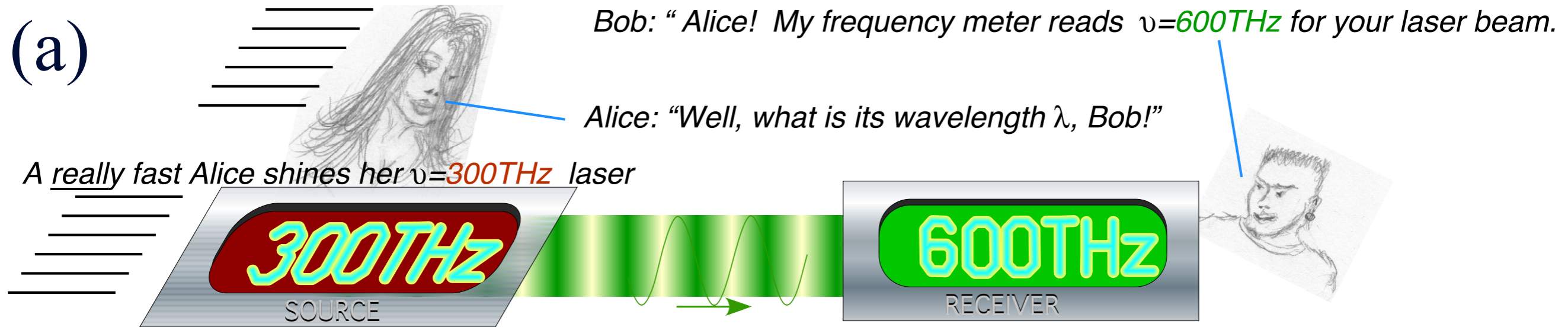


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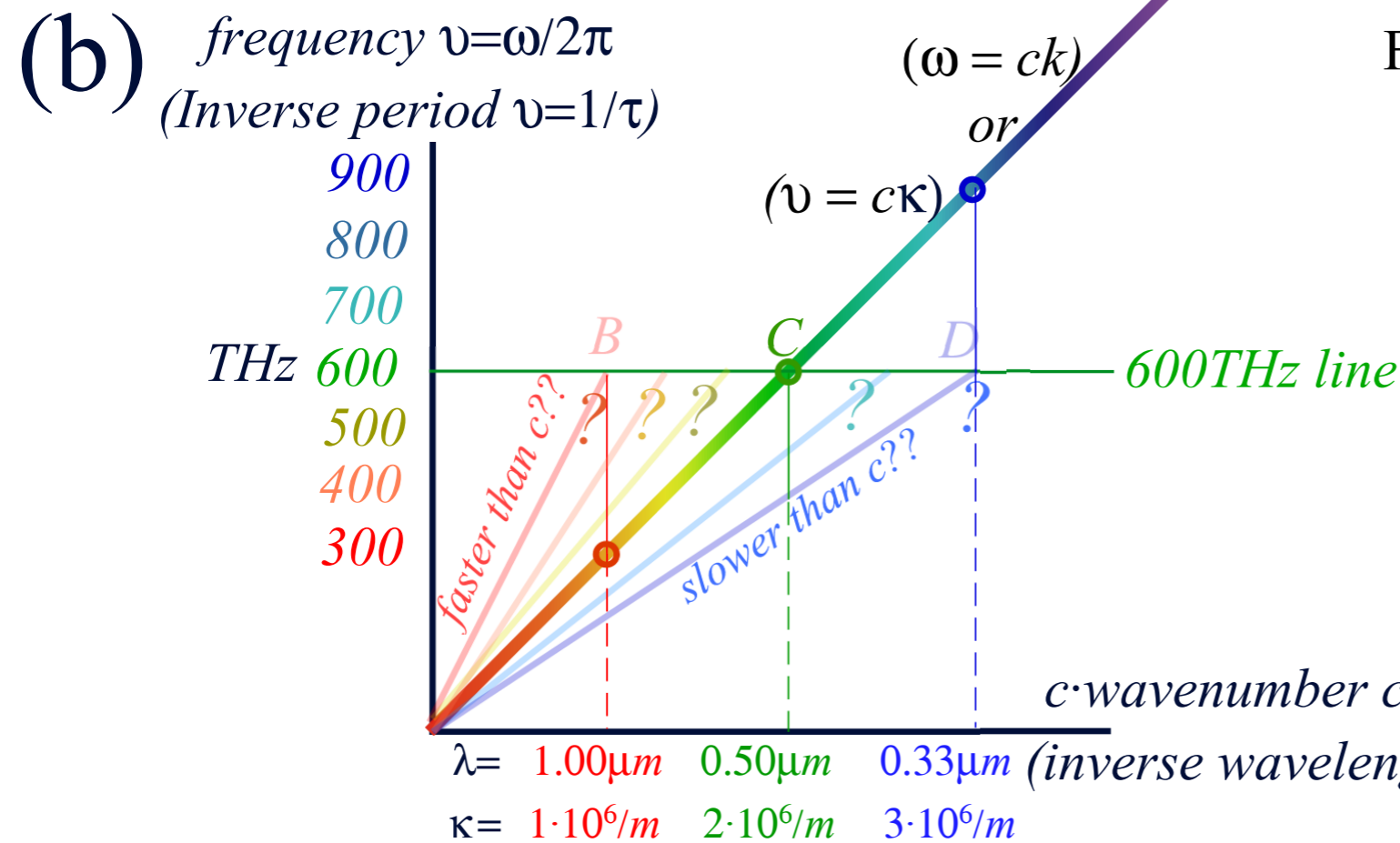
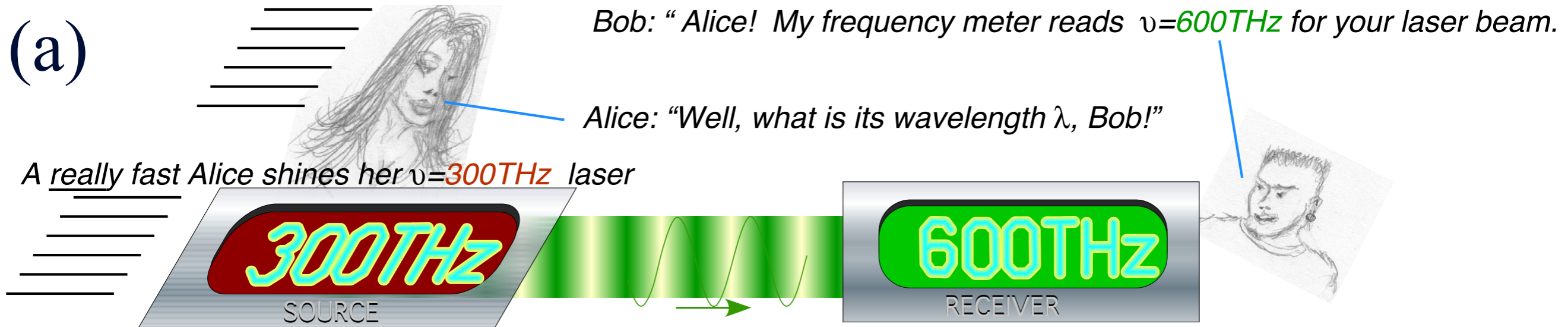


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Actually:  $2.99792458 \cdot 10^8 \text{ m} \cdot \text{s}^{-1}$

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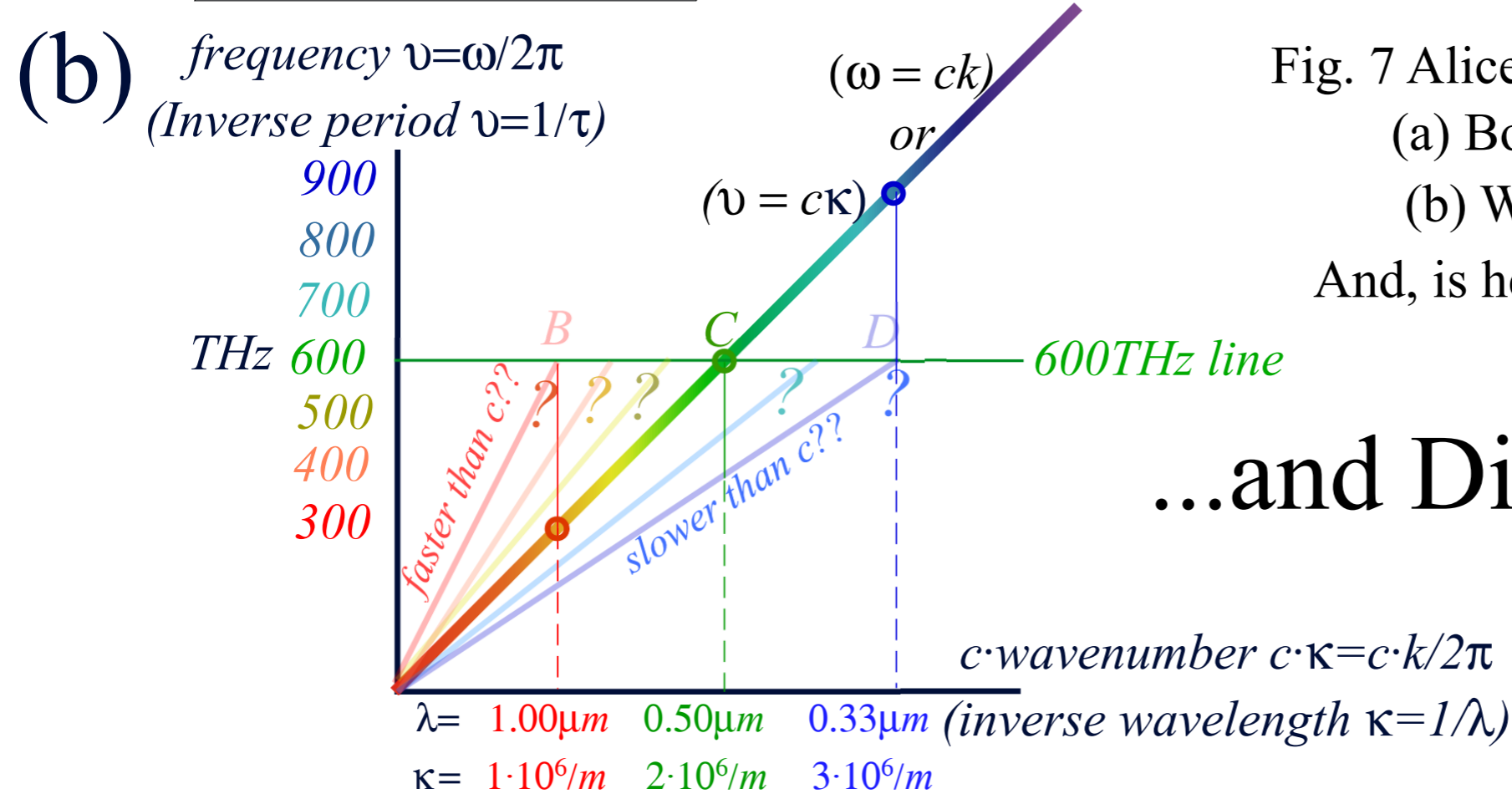
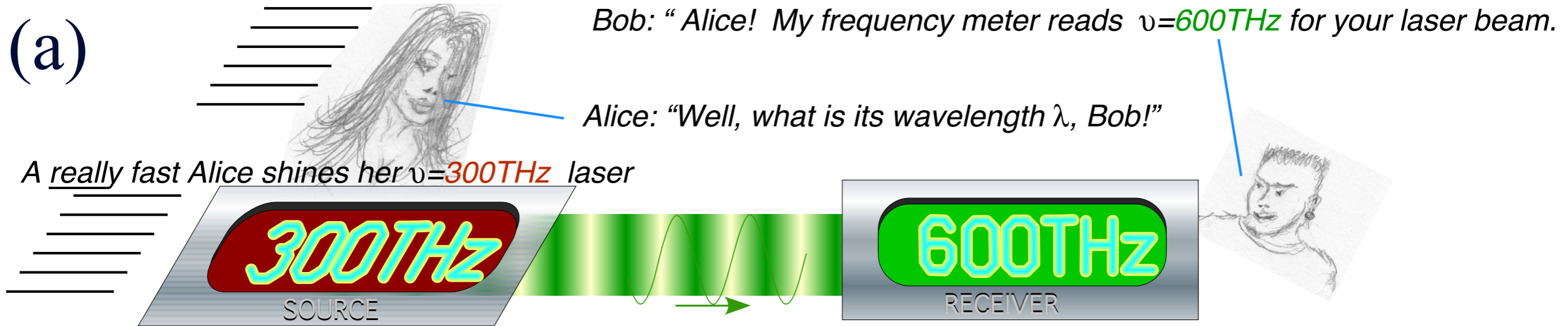


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## ...and Dispersion-Free!

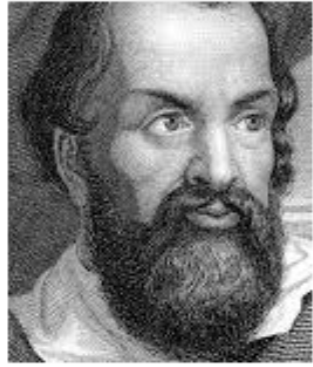
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Galileo Galilei



1564-1642

**Galileo's Revenge (part 1)**

*Rapidity adds just like  
Galilean velocity*

*Why Men in Black shot little Suzie... Learning about **sin!**, **cos** and... Trigonometric road maps*

Hyper-Trigonometric *Relativity* geometry and Euler exponential algebra

1CW wavefunctions and phasors

Per-space-per-time vs Space-time

Wave velocity formulas

Introducing Doppler shifting

Why is  $c$  so constant?!

➔ Introducing Doppler Arithmetic and *Rapidity*  $\rho$

Optical interference “baseball-diamond” displays *phase* and *group* velocity

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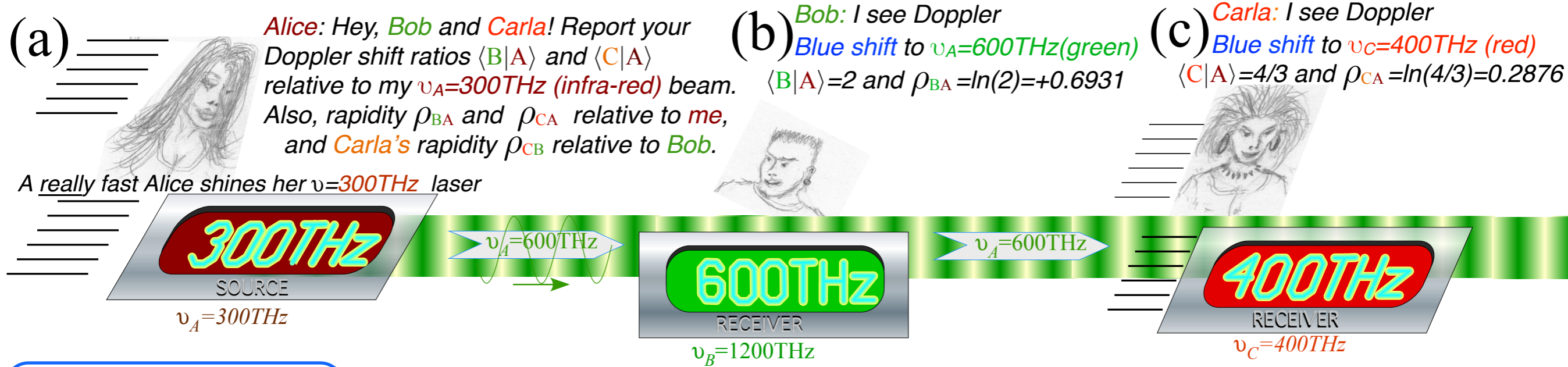
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“Occams Sword” and geometry of 16 parameter functions of  $\rho$  and  $\sigma$

Application to TE-Waveguide modes and synchrotron beam relativity



Doppler ratio:

$$\langle R|S \rangle = \frac{\nu_{RECEIVER}}{\nu_{SOURCE}}$$

$$\rho_{RS} = \ln \langle R|S \rangle$$

or:

$$\langle R|S \rangle = e^{\rho_{RS}} = e^{-\rho_{SR}}$$

Definition of Rapidity  $\rho_{RS}$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{\nu_B}{\nu_A} = \frac{600}{300} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \ln \langle B|A \rangle = \ln \frac{2}{1} = 0.6931$$

$$\rho_{AB} = \ln \langle A|B \rangle = \ln \frac{1}{2} = -0.6931 = -\rho_{BA}$$

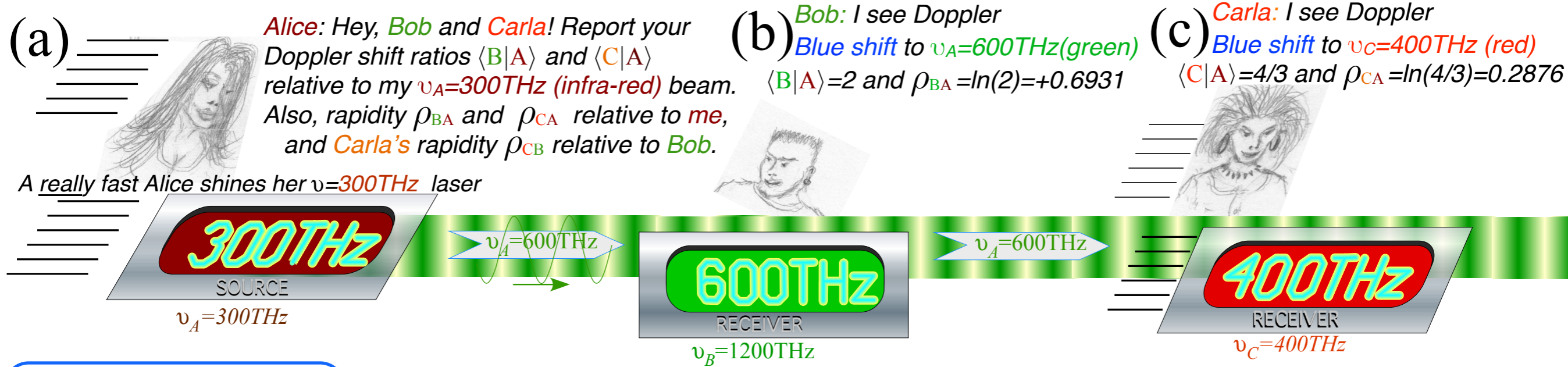
Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{\nu_C}{\nu_A} = \frac{400}{300} = \frac{4}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \ln \langle C|A \rangle = \ln \frac{4}{3} = 0.2876$$

Introducing Doppler Arithmetic and rapidity  $\rho$



Doppler ratio:  
 $\langle R|S \rangle = \frac{\nu_{RECEIVER}}{\nu_{SOURCE}}$   
 $\rho_{RS} = \ln \langle R|S \rangle$   
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*Definition of Rapidity*  
 $\rho_{RS}$

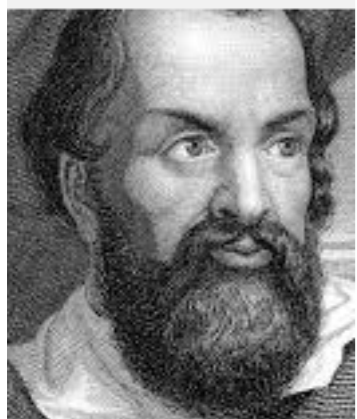
Bob-Alice Doppler ratio:  
 $\langle B|A \rangle = \frac{\nu_B}{\nu_A} = \frac{600}{300} = 2$   
 Bob-Alice rapidity:  
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Carla-Bob Doppler ratio:  
 $\langle C|B \rangle = \frac{\nu_C}{\nu_B} = \frac{\nu_C}{\nu_A} \frac{\nu_A}{\nu_B} = \langle C|A \rangle \langle A|B \rangle = \frac{4}{3} \frac{1}{2} = \frac{2}{3}$

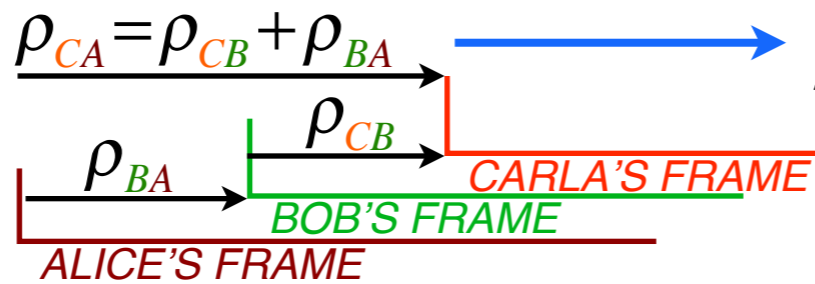
Carla-Bob rapidity:  
 $e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} = e^{\rho_{CA} + \rho_{AB}}$   
 $\rho_{CB} = \rho_{CA} + \rho_{AB} = 0.2876 - 0.6931 = -0.4055$   
 $= \ln \frac{4}{3} + \ln \frac{1}{2} = \ln \frac{2}{3}$

Galileo Galilei

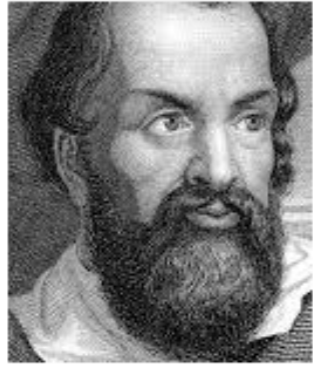


1564-1642

**Galileo's Revenge (part 1)**  
 Rapidity adds just like Galilean velocity



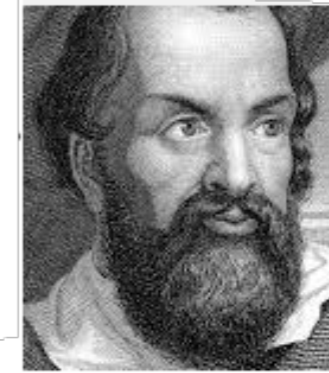
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**Galileo's Revenge (part 2)**

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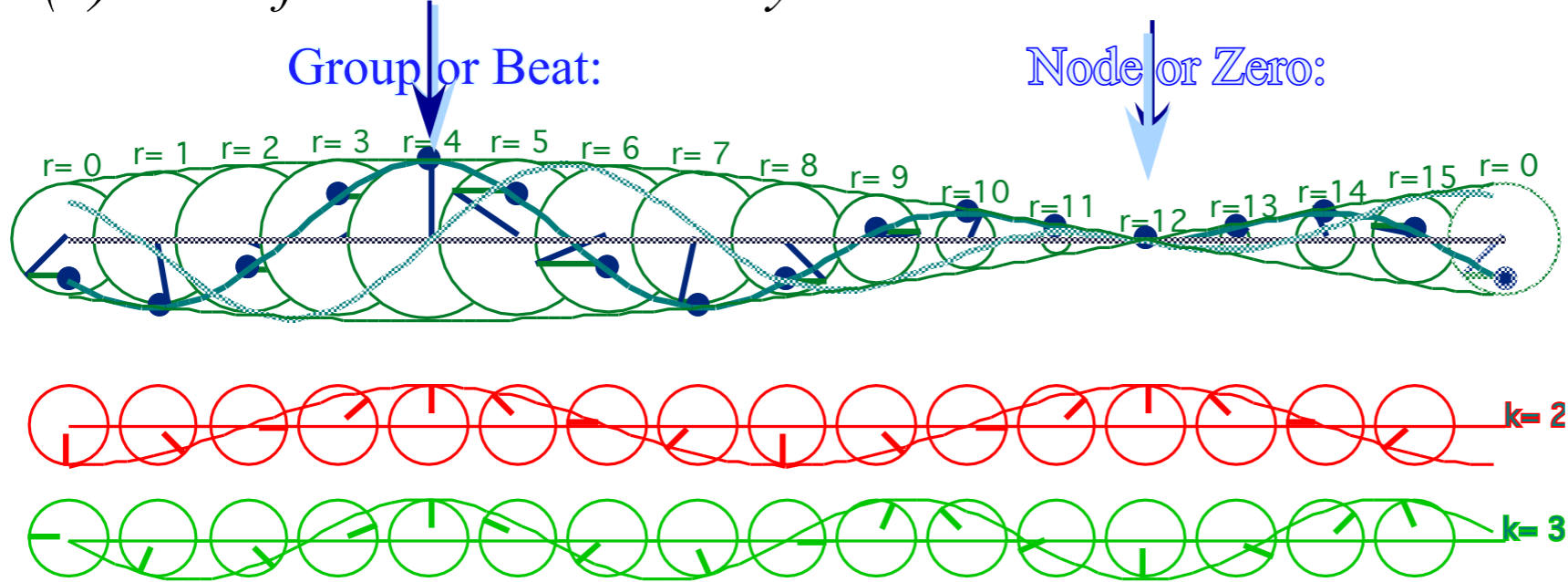
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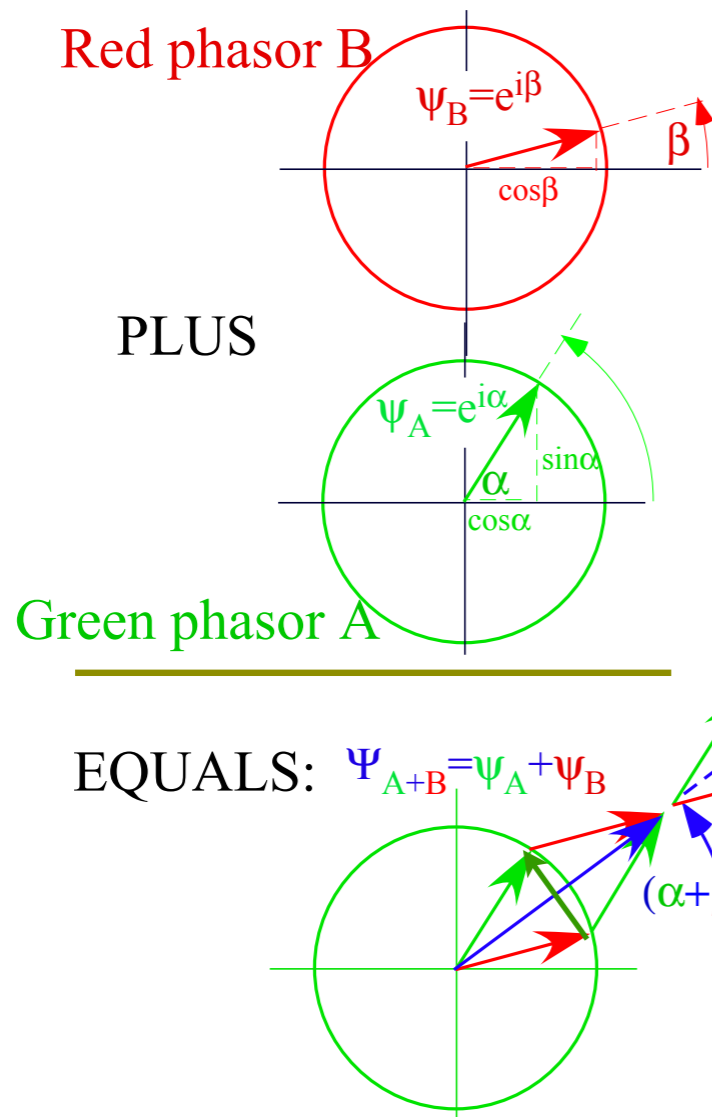
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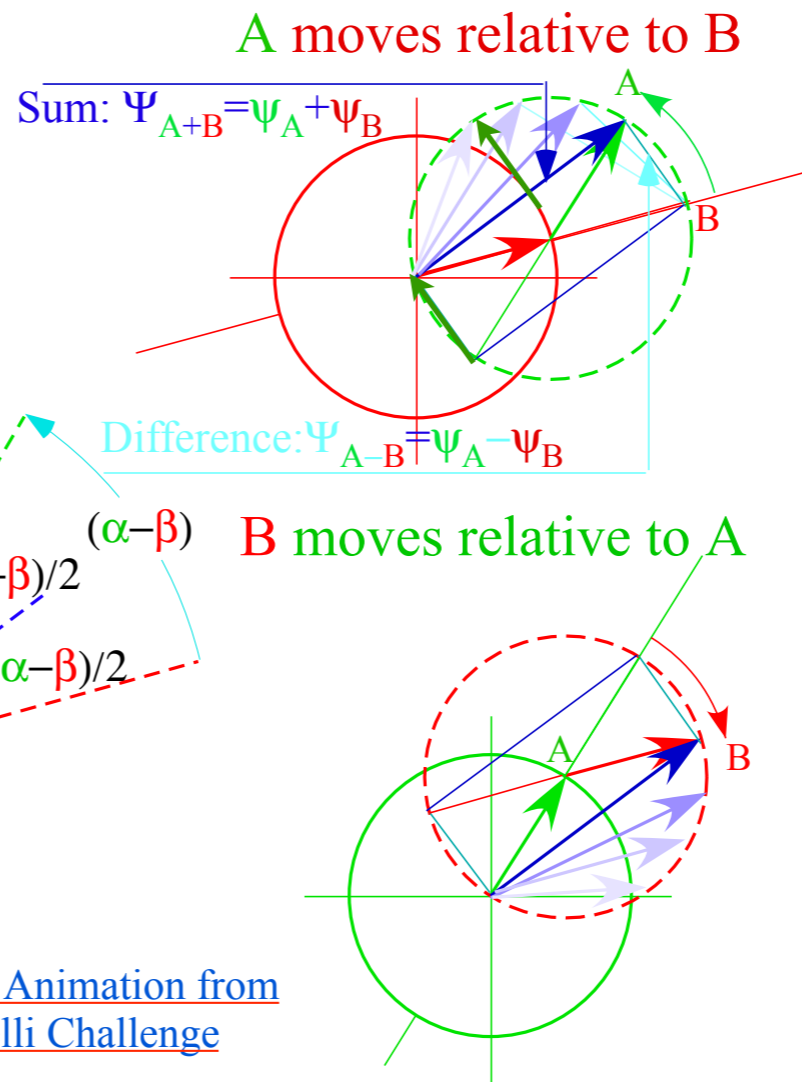
(a) Sum of Wave Phasor Array



(b) Typical Phasor Sum:

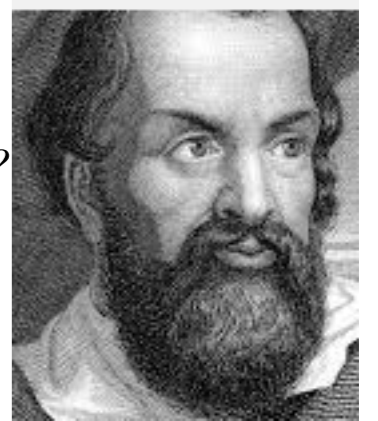


(c) Phasor-relative views



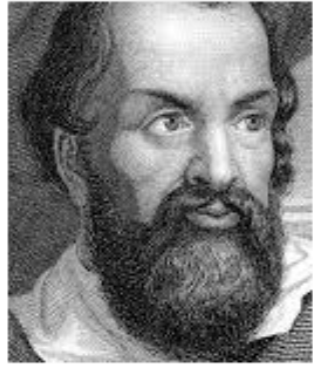
Geometry of the Half-sum Phase and Half-difference Group

Happy now?



Galileo's Revenge (part 2)  
Phasor angular velocity adds just like Galilean velocity

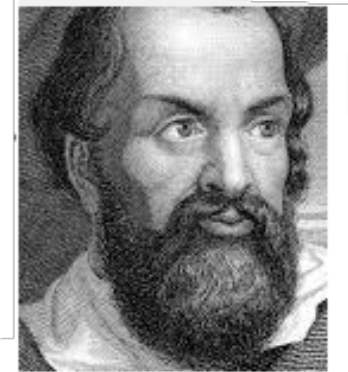
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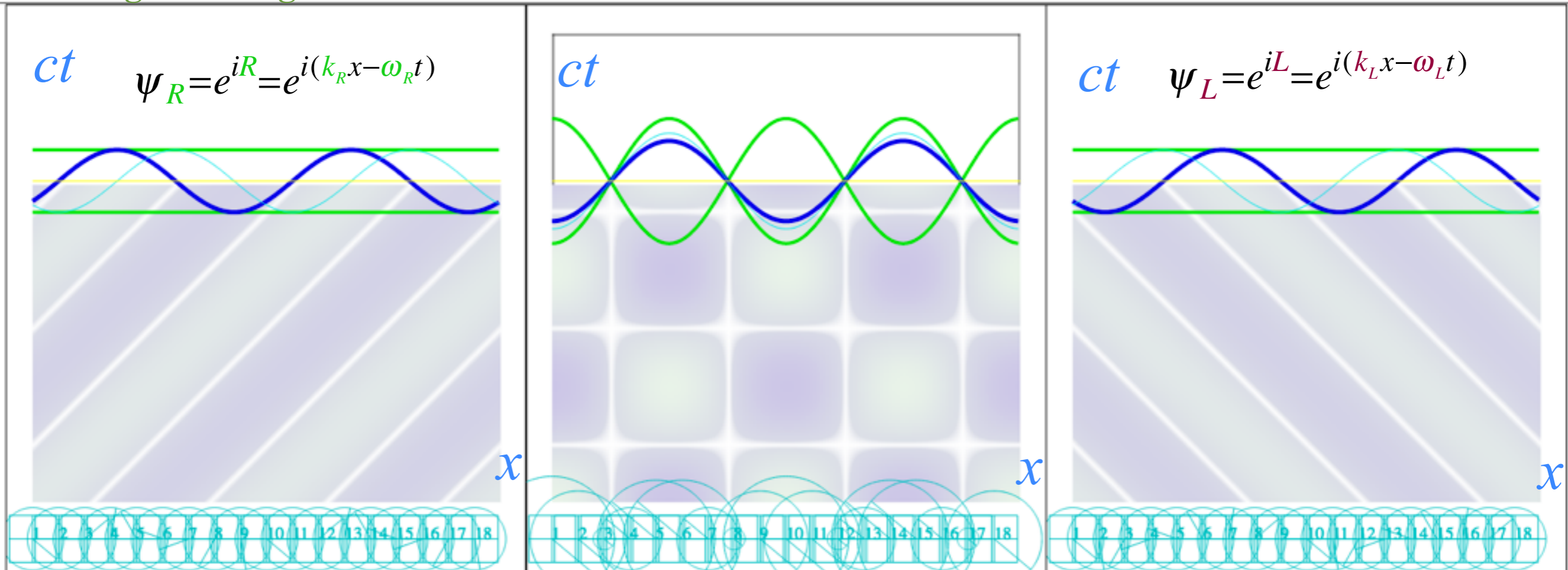
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right-moving CW laser

Colliding 2CW laser beams

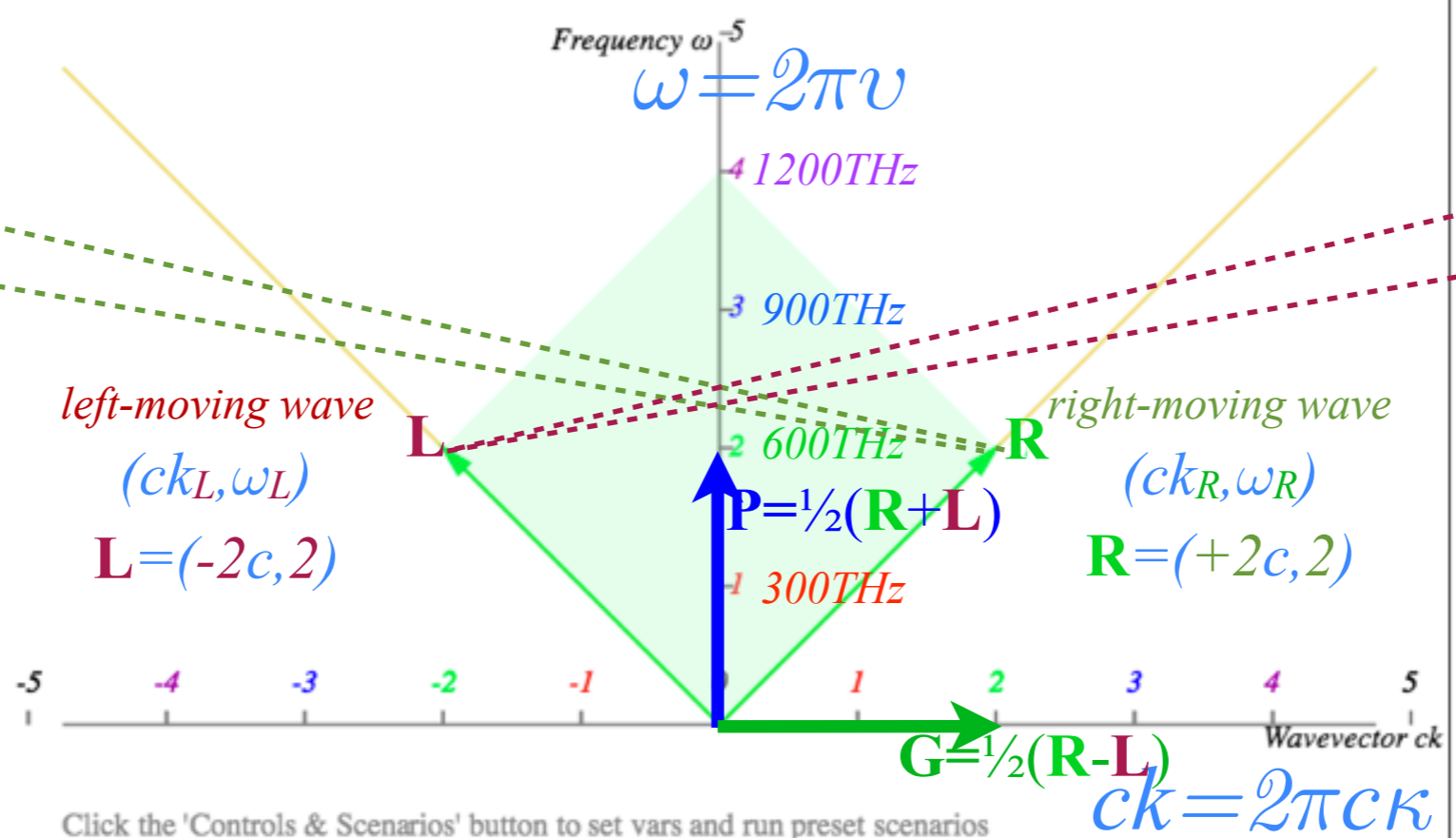
left-moving CW laser



right-moving wave  
Spacetime (x, ct)

left-moving wave  
Spacetime (x, ct)

Per-Spacetime  
(ck, ω)



BohrIt Web Simulation 2  
CW ct vs x Plot (ck = ±2)

Click the 'Controls & Scenarios' button to set vars and run preset scenarios  
Set the right & left-ward k values with clicks near the dispersion curve or ck axis.



## Parameters

BohrIt Panel 3x1: w/ k-Phasors

Configurations

Use Old ST  Use Old Phasor Canvas

Canvas

Time Behavior Loop back to t=0

Retain Space-Time Plot

Align k-Phasors for Reset T=0

x-Phasor Locations Fixed at Bottom

Type of KE Photon:  $\omega(k)=k$

Points per Well =

Space-Time Pixels per Phasor

## Display

E Phasor Scale

X Phasor Scale

$\psi$  Scale

Propagate Mouse Scale Changes

$|\psi|$  Line Width

Re( $\psi$ ) Line Width

Im( $\psi$ ) Line Width

Phasor Line Width

Zero Tracer Line Width

Trace Group Zeros  Trace Phase Zeros

Extra Coordinate Grid Axes w/ P & G Vectors

Background ST Plot Re( $\psi$ )

Zero enhancement Threshold =

Crest-Trough distinction term =

Group & Phase Vectors Both

Right & Left **K** Vectors Both

Shaded Regions Show Both

Axis Titles - Horizontal:  Vertical:

Axis Labels - Horizontal:  Vertical:

## Colors

Color Scheme Journal Color

Global alpha =

Space-Time background alpha =

Peak: Hue=  Val=

Trough: Hue=  Val=

Zero: Hue=  Val=

## Persistent Parameters

Default Space-Time Granularity

Best for tweaking the responsiveness, also vary/set persistent space-time granularity below

## Matter Wave: Bohr-Schrödinger Approximation

Bohr-Schrödinger {Quadratic dispersion}

CW
k=+1,+2
CW
k=+2,+3
CW k=-1,+2

## RelaWavity Scenarios

Dispersion Plot (300 THz Scale)

Make these two lower for finer/narrower zero lines. Note: they do vary with above settings

CW with ck = 2
CW with ck = 4
2 CW with ck = $\pm 2$ RelaWavity $\beta = 0.0$ , $v = 600$ THz
2 PW with ck = $\pm 2$ RelaWavity $\beta = 0.0$ , $v = 600$ THz
2 CW with ck = -1, 4 RelaWavity $\beta = 0.6$ , $v = \sqrt{(300*1200)} = 600$ THz

k-Phasor Plot (100 THz Scale)

CW with ck = 3
CW with ck = -3
CW with ck = 6
CW with ck = 12
2 CW with ck = $\pm 6$ RelaWavity $\beta = 0.0$ , $v = 600$ THz
2 CW with ck = -3, 12 RelaWavity $\beta = 0.6$ , $v = \sqrt{(300*1200)} = 600$ THz

In the APP, right click on a scenario button to expose the actual scenario string

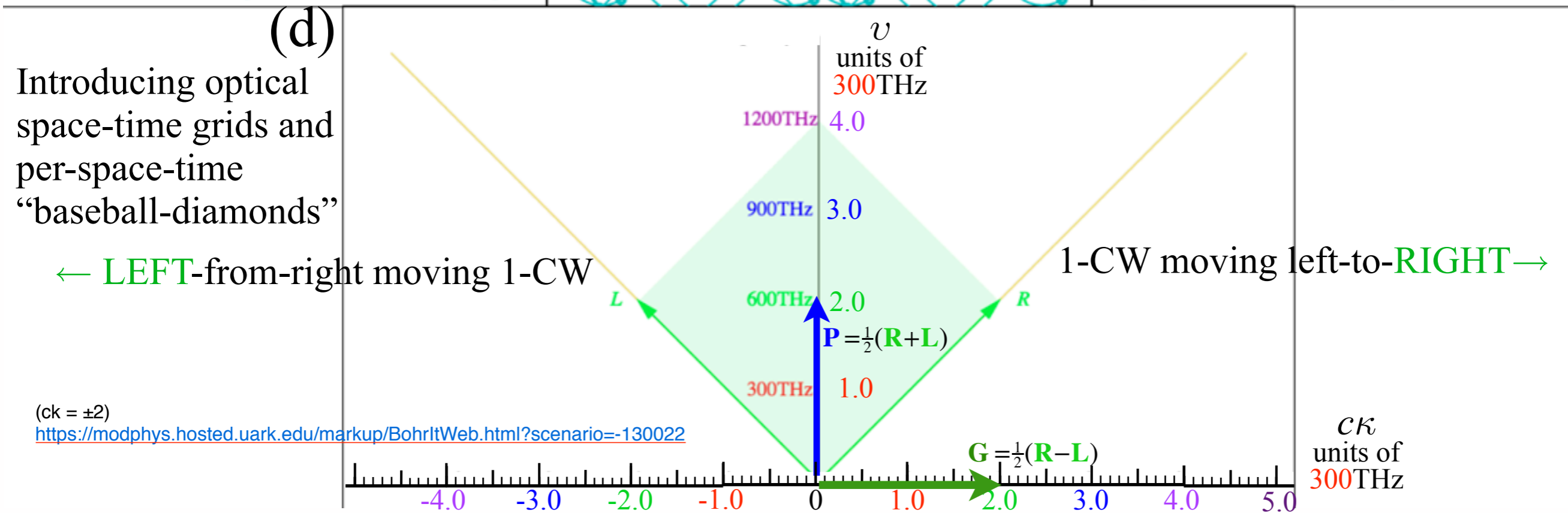
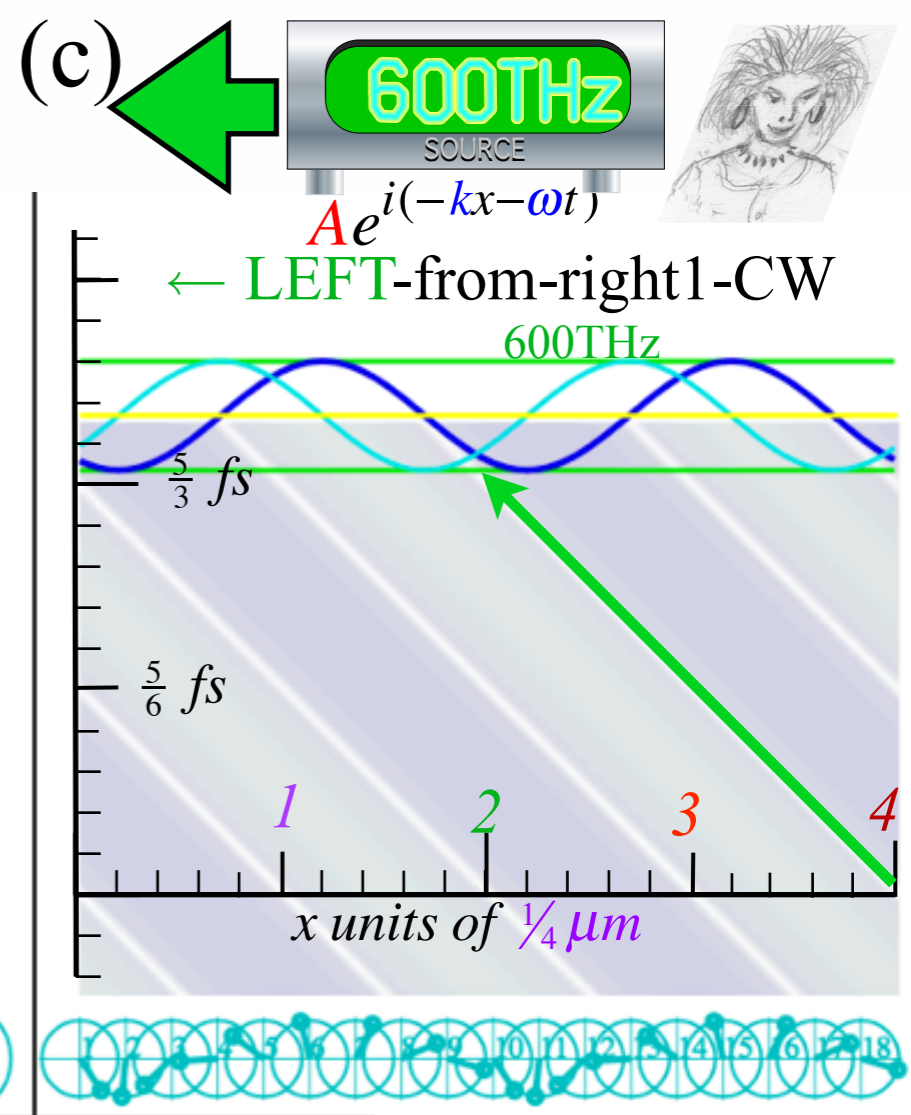
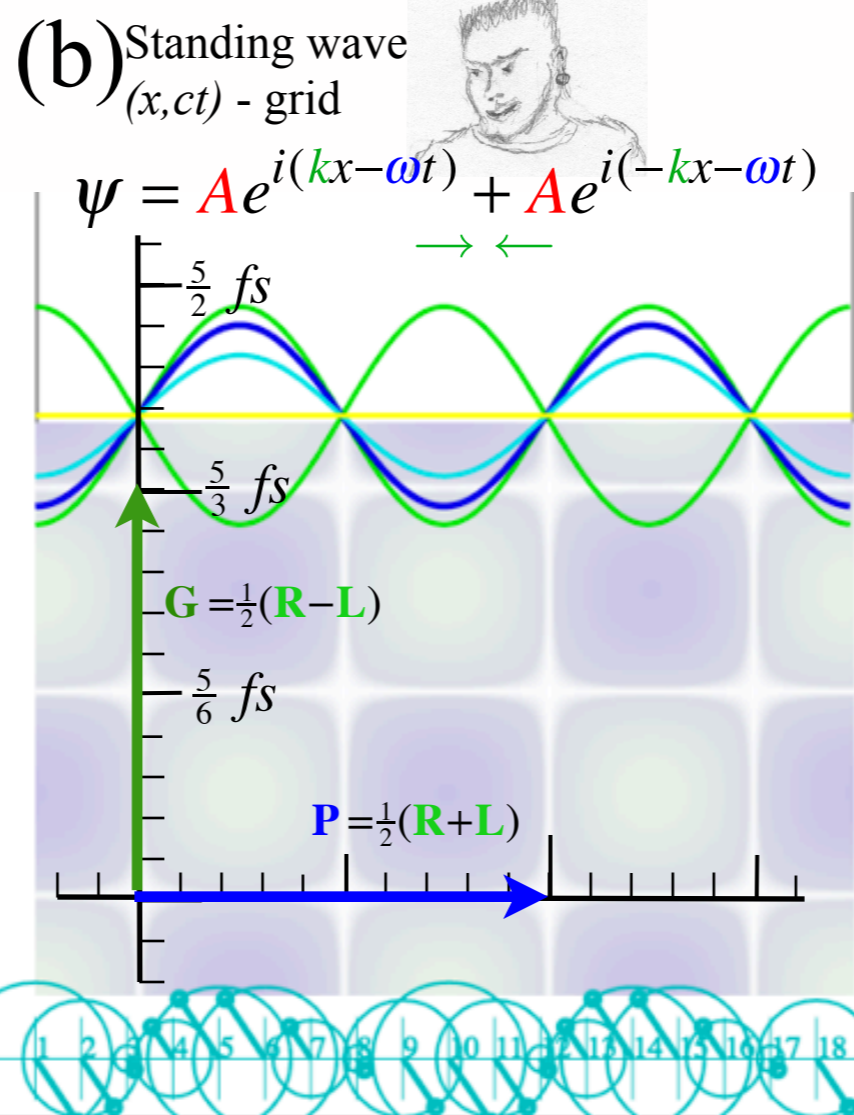
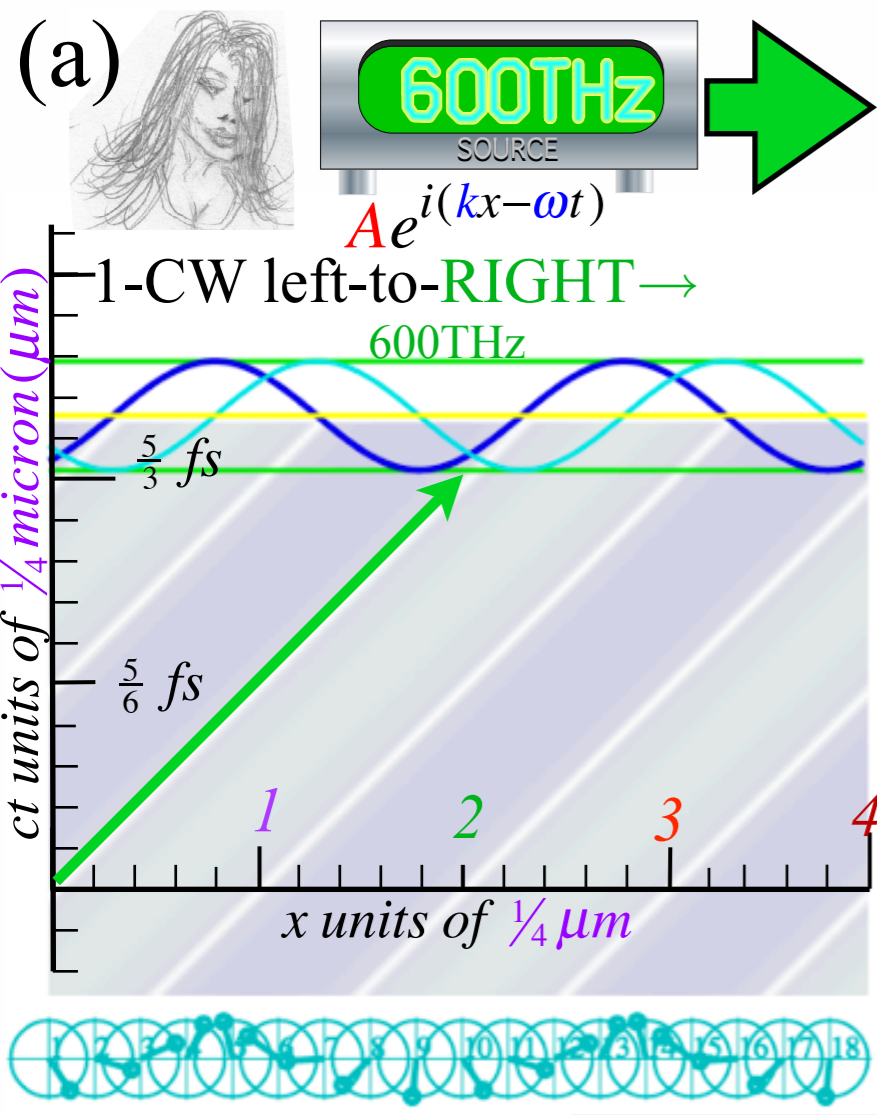
## Scenarios

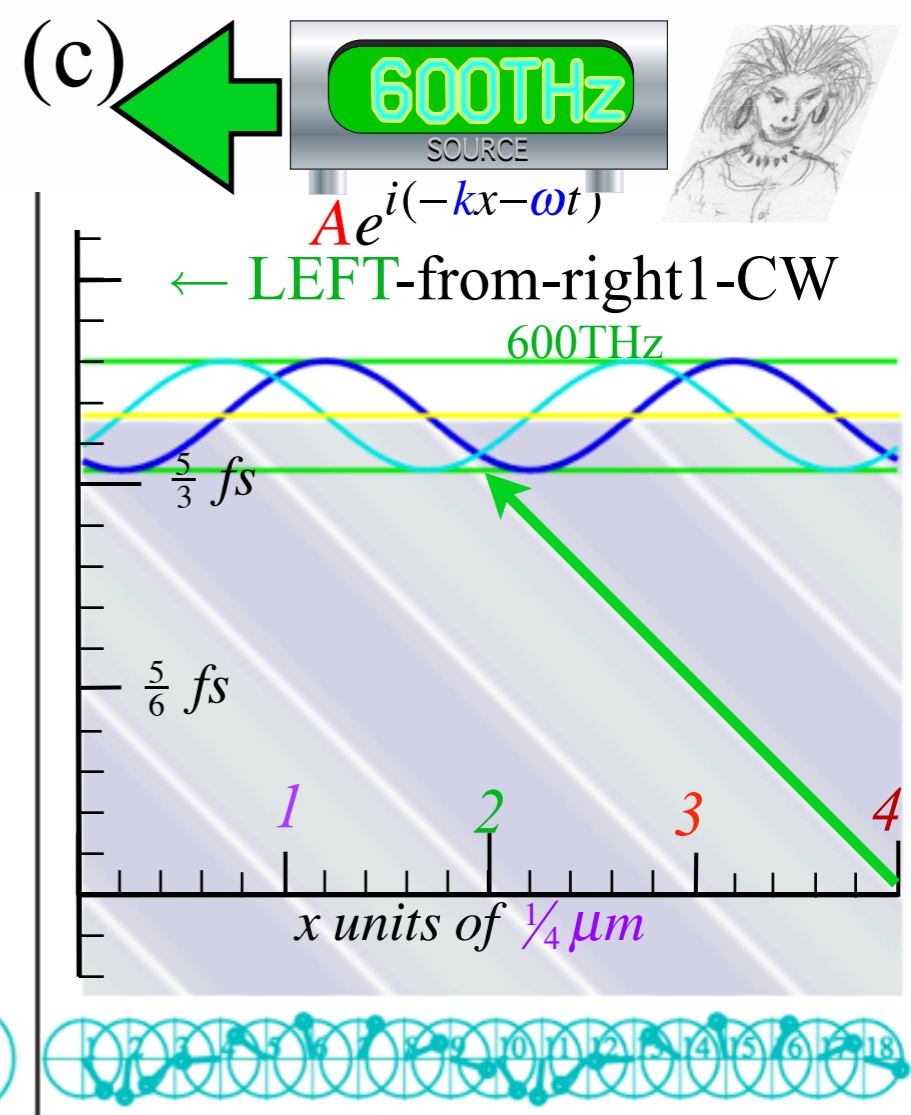
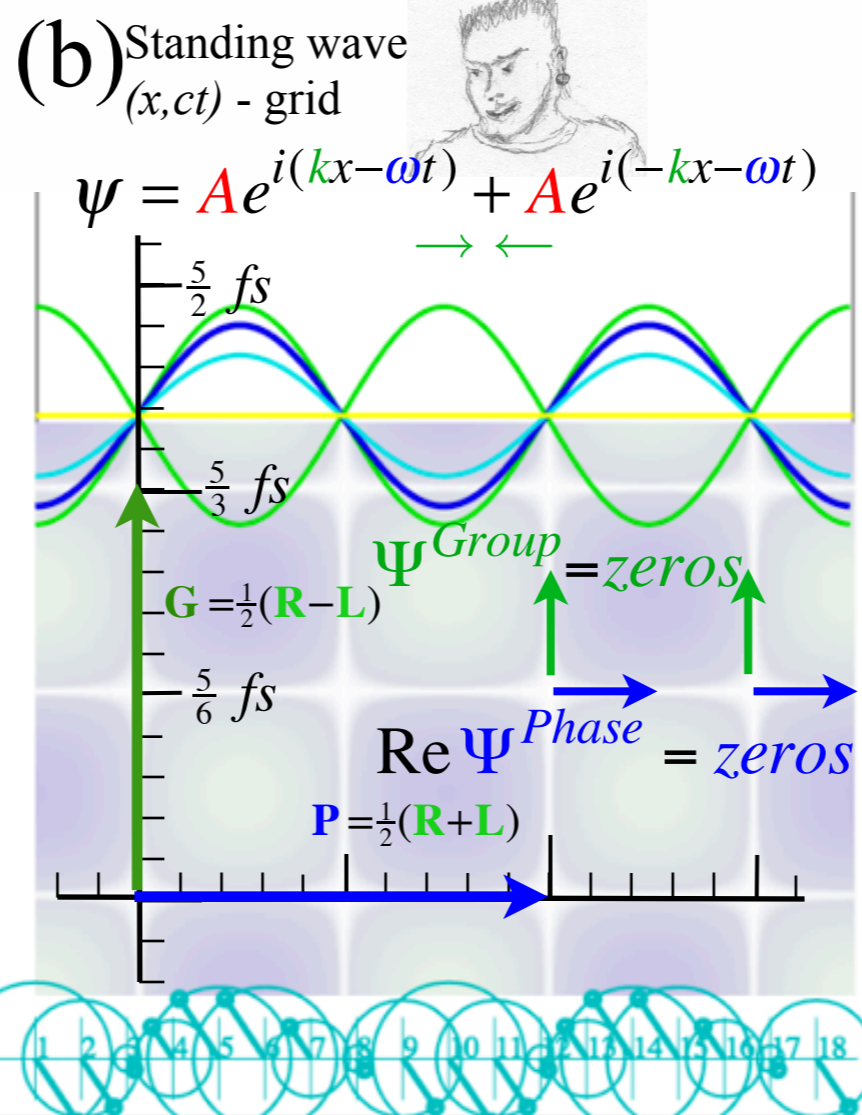
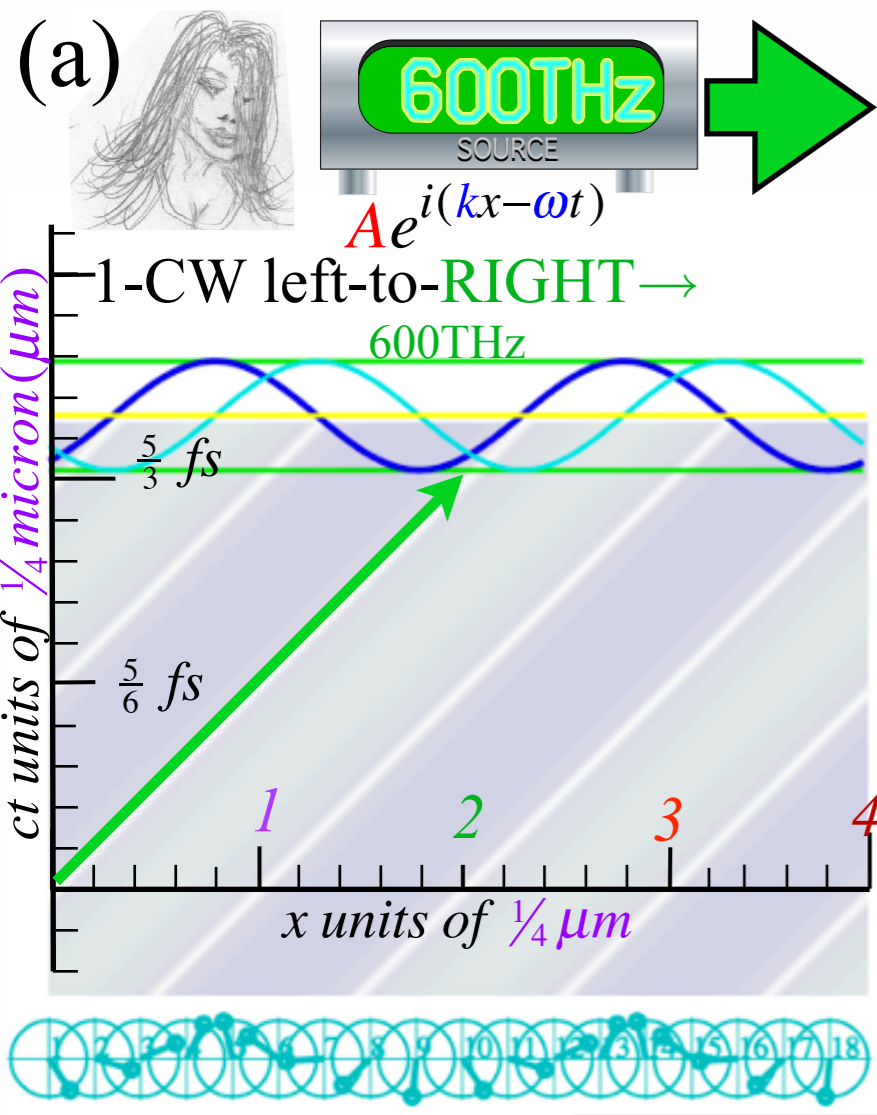
<	Basic CW +1 >	
-1..	CW Light $\pm 1$	+1..
>		
<	PW Lite $\pm 1$	+1..
-1..		>
	CW Light $\pm 2$	
<	PW Lite $\pm 2$	+2..
-2..		>
	CW Light -1 $\diamond$ +4	

-4..		>
	CW Light -2 $\diamond$ +8	
	PW Lit3 -2 $\diamond$ +8	

4

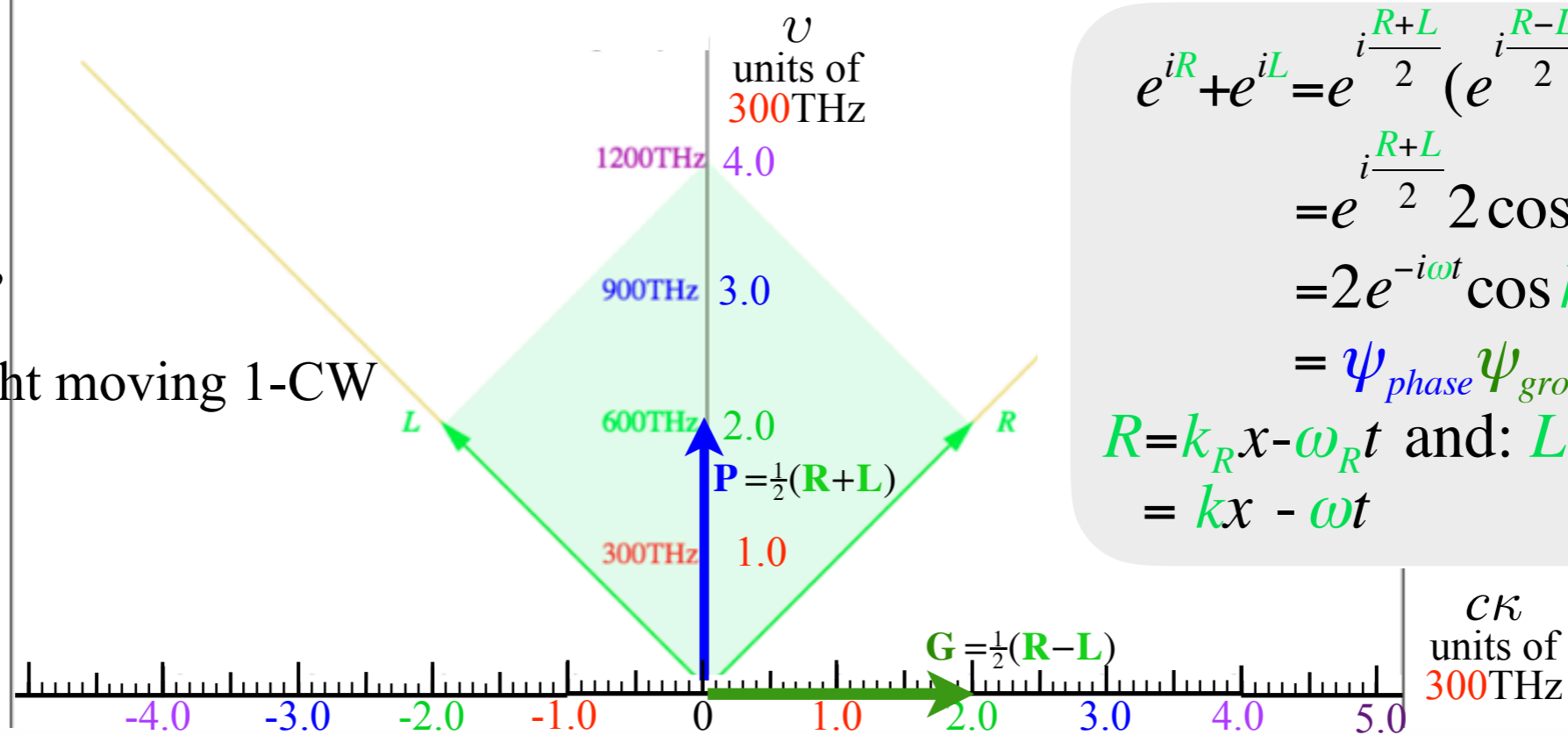






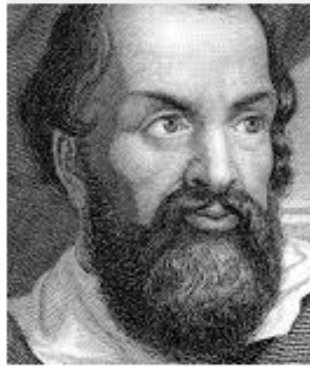
(d) Introducing optical space-time grids and per-space-time “baseball-diamonds”

$\leftarrow$  LEFT-from-right moving 1-CW



$$\begin{aligned}
 e^{iR} + e^{iL} &= e^{i\frac{R+L}{2}} (e^{i\frac{R-L}{2}} + e^{-i\frac{R-L}{2}}) \\
 &= e^{i\frac{R+L}{2}} 2 \cos \frac{R-L}{2} \\
 &= 2e^{-i\omega t} \cos kx \\
 &= \psi_{phase} \psi_{group} \\
 R &= k_R x - \omega_R t \text{ and: } L = -k_L x - \omega_L t \\
 &= kx - \omega t \qquad \qquad \qquad = -kx - \omega t
 \end{aligned}$$

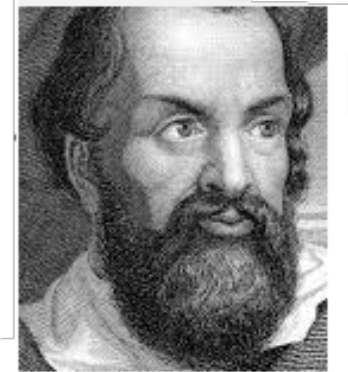
Galileo Galilei



1564-1642

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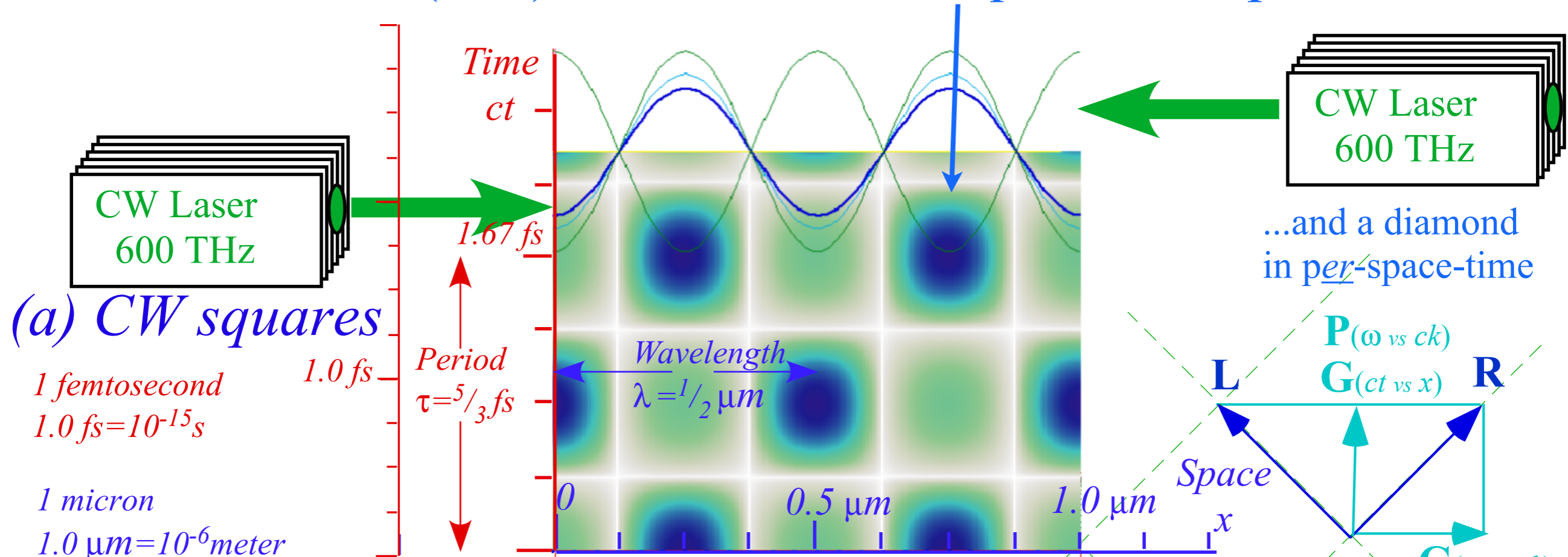
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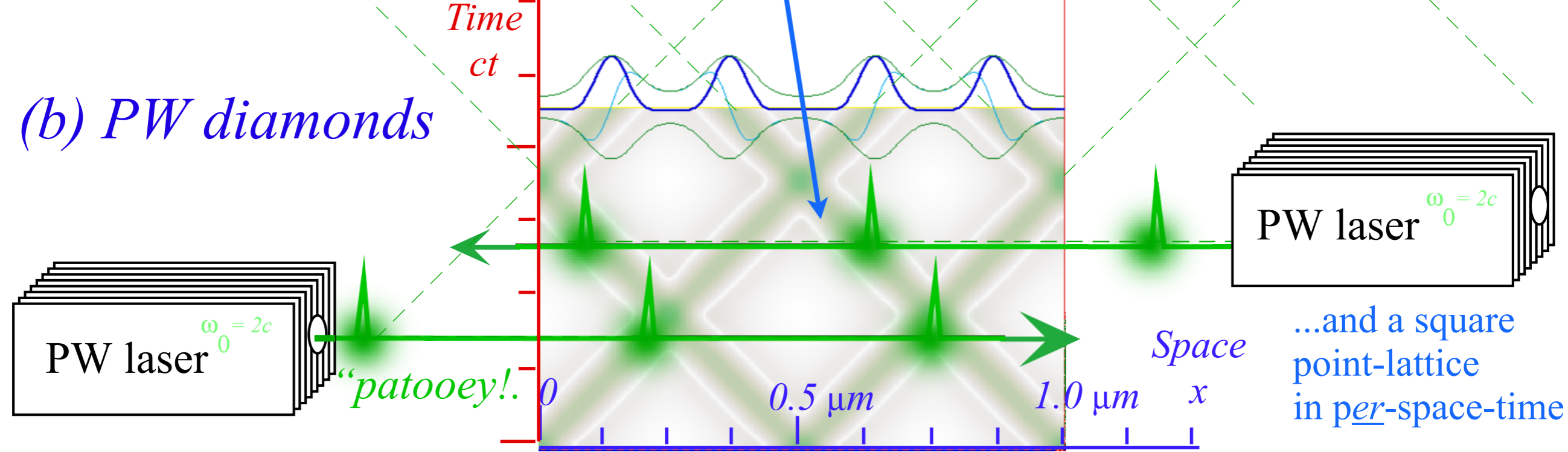
Application to TE-Waveguide modes and synchrotron beam relativity



# Continuous Waves (CW) trace “Cartesian squares” in space-time

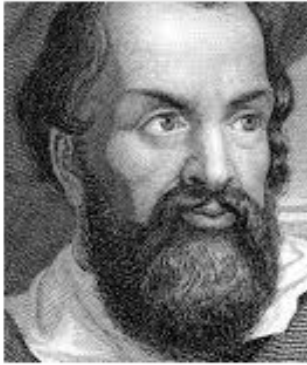


# Pulse Waves (PW) trace “baseball diamonds” in space-time





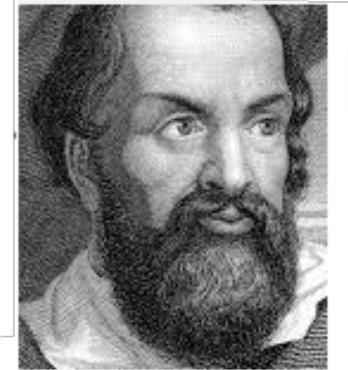
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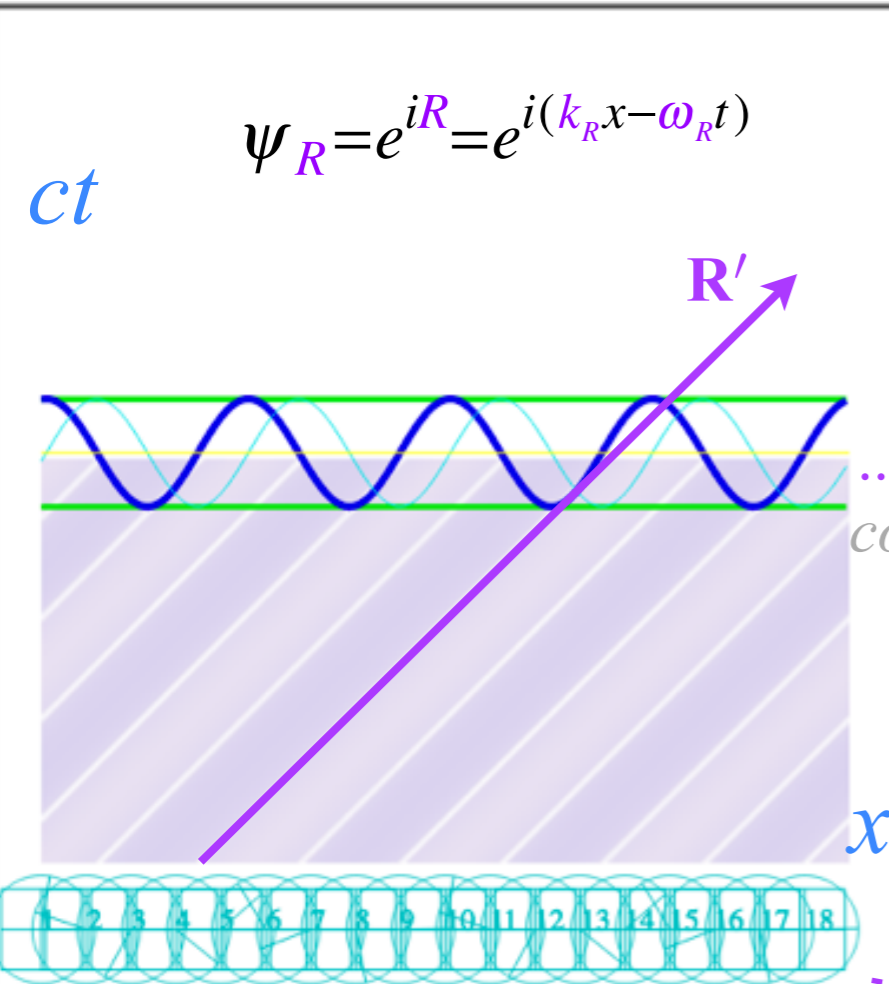
*Application to TE-Waveguide modes and synchrotron beam relativity*

right-moving Doppler blue shifted wave

left-moving Doppler red shifted wave



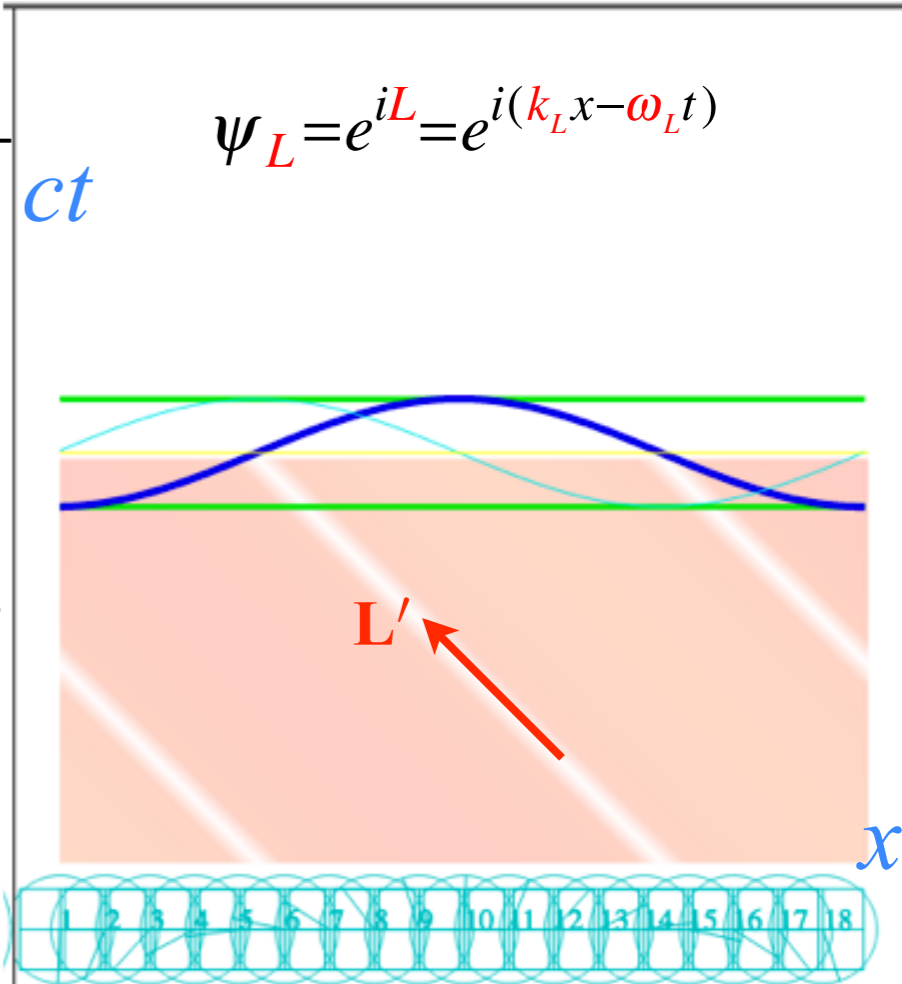
Rapidly moving Bob sees...



$$\psi_R = e^{iR} = e^{i(k_R x - \omega_R t)}$$

...Blue shifted wave coming at him and...

...Red shifted wave behind him.

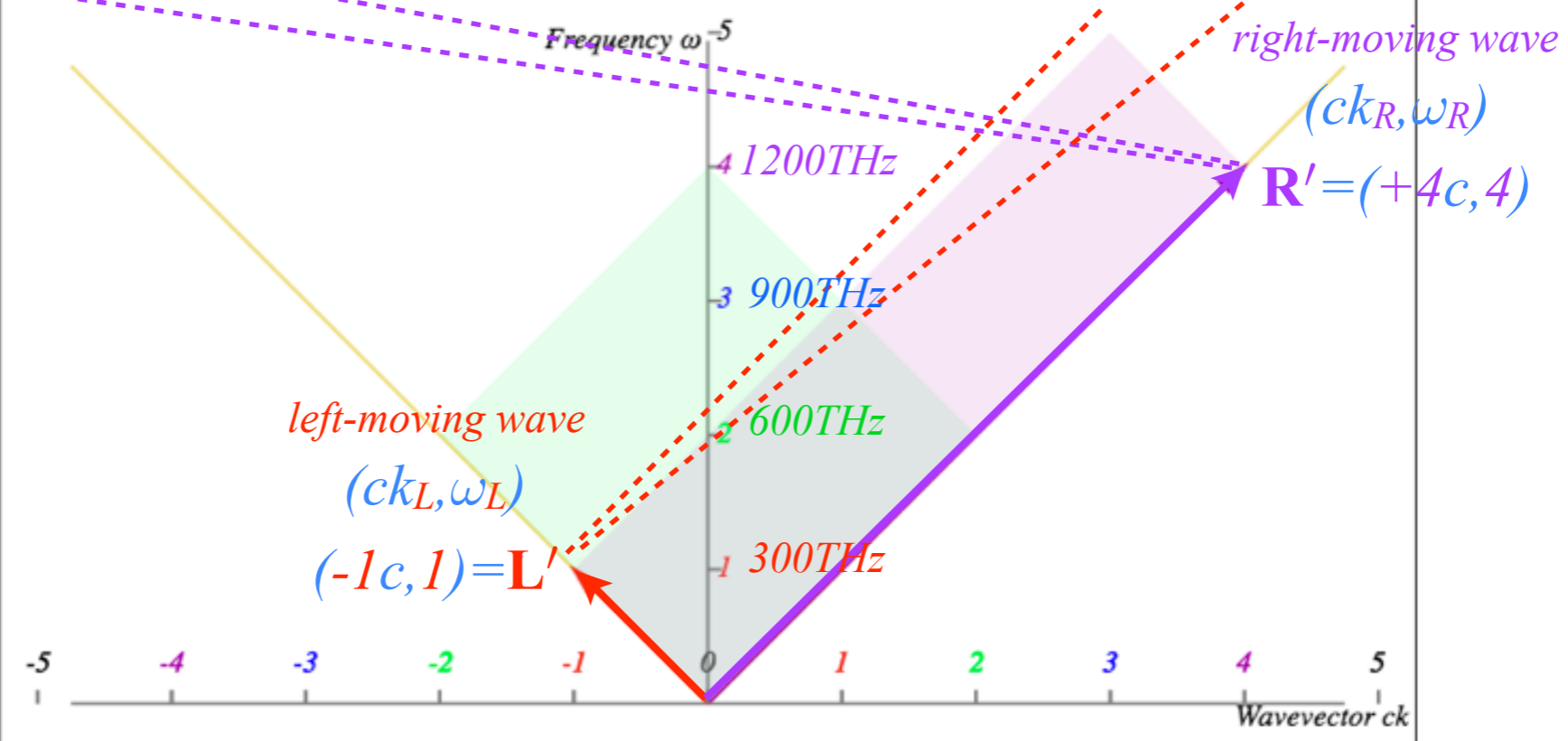


$$\psi_L = e^{iL} = e^{i(k_L x - \omega_L t)}$$

BohrIt Web 1 CW Simulations

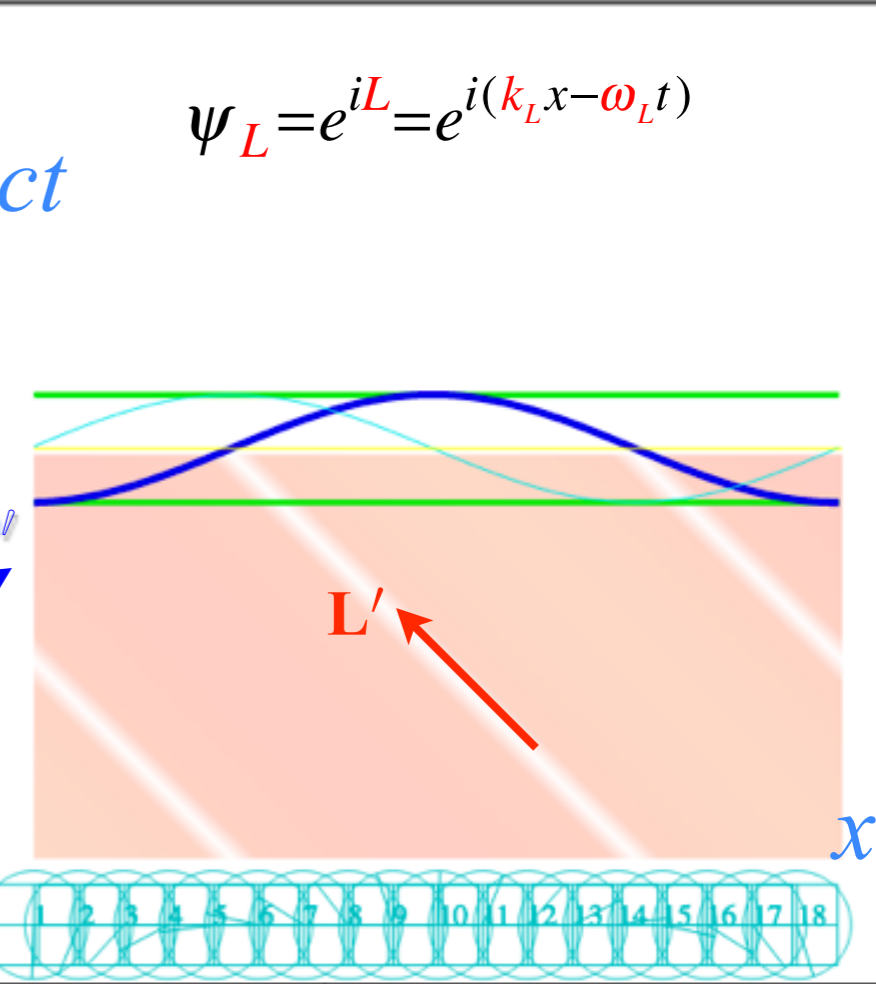
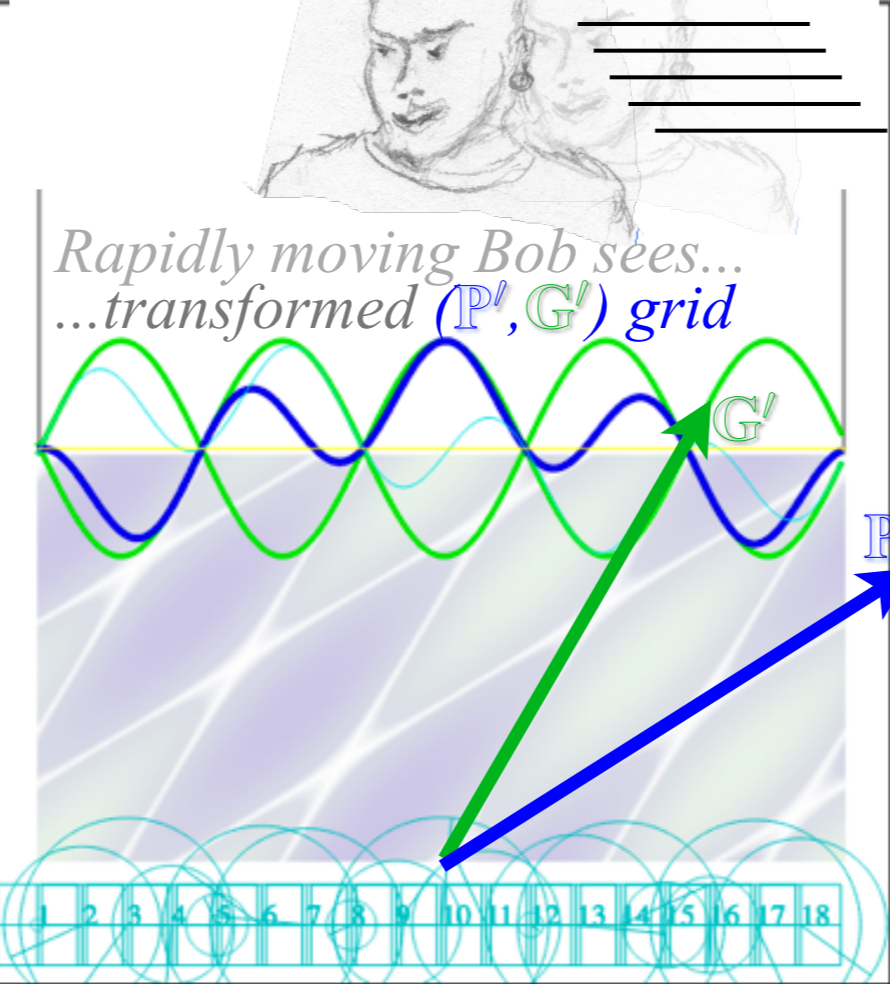
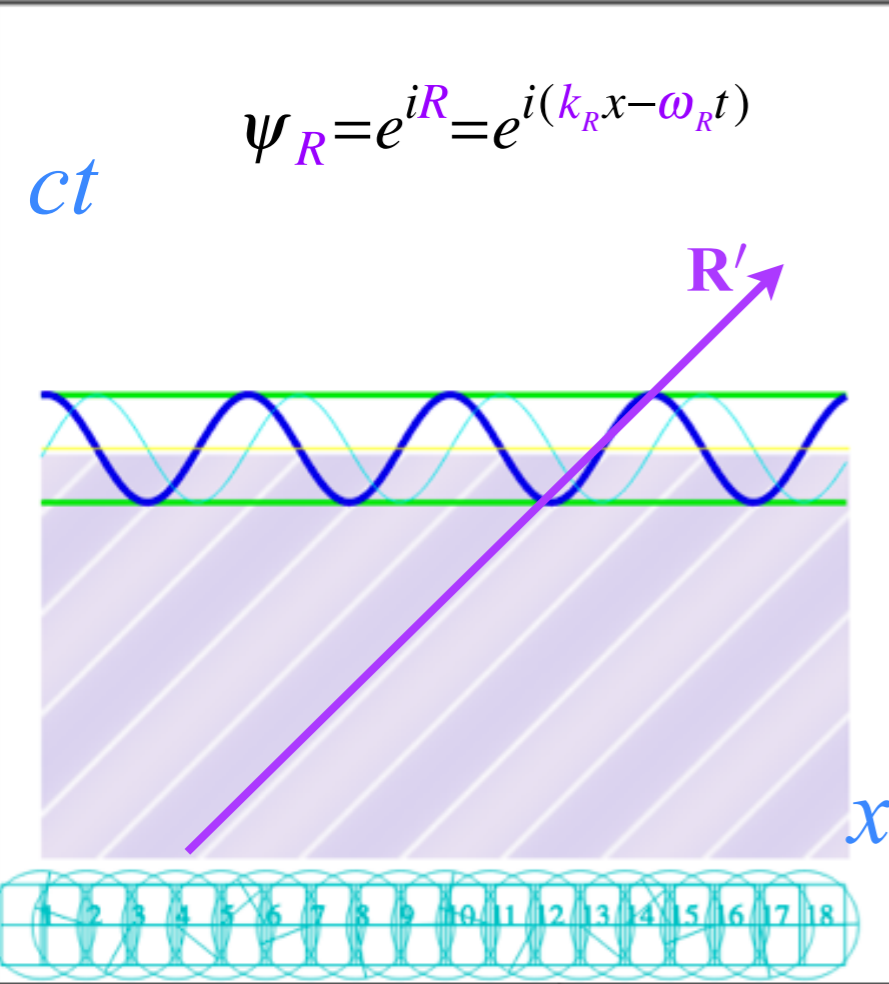
1 CW ct vs x Plot  
( $ck = +4$ )

1 CW ct vs x Plot  
( $ck = -1$ )



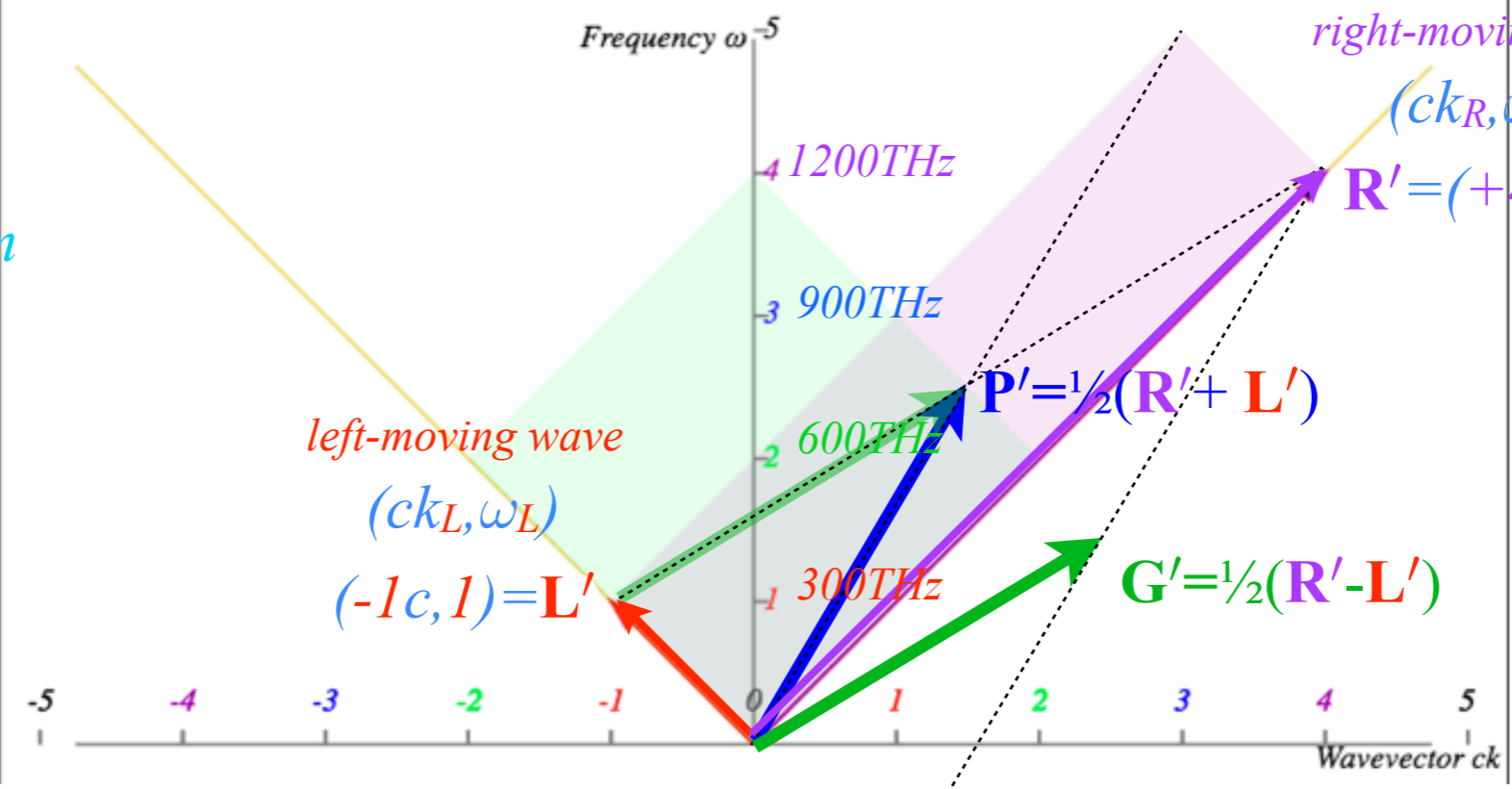
right-moving Doppler blue shifted wave

left-moving Doppler red shifted wave



...Doppler shifts give Lorentz transformation of both these graphs

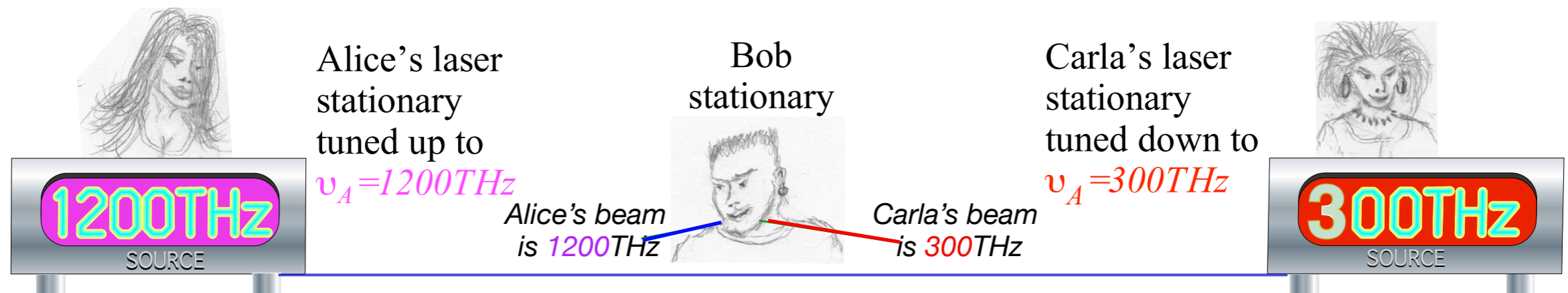
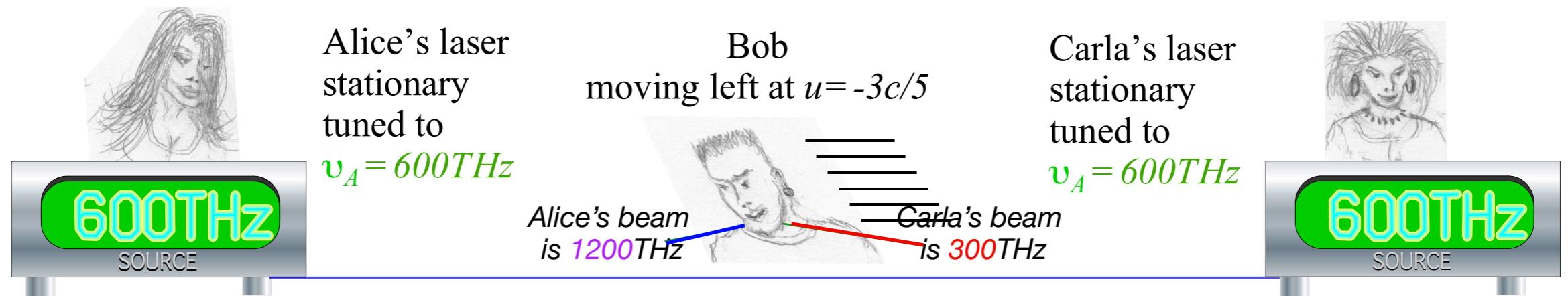
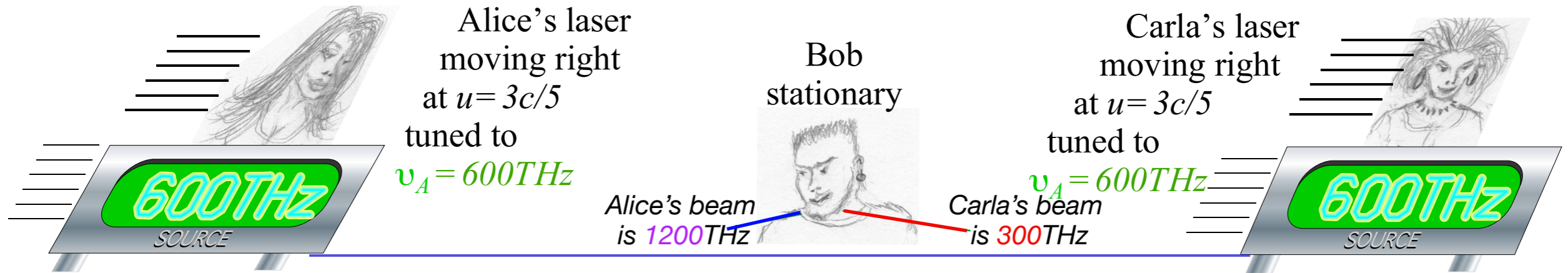
Per-Spacetime  $(ck, \omega)$



BohrIt Web Simulation  
2 CW Minkowski Plot  
 $(ck = -1, +4)$



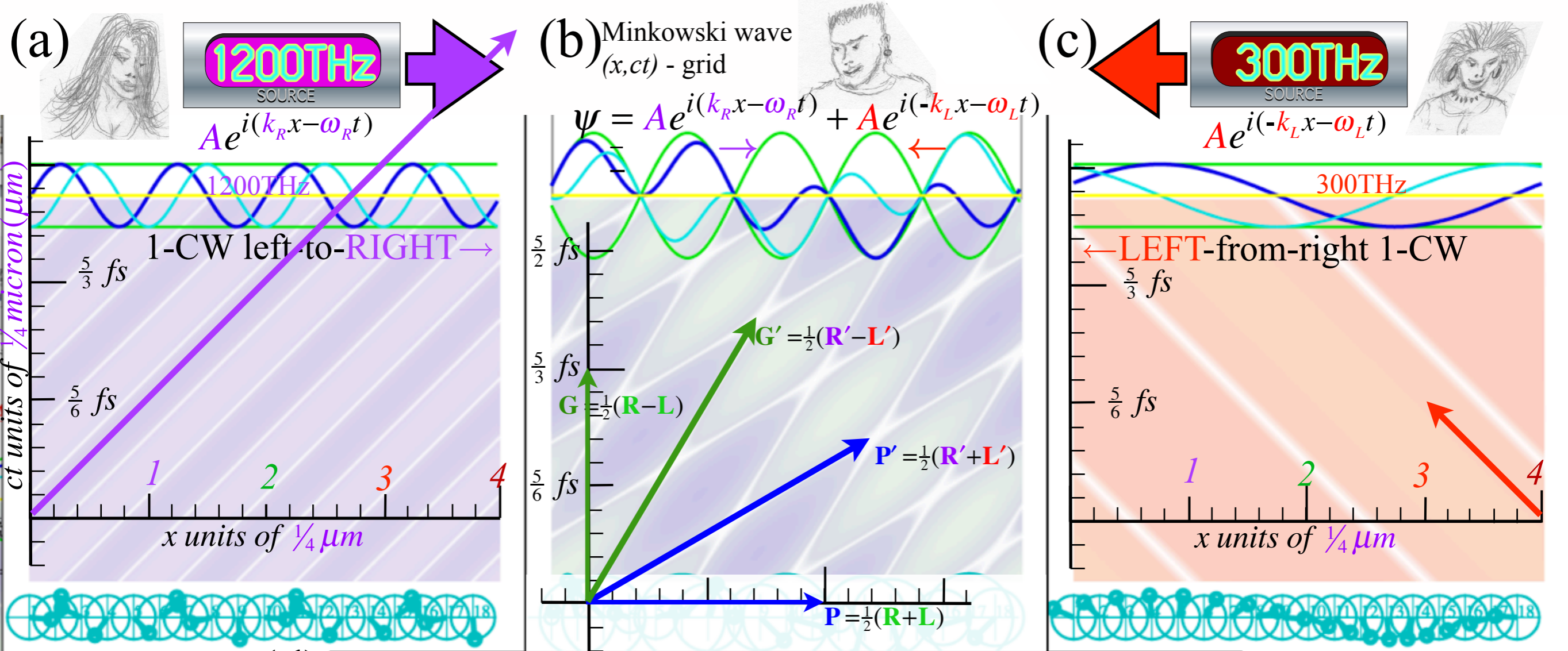
# Three scenarios that look the same to Bob



*Much cheaper (and safer) to do the 3<sup>rd</sup> scenario!\$!*







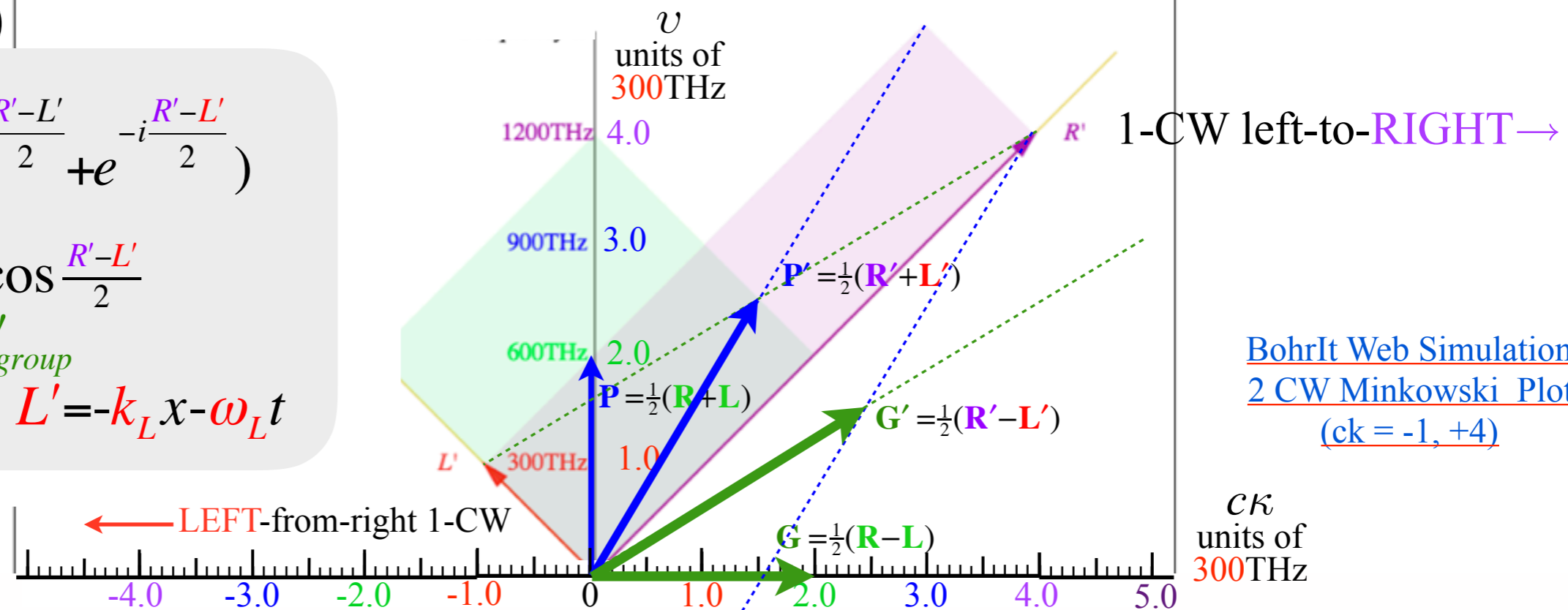
**(d)**

$$e^{iR'} + e^{iL'} = e^{i\frac{R'+L'}{2}} (e^{i\frac{R'-L'}{2}} + e^{-i\frac{R'-L'}{2}})$$

$$= e^{i\frac{R'+L'}{2}} 2 \cos \frac{R'-L'}{2}$$

$$= \psi'_{phase} \psi'_{group}$$

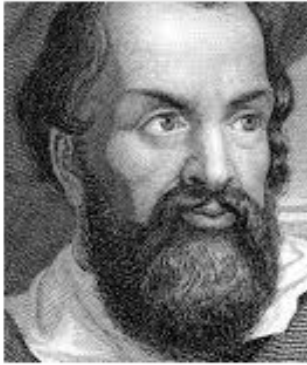
$R' = k_R x - \omega_R t$  and:  $L' = -k_L x - \omega_L t$



[BohrIt Web Simulation](#)  
[2 CW Minkowski Plot](#)  
 (ck = -1, +4)



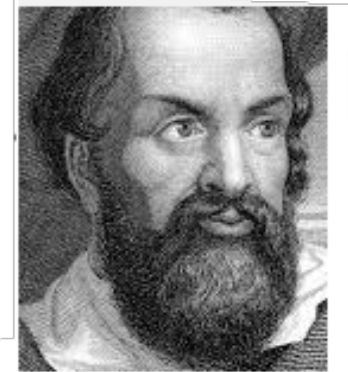
Galileo Galilei



1564-1642

**Galileo's Revenge (part 1)**

*Rapidity adds just like  
Galilean velocity*



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*Phasor angular velocity  
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Optical interference “baseball-diamond” displays *phase* and *group* velocity

Details of 2CW wavefunctions in rest frame

Pulse waves (PW) versus Continuous Waves (CW)

Doppler shifted “baseball-diamond” displays Lorentz frame transformation

➔ Analyzing wave velocity by *per-space-per-time* and *space-time* graphs

16 coefficients of relativistic 2CW interference

Two “famous-name” coefficients and the Lorentz transformation

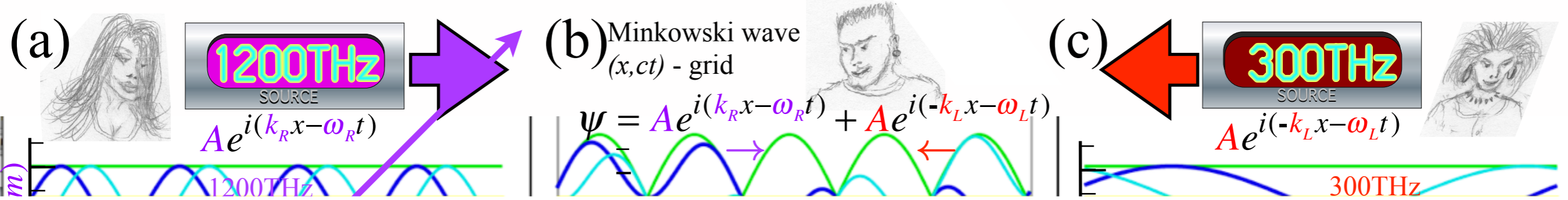
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*Rapidity  $\rho$  related to stellar aberration angle  $\sigma$  and L. C. Epstein's approach to relativity*

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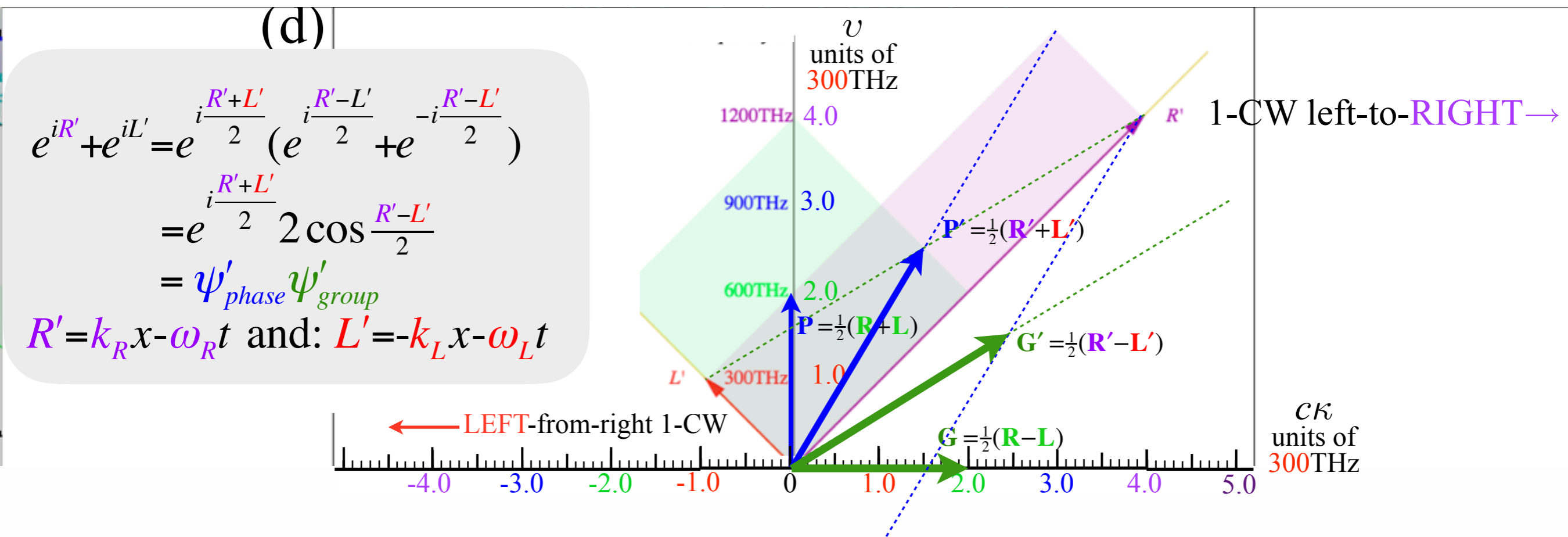
*“Occams Sword” and geometry of 16 parameter functions of  $\rho$  and  $\sigma$*

*Application to TE-Waveguide modes and synchrotron beam relativity*



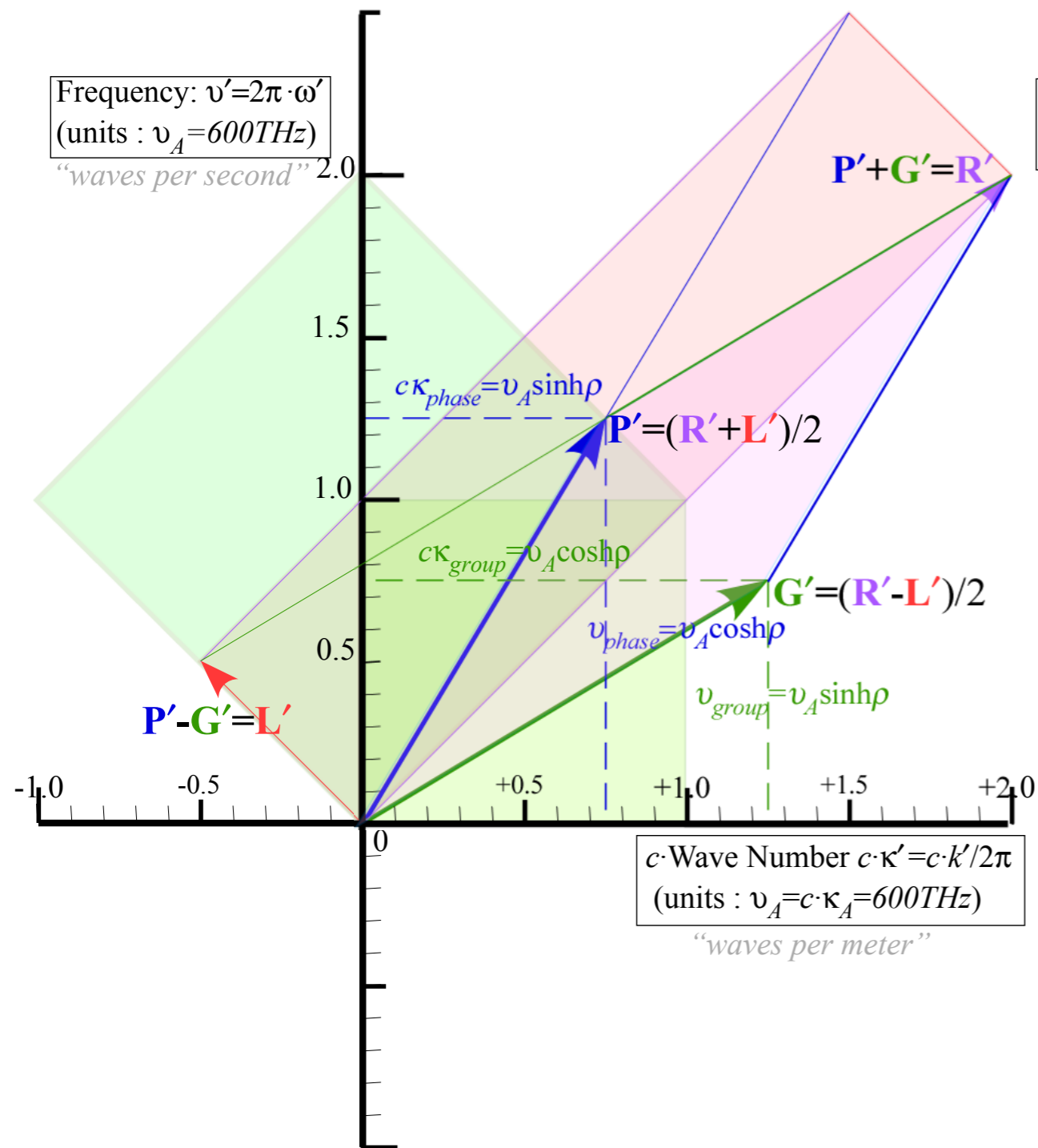
$$\mathbf{P}' = \begin{pmatrix} v'_{phase} \\ cK'_{phase} \end{pmatrix} = \frac{1}{2}(\mathbf{R}' + \mathbf{L}') = v_A \begin{pmatrix} \frac{1}{2}(e^\rho + e^{-\rho}) \\ \frac{1}{2}(e^\rho - e^{-\rho}) \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} \frac{5}{2} \\ \frac{3}{2} \end{pmatrix} \text{Bob's View} \quad \text{or: } v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \text{Alice's View}$$

$$\mathbf{G}' = \begin{pmatrix} v'_{group} \\ cK'_{group} \end{pmatrix} = \frac{1}{2}(\mathbf{R}' - \mathbf{L}') = v_A \begin{pmatrix} \frac{1}{2}(e^\rho - e^{-\rho}) \\ \frac{1}{2}(e^\rho + e^{-\rho}) \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} \frac{3}{2} \\ \frac{5}{2} \end{pmatrix} \text{Bob's View} \quad \text{or: } v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \text{Alice's View}$$

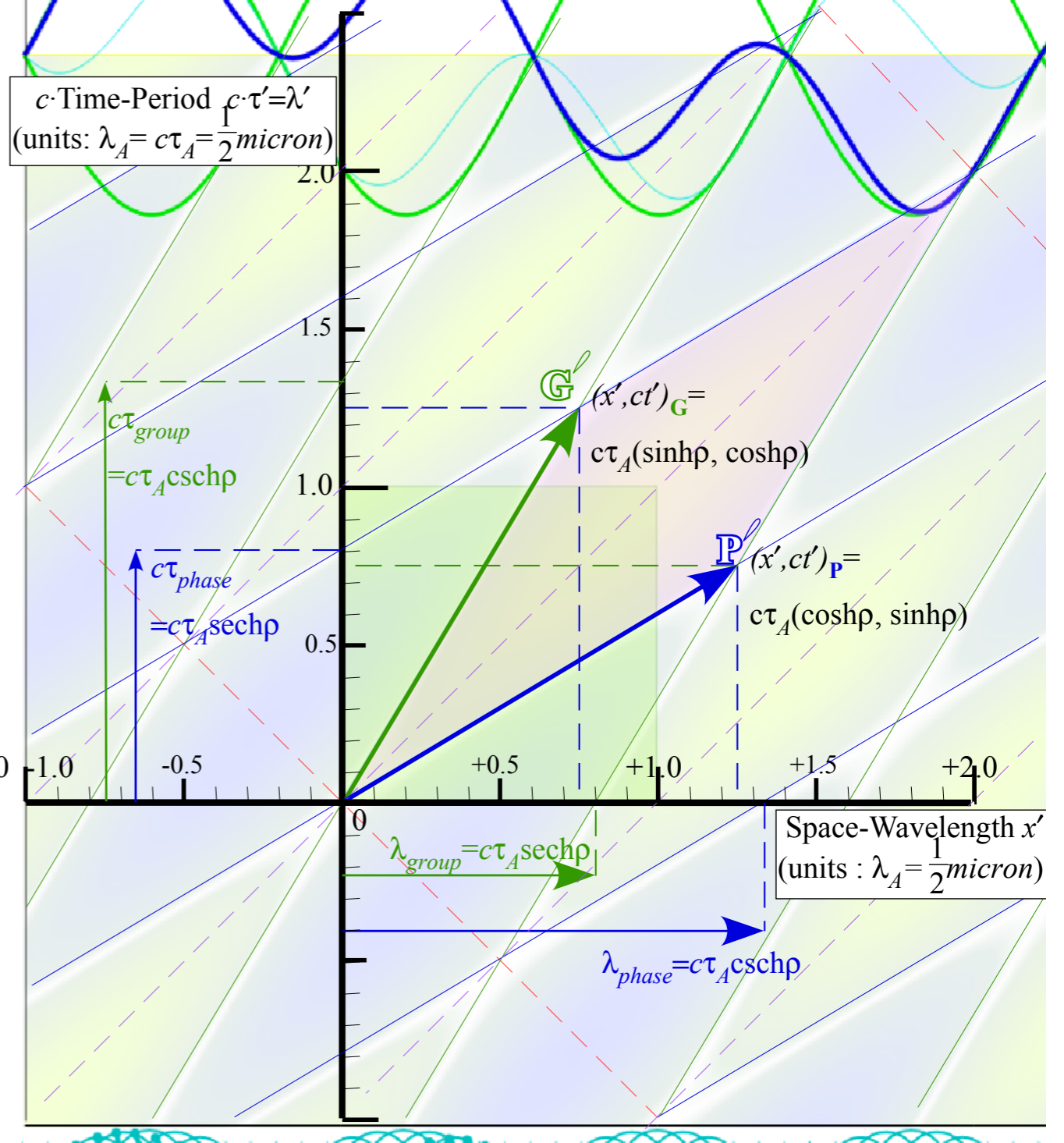




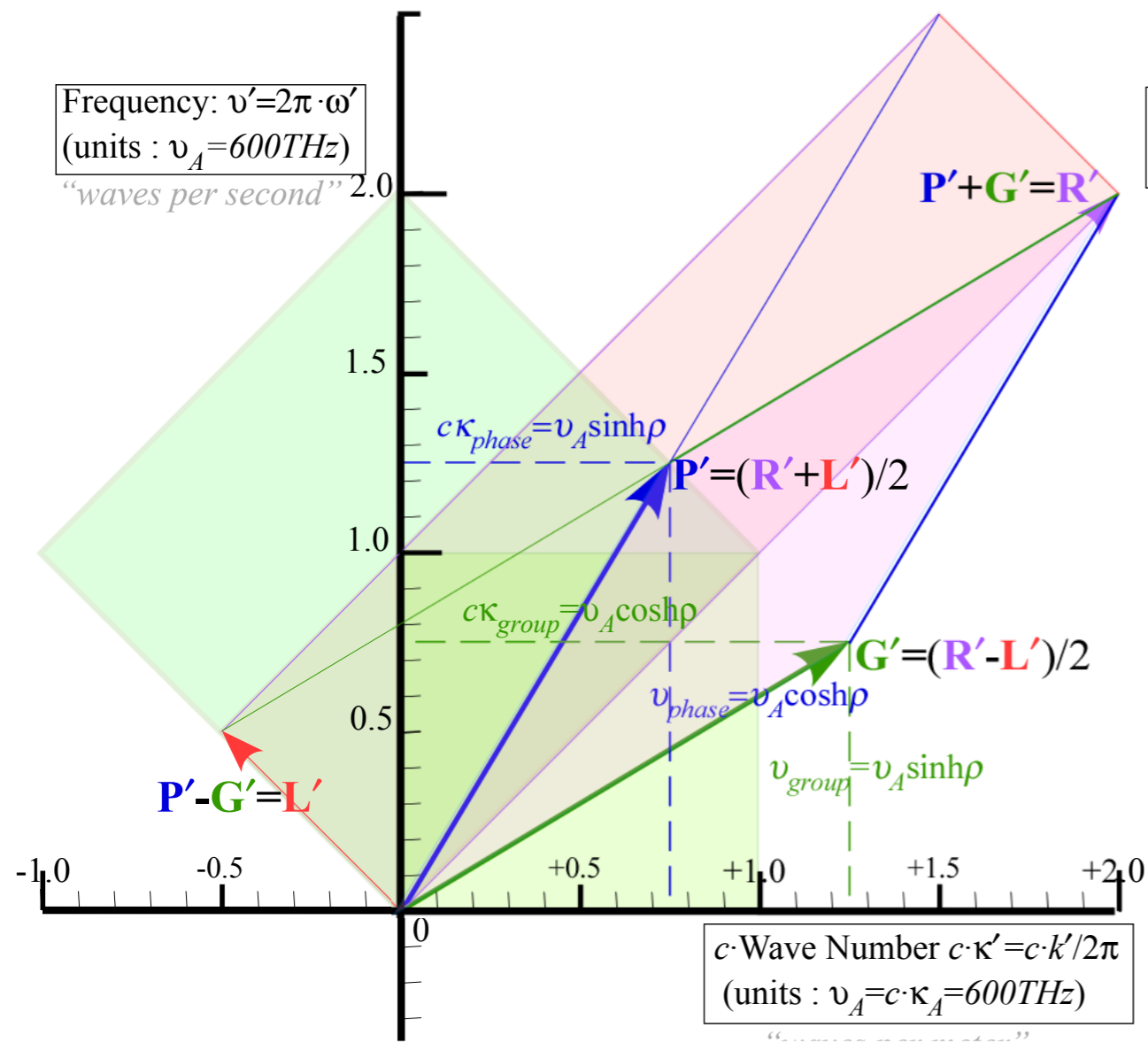
(a) Per-space-time  $(\nu', c\kappa')$  geometry of 2-CW vectors



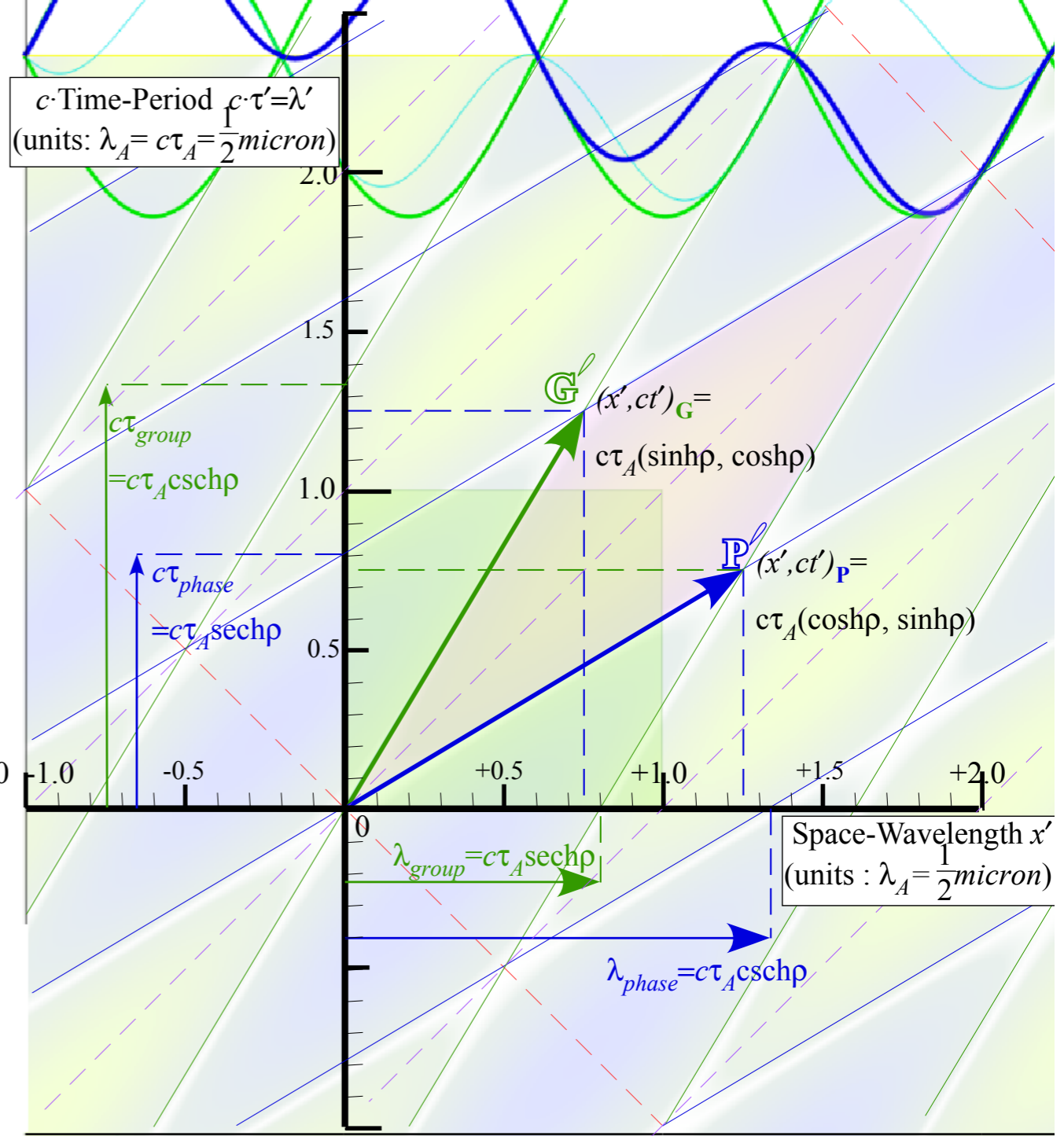
(b) Space-time  $(c\tau', x')$  geometry of 2-CW paths



(a) Per-space-time  $(v', c\kappa')$  geometry of 2-CW vectors



(b) Space-time  $(c\tau', x')$  geometry of 2-CW paths

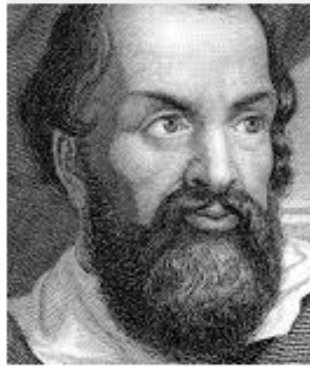


The slope of Bob's group vector  $\mathbf{G}'$  in  $(c\kappa, v)$ -plot is actual group wave velocity in  $c$ -units.

$$\frac{V^{group}}{c} = \frac{v'_{group}}{c\kappa'_{group}} = \frac{\sinh \rho}{\cosh \rho} = \tanh \rho = \frac{\frac{3}{2}}{\frac{5}{2}} = \frac{3}{5} \equiv \frac{u}{c} \equiv \beta$$

Group vector  $\mathbf{G}'$  in  $(x, ct)$ -plot has 3/5 slope relative to time axis

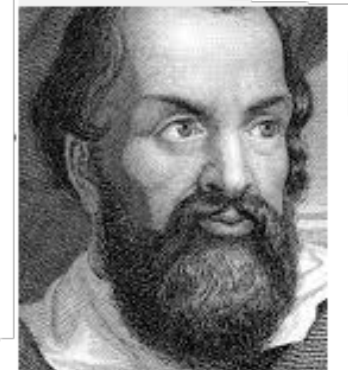
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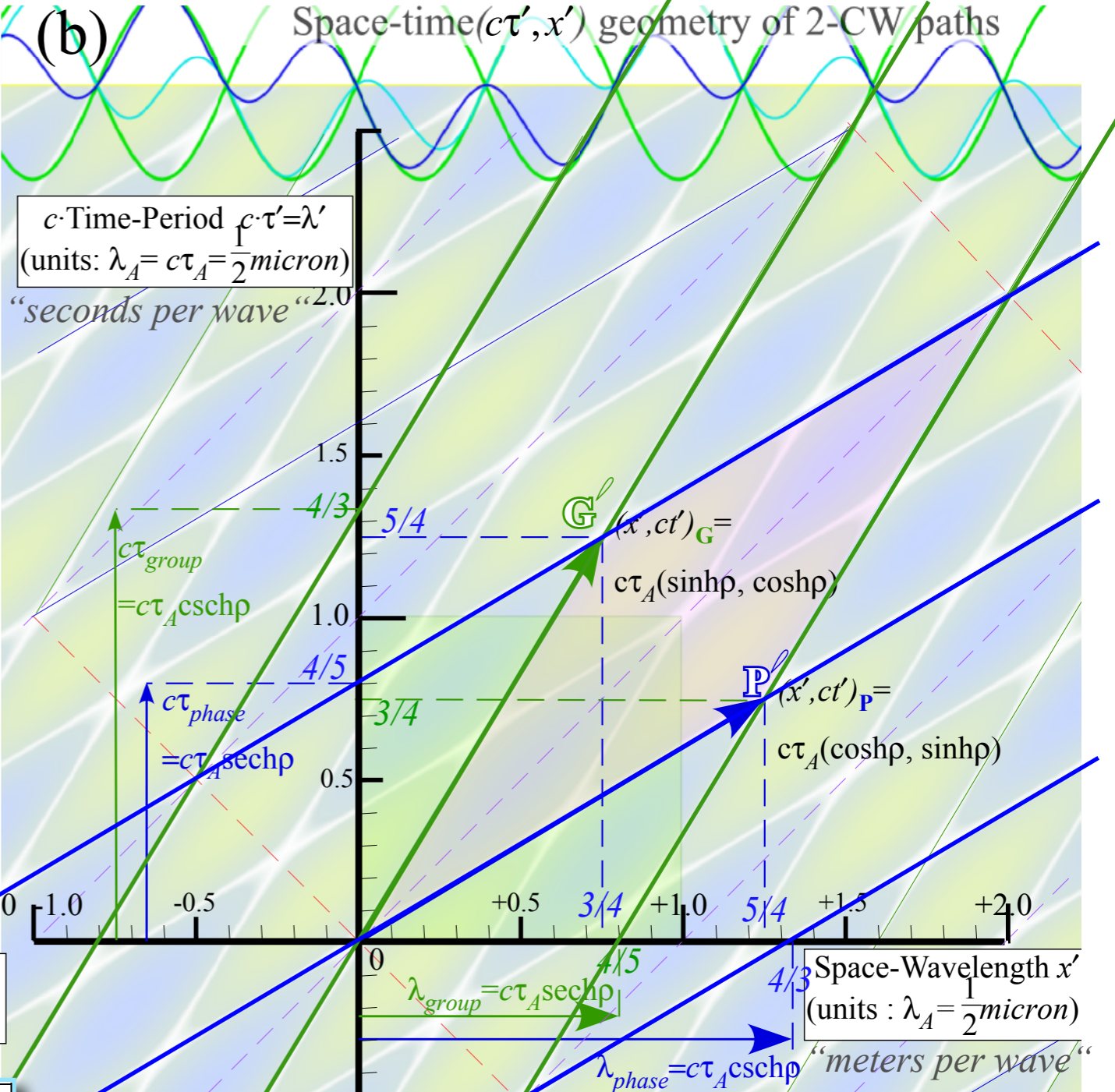
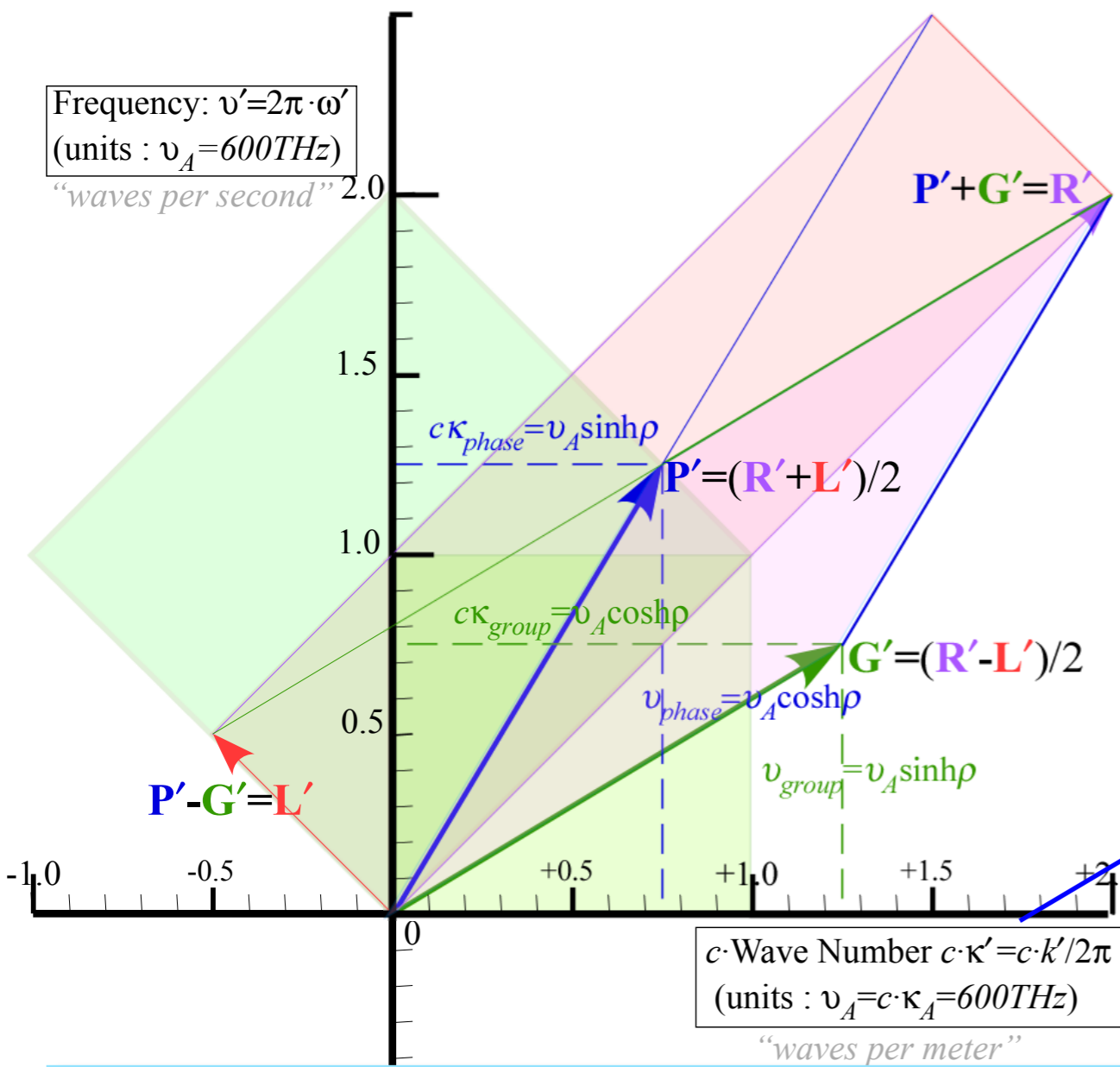
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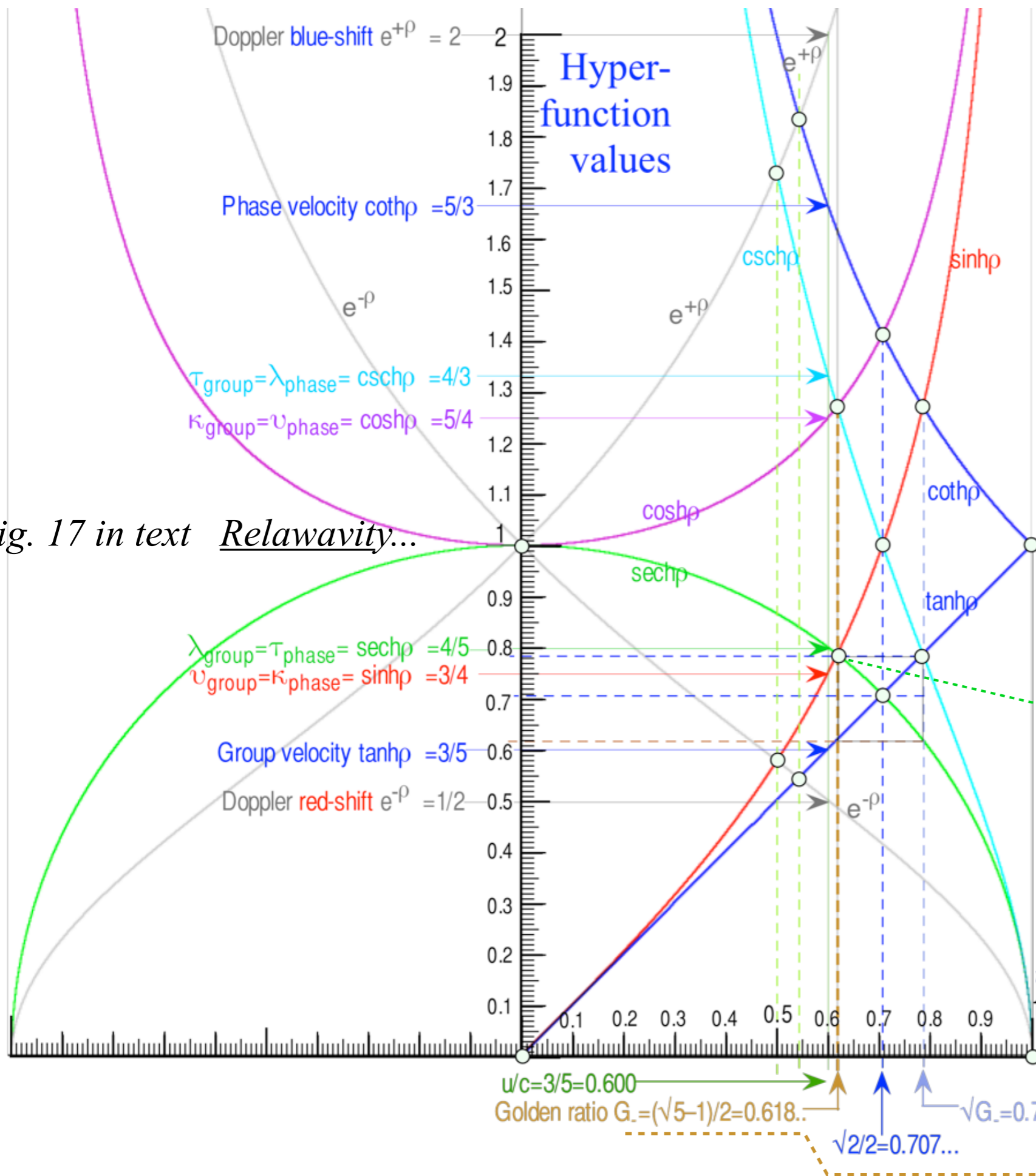


(a) Per-space-time  $(v', c\kappa')$  geometry of 2-CW vectors



group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Fig. 11 in text *Relativity...*



If  $\frac{u}{c} = \tanh \rho = 0.618\dots$  (Golden-Mean  $G_-$ )

two parameters become *exactly* equal :

$$\frac{ct'_P}{c\tau_A} = \sinh \rho = \frac{\lambda_{\text{group}}}{\lambda_A} = \frac{\tau_{\text{phase}}}{\tau_A} = \text{sech} \rho = 0.786\dots = \sqrt{G_-}$$

and

$$\frac{x'_P}{\lambda_A} = \cosh \rho = \frac{\lambda_{\text{phase}}}{\lambda_A} = \frac{\tau_{\text{group}}}{\tau_A} = \text{csch} \rho = 1.272\dots = 1/\sqrt{G_-}$$

Fig. 17 in text Relativity...

Solve :

$$\text{sech} \rho = \sinh \rho$$

or:

$$\sinh \rho \cosh \rho = 1$$

or:

$$\sinh 2\rho = 2$$

$$\rho = \frac{1}{2} \sinh^{-1} 2 = 0.7218\dots$$

$$\tanh \rho = 0.618\dots = \frac{\sqrt{5}-1}{2}$$



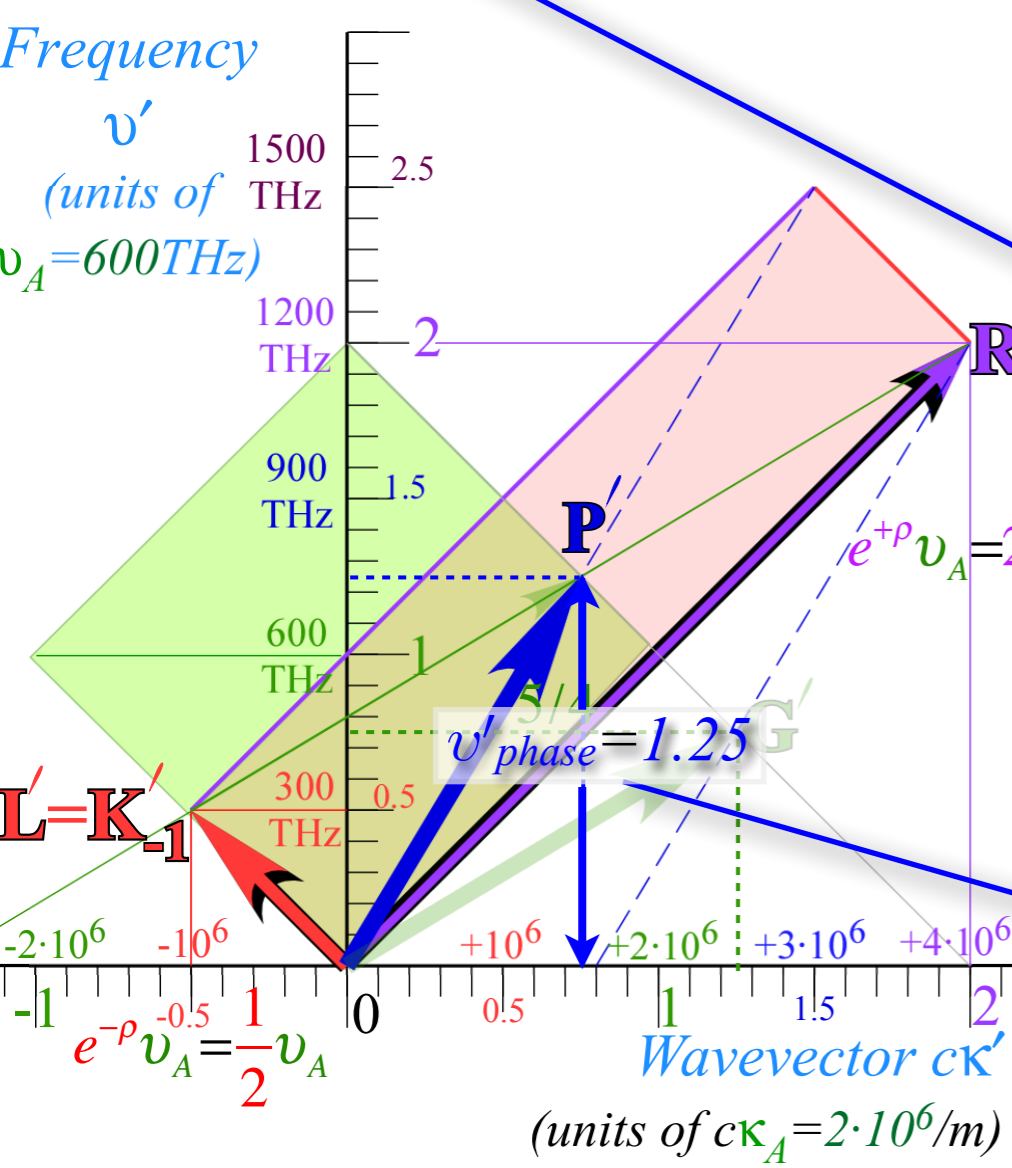
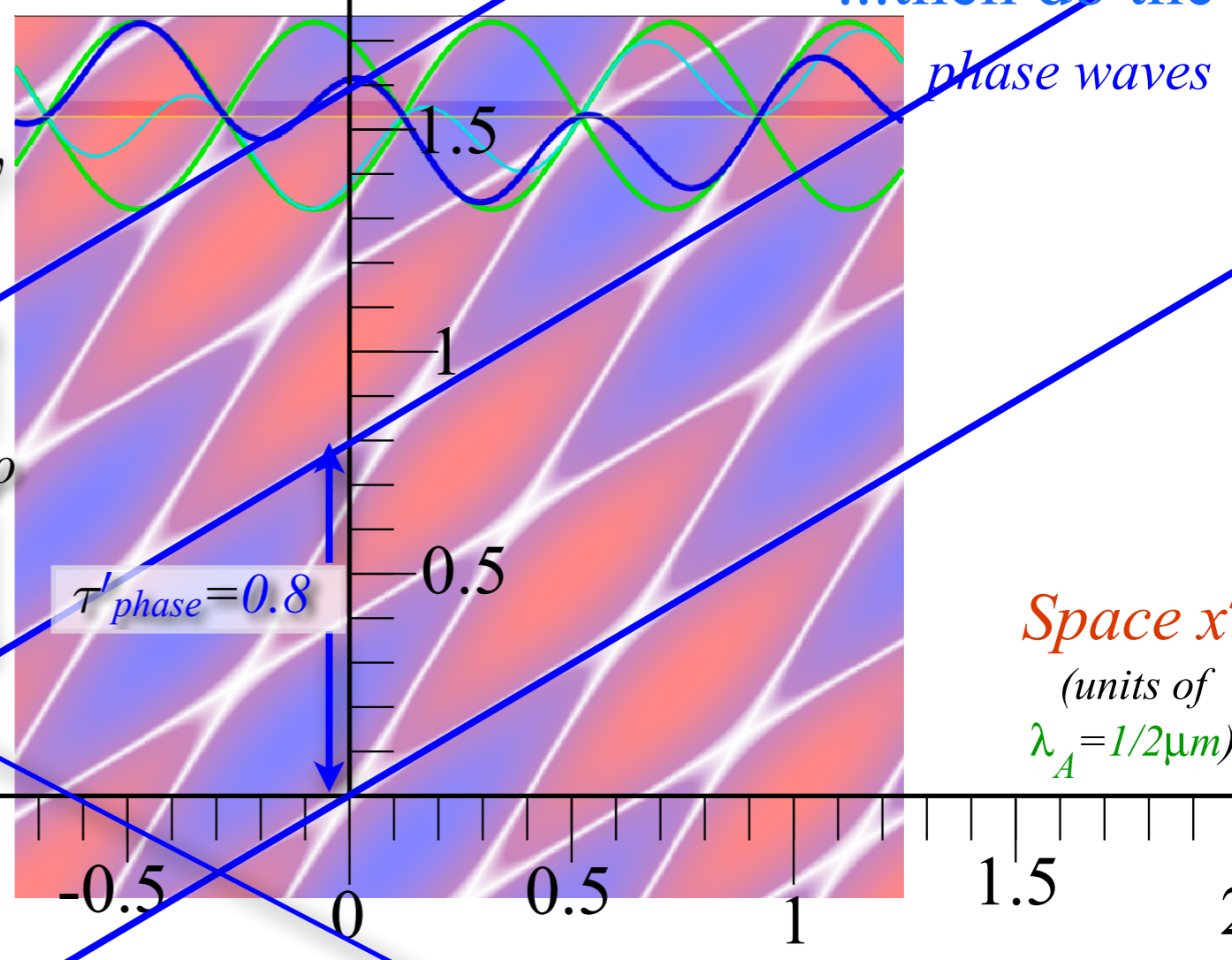
# The 16 dimensions of 2CW interference

*Time  $ct'$*   
(units of  $\lambda_A = 1/2\mu\text{m}$ )

Start with the *Dopplers*  
...then do the *phase waves*

$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

Phase frequency  $v'_{phase} = v_A \cosh \rho = 5/4 = 1.25$  flips to Phase period  $\tau'_{phase} = \tau_A \text{sech} \rho = 4/5 = 0.8$



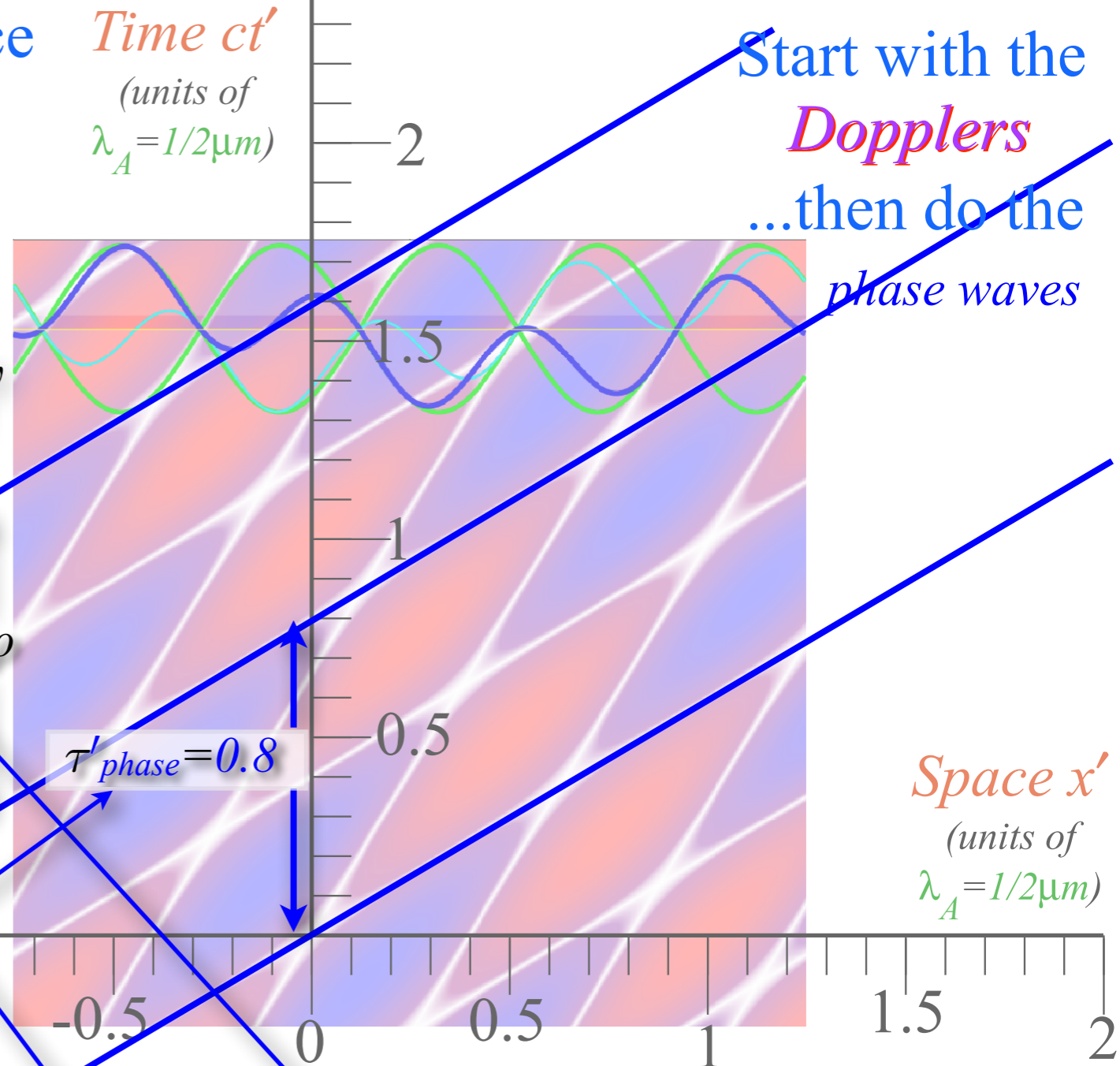
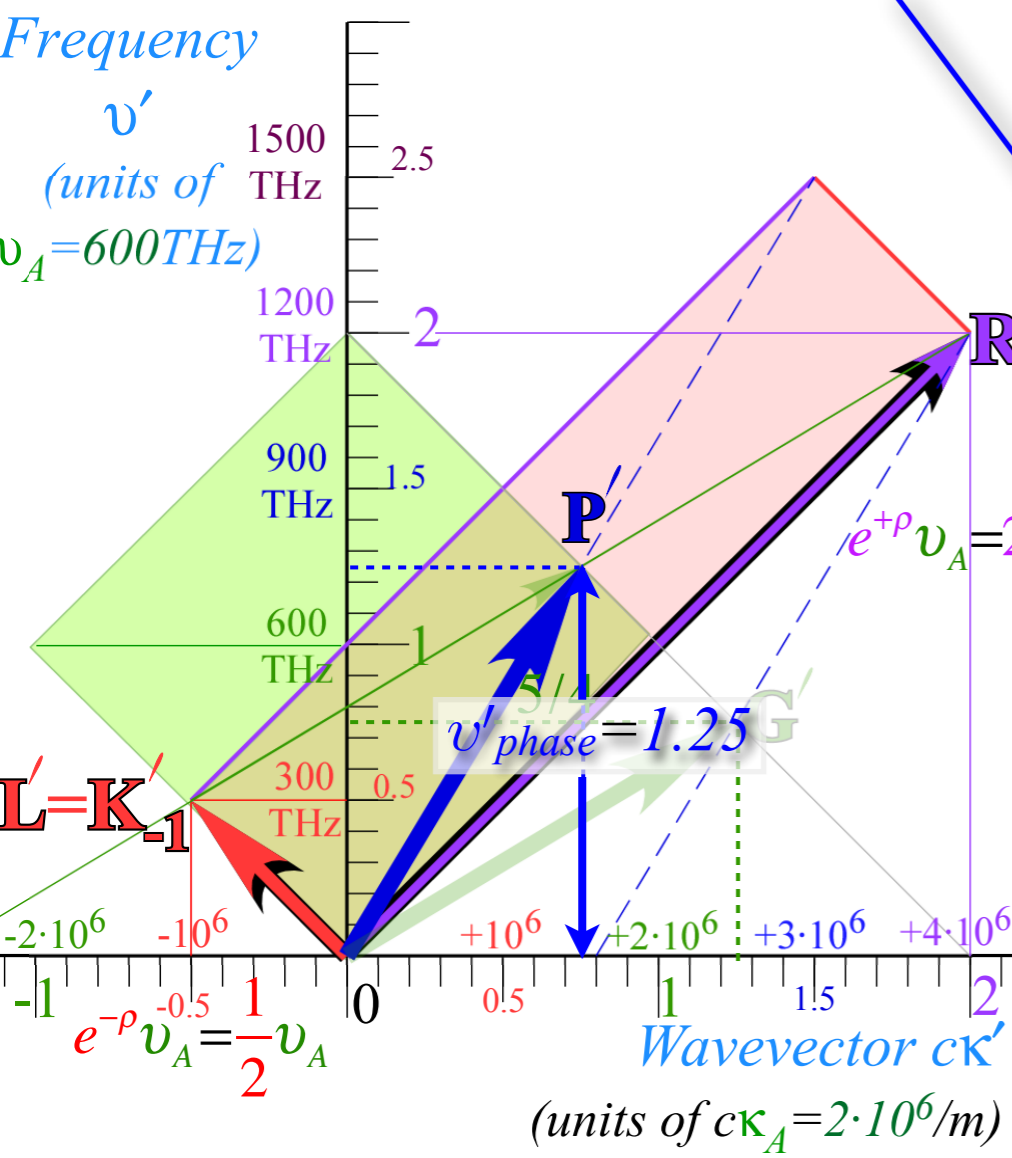
phase	$b_{\text{Doppler RED}}$	$\frac{c}{v_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{\text{Doppler BLUE}}$
group	1	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{\text{Doppler RED}}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech} \rho$	$\cosh \rho$	$\text{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$



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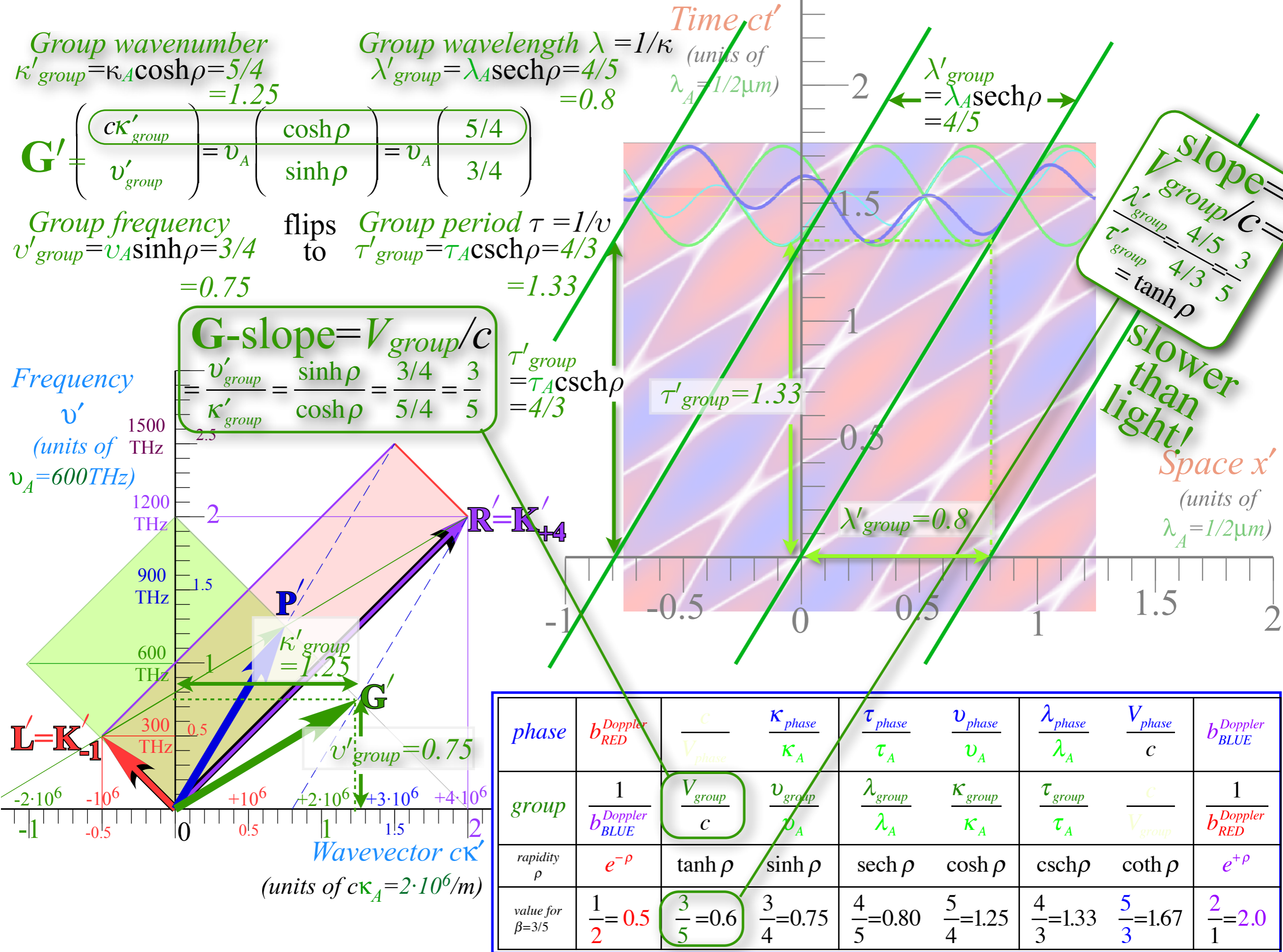


Start with the Dopplers ... then do the phase waves

phase	$b_{RED}^{Doppler}$	$\frac{v_{phase}}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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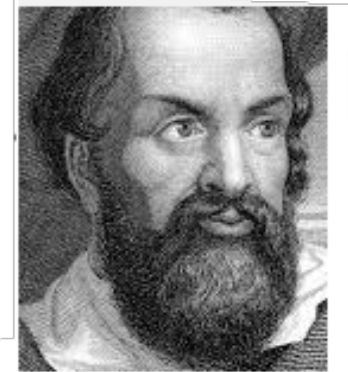
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# Lorentz transformations...

write  $\mathbf{G}'$  and  $\mathbf{P}'$  in terms of  $\mathbf{G}$  and  $\mathbf{P}$  using  $\cosh \rho$  and  $\sinh \rho$

$$\mathbf{G}' = \begin{pmatrix} c\mathbf{K}'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

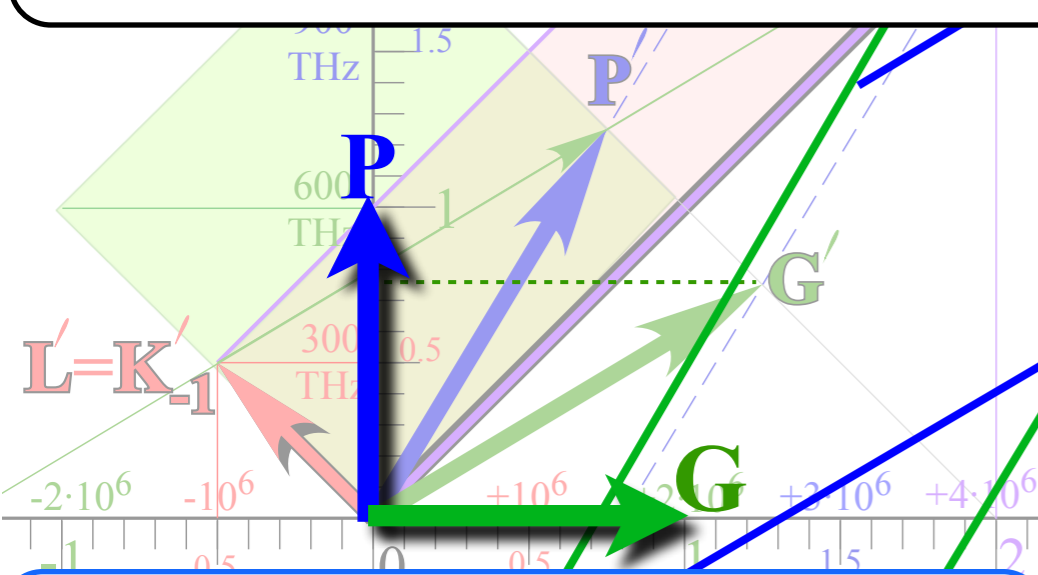
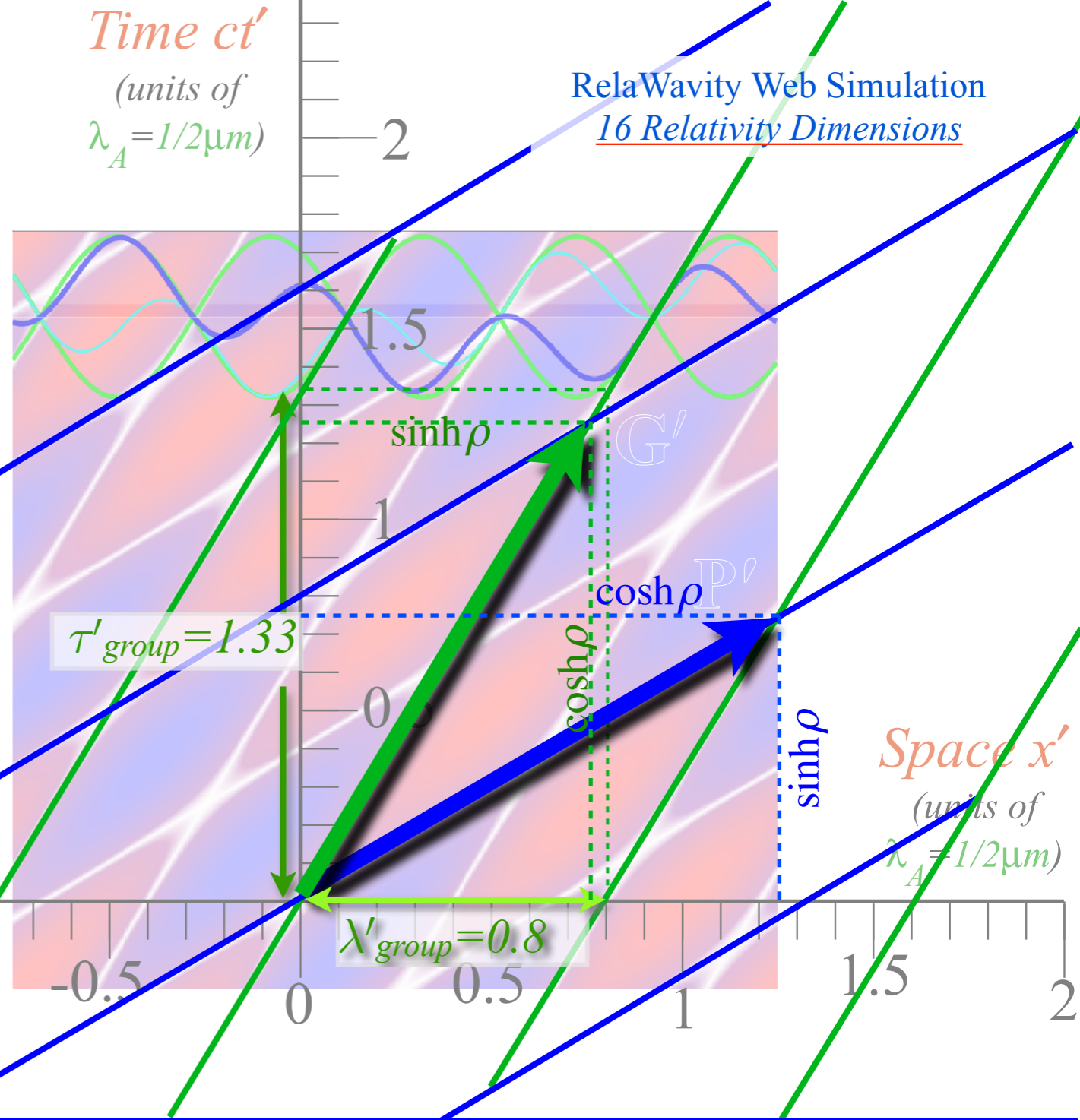
$$= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \cosh \rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \sinh \rho$$

$$\mathbf{G}' = \mathbf{G} \cosh \rho + \mathbf{P} \sinh \rho$$

$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

$$= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \sinh \rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \cosh \rho$$

$$\mathbf{P}' = \mathbf{G} \sinh \rho + \mathbf{P} \cosh \rho$$



phase	$b_{Doppler RED}$	$\frac{c}{V_{phase}}$	$\frac{\mathbf{K}_{phase}}{\mathbf{K}_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
group	$\frac{1}{b_{Doppler BLUE}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\mathbf{K}_{group}}{\mathbf{K}_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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$$\begin{pmatrix} \cosh \rho & \sinh \rho \\ \sinh \rho & \cosh \rho \end{pmatrix} \text{ Lorentz transform matrix}$$

# Two Famous-Name Coefficients

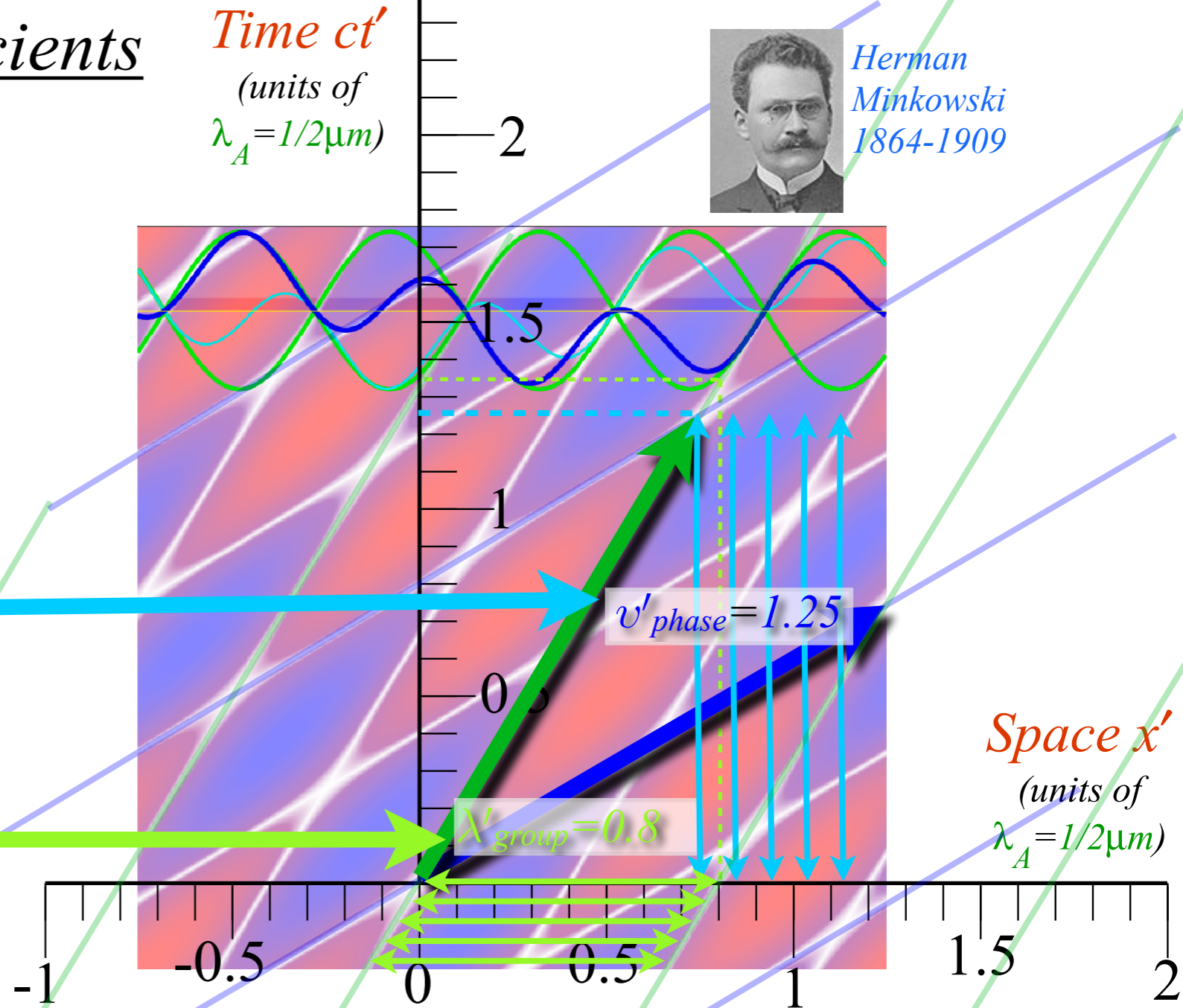
Time  $ct'$   
(units of  
 $\lambda_A = 1/2\mu m$ )



Herman  
Minkowski  
1864-1909

This number  
is called an: **Einstein  
time-dilation**  
(dilated by 25% here)

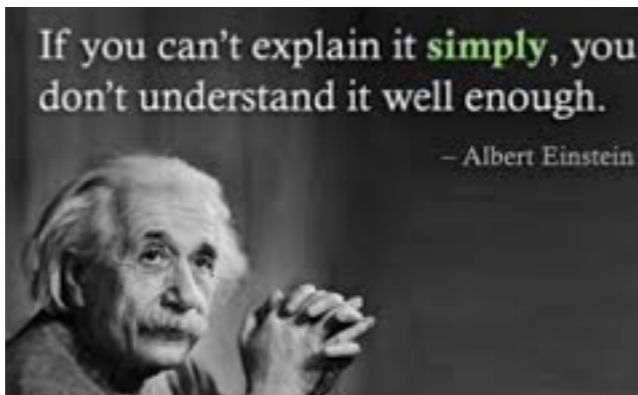
This number  
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length-contraction**  
(contracted by 20% here)





# Two Famous-Name Coefficients

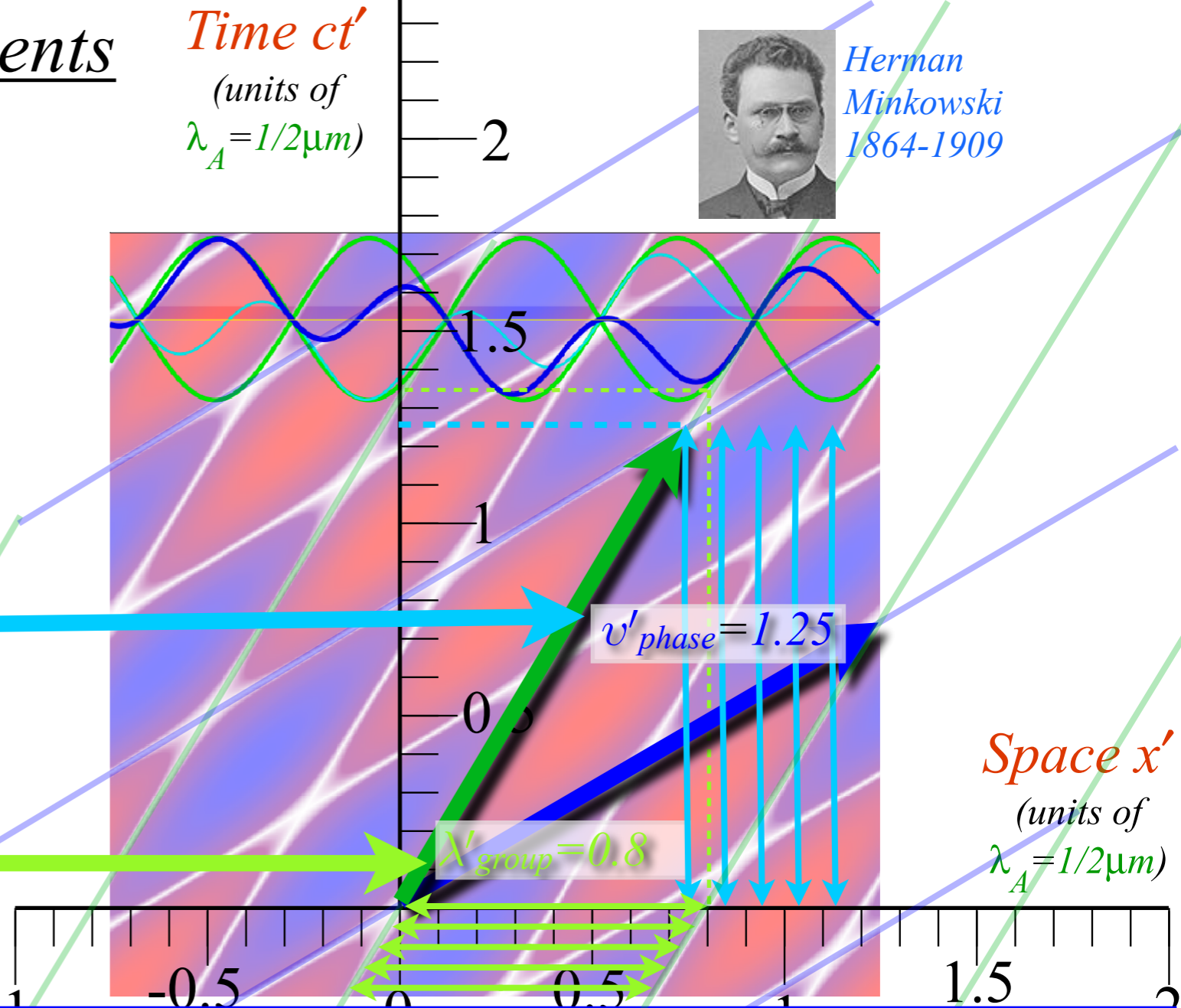
Albert Einstein  
1859-1955



Time  $ct'$   
(units of  $\lambda_A = 1/2\mu\text{m}$ )



Herman Minkowski  
1864-1909



This number is called an: **Einstein time-dilation**  
(dilated by 25% here)

This number is called a: **Lorentz length-contraction**  
(contracted by 20% here)



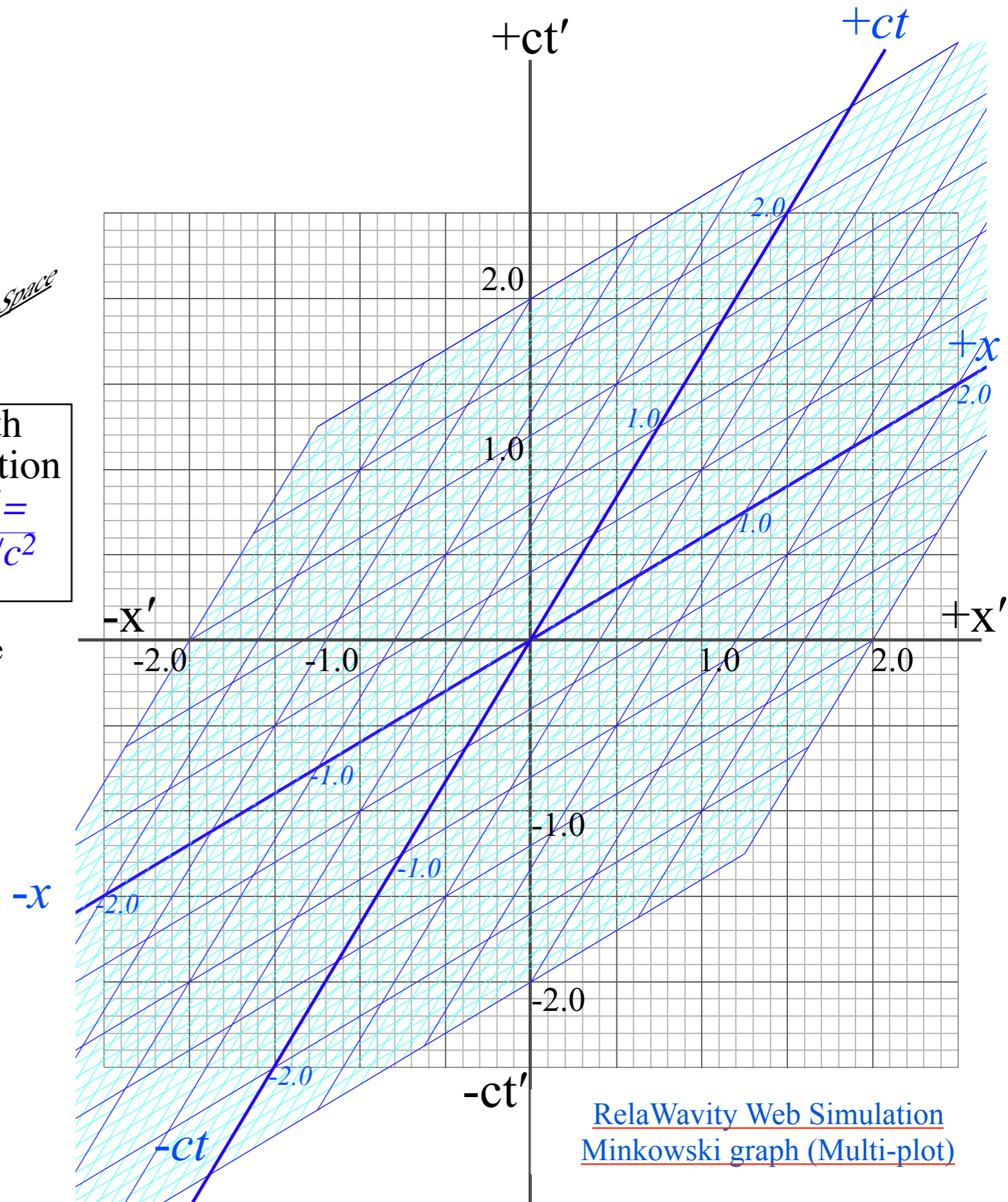
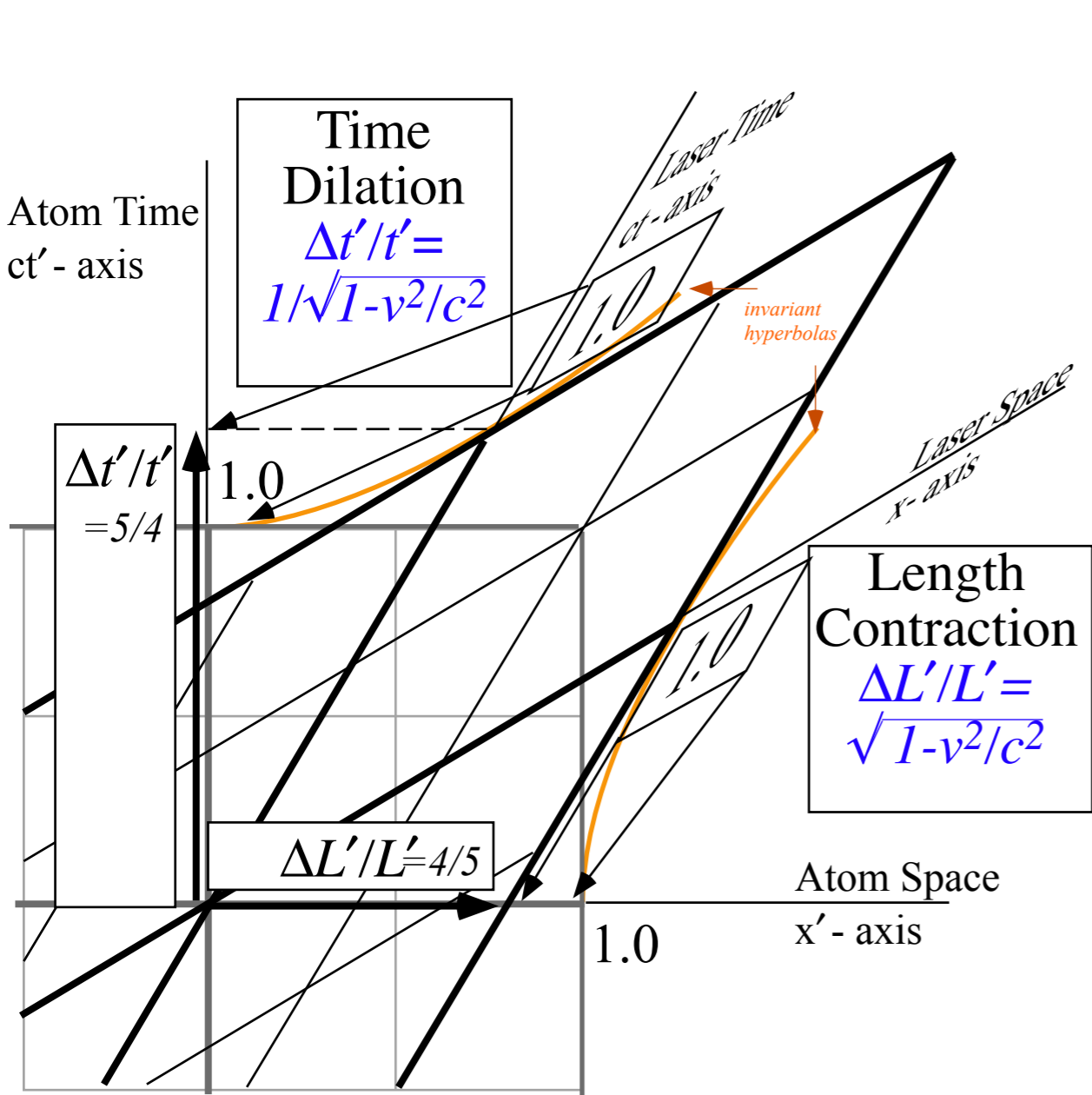
Hendrik A. Lorentz  
1853-1928

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

## Old-Fashioned Notation

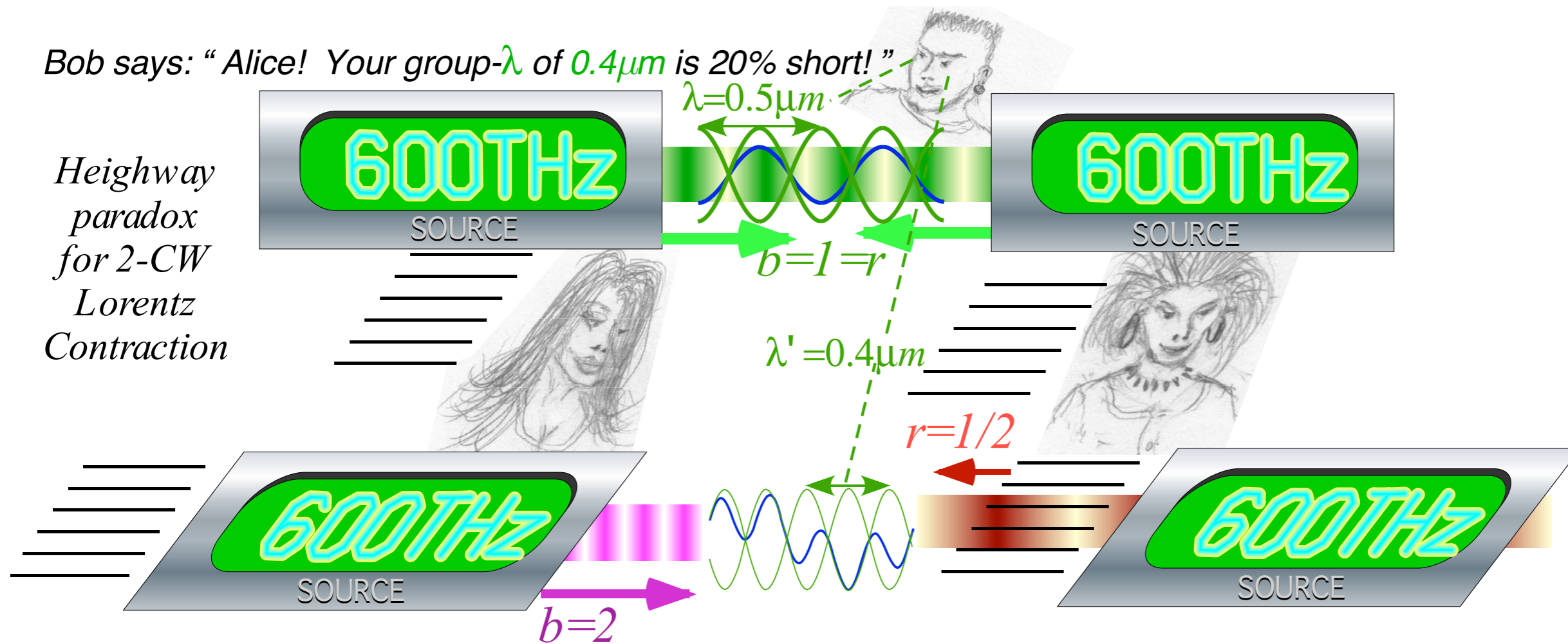
RelaWavity Web Simulation  
[Relativistic Terms \(Expanded Table\)](#)

# Reading Minkowski graph plots



Bob says: "Alice! Your group- $\lambda$  of  $0.4\mu\text{m}$  is 20% short!"

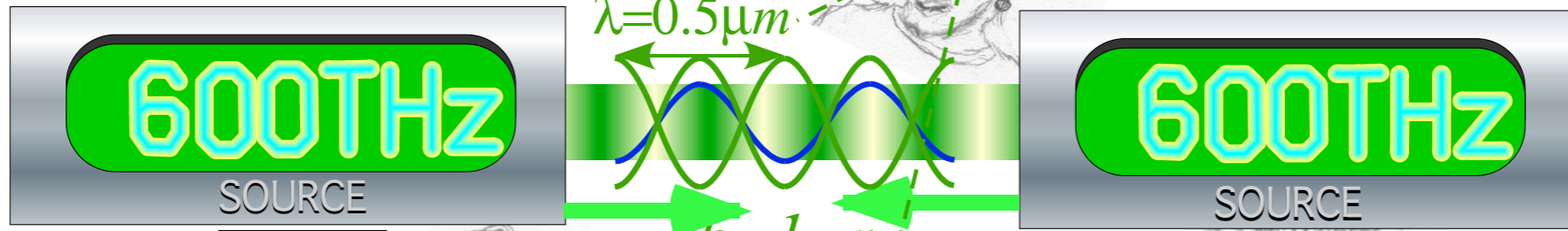
Heighway  
paradox  
for 2-CW  
Lorentz  
Contraction



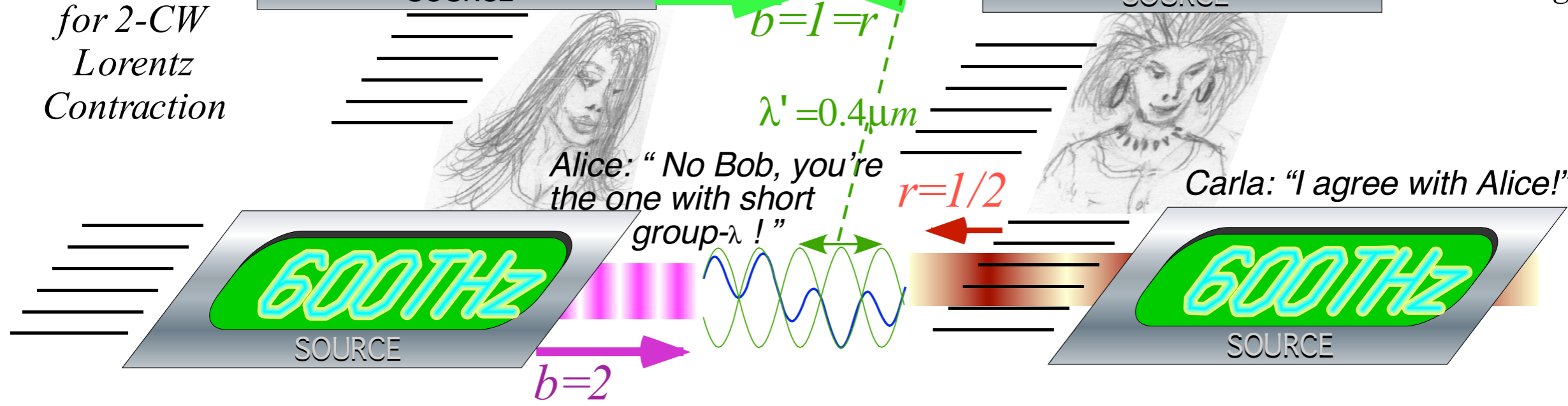


Bob says: "Alice! Your group- $\lambda$  of  $0.4\mu m$  is 20% short!"

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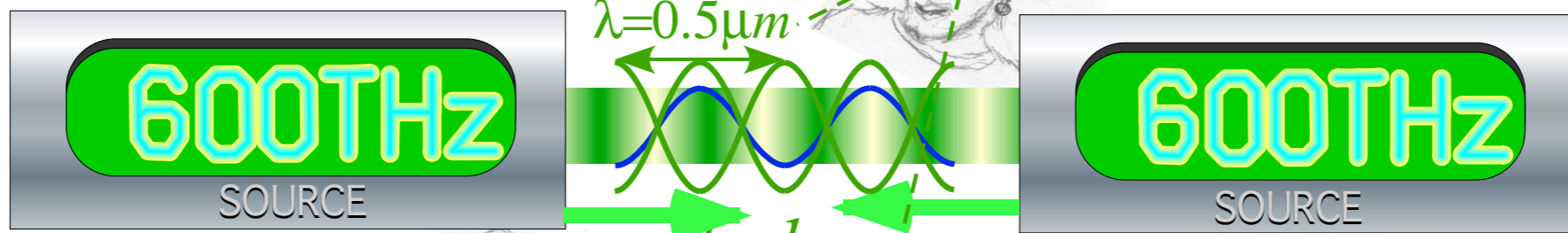
(Seems we have a most terrible lovers' quarrel...  
...both are *right*!)



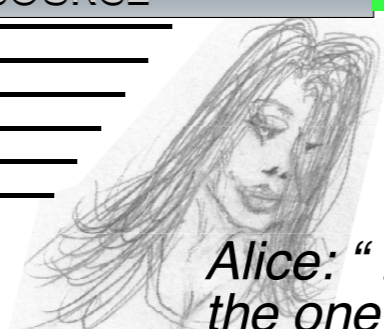
Carla: "I agree with Alice!"

Bob says: "Alice! Your group- $\lambda$  of  $0.4\mu\text{m}$  is 20% short!"

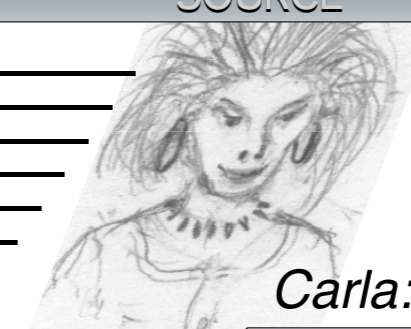
Heighway paradox for 2-CW Lorentz Contraction



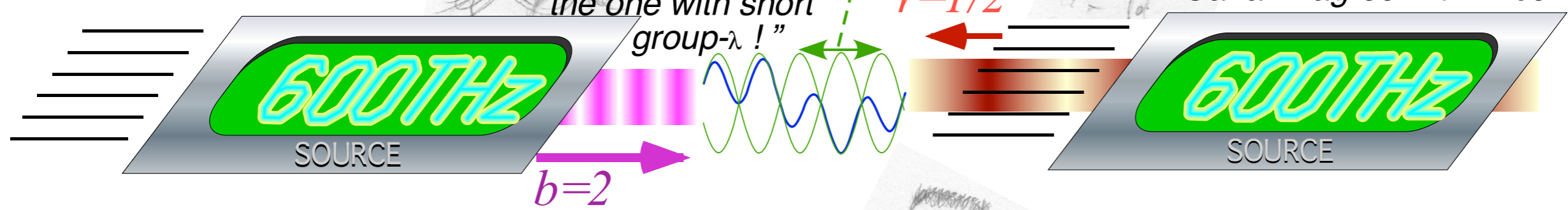
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Alice: "No Bob, you're the one with short group- $\lambda$ !"

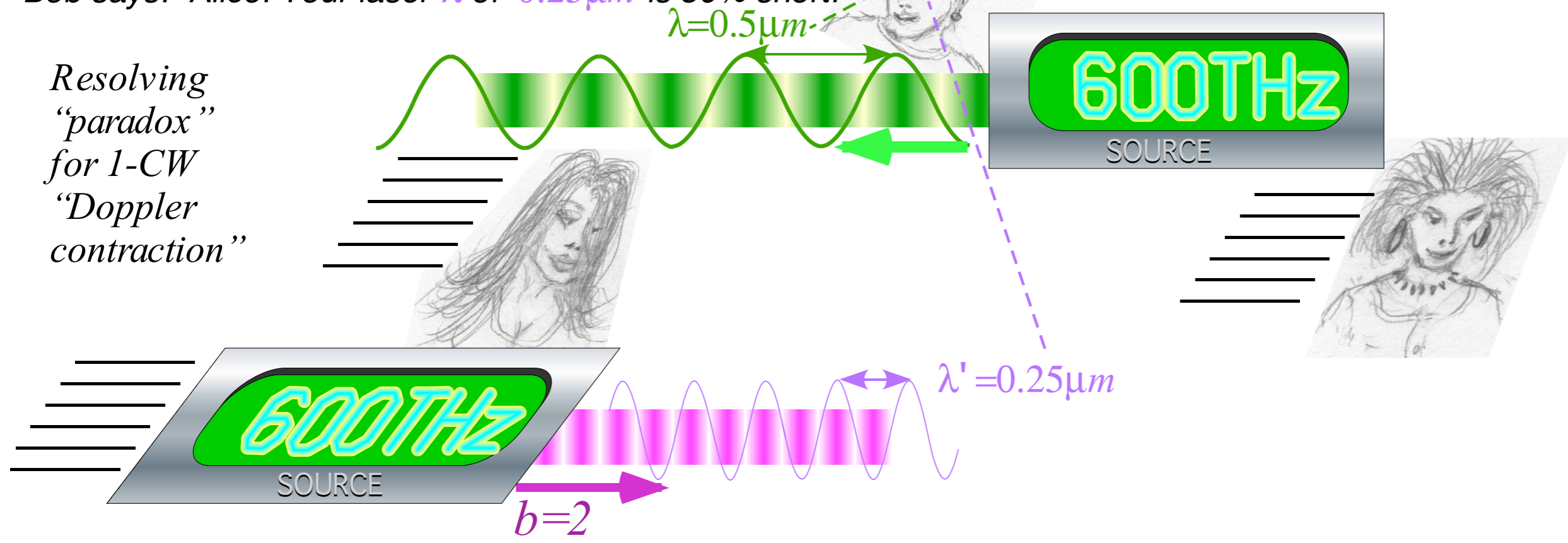


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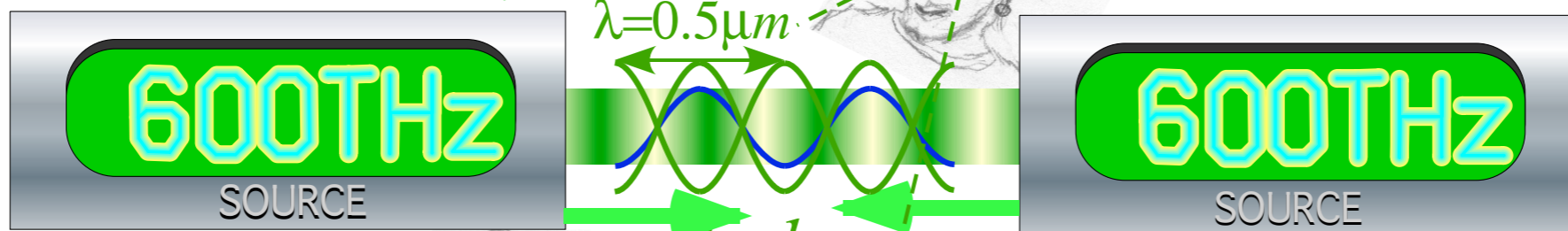
Bob says: "Alice! Your laser- $\lambda$  of  $0.25\mu\text{m}$  is 50% short!"

Resolving "paradox" for 1-CW "Doppler contraction"



Bob says: "Alice! Your group- $\lambda$  of  $0.4\mu\text{m}$  is 20% short!"

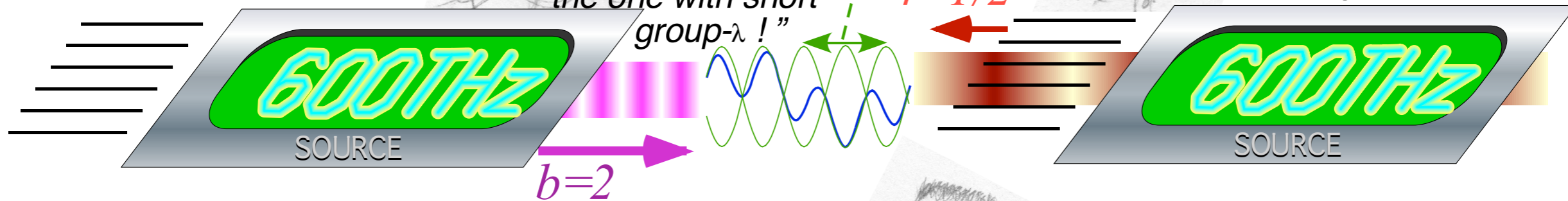
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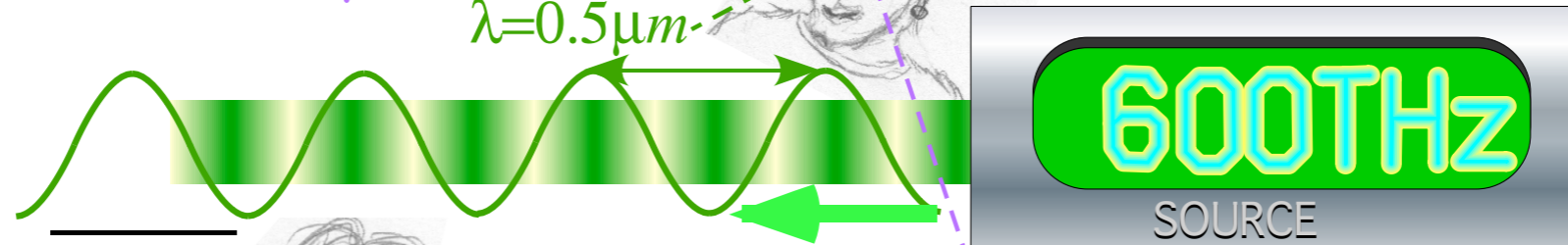
Alice: "No Bob, you're the one with short group- $\lambda$ !"

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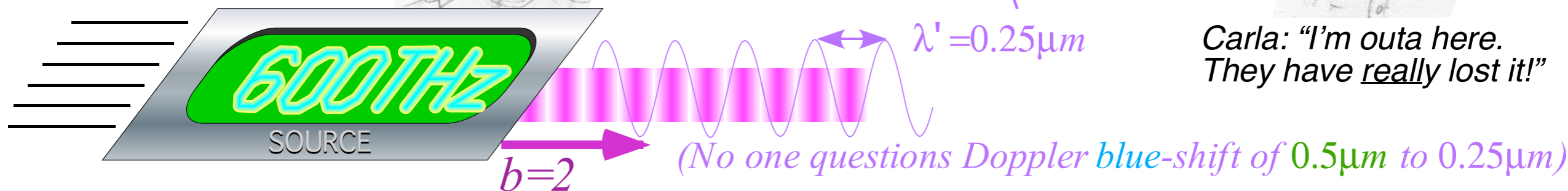
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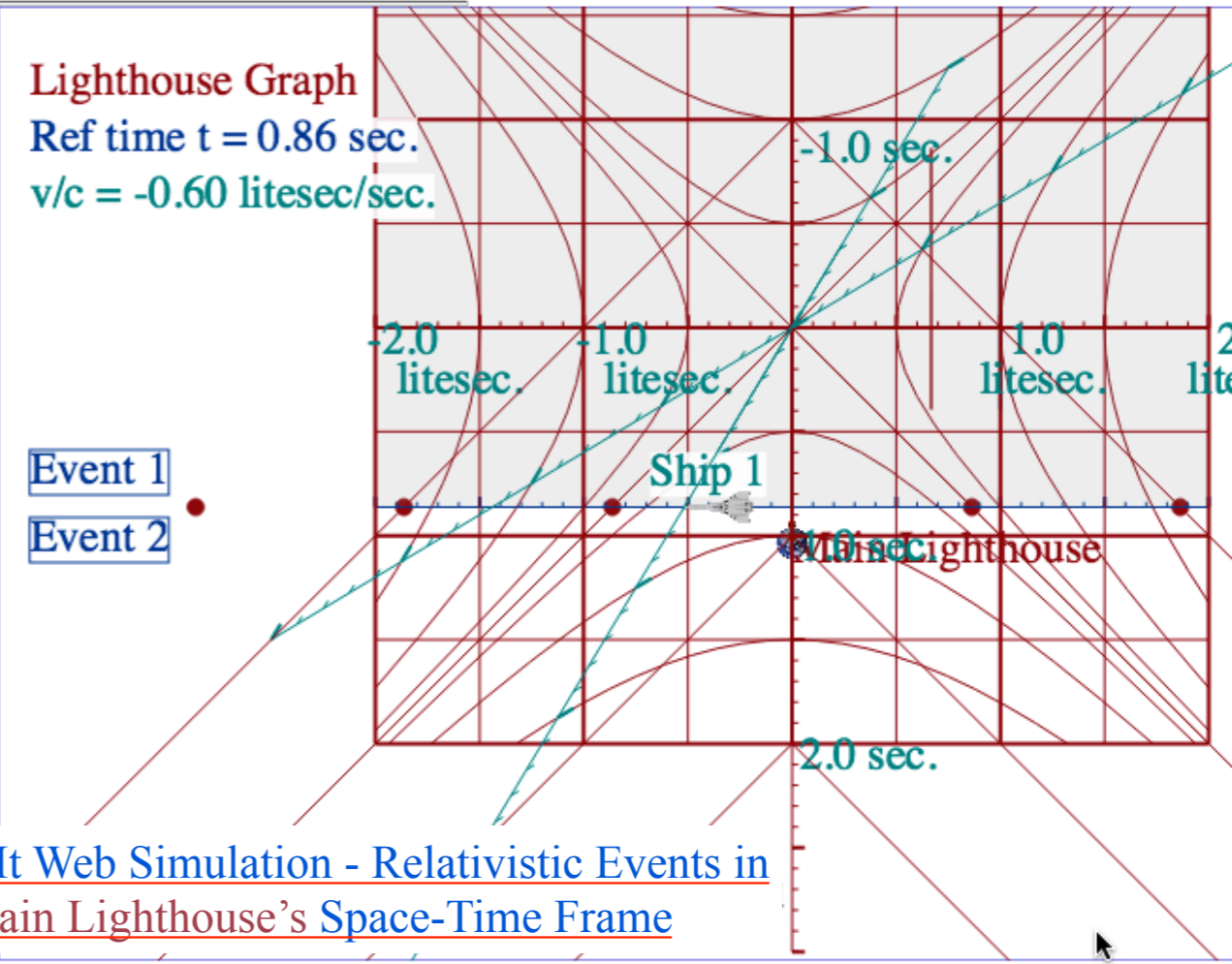
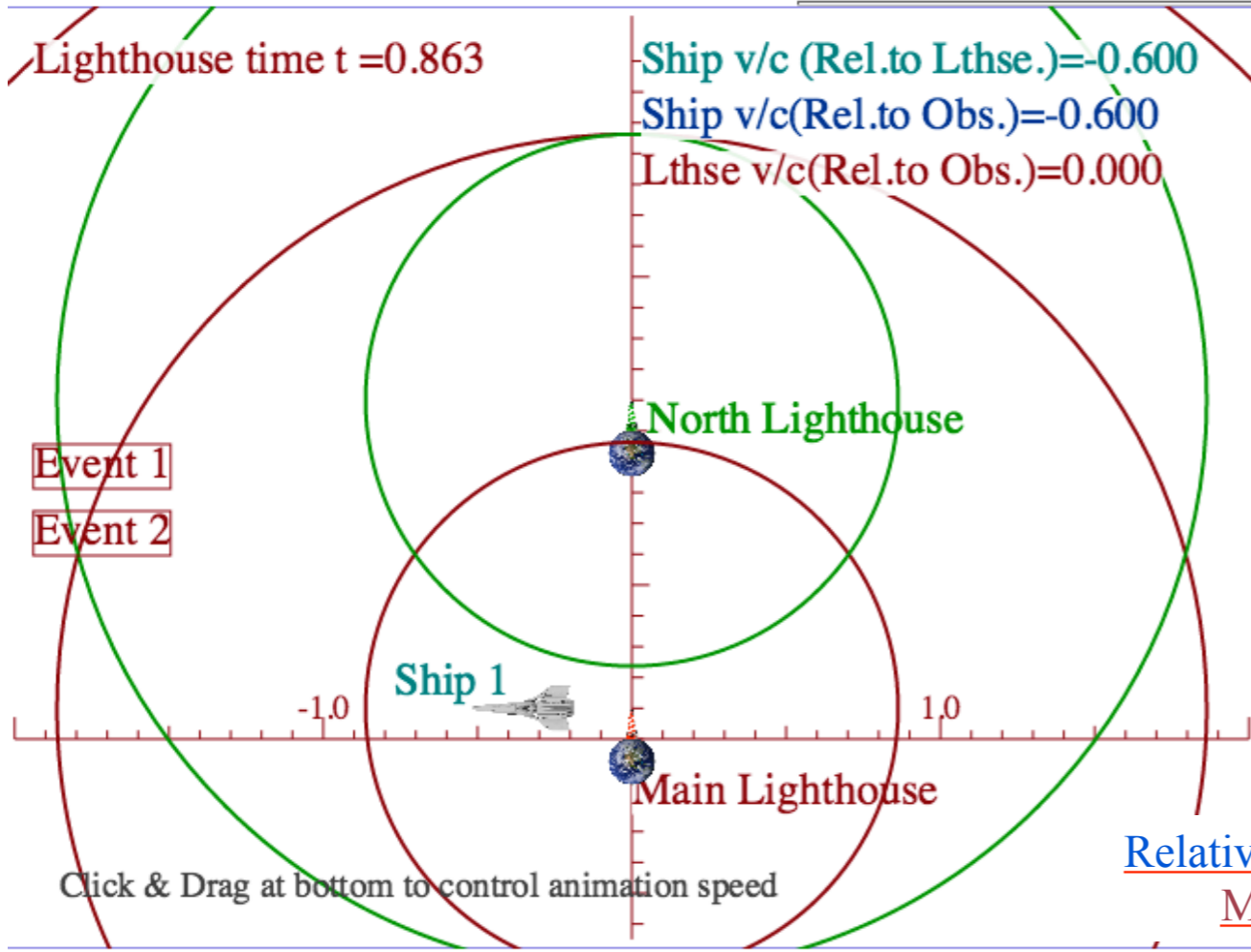
Alice: "No Bob, you're the one with short laser- $\lambda$ !"

Carla: "I'm outa here. They have really lost it!"

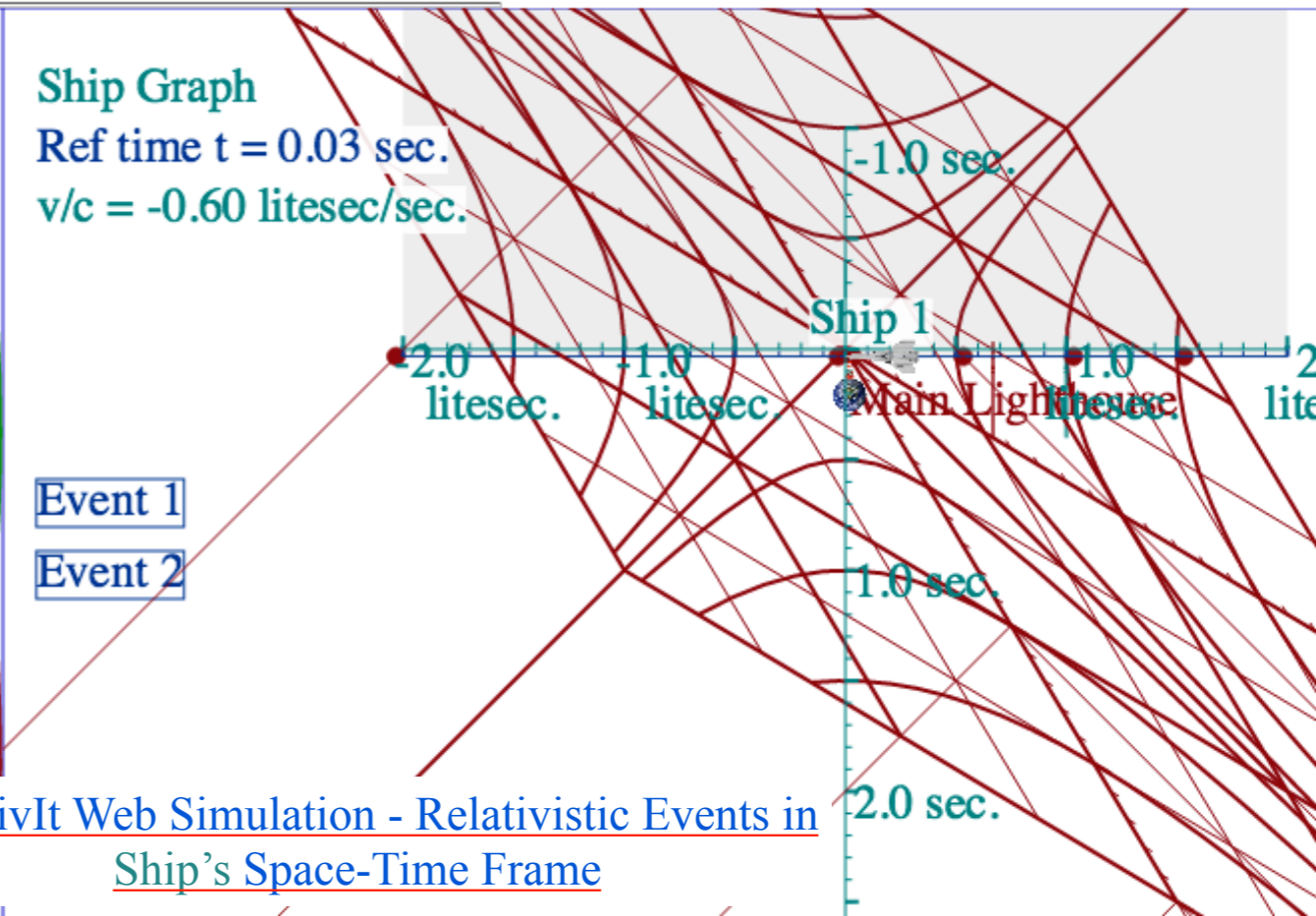
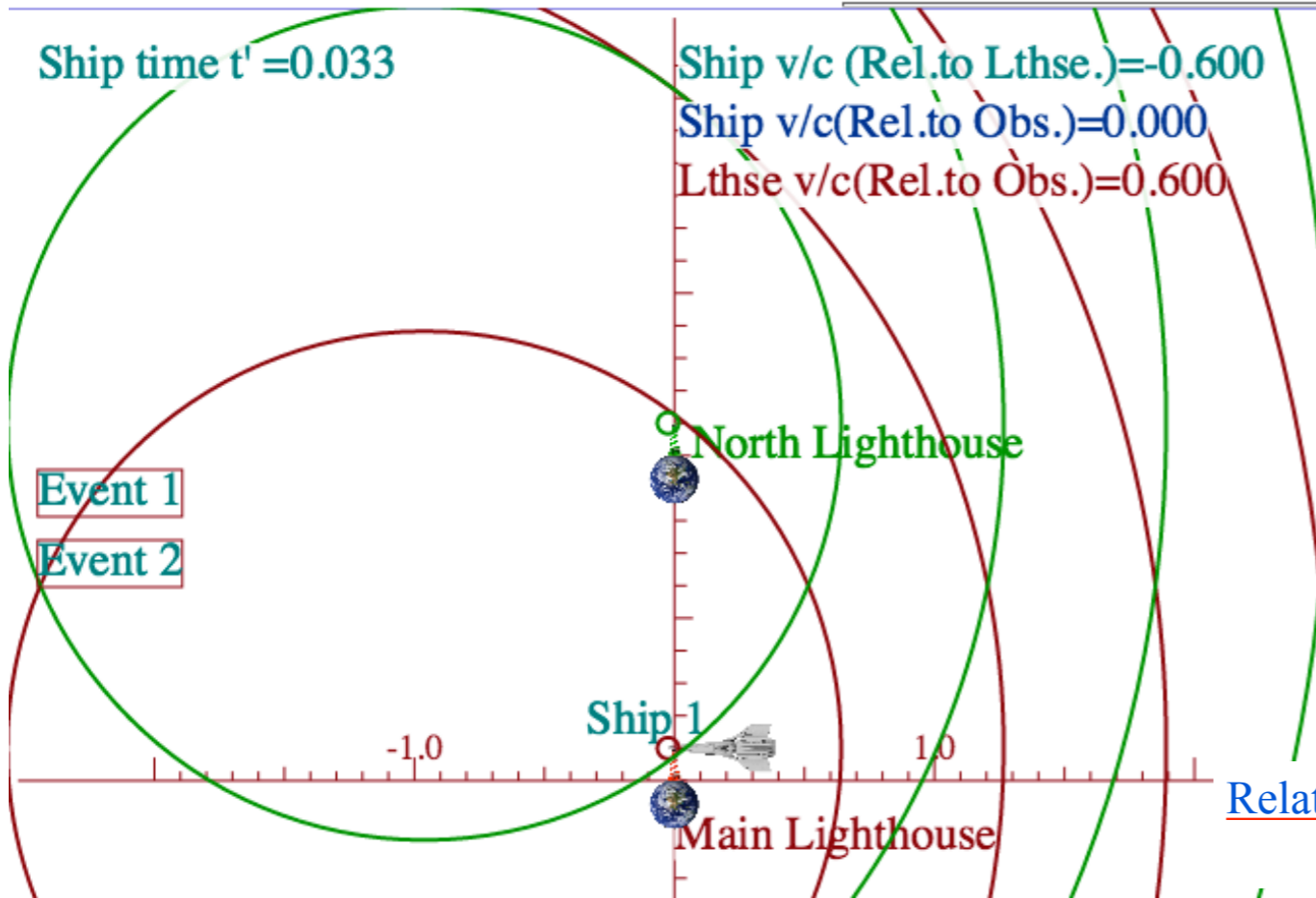




<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\nabla$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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<i>effects</i>	$b_{RED}^{Doppler}$	$V_{group}$	<i>past-future asymmetry</i> (off-diagonal Lorentz-transform)	<i>x-contraction</i> <sup>(Lorentz)</sup> $\tau_{phase}$ -contraction	<i>t-dilation</i> <sup>(Einstein)</sup> $v_{phase}$ -dilation (on-diagonal Lorentz-transform)	<i>inverse asymmetry</i>	$V_{phase}$	$b_{BLUE}^{Doppler}$

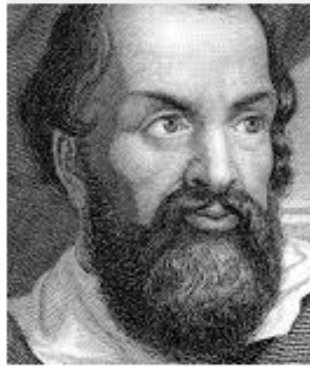


RelativIt Web Simulation - Relativistic Events in Main Lighthouse's Space-Time Frame



RelativIt Web Simulation - Relativistic Events in Ship's Space-Time Frame

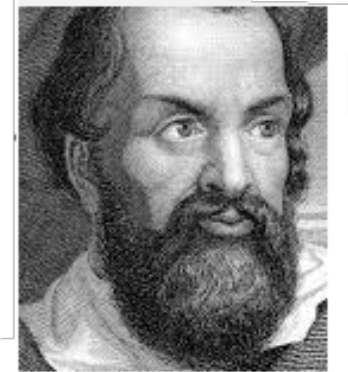
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1564-1642

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# Doppler Jeopardy

$$\nu_R = 600 \text{ THz}$$



$$\nu_L = 300 \text{ THz}$$

- (1.) To what velocity  $u_E$  must Bob accelerate so he sees beams with equal frequency  $\omega_E$ ?
- (2.) What is that frequency  $\omega_E$ ?

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Query (1.) has a Jeopardy-style answer-by-question: What is beam group velocity?

$$u_E = V_{group} = \frac{\omega_{group}}{k_{group}} = \frac{\omega_R - \omega_L}{k_R - k_L} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L}$$

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Query (2.) similarly: What  $\omega_E$  is blue-shift  $b\omega_L$  of  $\omega_L$  and red-shift  $\omega_R/b$  of  $\omega_R$ ?

$$\omega_E = b\omega_L = \omega_R/b \quad \Rightarrow \quad b = \sqrt{\omega_R / \omega_L} \quad \Rightarrow \quad \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

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$$\sqrt{6 \cdot 3} = 3\sqrt{2} = 4.24$$

$$\omega_E = b\omega_L = \omega_R/b \Rightarrow b = \sqrt{\omega_R / \omega_L} \Rightarrow \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

$$\begin{aligned} \omega_E &= \sqrt{\omega_R \cdot \omega_L} \\ &= \sqrt{180000} \\ &= 424 \end{aligned}$$

Geometric mean

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$V_{group}/c$  is ratio of difference mean  $\omega_{group} = \frac{\omega_R - \omega_L}{2}$  to arithmetic mean  $\omega_{phase} = \frac{\omega_R + \omega_L}{2}$ . Frequency  $\omega_E = B$  is the **geometric mean**  $\sqrt{\omega_R \cdot \omega_L}$  of left and right-moving frequencies defining the geometry

Geometric mean



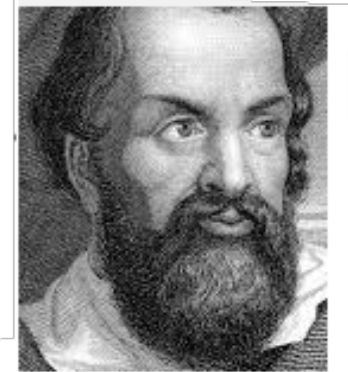
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1564-1642

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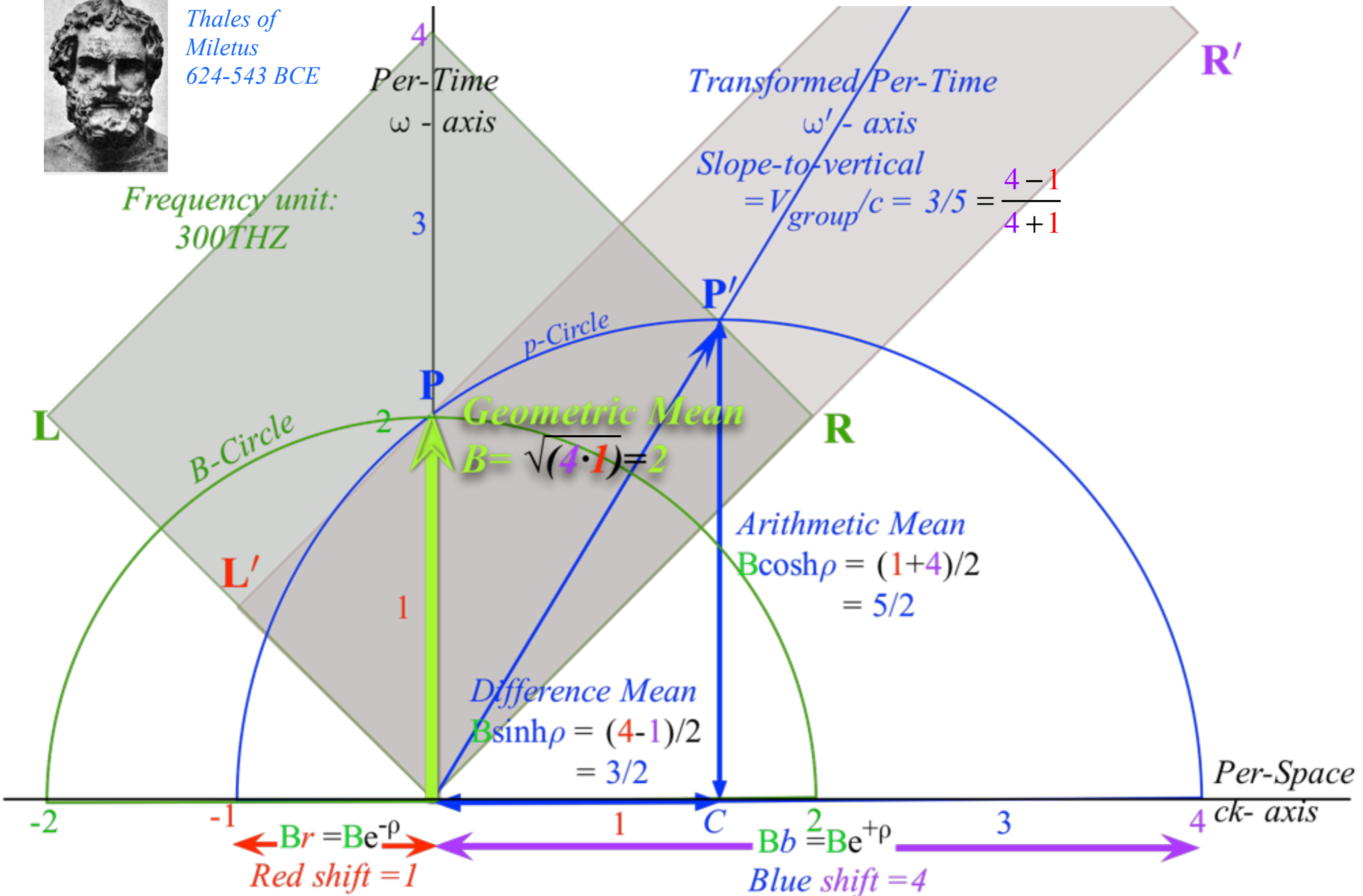
# Thales Mean Geometry (600BCE)

helps “Relativity”



Thales of Miletus  
624-543 BCE

Frequency unit:  
300THZ



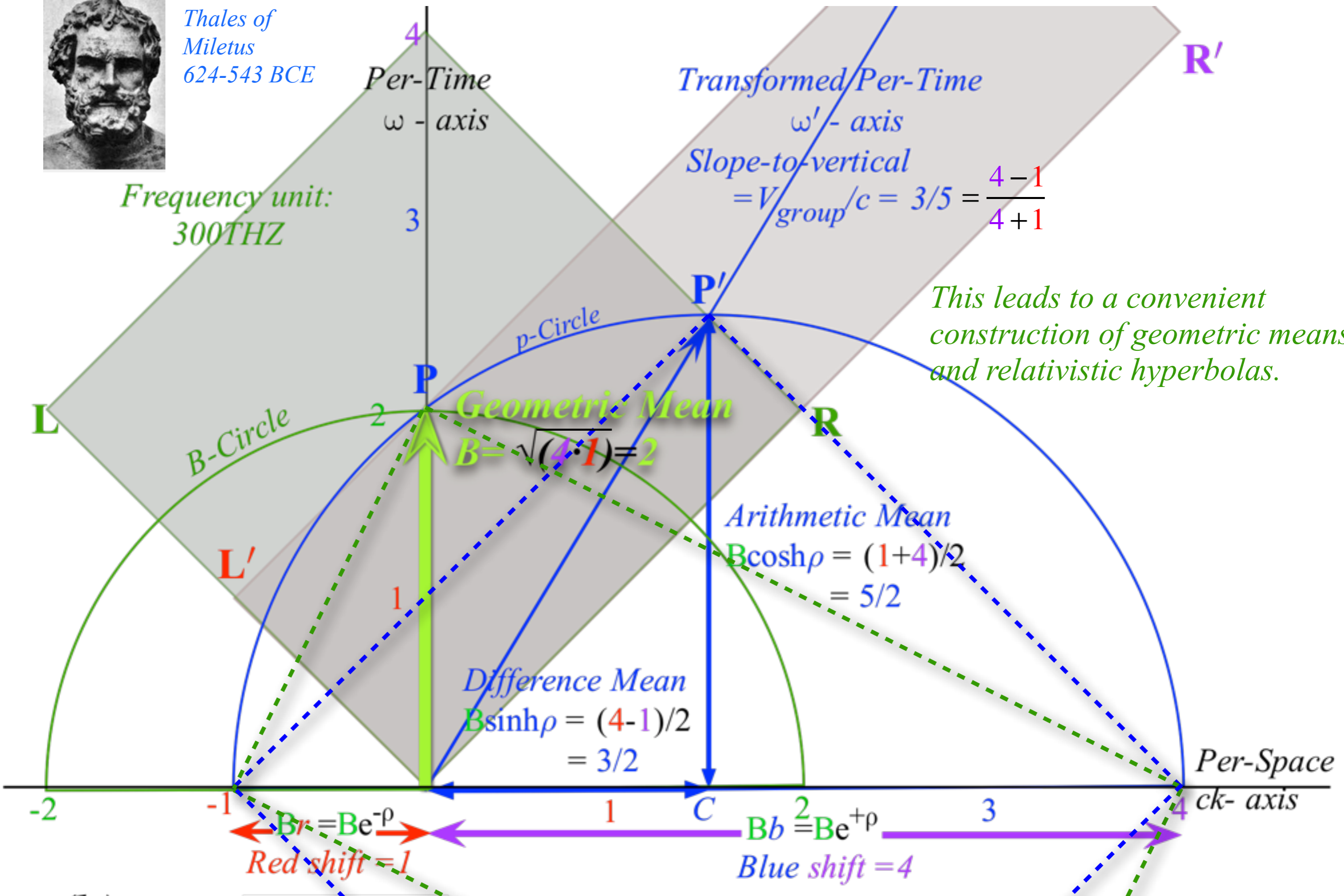
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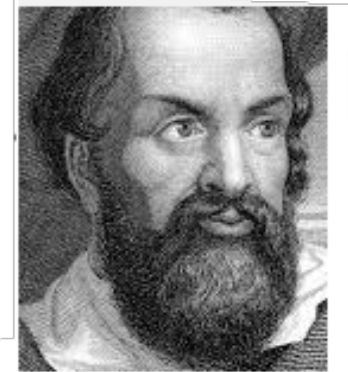
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Thales geometry of Lorentz transformation ➔...and invariant hyperbolas ←

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Frequency unit:  
300THZ

Per-Time  
 $\omega$  - axis

Transformed/Per-Time  
 $\omega'$  - axis

Slope-to-vertical

$$= V_{\text{group}}/c = 3/5 = \frac{4-1}{4+1}$$

equilateral hyperbola  
 $r \cdot b = 2$

R'

This leads to a convenient construction of geometric means and relativistic hyperbolas.

L

B-Circle

Geometric Mean

$$B = \sqrt{(4 \cdot 1)} = 2$$

R

Arithmetic Mean

$$B \cosh \rho = (1+4)/2 = 5/2$$

Difference Mean

$$B \sinh \rho = (4-1)/2 = 3/2$$

Per-Space  
ck- axis

-2

-1

1

2

3

4

$$Br = Be^{-\rho}$$

Red shift = 1

$$Bb = Be^{+\rho}$$

Blue shift = 4

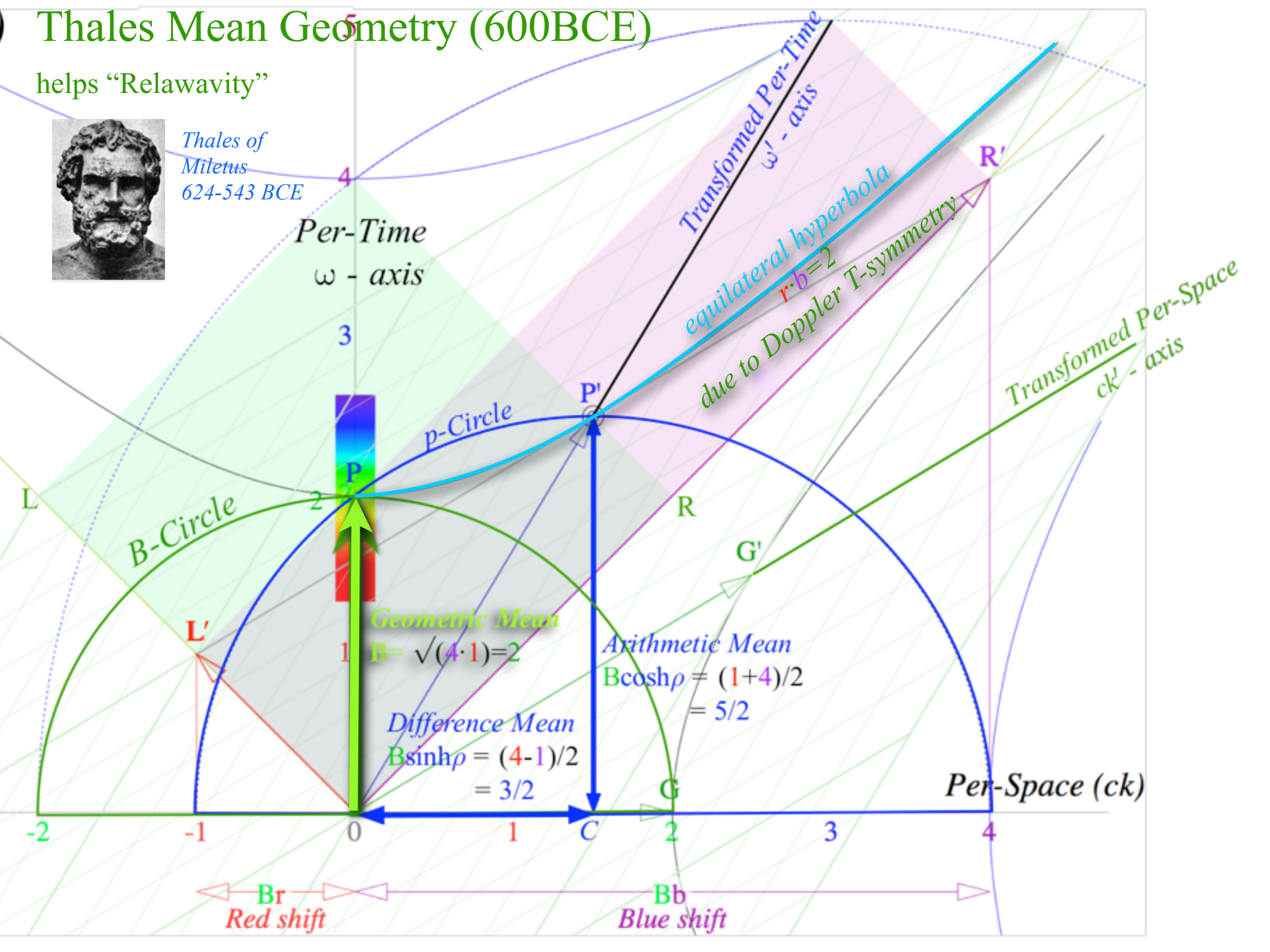


# Thales Mean Geometry (600BCE)

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Per-Time ( $\omega$ )

Acoustical base frequency =  $B = 600\text{Hz}$

Hi freq = 1200.000

Lo freq = 300.000

Laser base frequency =  $B = 600\text{THz}$

Doppler blue shift factor =  $b = 2.000$

Doppler red shift factor =  $r = 0.500$

$q = 0.693$

CW Light Axioms

All colors go c:  $\omega/k = c$  or L&R on diagonals

Time Reversal ( $r \leftrightarrow b$ ):  $r = 1/b$

$$G' = G \cosh(q) + P \sinh(q)$$

$$P' = G \sinh(q) + P \cosh(q)$$

$$G = G' \cosh(q) - P' \sinh(q)$$

$$P = -G' \sinh(q) + P' \cosh(q)$$

H. sapiens Visual Best = 600THz

600Hz = Auditory Base

Visual Min = 400THz

20Hz = Auditory Min

-2.0 -1.5 -1.0 -0.5 0.0 0.5 1.0 1.5 2.0 Per-Space ( $ck$ )

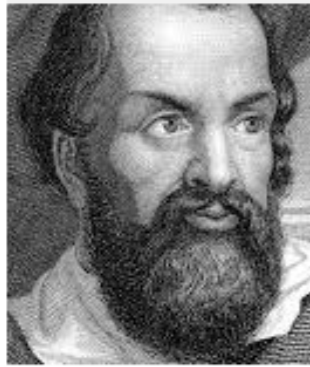
[RelaWavity Web Simulation](#)  
[Detailed Thales Geometry](#)

$B e^{-q}$   
Red shift

$B e^{+q}$   
Blue shift

Select from the top menus to choose the view type and sub-type.  
Click the 'Controls' button to set shared model & display vars.  
Set parameters with click (& drag) near the ck axis: r,b; the green semi-circle:  $\sigma$ ; the hyperbolae:  $v$   
Right (or CTRL+) click figure to set plot specific vars.

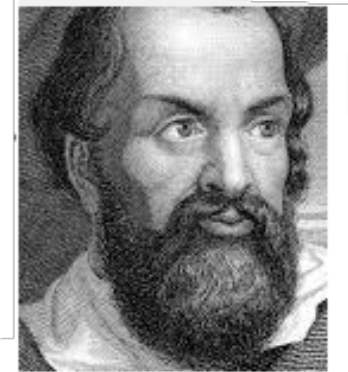
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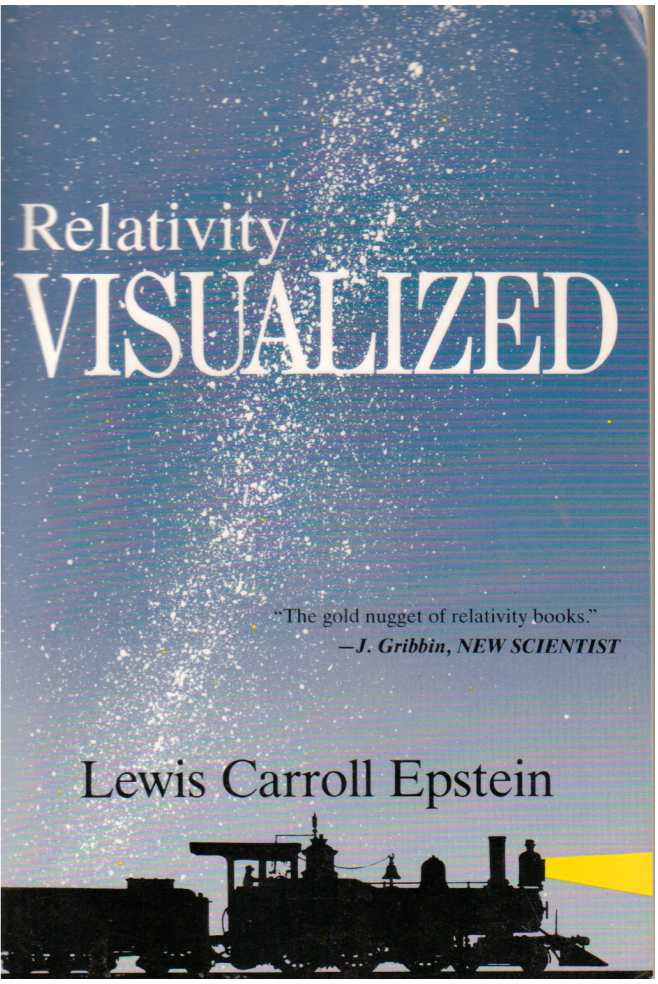
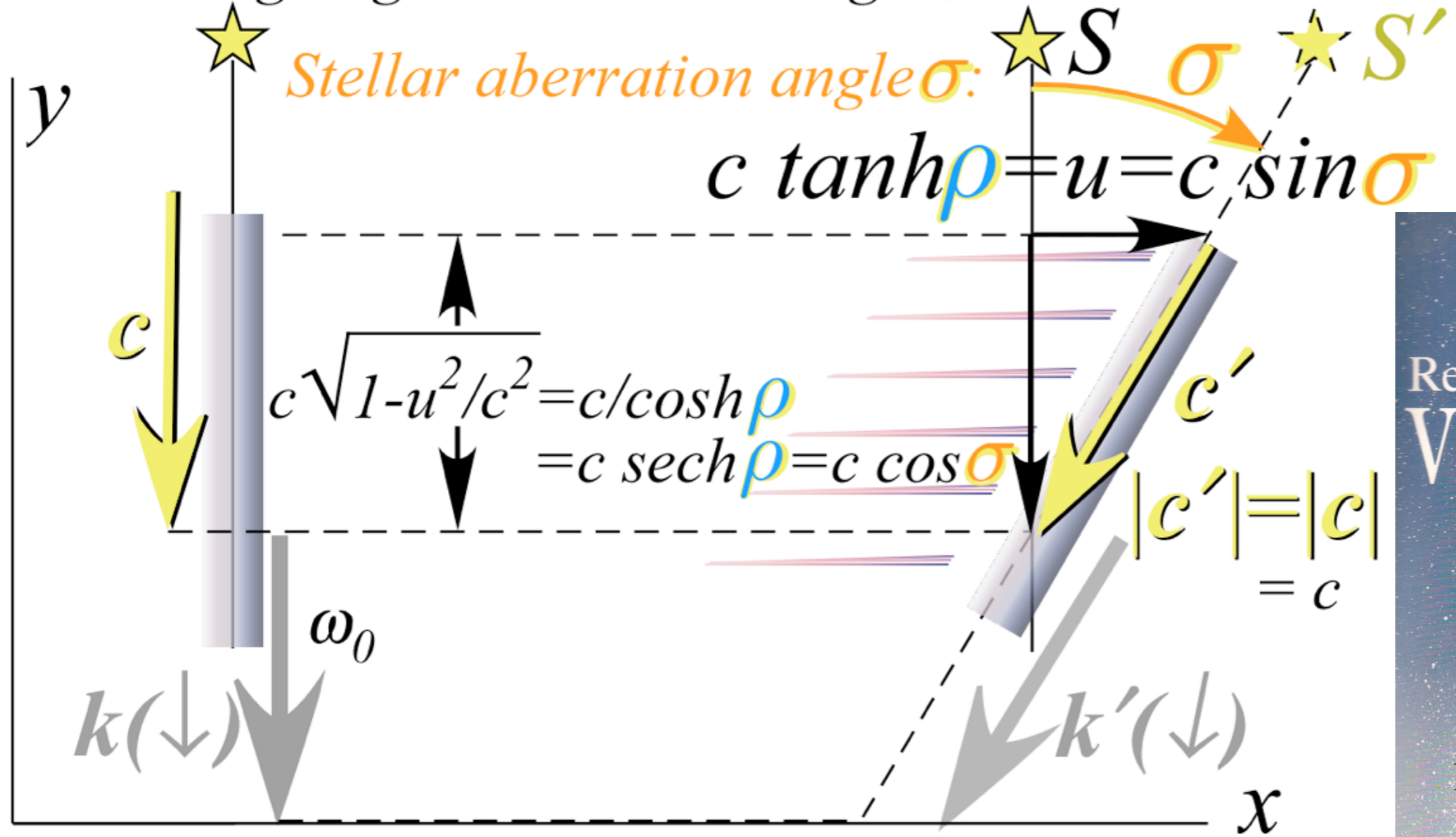
# Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

to a Transverse\*relativity parameter: Stellar aberration angle  $\sigma$

\*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.  
 Observer going  $u$  sees star at angle  $\sigma$  in  $u$  direction.

We used notation  $\sigma$  for stellar-ab-angle, (a “flipped-out”  $\rho$ ). Epstein not interested in  $\rho$  analysis or in relation of  $\sigma$  and  $\rho$ .



Purchase at:



# Details of stellar aberration angle $\sigma$ of K-vexctor rotation

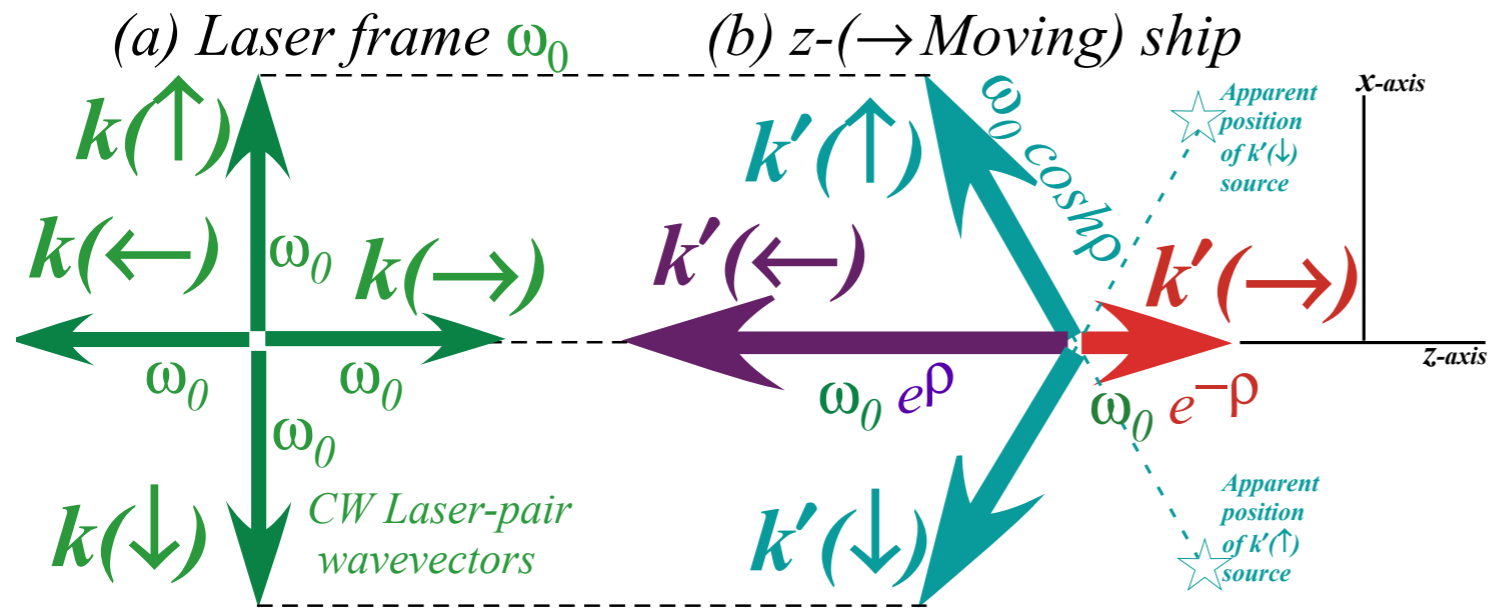
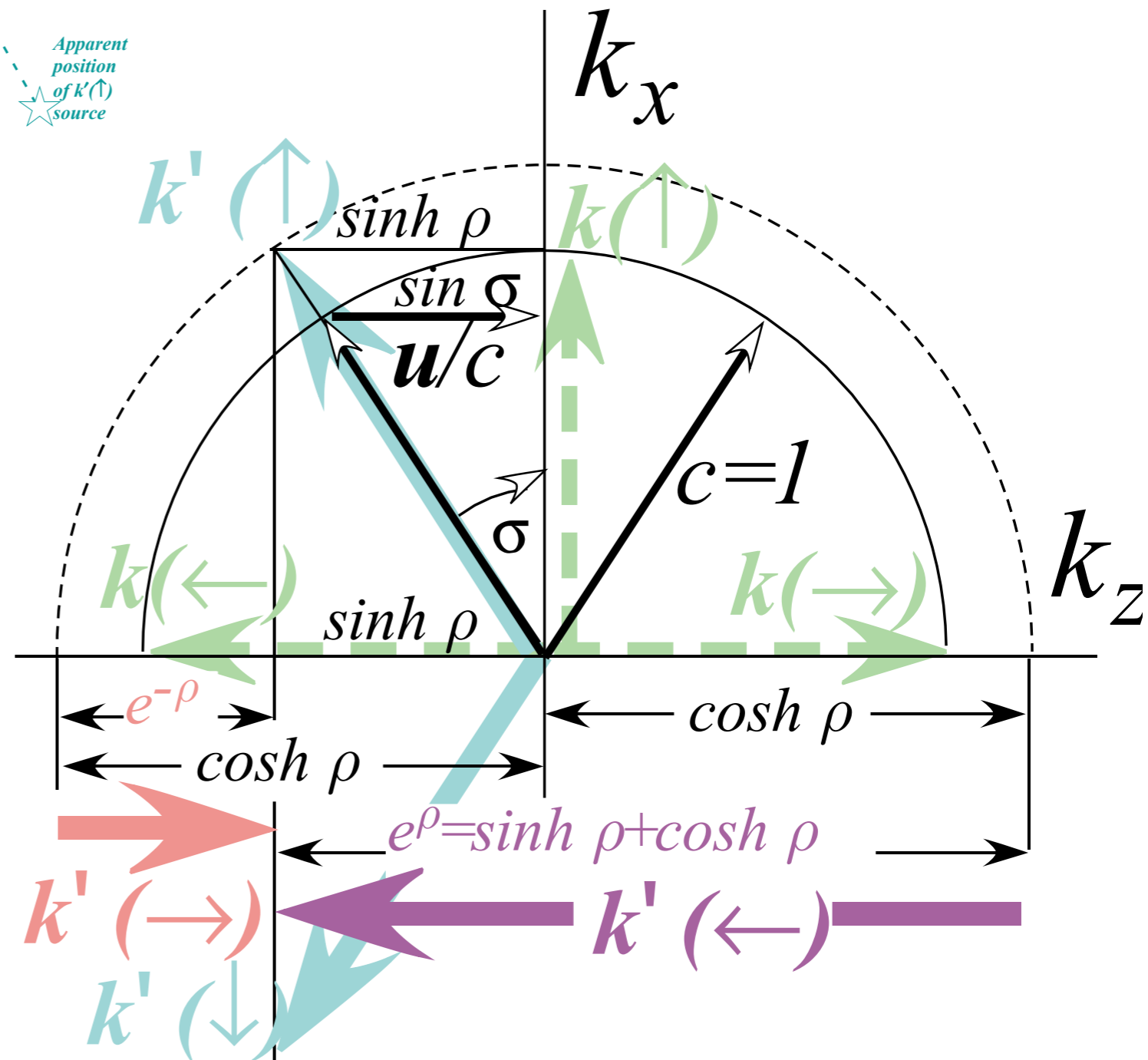


Fig. 8.5.7  
Unit 8

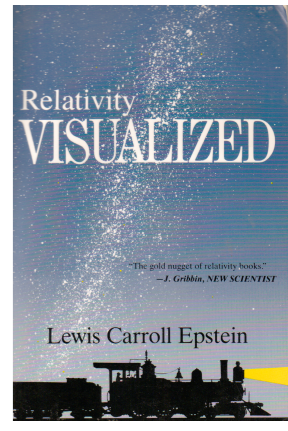
Fig. 8.5.10  
Unit 8



Comparing Longitudinal relativity parameter: Rapidity  $\rho = \log_e(\text{Doppler Shift})$

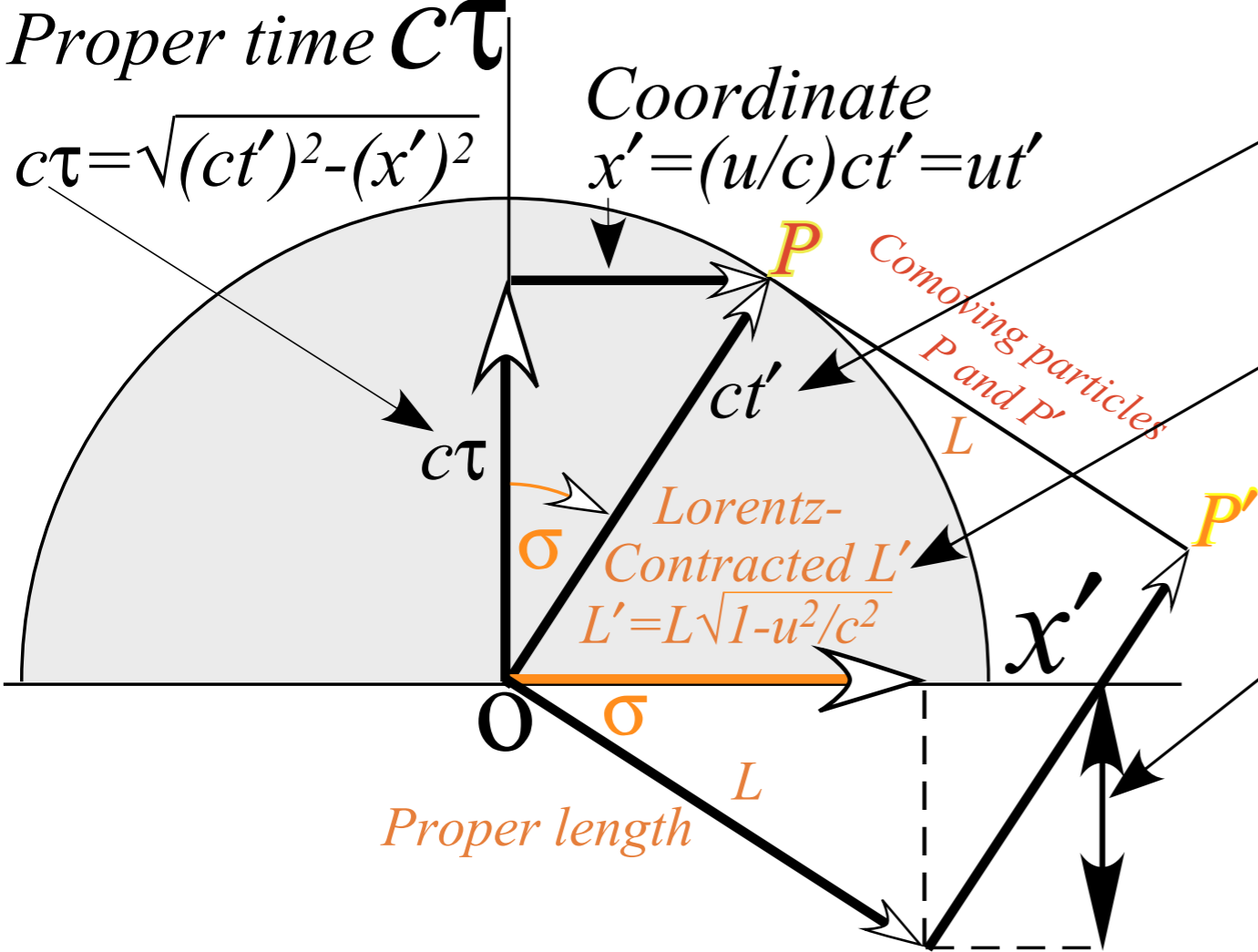
to a Transverse\*relativity parameter: Stellar aberration angle  $\sigma$

\*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-



Proper time  $c\tau$  vs. coordinate space  $x$  - (L. C. Epstein's "Cosmic Speedometer")

Particles  $P$  and  $P'$  have speed  $u$  in  $(x', ct')$  and speed  $c$  in  $(x, c\tau)$



Einstein time dilation:  
 $ct' = c\tau \sec\sigma = c\tau \cosh\rho = c\tau / \sqrt{1-u^2/c^2}$

Lorentz length contraction:  
 $L' = L \operatorname{sech}\rho = L \cos\sigma = L \cdot \sqrt{1-u^2/c^2}$

Proper Time asimultaneity:  
 $c \Delta\tau = L' \sinh\rho = L \cos\sigma \sinh\rho$   
 $= L \cos\sigma \tan\sigma$   
 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$





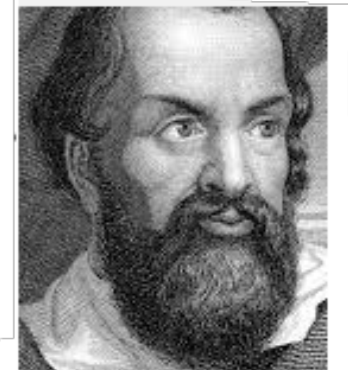
Galileo Galilei



1564-1642

**Galileo's Revenge (part 1)**

*Rapidity adds just like  
Galilean velocity*



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➔ “Occams Sword” and geometry of 16 parameter functions of  $\rho$  and  $\sigma$

Application to TE-Waveguide modes and synchrotron beam relativity

This map has circle sector arc-area  $\sigma = 0.6435$

set to angle  $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

$$\begin{aligned} \sin(\sigma) &= 0.6000 &= \tanh(\rho) &= 3/5 \\ \tan(\sigma) &= 0.7500 &= \sinh(\rho) &= 3/4 \\ \sec(\sigma) &= 1.2500 &= \cosh(\rho) &= 5/4 \\ \cos(\sigma) &= 0.8000 &= \operatorname{sech}(\rho) &= 4/5 \\ \cot(\sigma) &= 1.3333 &= \operatorname{csch}(\rho) &= 4/3 \\ \csc(\sigma) &= 1.6667 &= \operatorname{coth}(\rho) &= 5/3 \end{aligned}$$

$$\cosh(\rho) + \sinh(\rho) = \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho}$$

$$\cosh(\rho) - \sinh(\rho) = \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho}$$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2} \quad \text{Half-Sum-}$$

Half-Difference

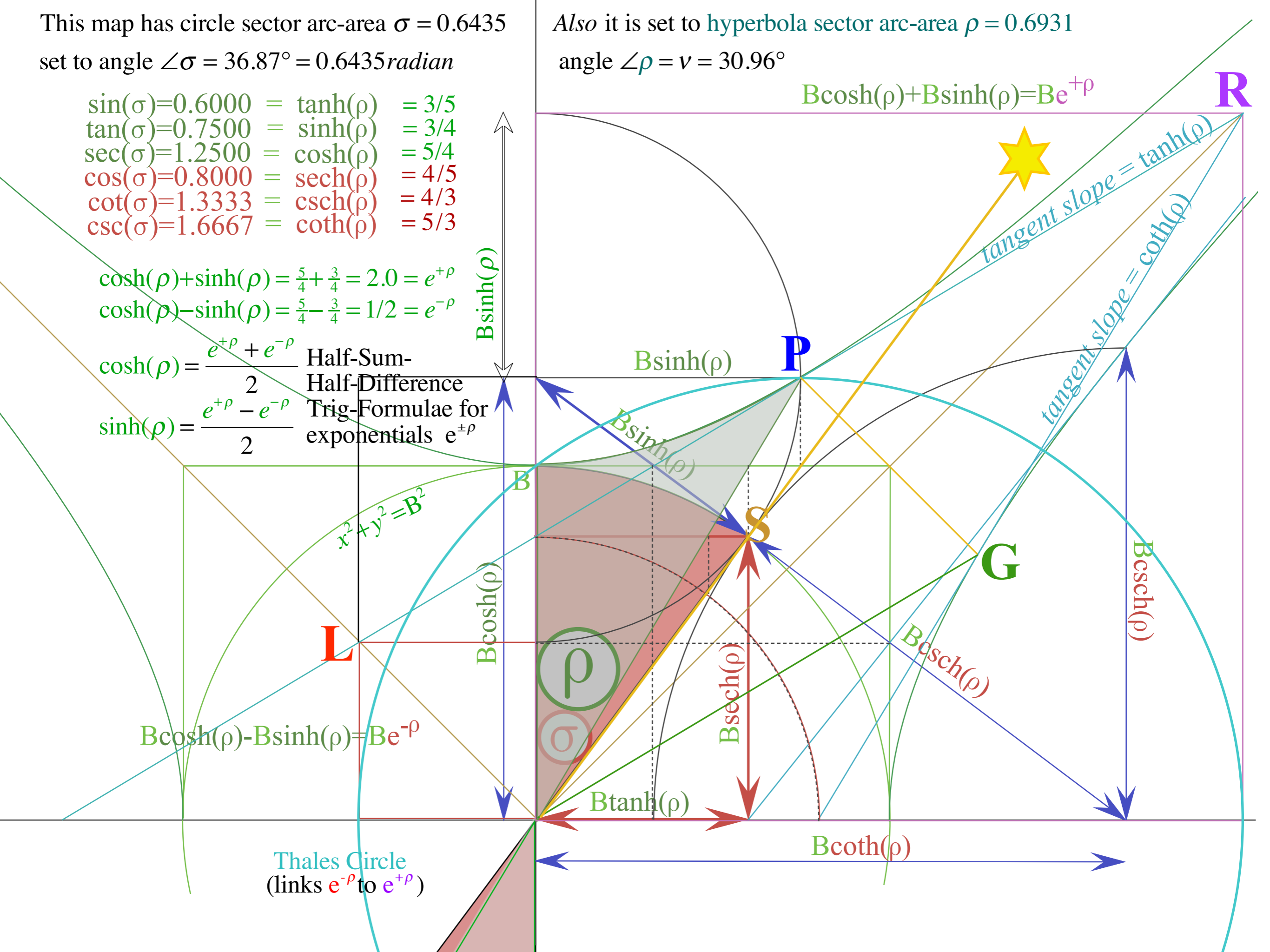
$$\sinh(\rho) = \frac{e^{+\rho} - e^{-\rho}}{2} \quad \text{Trig-Formulae for}$$

exponentials  $e^{\pm\rho}$

Also it is set to hyperbola sector arc-area  $\rho = 0.6931$

angle  $\angle\rho = \nu = 30.96^\circ$

$$B\cosh(\rho) + B\sinh(\rho) = B e^{+\rho}$$



$$B\cosh(\rho) - B\sinh(\rho) = B e^{-\rho}$$

Thales Circle  
(links  $e^{-\rho}$  to  $e^{+\rho}$ )

$\rho$

R

P

S

G

L

$B\sinh(\rho)$

$B\sinh(\rho)$

$B\sinh(\rho)$

$B\cosh(\rho)$

$B\operatorname{sech}(\rho)$

$B\tanh(\rho)$

$B\operatorname{coth}(\rho)$

$B\operatorname{sch}(\rho)$

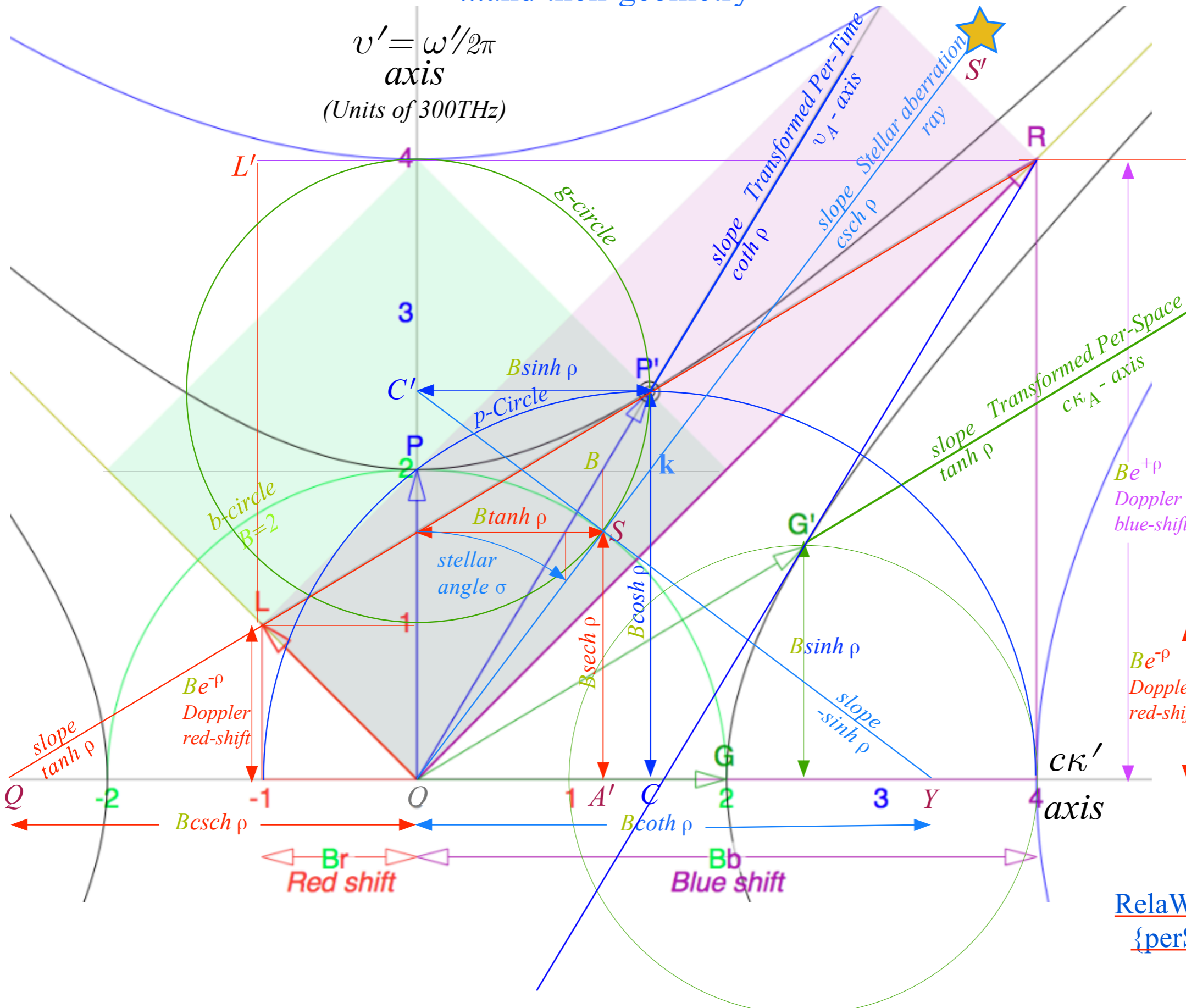
$B\operatorname{sch}(\rho)$

tangent slope =  $\tanh(\rho)$

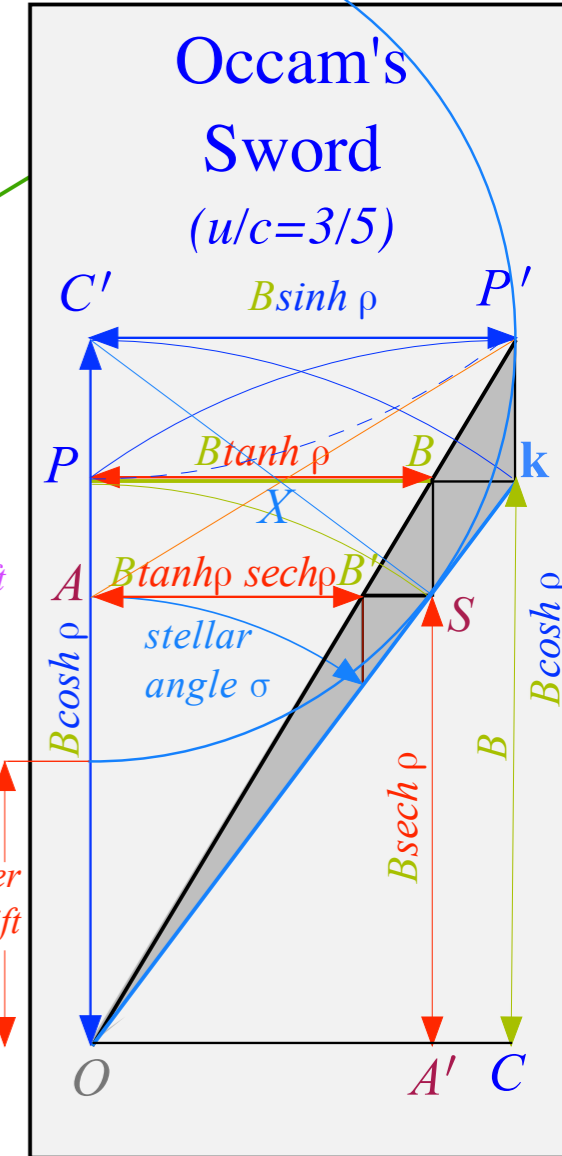
tangent slope =  $\operatorname{coth}(\rho)$

# Summary of optical wave parameters for relativity and QM

...and their geometry



An aid to pattern recognition:

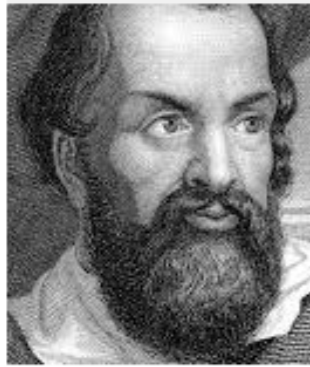


[RelaWavity Web Simulation](#)  
 {perSpace - perTime All}





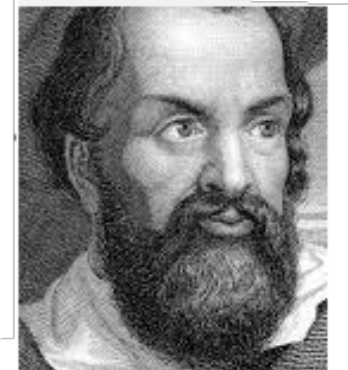
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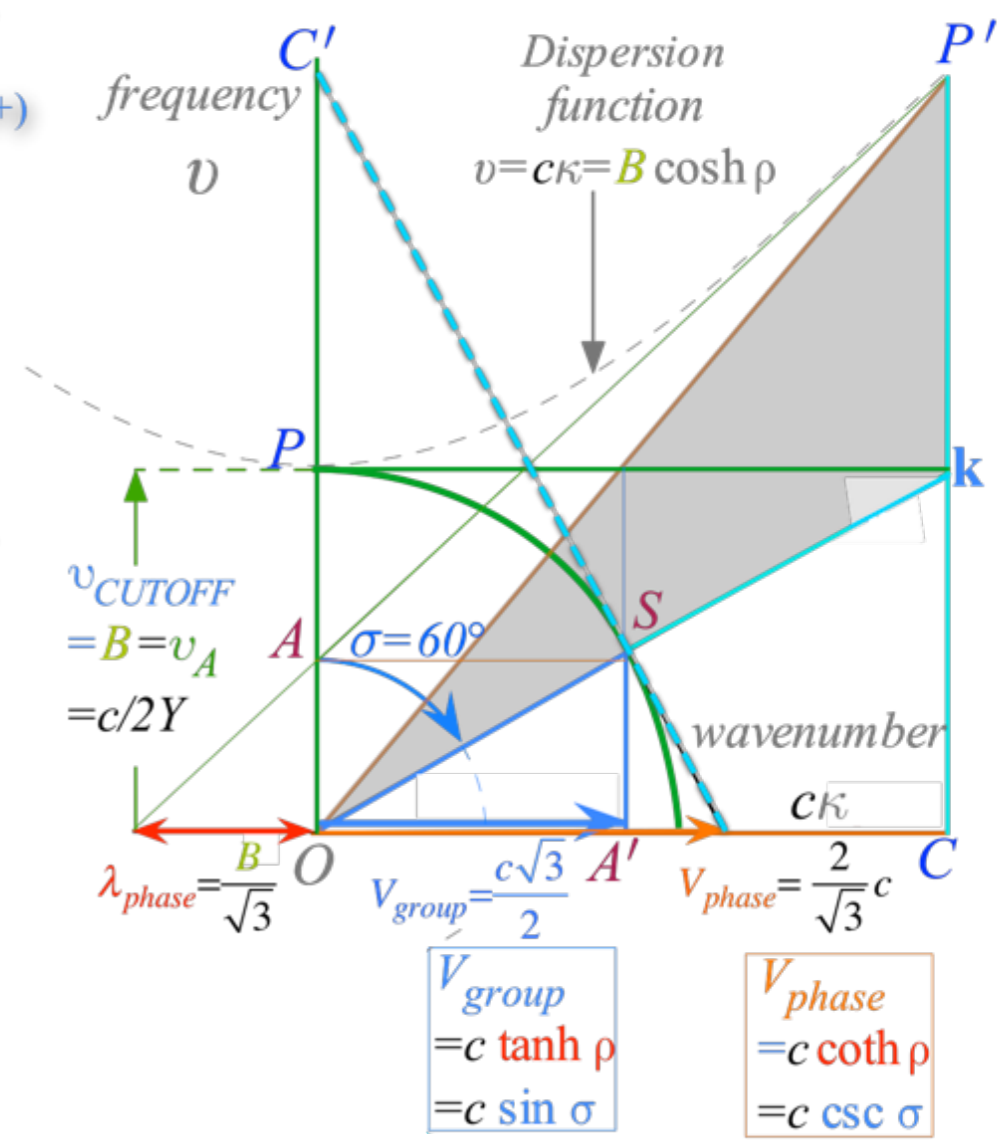
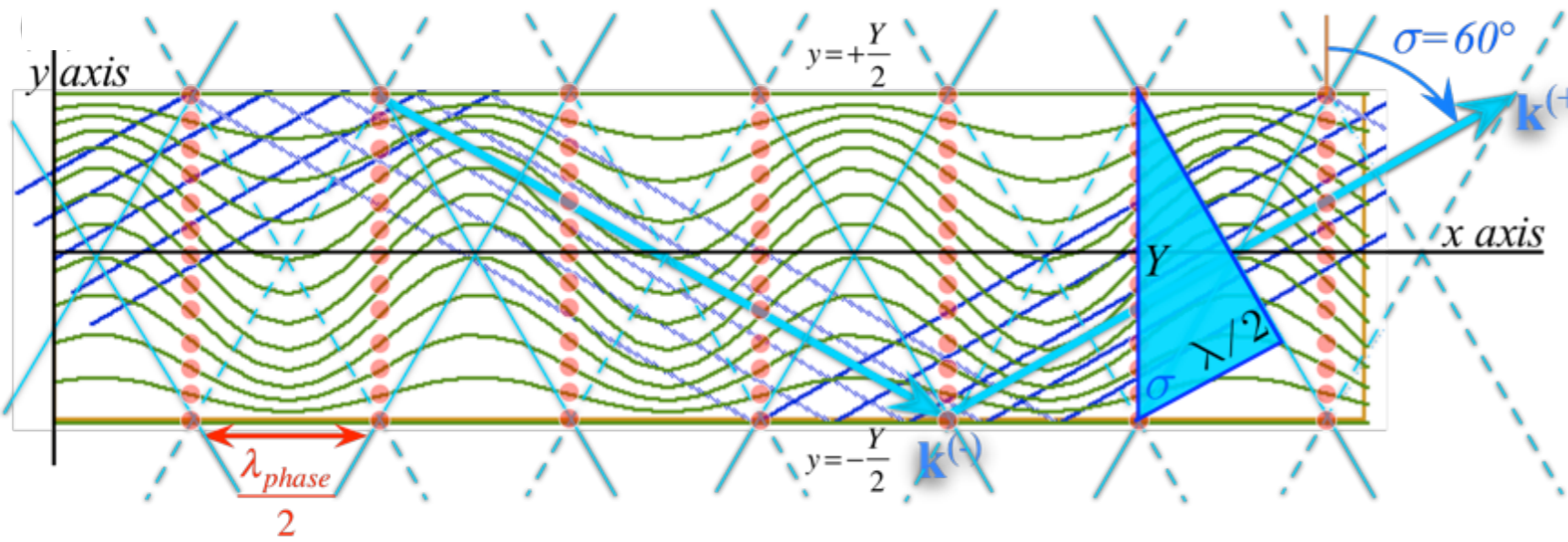
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➔ Application to TE-Waveguide modes and synchrotron beam relativity ←

# Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to  $(x,y)$  space-space  
 to  $(k_x, k_y)$  per-space-per-space  
 to  $(x, ct)$  space-time

Relativistic mode with near-c  $V_{group}=c/2$  and  $V_{phase}=2c$ . (Low dispersion.)



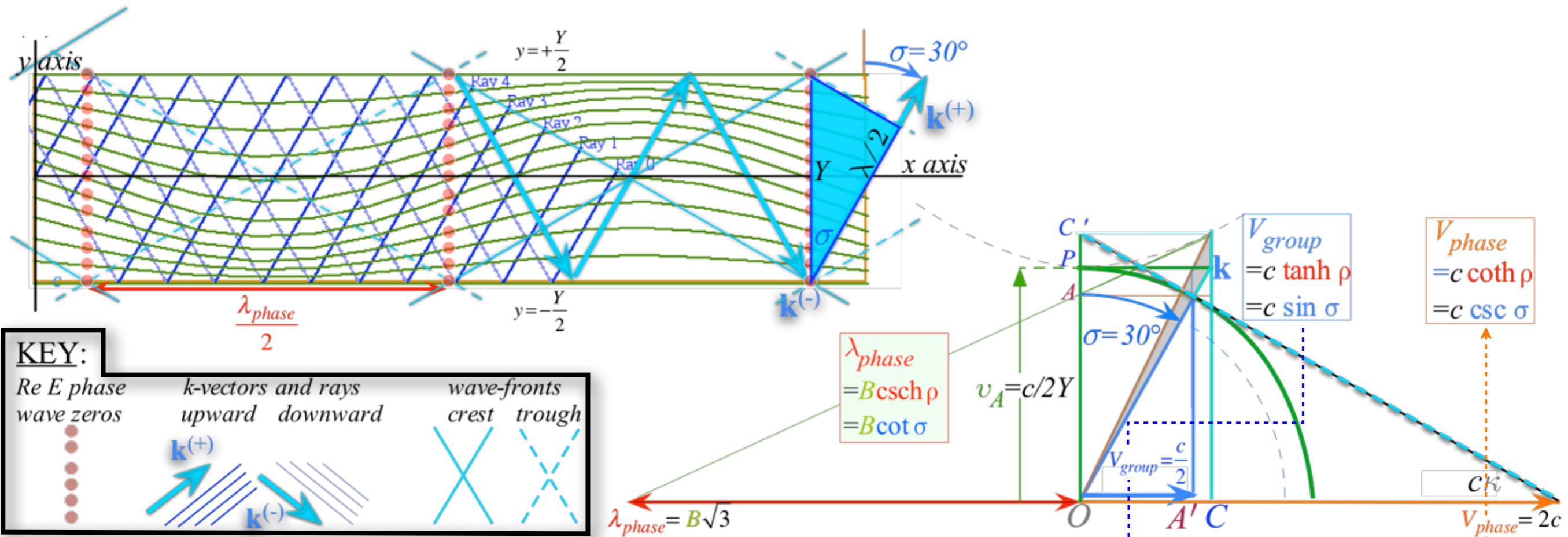
**KEY:**

<p>Re E phase wave zeros</p>	<p>k-vectors and rays upward downward</p>	<p>wave-fronts crest trough</p>
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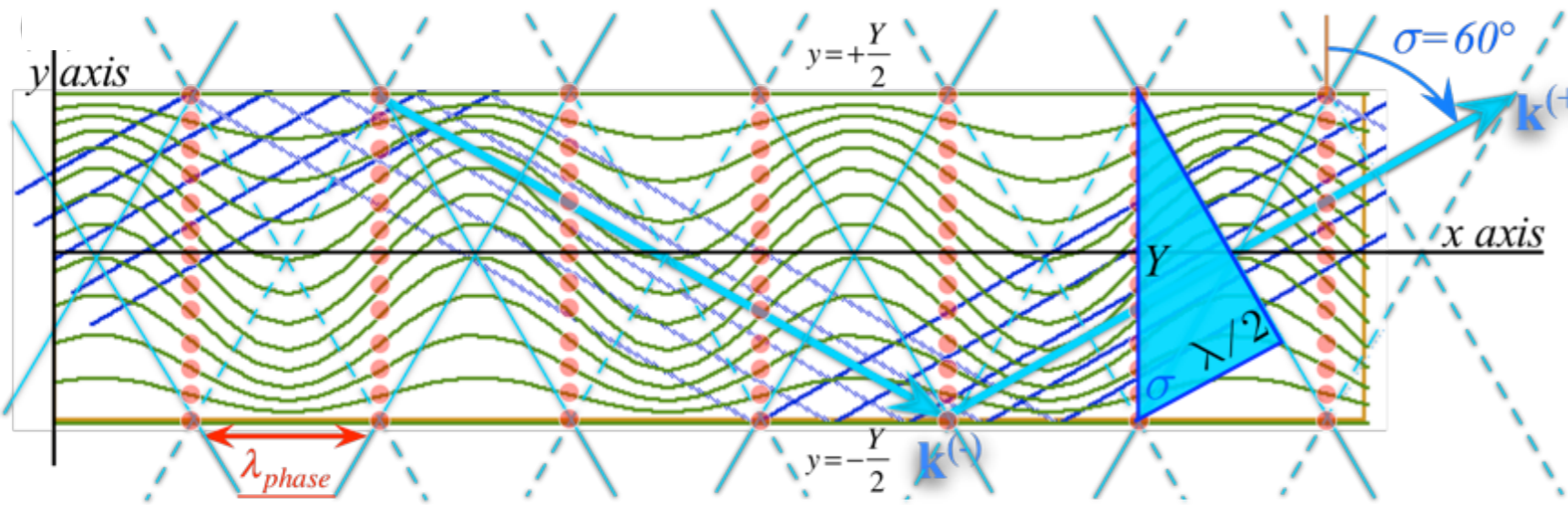


Example of near-cut-off mode with low  $V_{group} = c/2$  and high  $V_{phase} = 2c$ . (High dispersion.)

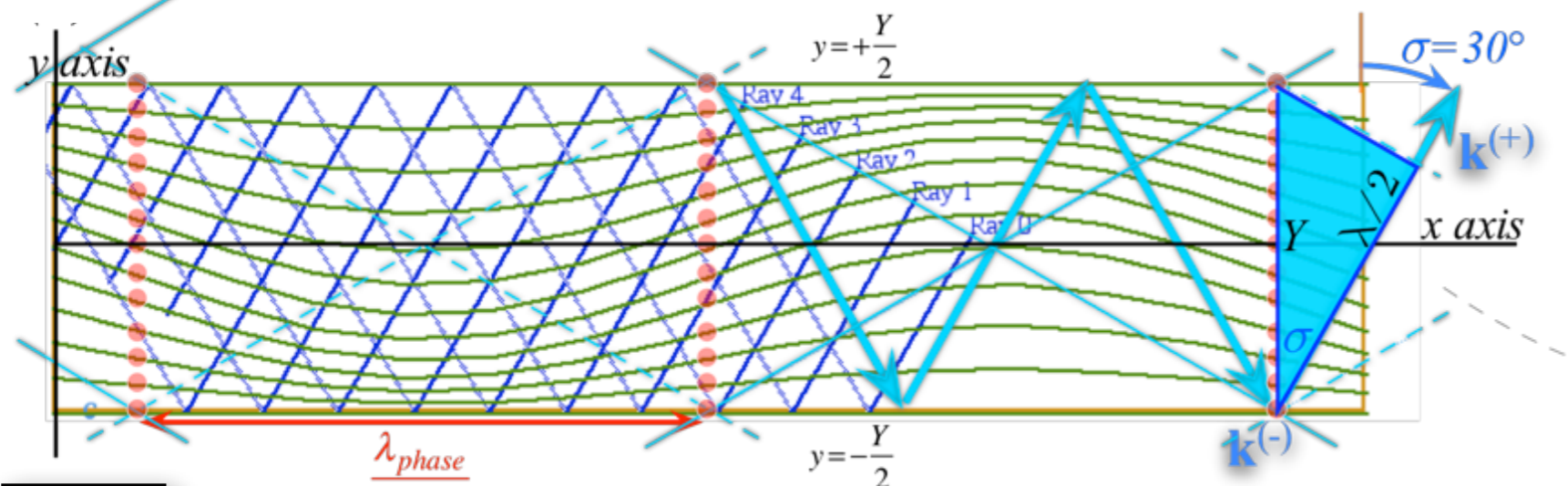
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GuidIt Web Simulation:  $\sigma = 60^\circ$

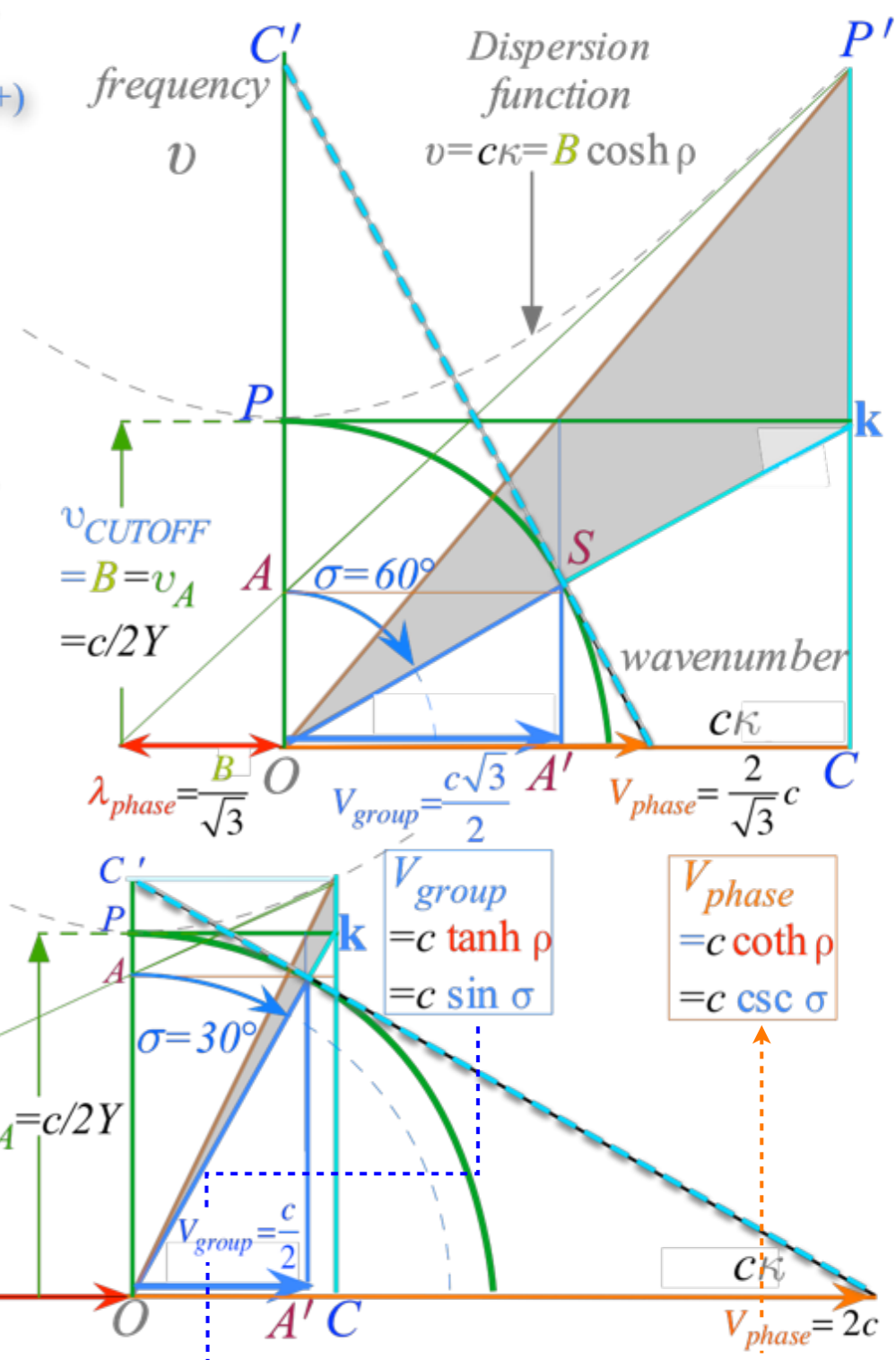


GuidIt Web Simulation:  $\sigma = 30^\circ$

**KEY:**

Re E phase wave zeros	k-vectors and rays upward downward	wave-fronts crest trough

$k^{(+)}$   $k^{(-)}$



Example of near-cut-off mode with low  $V_{group}=c/2$  and high  $V_{phase}=2c$ . (High dispersion.)



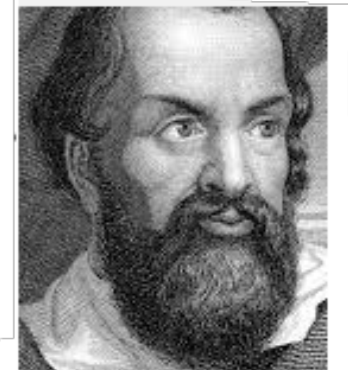
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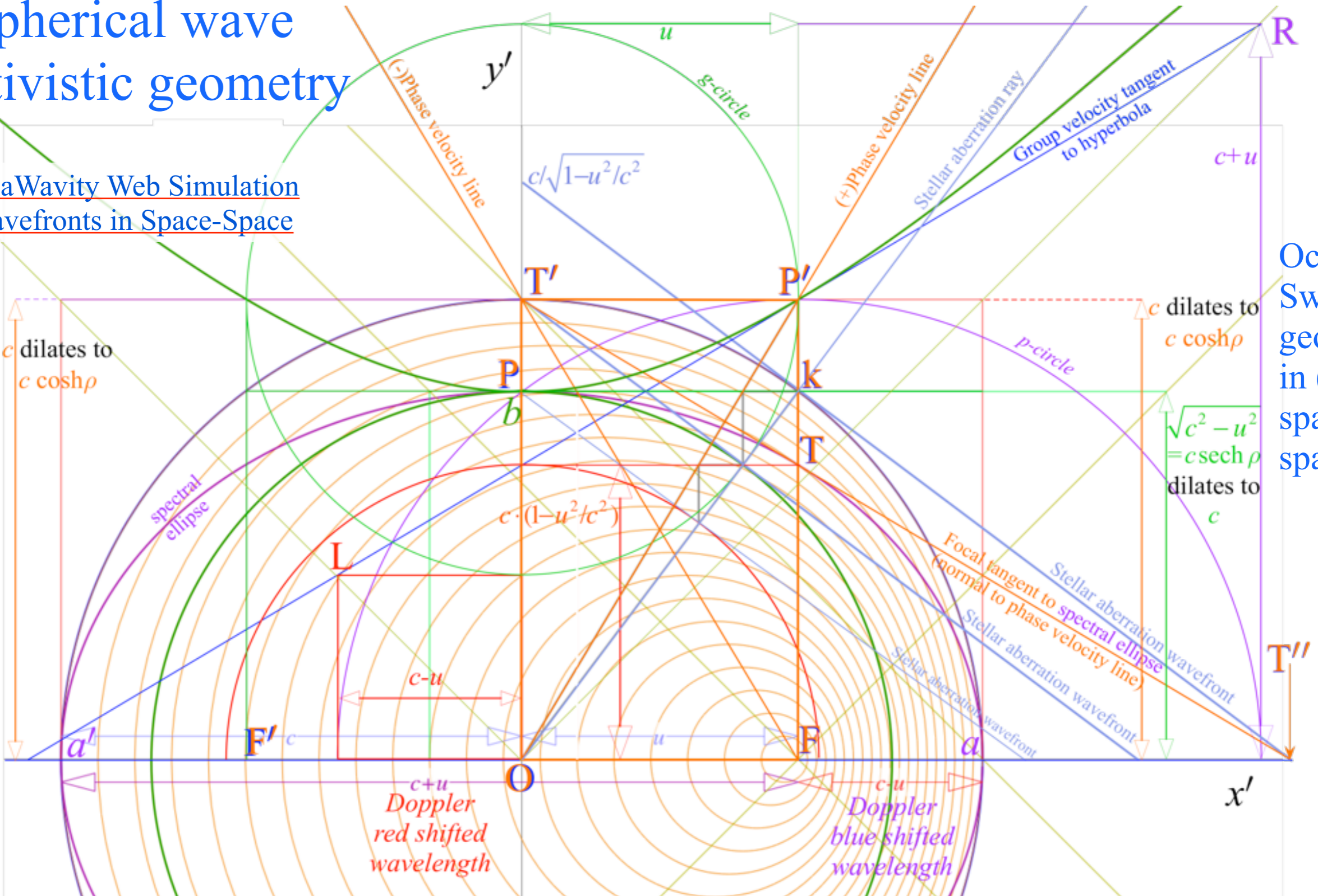






# Spherical wave relativistic geometry

[RelaWavity Web Simulation](#)  
[Wavefronts in Space-Space](#)



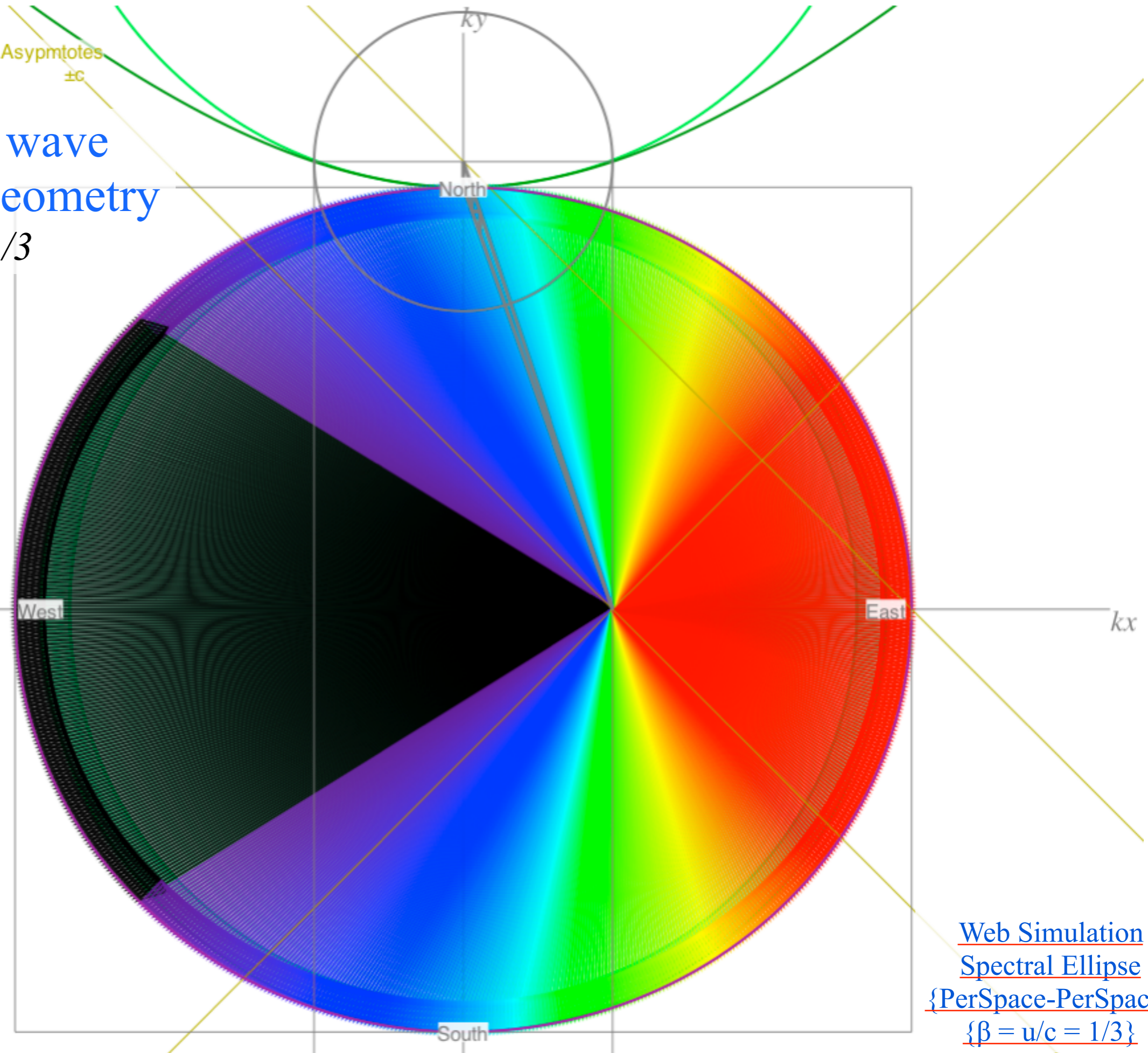
Occam  
Sword  
geometry  
in (x,y)  
space-space

<p>Doppler Red <math>\lambda=c+u</math>  dilates to: <math>(c+u) \cosh \rho = c \sqrt{\frac{c+u}{c-u}} = ce^{+\rho}</math></p> <p>ellipse major radius <math>a=OFa=c</math>  dilates to: <math>c \cosh \rho = c/\sqrt{1-u^2/c^2}</math></p>	<p>Applications of Einstein dilation factor:  <math>\gamma = \cosh \rho = 1/\sqrt{1-u^2/c^2}</math></p>	<p>ellipse focal length <math>FO = u = c \tanh \rho</math>  dilates to: <math>u \cosh \rho = c \sinh \rho</math></p> <p>ellipse latus radius <math>FT = c(1-u^2/c^2)</math>  dilates to: <math>c(1-u^2/c^2) \cosh \rho = c\sqrt{1-u^2/c^2} = c \operatorname{sech} \rho</math></p>	<p>Doppler Blue <math>\lambda=c-u</math>  dilates to: <math>(c-u) \cosh \rho = c \sqrt{\frac{c-u}{c+u}} = ce^{-\rho}</math></p> <p>Base height <math>FTk = \sqrt{c^2 - u^2}</math>  dilates to: <math>\sqrt{c^2 - u^2} \cosh \rho = c</math>  (equal to ellipse minor radius <math>b</math>)</p>
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Spherical wave  
relativistic geometry

$u/c = 1/3$

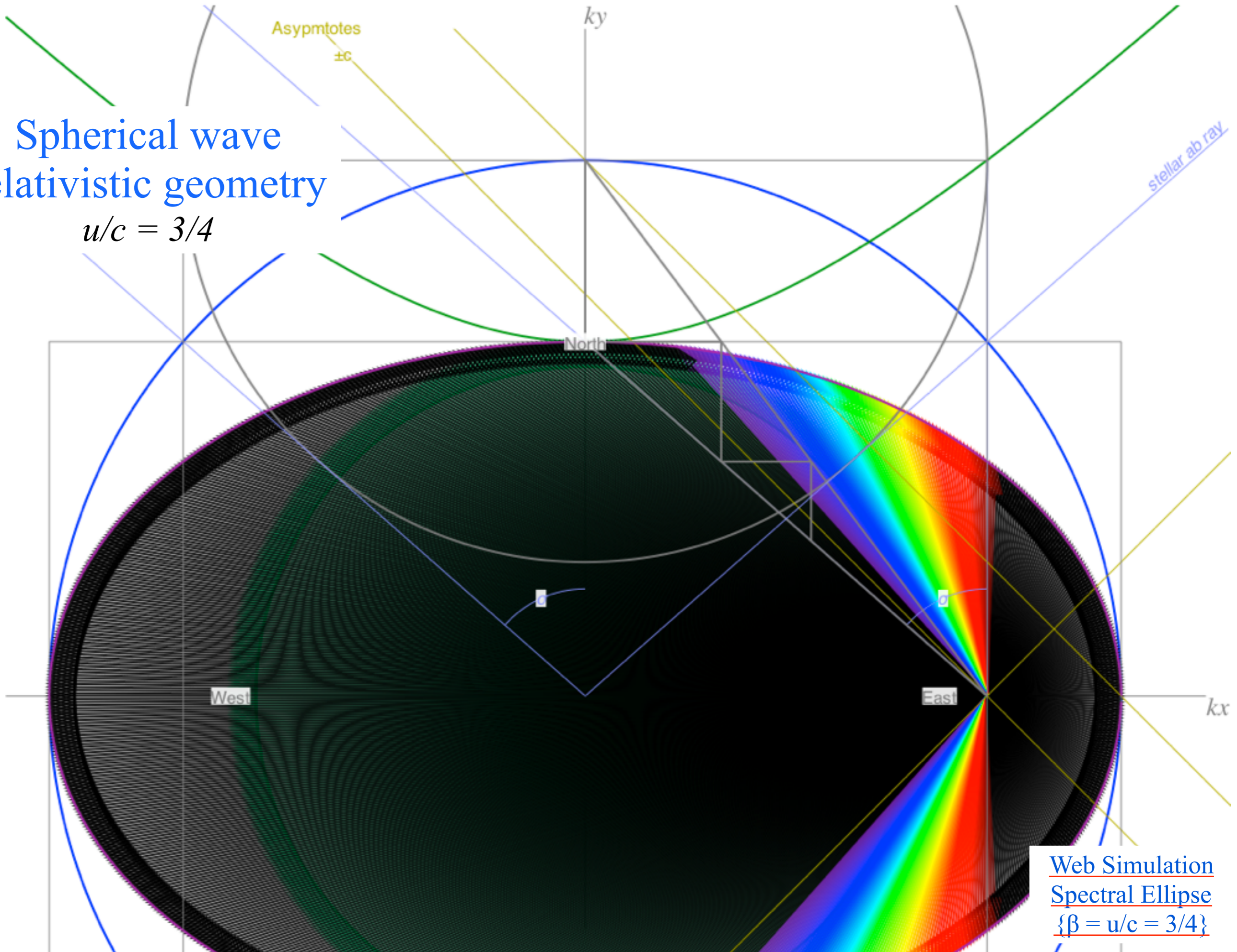


[Web Simulation](#)  
[Spectral Ellipse](#)  
[{PerSpace-PerSpace}](#)  
[{ \$\beta = u/c = 1/3\$ }](#)

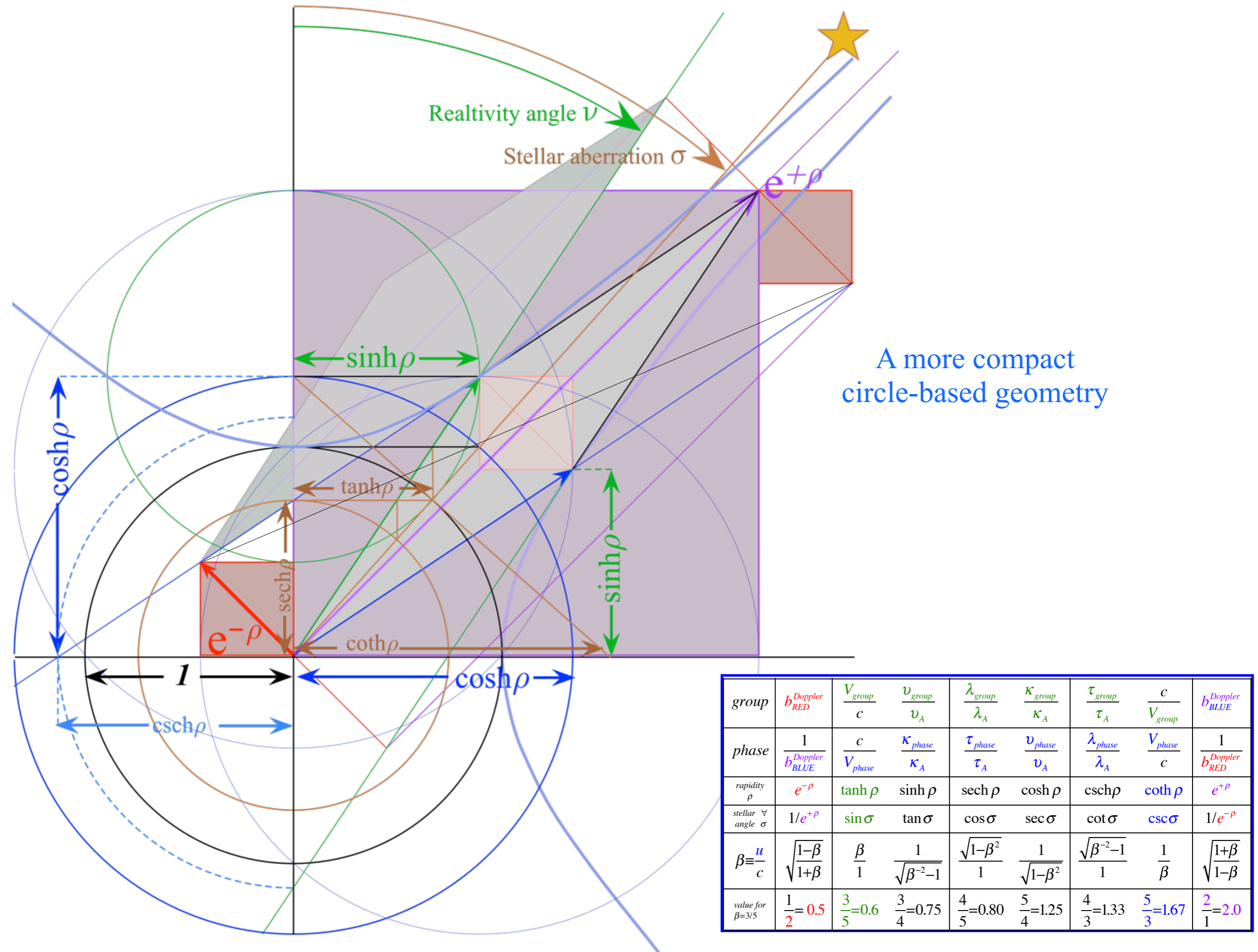


Spherical wave  
relativistic geometry

$$u/c = 3/4$$



Web Simulation  
Spectral Ellipse  
{ $\beta = u/c = 3/4$ }



A more compact circle-based geometry

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$