

# Lecture 30

Wed. 12.05.2018

## *Relawavity : a novel introduction to relativistic mechanics III.*

([CMwBang! Unit 8](#) , [AMOP Ch.0](#) , )

Review: Relawavity  $\rho$  functions and plots vs.  $\rho$

Derive relawavity parameters and Minkowski coordinates for  $\nu_R=2.5\text{THz}$  and  $\nu_L=0.5\text{THz}$

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

Relativistic action and Lagrangian-Hamiltonian relations

Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization "photon" number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relawavity* in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# *A running collection of links to course-relevant sites and articles*

## *Physics Web Resources*

[Comprehensive Harter-Soft Resource Listing](#)

[UAF Physics YouTube channel](#)

[LearnIt Physics Web Applications](#)

Neat external material to start the class:

[AIP publications](#)

[AJP article on superball dynamics](#)

[AAPT summer reading](#)

These are hot off the presses:

[Sorting ultracold atoms in a 3D optical lattice in a realization of Maxwell's demon - Kumar-Nature-Letters-2018](#)

[Synthetic three-dimensional atomic structures assembled atom by atom - Berredo-Nature-Letters-2018](#)

Slightly Older ones:

[Wave-particle duality of C60 molecules](#)

[Optical vortex knots – One Photon at a Time](#)

## *Older Links from Lectures 14-20*

<http://thearmchaircritic.blogspot.com/2011/11/punkin-chunkin.html>

<http://www.sussexcountyonline.com/news/photos/punkinchunkin.html>

[Shooting-range-for-medieval-siege-weapons-Anybody-knows](#)

<https://modphys.hosted.uark.edu/markup/TrebuchetWeb.html>

<https://modphys.hosted.uark.edu/markup/TrebuchetWeb.html?scenario=MontezumasRevenge>

<https://modphys.hosted.uark.edu/markup/TrebuchetWeb.html?scenario=SeigeOfKenilworth>

[The trebuchet, Chevedden, Sci Am 1995](#)

'Simple' Pendulum Sim: <https://modphys.hosted.uark.edu/markup/PendulumWeb.html>

'Cycloid' Pendulum: <https://modphys.hosted.uark.edu/markup/CycloidulumWeb.html>

[Google search on: "Satelite view of Patricia" \(Images\)](#)

[Physics Girl Channel - Fun with Vortex Rings in the Pool](#)

[iBall demo - Quasi-periodicity: https://youtu.be/\\_jntDtULxDe](#)

<https://modphys.hosted.uark.edu/markup/CoulltWeb.html?scenario=SynchrotronMotion>

<https://modphys.hosted.uark.edu/markup/CoulltWeb.html?scenario=SynchrotronMotion2>

[Mechanical Analog to EM Motion \(YouTube video\) - https://youtu.be/hTd5FTJ-vRk](#)

[Coullt Web Simulation: Bound-state motion in parabolic coordinates](#)

[Coullt Web Simulation: Bound-state motion in hyperbolic coordinates](#)

[Oscillt Web App: Simulations of various types of resonance: 18, 27, 31, 35, 38, 39](#)

[Smith Chart](#)

<http://nobelprize.org/>

*AnalyIt Web Application*, posted 10/22/2018 in our *testing area*:

<https://modphys.hosted.uark.edu/testing/markup/AnalyItBJS.html>

## *"Texts"*

[Classical Mechanics with a Bang!](#)

[Quantum Theory for the Computer Age](#)

[Principles of Symmetry, Dynamics, and Spectroscopy](#)

[Modern Physics and its Classical Foundations](#)

[AMOP Detailed Development of Relativity](#)

[AMOP Ch 0 Space-Time Symmetry - 2019](#)

[Seminar at Rochester Institute of Optics, Auxiliary slides, June 19, 2018](#)

[Springer AMO Handbook - Ch 32 - Harter-Reimer-2019](#)

"Relativity" and quantum basis of *Lagrangian & Hamiltonian* mechanics:

[2-CW laser wave - BohrIt Web App](#)

[Lagrangian vs Hamiltonian - RelaWavity Web App](#)

## *Classes*

[2014 AMOP](#)

[2017 Group Theory for QM](#)

[2018 AMOP](#)

[2018 Adv Mechanics](#)

## *Older Links from Lectures 21-23*

Advanced Atomic and Molecular Optical Physics 2018 Class #9, pages: [5](#), [61](#)

[BoxIt Web Simulations](#)

[Pure A-Type w/Cosine](#)

[Pure B-Type w/Cosine](#)

[Pure B-Type w/Freq ratios](#)

[Mixed AB-Type 2:1 Freq ratio](#)

[Pure A-Type A=4.9, B=0, C=0, & D=4.0](#)

[Pure B-Type: A=4.0, B=-0.2, C=0, & D=4.0](#)

[Pure C-Type A,D=4.055, B=0, C=0.1](#)

[Mixed AB-Type w/Cosine](#)

[Mixed AB Type A=4.0, BU2=0.866..., CU2=0, & D=1.0 w/Stokes & Freq rats](#)

[Mixed AB Type A=5.086 B=-0.27 C=0 D=2.024 w/Stokes plot](#)

[Mixed ABC Type A=4.833 B=0.2403 C=0.4162 D=4.277 w/Stokes plot](#)

[Recent mixed ABC Type A=0.325 B=0.375 C=0.825 D=0.05 w/Stokes plot](#)

[Classical Mechanics with a Bang! 2018](#)

[Lectures 8, 9, 23 page 93](#)

[Text Unit 6, page=27](#)

[ColorU2 for the Web - in development](#)

[Group Theory for Quantum Mechanics - 2017 Lectures: \[6\]\(#\), \[7\]\(#\), \[8\]\(#\), and the combined \[9-10\]\(#\)](#)

[Quantum Theory for the Computer Age Unit 3 Ch.7-10, page=90](#)

[Web based 3D & XR \( \$x \in \{A, M, V\}\$ , R=Reality\) <https://www.babylonjs.com/>](#)

[Web based 3D graphics WebGL API \(Graphics Layer modeled after OpenGL\)](#)

[Wiki on Pafnuty Chebyshev](#)

*continued* ↘

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## *Classes*

[2014 AMOP](#)

[2017 Group Theory for QM](#)

[2018 AMOP](#)

[2018 Adv Mechanics](#)

←————— Repeated from previous page —————→

## *Older Links from Lectures 24-27*

[JerkIt Web App: 2-, 2+, Amp50Omega147-, Amp50Omega296, Amp50Omega602, Gap\(1\)](#)

[MolVibes Web App: C3vN3](#)

[Wavelt Web App:](#)

[Dim = 3 w/Wave Components;](#)

[Static Char Table: 6, 12, 12\(b\), 16, 36, 256](#)

[Quantum Carpet with N=20: Gaussian, Boxcar](#)

[Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-CPL-2015](#)

[QTCA Unit 5 Ch14 2013](#)

[Lester. R. Ford, Am. Math. Monthly 45,586\(1938\)](#)

[John Farey, Phil. Mag.\(1816\) Wolfram](#)

[Harter, J. Mol. Spec. 210, 166-182 \(2001\)](#)

[Harter, Li IMSS \(2013\)](#)

[Li, Harter, Chem.Phys.Letters \(2015\)](#)

[OscillatorPE Web App: IHO Scenario 2, Coulomb Scenario 3](#)

[RelaWavity Web App/Simulator/Calculator: Elliptical - IHO orbits](#)

[Coultt Web App Simulations: p19, p32, p72, p73, p92, R=-0.375, R=+0.5, Rutherford](#)

## *Older Links from Lectures 28*

[CMwBang Text 2012 Unit 6 page=5](#)

[Bouncelt Web App/Scenarios:](#)

[5002, 5003](#)

[Coultt Web App/Scenarios:](#)

[TwoParticleCollision LToR, TwoParticleCollision LToR CM, TwoParticleOrbit Coulomb,](#)

[TwoParticleOrbit Coulomb CM, TwoParticleOrbit Hooke, TwoParticleOrbit Hooke CM](#)

[Singular Motion of Asymmetric Rotators AJP 44, 11 p1080 Harter-Kim-1976](#)

[Molecular Eigensolution Symmetry Analysis and Fine Structure - Int.J.MolSci1.4.13 Harter-Mitchell-IJMS-2013](#)

[Lenz Vector and Orbital Analog Computers - AJP 44 p348 1976](#)

[Some Geometric Aspects of Classical Coulomb Scattering AJP 40 4 p1852 1972](#)

[How Molecules do Self-NMR - Harter-Mitchell-Columbus-2009](#)

[Classical Mechanics with a Bang! - Asymmetric Top Demo](#)

[Allbookstores.com - Compare for Heller's SemiClassical Way - 0691163731](#)

["My Bomerang Won't Come Back" \(YouTube: Playlist\)](#)

[Rotating Solid Bodies in Microgravity \(YouTube\)](#)

[Dancing T-handle in zero-g \(YouTube\)](#)

## *Supplemental Links for Lectures 29-30*

### *RelaWavity*

[AMOP Chapter 0: Space-Time Symmetry](#)

[AMOP Detailed Development of RelaWavity](#)

[2018 Rochester Talk \(Auxiliary Slides\)](#)

[Special Relativity and Quantum Theory by Ruler and Compass](#) - Earlier, expanded draft

[Ruler & Compass - Relawavity Exercise](#)

[2018 RelaWavity Portal Page](#)

[Pirelli Relativity Challenge Web Site:](#)

[Clocks 12 hr, Clocks 24 hr QT, Phasors Addition, Quantized 1](#)

[BohrIt Web App/Simulations: -130022; -30001; -30104; 30004; 30022](#)

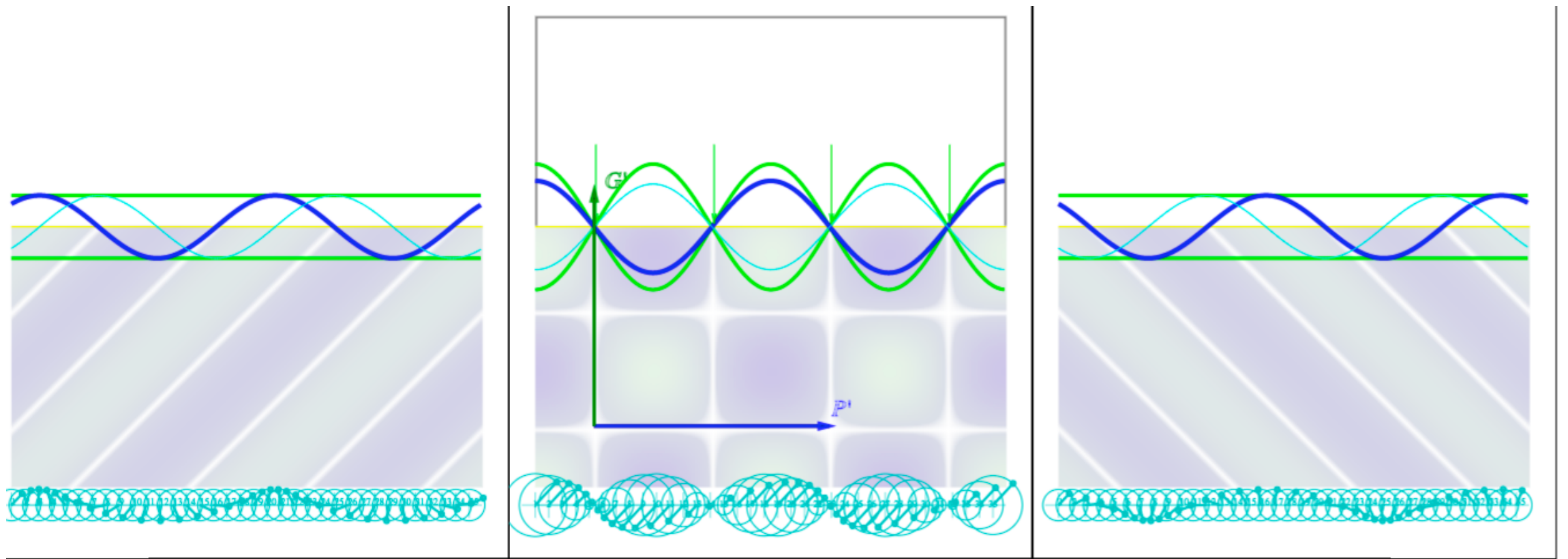
[Guidelt Web App/Scenarios: 230; 260](#)

[RelativIt Web App/Scenarios: 22; 24; 69](#)

[RelaWavity Web App/Scenarios: 0,9; 3,6; 3,6 NoMink; 4,8; 5,6a; 6,1; 6,3a; 6,3b; 6.3c; 7,1;](#)

[7,2,1; 7,2,2; 7,2,3; 7,2,7; 8,3; 8,5; 8,6; 8,7; 8,8](#)

[2012 ModPhys Lect 35](#)



$$\begin{aligned}
 e^{iR} + e^{iL} &= e^{i\frac{R+L}{2}} \left( e^{i\frac{R-L}{2}} + e^{-i\frac{R-L}{2}} \right) \\
 &= e^{i\frac{R+L}{2}} 2 \cos \frac{R-L}{2} \\
 &= 2e^{-i\omega t} \cos kx \\
 &= \psi_{\text{phase}} \psi_{\text{group}}
 \end{aligned}$$

$$\begin{aligned}
 R &= k_R x - \omega_R t \quad \text{and:} \quad L = -k_L x - \omega_L t \\
 &= kx - \omega t \quad \quad \quad = -kx - \omega t
 \end{aligned}$$

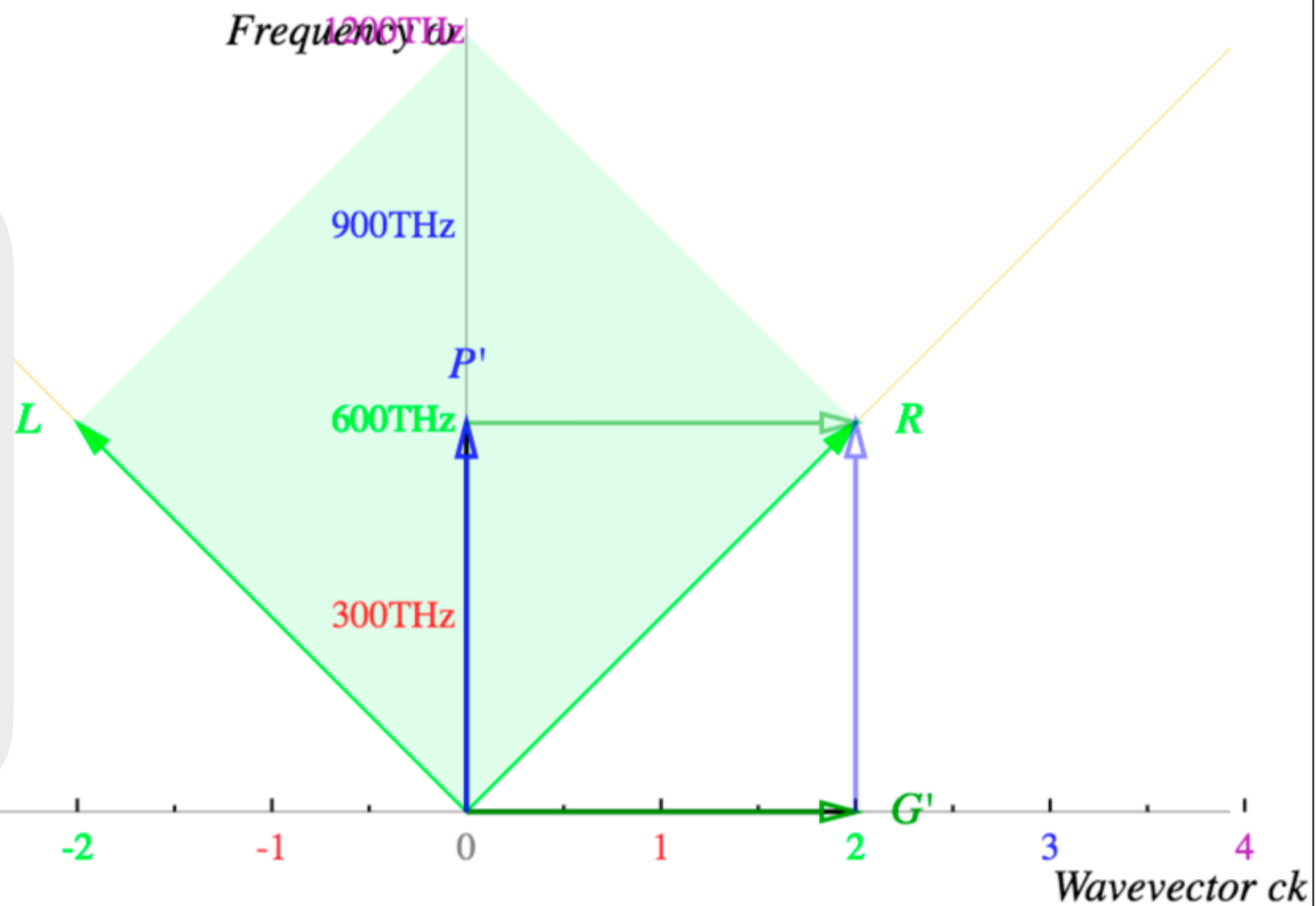
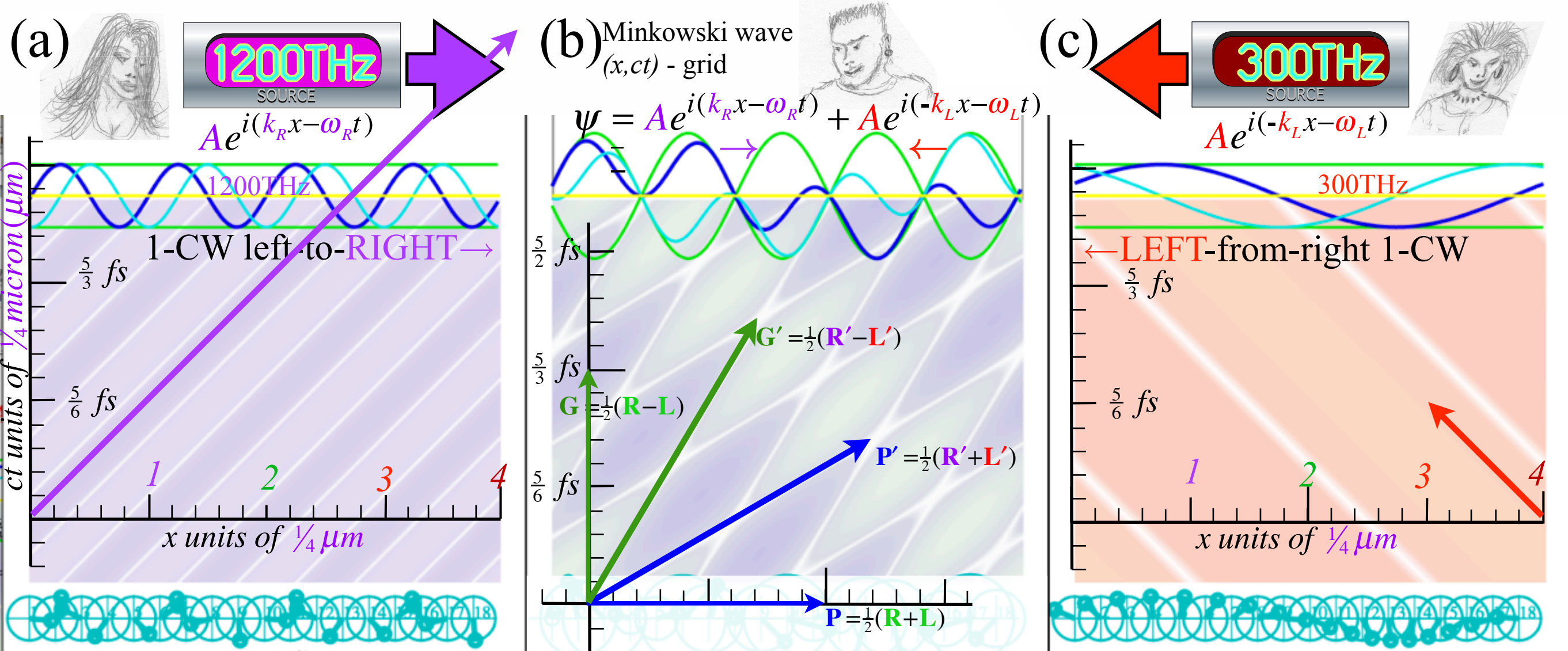


Fig. 9 in text





(d)

$$e^{iR'} + e^{iL'} = e^{i\frac{R'+L'}{2}} (e^{i\frac{R'-L'}{2}} + e^{-i\frac{R'-L'}{2}})$$

$$= e^{i\frac{R'+L'}{2}} 2 \cos \frac{R'-L'}{2}$$

$$= \psi'_{phase} \psi'_{group}$$

$R' = k_R x - \omega_R t$  and:  $L' = -k_L x - \omega_L t$

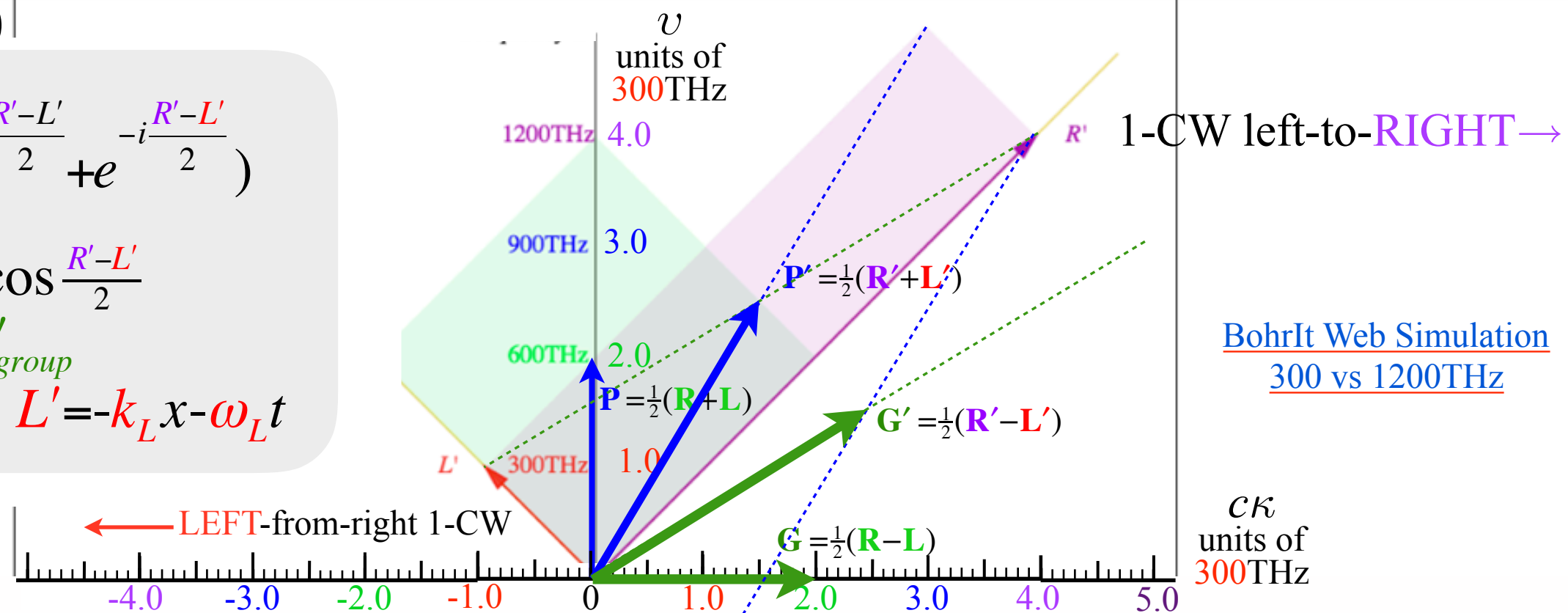
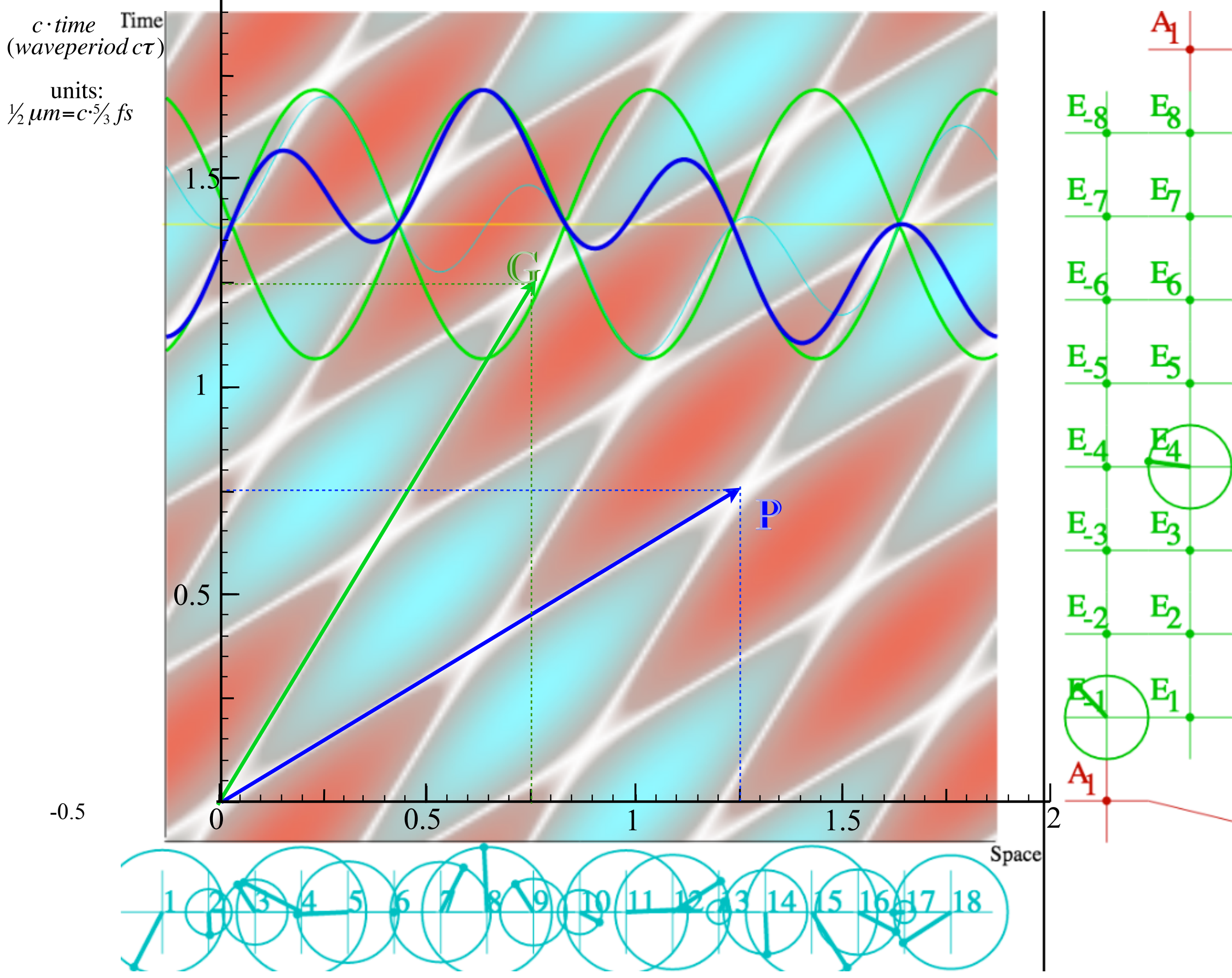
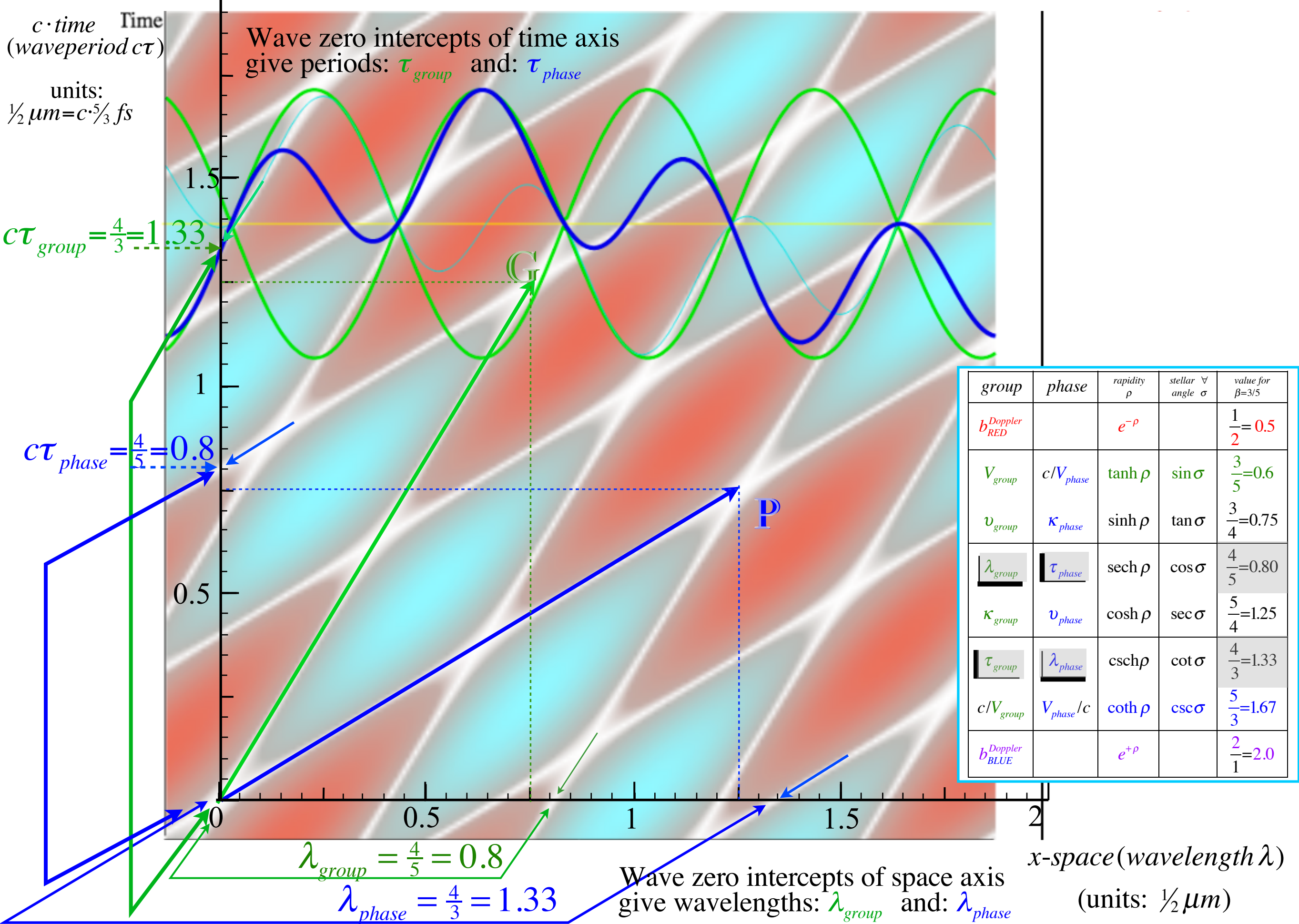
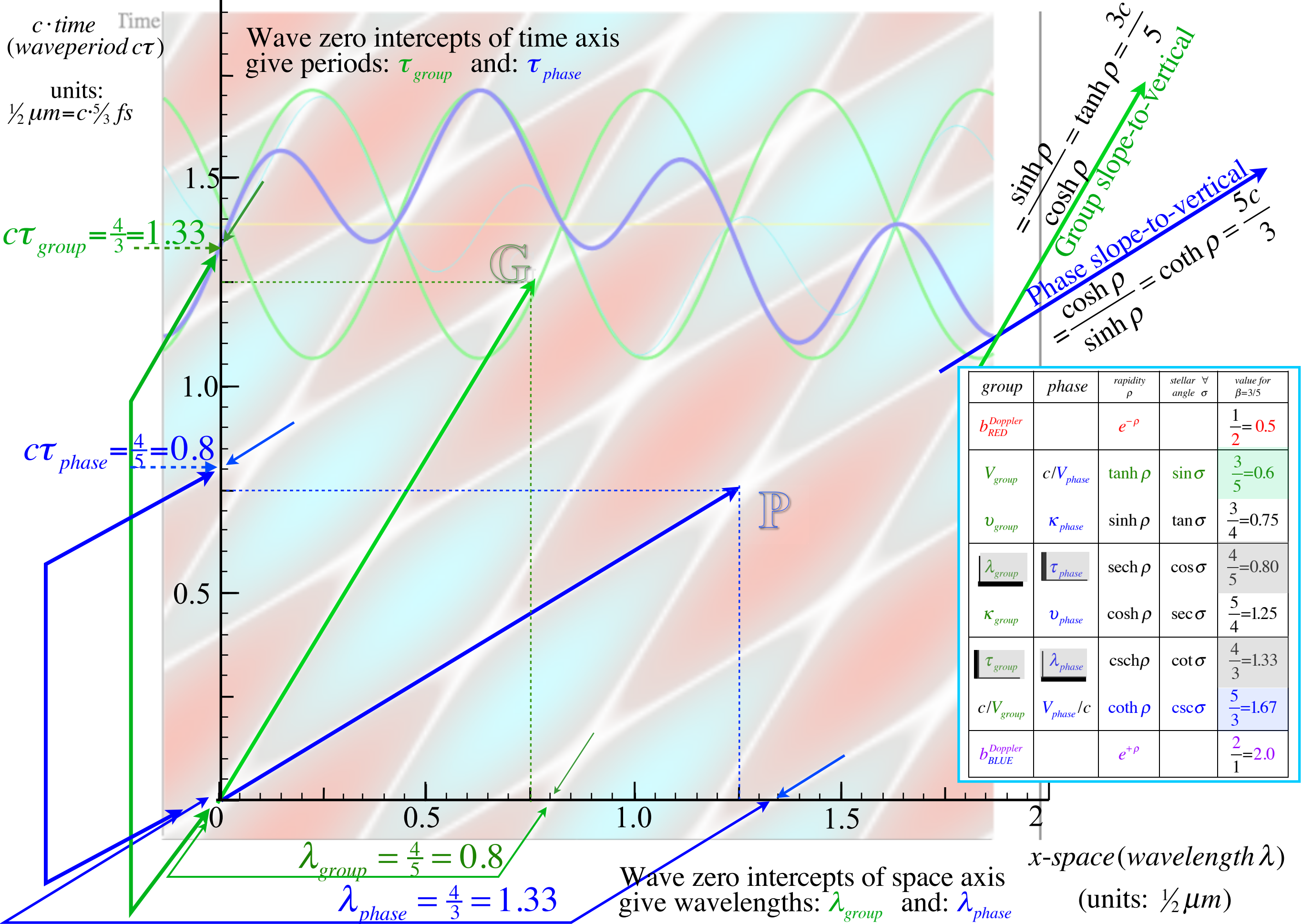


Fig. 3 in Ch.0 intro

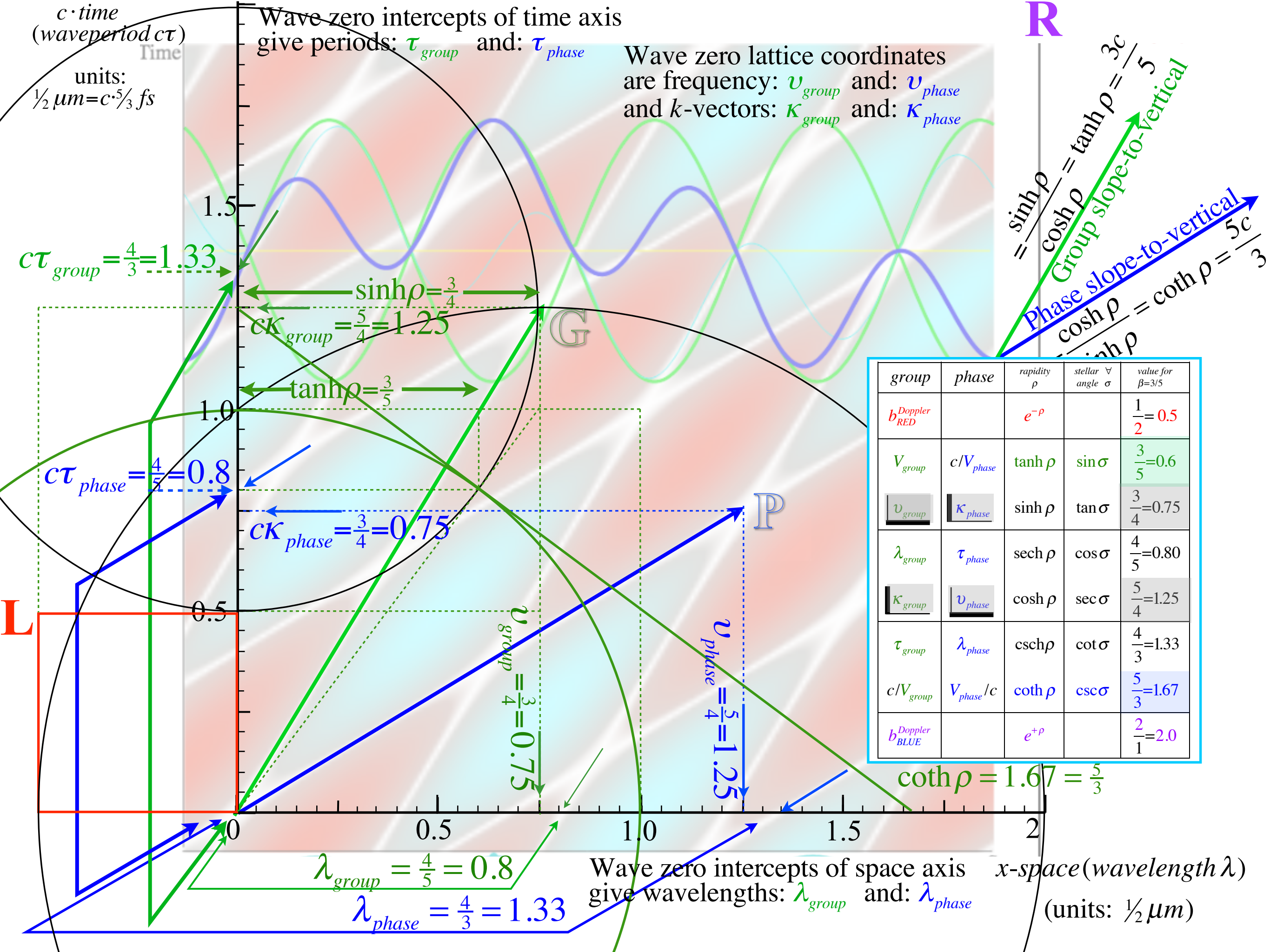












This map has circle sector arc-area  $\sigma = 0.6435$

set to angle  $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

$$\begin{aligned} \sin(\sigma) &= 0.6000 &= \tanh(\rho) &= 3/5 \\ \tan(\sigma) &= 0.7500 &= \sinh(\rho) &= 3/4 \\ \sec(\sigma) &= 1.2500 &= \cosh(\rho) &= 5/4 \\ \cos(\sigma) &= 0.8000 &= \operatorname{sech}(\rho) &= 4/5 \\ \cot(\sigma) &= 1.3333 &= \operatorname{csch}(\rho) &= 4/3 \\ \csc(\sigma) &= 1.6667 &= \operatorname{coth}(\rho) &= 5/3 \end{aligned}$$

$$\cosh(\rho) + \sinh(\rho) = \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho}$$

$$\cosh(\rho) - \sinh(\rho) = \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho}$$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2} \quad \text{Half-Sum-}$$

$$\sinh(\rho) = \frac{e^{+\rho} - e^{-\rho}}{2} \quad \text{Half-Difference}$$

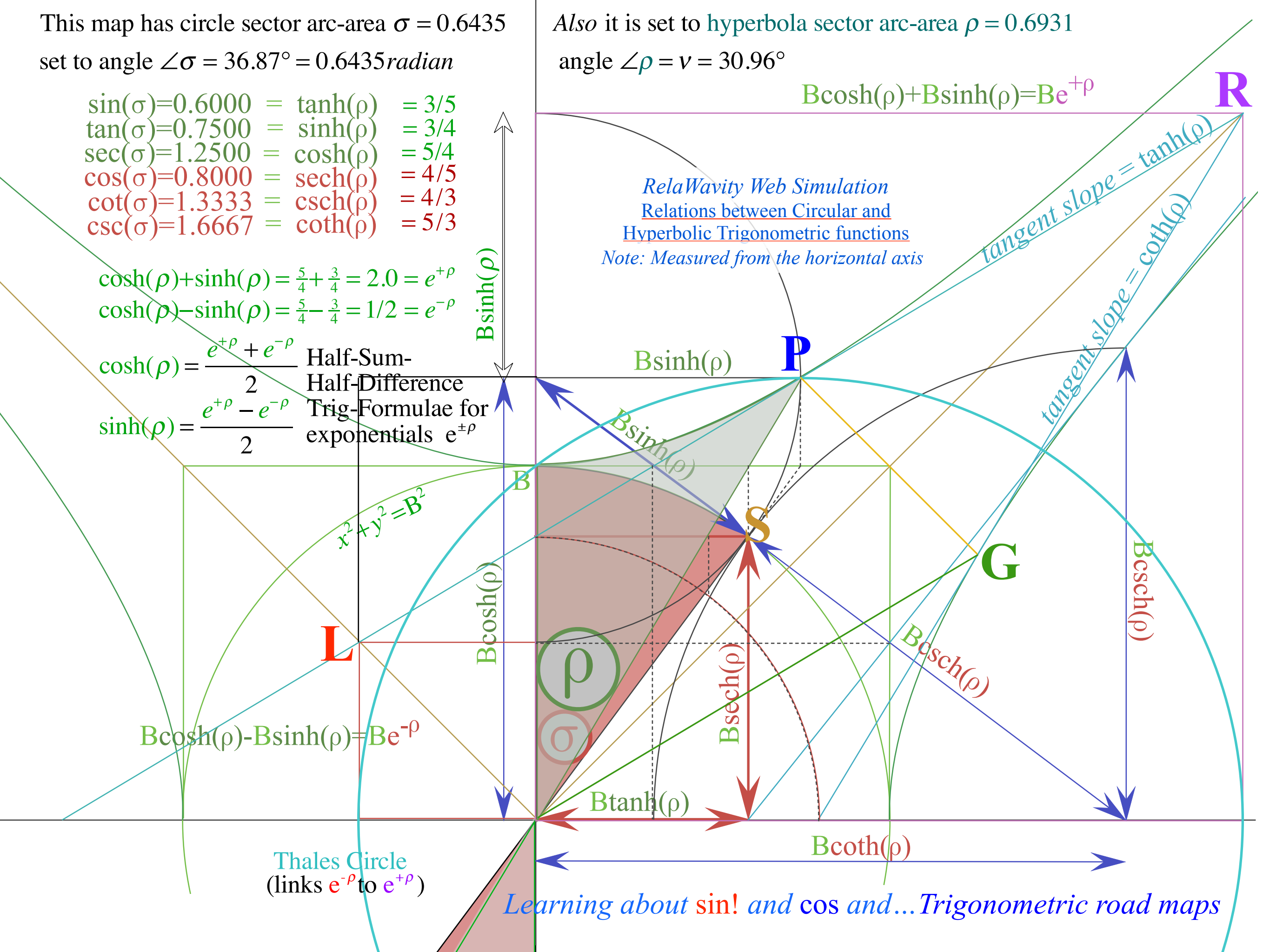
Trig-Formulae for  
exponentials  $e^{\pm\rho}$

Also it is set to hyperbola sector arc-area  $\rho = 0.6931$

angle  $\angle\rho = \nu = 30.96^\circ$

$$B\cosh(\rho) + B\sinh(\rho) = B e^{+\rho}$$

*RelaWavity Web Simulation*  
Relations between Circular and  
Hyperbolic Trigonometric functions  
*Note: Measured from the horizontal axis*



$$B\cosh(\rho) - B\sinh(\rho) = B e^{-\rho}$$

Thales Circle  
(links  $e^{-\rho}$  to  $e^{+\rho}$ )

Learning about **sin!** and **cos** and... Trigonometric road maps

# Derive relativity parameters and Minkowski coordinates for $\nu_R=2.5\text{THz}$ and $\nu_L=0.5\text{THz}$

Per-space-time  $(\nu', c\kappa')$  geometry of 2-CW vectors

Space-time  $(c\tau', x')$  geometry of 2-CW paths

Time-inversion symmetry requires:  
 (Red-shift factor)(Blue-shift factor) = 1

$$(Be^{-\rho})(Be^{\rho}) = B^2 = (0.5)(2.5) = \frac{1}{2} \cdot \frac{5}{2} = \frac{5}{4}$$

$$B = \frac{\sqrt{5}}{2} = 1.12..$$

$$\mathbf{R} = B \begin{pmatrix} e^{\rho} \\ e^{\rho} \end{pmatrix} \approx \begin{pmatrix} 5/2 \\ 5/2 \end{pmatrix}$$

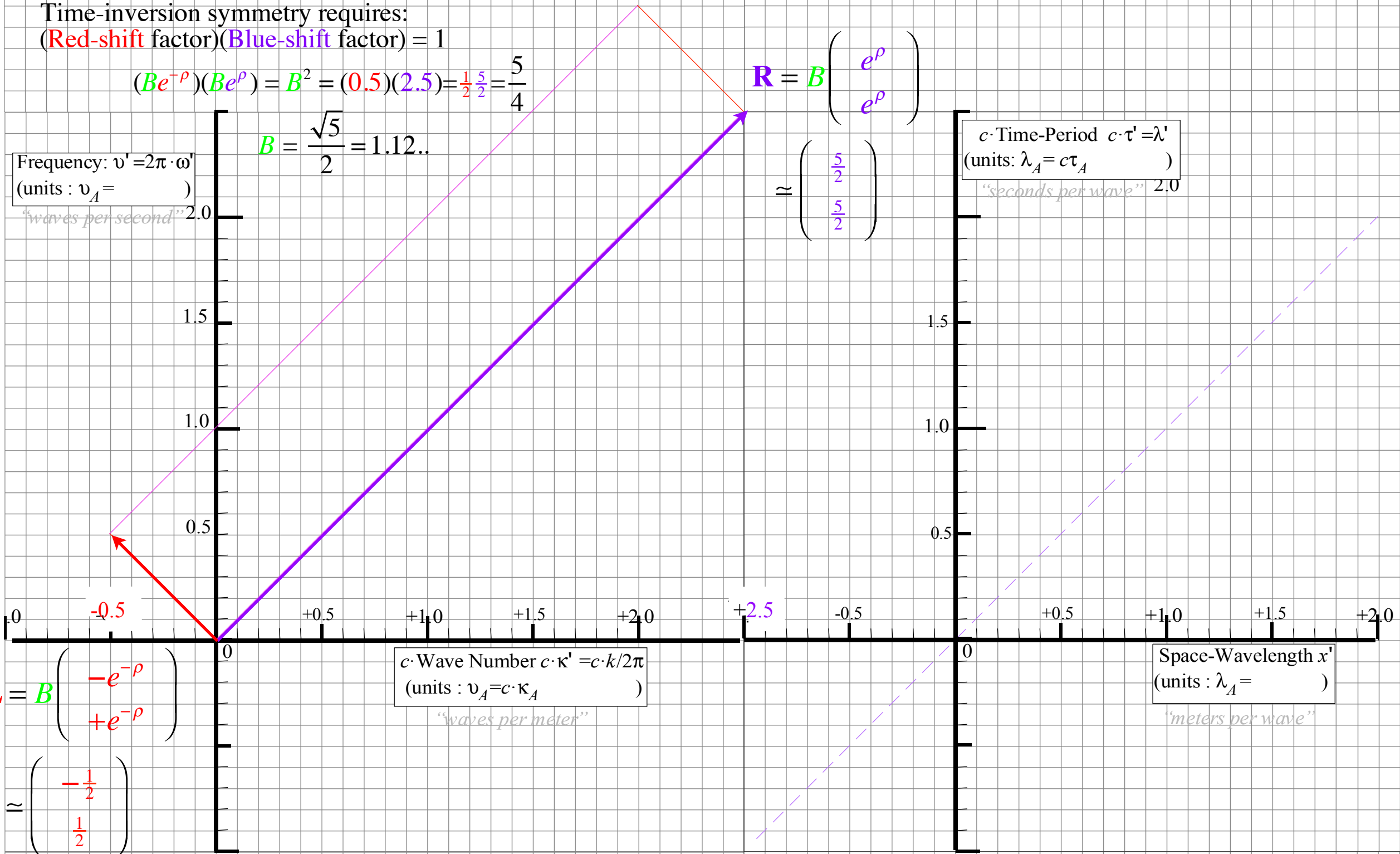
Frequency:  $\nu' = 2\pi \cdot \omega'$   
 (units:  $\nu_A =$  )  
 "waves per second"

$c \cdot \text{Time-Period } c \cdot \tau' = \lambda'$   
 (units:  $\lambda_A = c\tau_A$  )  
 "seconds per wave"

$$\mathbf{L} = B \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix} \approx \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix}$$

$c \cdot \text{Wave Number } c \cdot \kappa' = c \cdot k / 2\pi$   
 (units:  $\nu_A = c \cdot \kappa_A$  )  
 "waves per meter"

Space-Wavelength  $x'$   
 (units:  $\lambda_A =$  )  
 "meters per wave"



# Derive relativity parameters and Minkowski coordinates for $\nu_R=2.5\text{THz}$ and $\nu_L=0.5\text{THz}$

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$c \cdot \text{Time-Period } c \cdot \tau' = \lambda'$   
(units:  $\lambda_A = c\tau_A$  )  
"seconds per wave"

$$\begin{pmatrix} \frac{1}{2} & \frac{5}{2} \\ \frac{1}{2} & \frac{5}{2} \end{pmatrix} + \begin{pmatrix} -\frac{1}{2} & \frac{1}{2} \\ \frac{1}{2} & \frac{1}{2} \end{pmatrix} \approx \frac{1}{2}(\mathbf{R} + \mathbf{L}) = \mathbf{P}$$

$$\approx \begin{pmatrix} 1 \\ \frac{3}{2} \end{pmatrix}$$

$$\mathbf{G} = \frac{1}{2}(\mathbf{R} - \mathbf{L})$$

$$\approx \frac{1}{2} \begin{pmatrix} \frac{5}{2} \\ \frac{5}{2} \end{pmatrix} - \frac{1}{2} \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix} = \begin{pmatrix} \frac{3}{2} \\ 1 \end{pmatrix}$$

$$\mathbf{L} = B \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix}$$

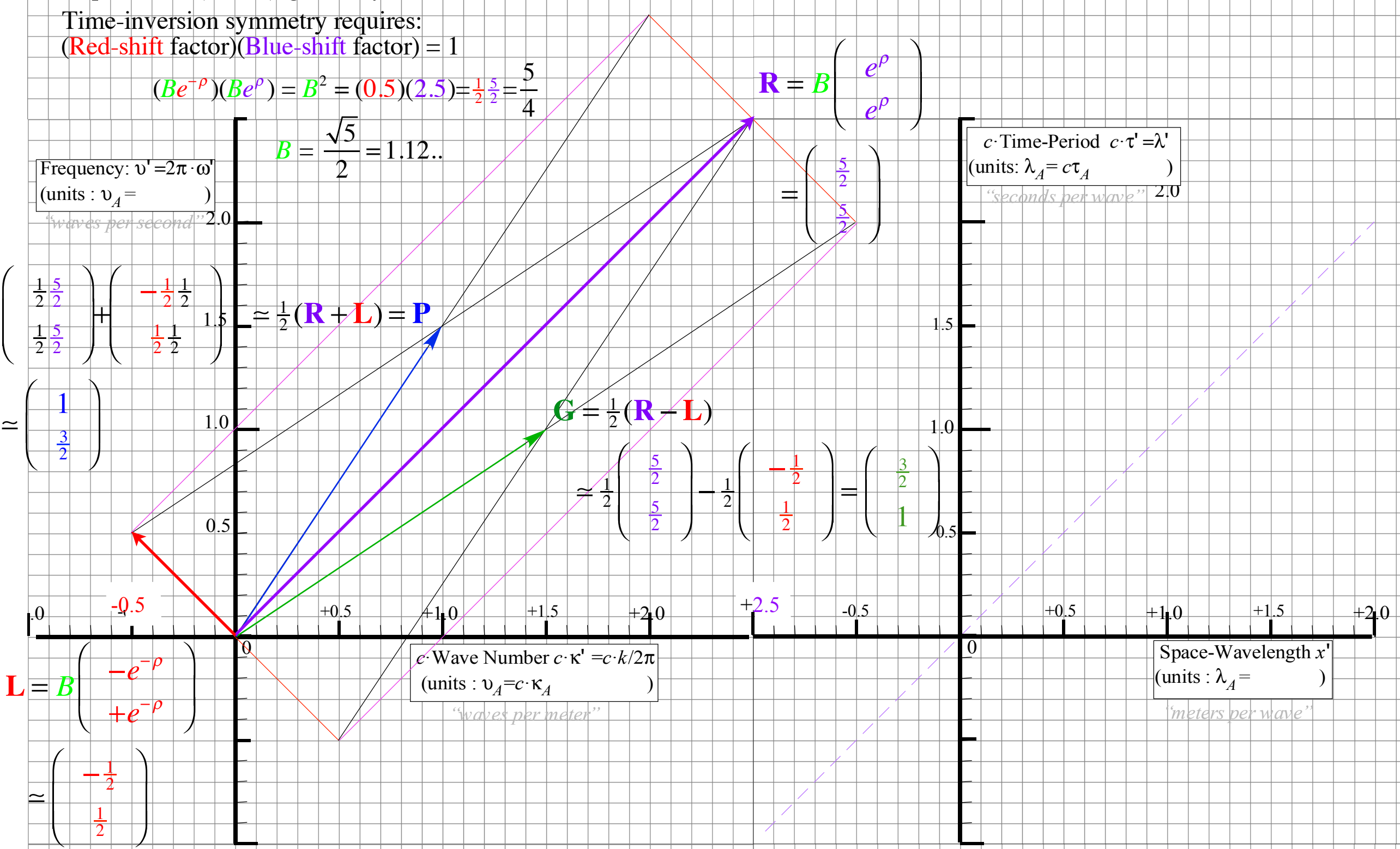
$$\approx \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix}$$

$$\mathbf{R} = B \begin{pmatrix} e^{\rho} \\ e^{\rho} \end{pmatrix}$$

$$= \begin{pmatrix} \frac{5}{2} \\ \frac{5}{2} \end{pmatrix}$$

$c \cdot \text{Wave Number } c \cdot \kappa' = c \cdot k / 2\pi$   
(units:  $\nu_A = c \cdot \kappa_A$  )  
"waves per meter"

Space-Wavelength  $x'$   
(units:  $\lambda_A =$  )  
"meters per wave"





# Derive relativity parameters and Minkowski coordinates for $\nu_R=2.5\text{THz}$ and $\nu_L=0.5\text{THz}$

Per-space-time  $(\nu', c\kappa')$  geometry of 2-CW vectors

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$$B = \frac{\sqrt{5}}{2} = 1.12..$$

$$\begin{pmatrix} 1 \\ \frac{3}{2} \end{pmatrix} = \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix}$$

$$= \frac{1}{2}(\mathbf{R} + \mathbf{L}) = \mathbf{P}$$

$$\mathbf{G} = \frac{1}{2}(\mathbf{R} - \mathbf{L}) = \begin{pmatrix} \frac{3}{2} \\ 1 \end{pmatrix}$$

$$\mathbf{R} = B \begin{pmatrix} e^{\rho} \\ e^{\rho} \end{pmatrix}$$

$$= \begin{pmatrix} 2.5 \\ 2.5 \end{pmatrix}$$

$$\mathbf{L} = B \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix}$$

$$= \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix}$$

$c \cdot$  Wave Number  $c \cdot \kappa' = c \cdot k / 2\pi$   
(units:  $\nu_A = c \cdot \kappa_A$ )

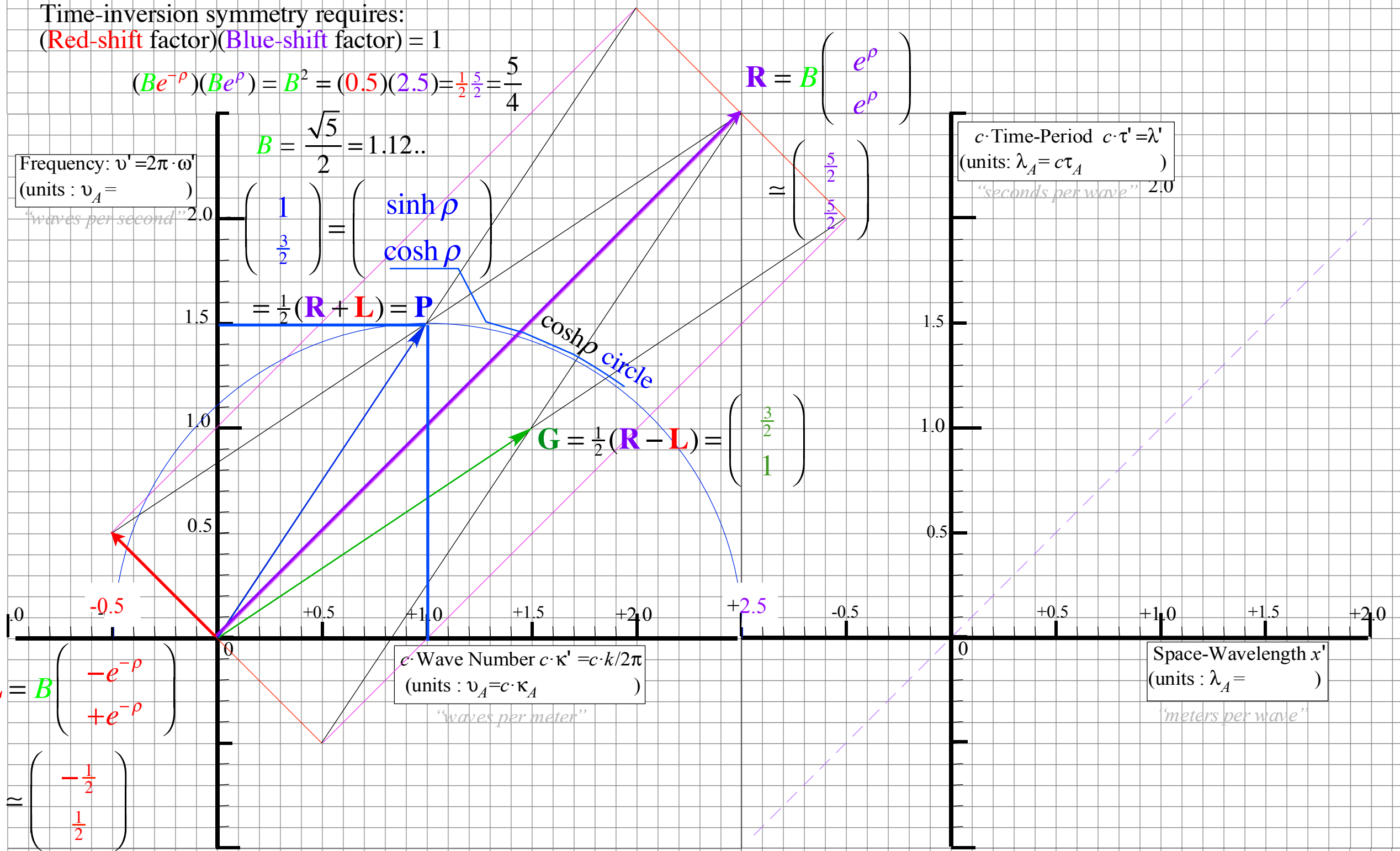
"waves per meter"

$c \cdot$  Time-Period  $c \cdot \tau' = \lambda'$   
(units:  $\lambda_A = c \tau_A$ )

"seconds per wave"  $\tau'$

Space-Wavelength  $x'$   
(units:  $\lambda_A =$ )

"meters per wave"



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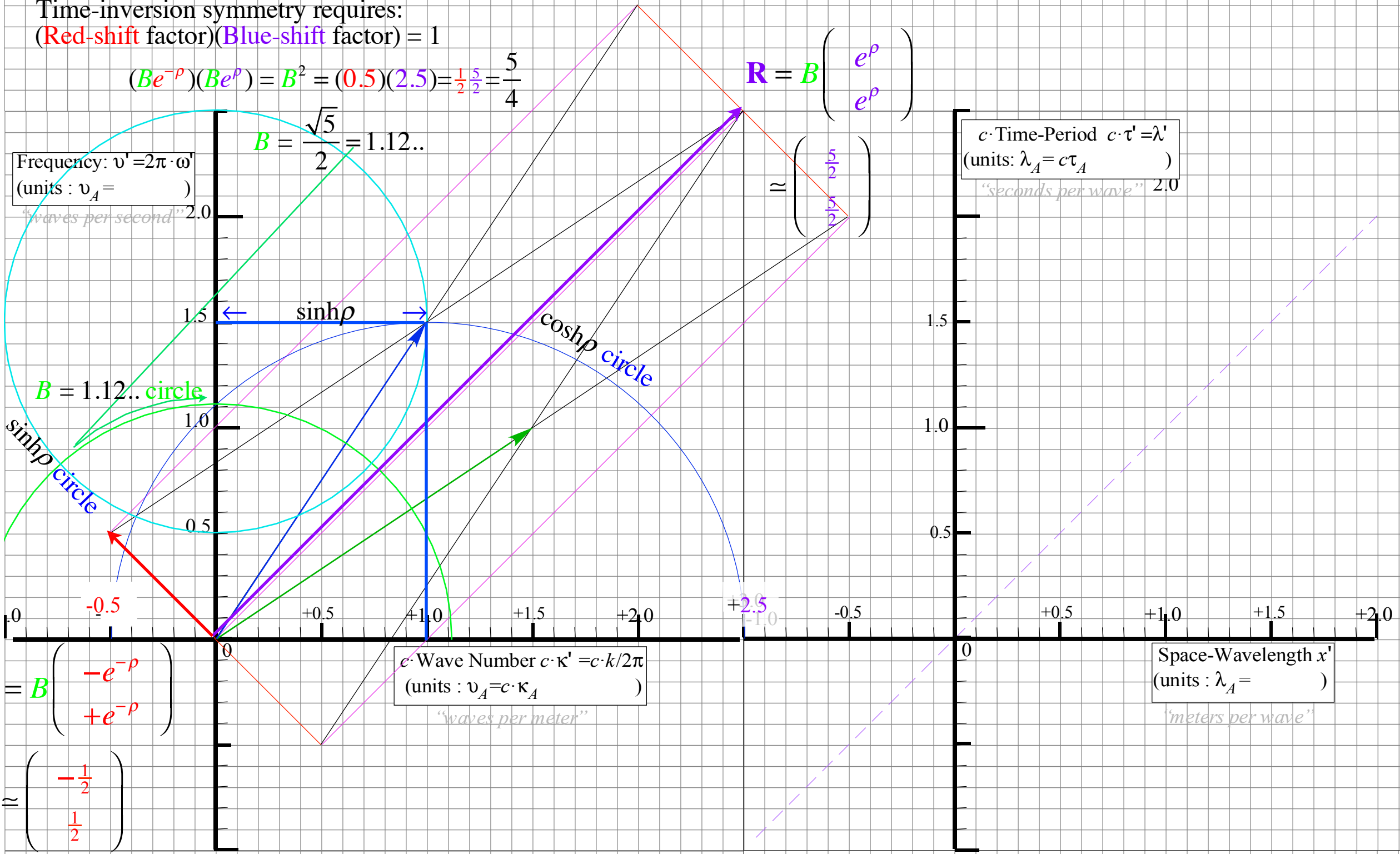
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Frequency:  $\nu' = 2\pi \cdot \omega'$   
 (units:  $\nu_A =$  )  
 "waves per second"

$c \cdot \text{Time-Period } c \cdot \tau' = \lambda'$   
 (units:  $\lambda_A = c\tau_A$  )  
 "seconds per wave" 2.0



$$L = B \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix}$$

$$\approx \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix}$$

$c \cdot \text{Wave Number } c \cdot \kappa' = c \cdot k / 2\pi$   
 (units:  $\nu_A = c \cdot \kappa_A$  )  
 "waves per meter"

Space-Wavelength  $x'$   
 (units:  $\lambda_A =$  )  
 "meters per wave"

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Frequency:  $\nu' = 2\pi \cdot \omega'$   
 (units:  $\nu_A =$  )  
 "waves per second"

$c \cdot$  Time-Period  $c \cdot \tau' = \lambda'$   
 (units:  $\lambda_A = c\tau_A$  )  
 "seconds per wave"

$c \cdot$  Wave Number  $c \cdot \kappa' = c \cdot k/2\pi$   
 (units:  $\nu_A = c \cdot \kappa_A$  )  
 "waves per meter"

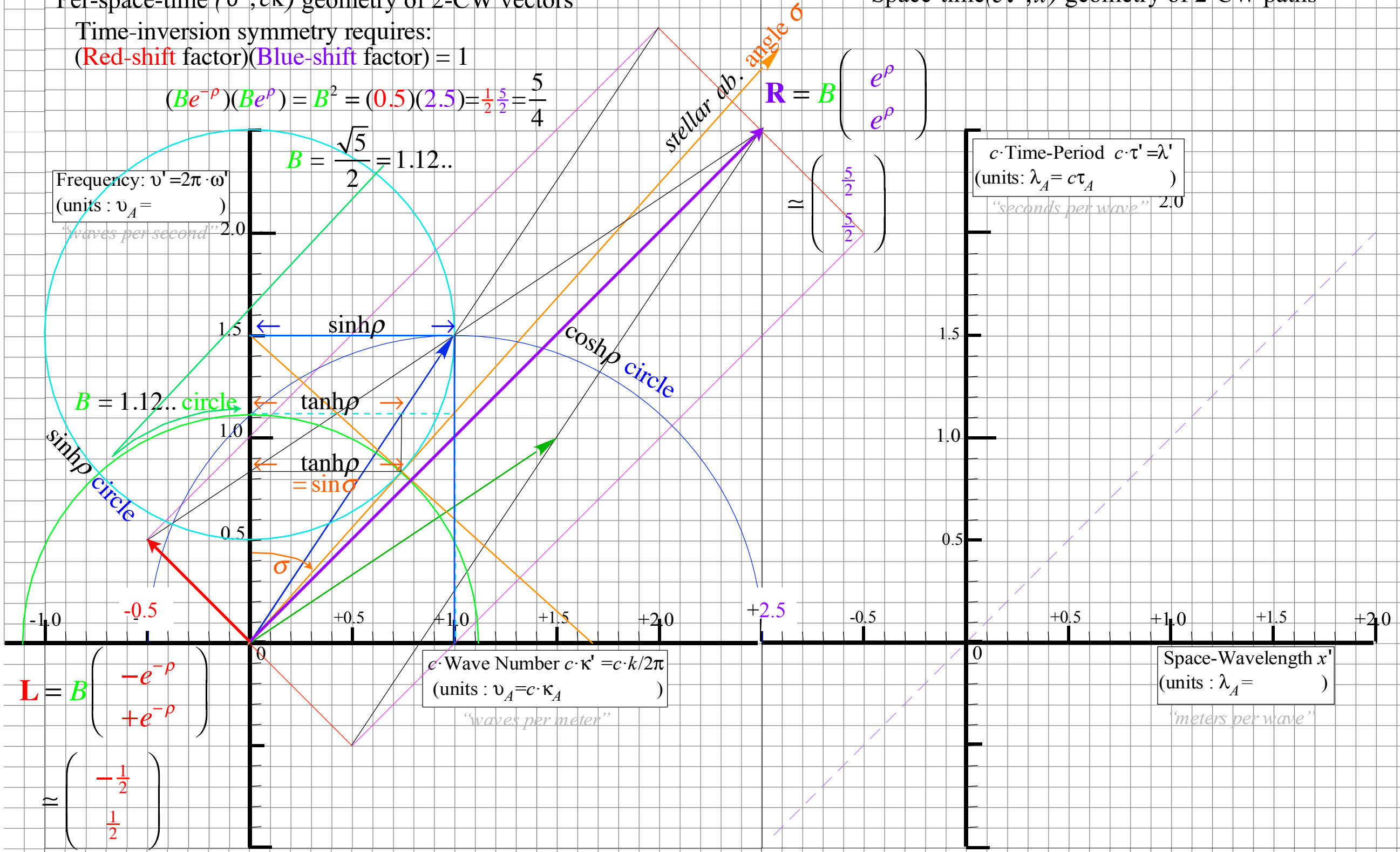
Space-Wavelength  $x'$   
 (units:  $\lambda_A =$  )  
 "meters per wave"

$$L = B \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix}$$

$$\approx \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix}$$

$$R = B \begin{pmatrix} e^{\rho} \\ e^{\rho} \end{pmatrix}$$

$$\approx \begin{pmatrix} \frac{5}{2} \\ \frac{5}{2} \end{pmatrix}$$



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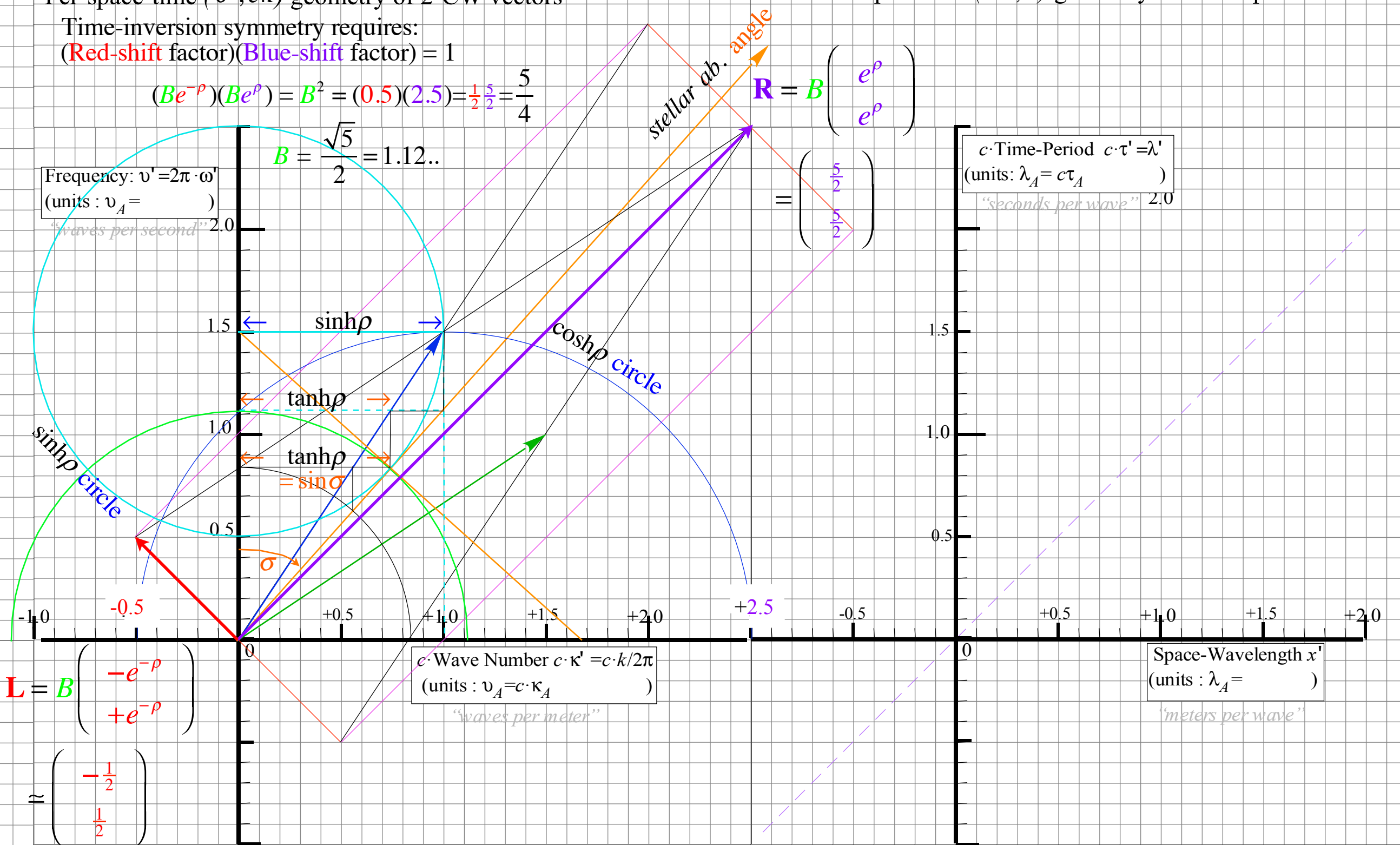
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$$B = \frac{\sqrt{5}}{2} = 1.12..$$

Frequency:  $\nu' = 2\pi \cdot \omega'$   
 (units:  $\nu_A =$  )  
 "waves per second"

Space-time  $(c\tau', x')$  geometry of 2-CW paths

$c \cdot \text{Time-Period } c \cdot \tau' = \lambda'$   
 (units:  $\lambda_A = c\tau_A$  )  
 "seconds per wave" 2.0





# Derive relativity parameters and Minkowski coordinates for $\nu_R=2.5\text{THz}$ and $\nu_L=0.5\text{THz}$

Per-space-time  $(\nu', c\kappa')$  geometry of 2-CW vectors

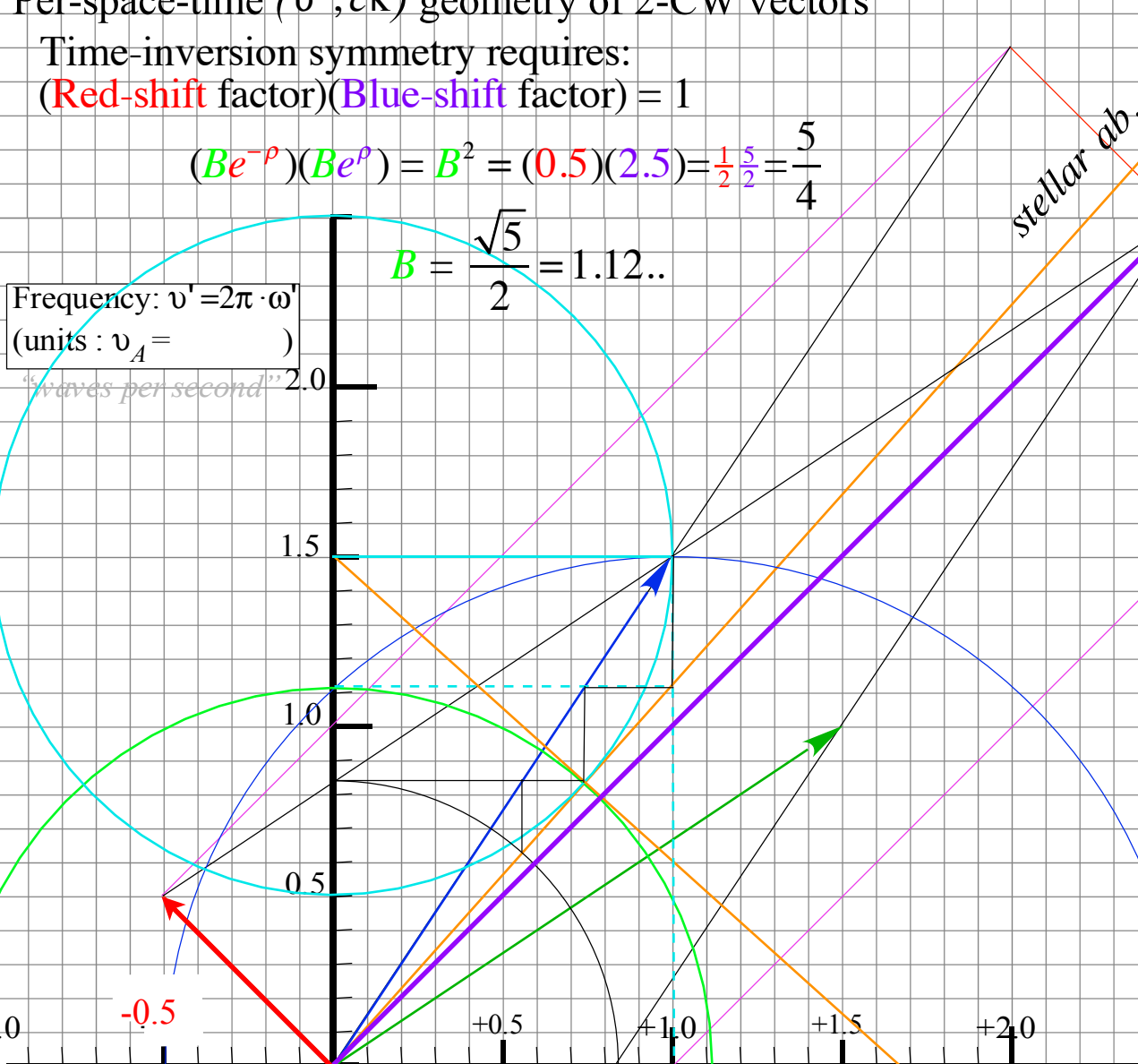
Time-inversion symmetry requires:

$$(\text{Red-shift factor})(\text{Blue-shift factor}) = 1$$

$$(Be^{-\rho})(Be^{\rho}) = B^2 = (0.5)(2.5) = \frac{1}{2} \cdot \frac{5}{2} = \frac{5}{4}$$

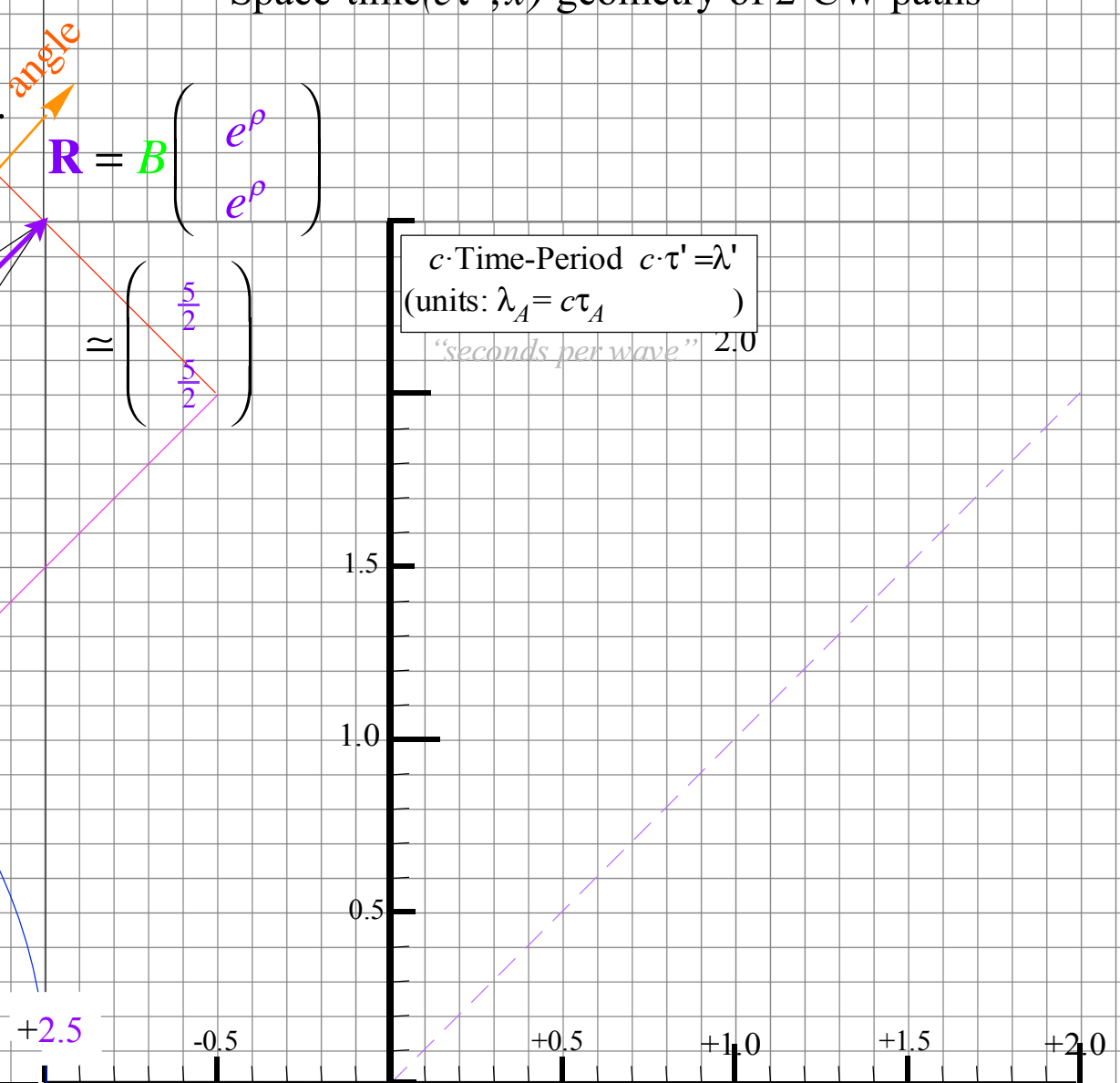
$$B = \frac{\sqrt{5}}{2} = 1.12..$$

Frequency:  $\nu' = 2\pi \cdot \omega'$   
(units:  $\nu_A = \text{waves per second}$ )



Space-time  $(c\tau', x')$  geometry of 2-CW paths

$c \cdot \text{Time-Period } c \cdot \tau' = \lambda'$   
(units:  $\lambda_A = c \tau_A$ )  
"seconds per wave" 2.0



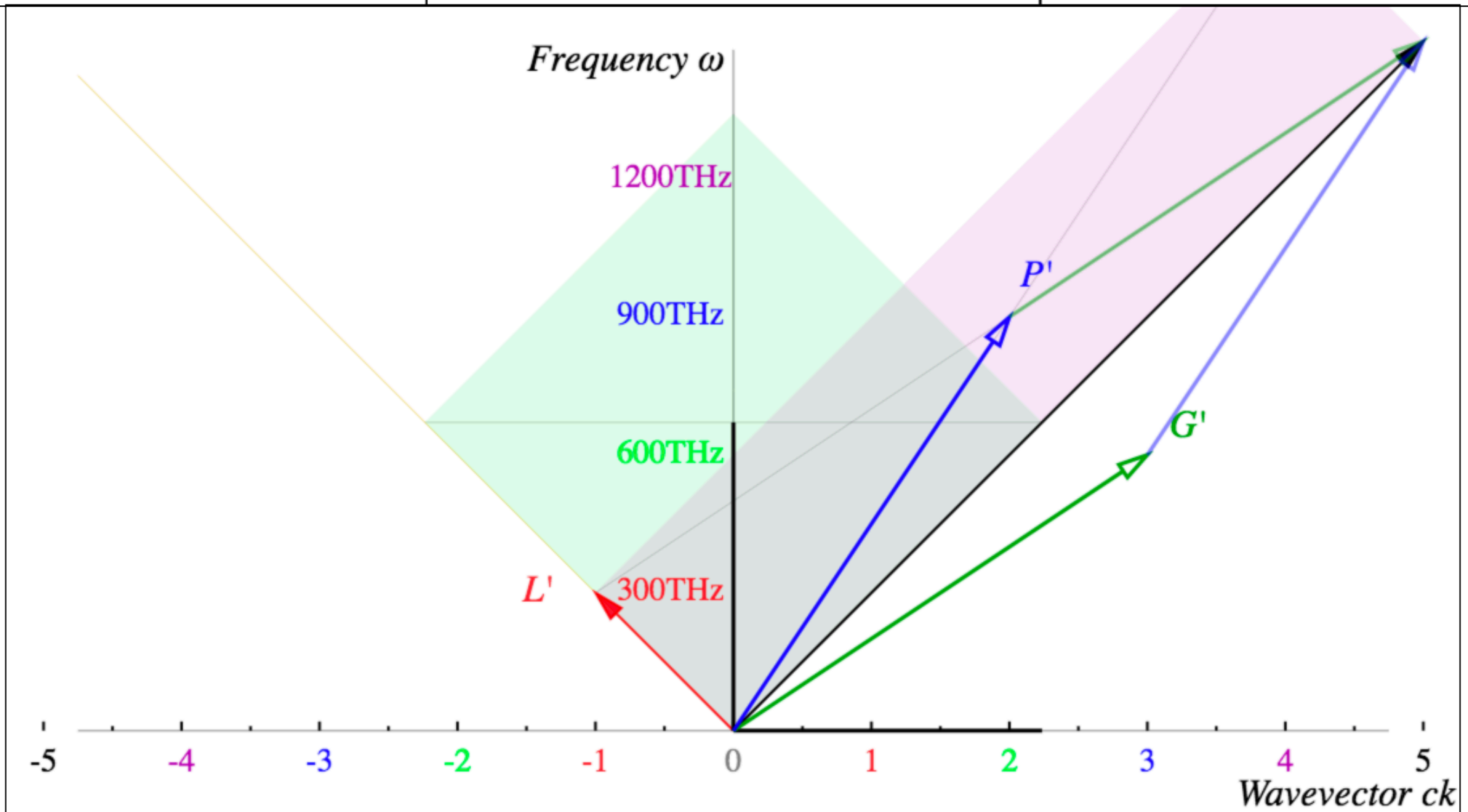
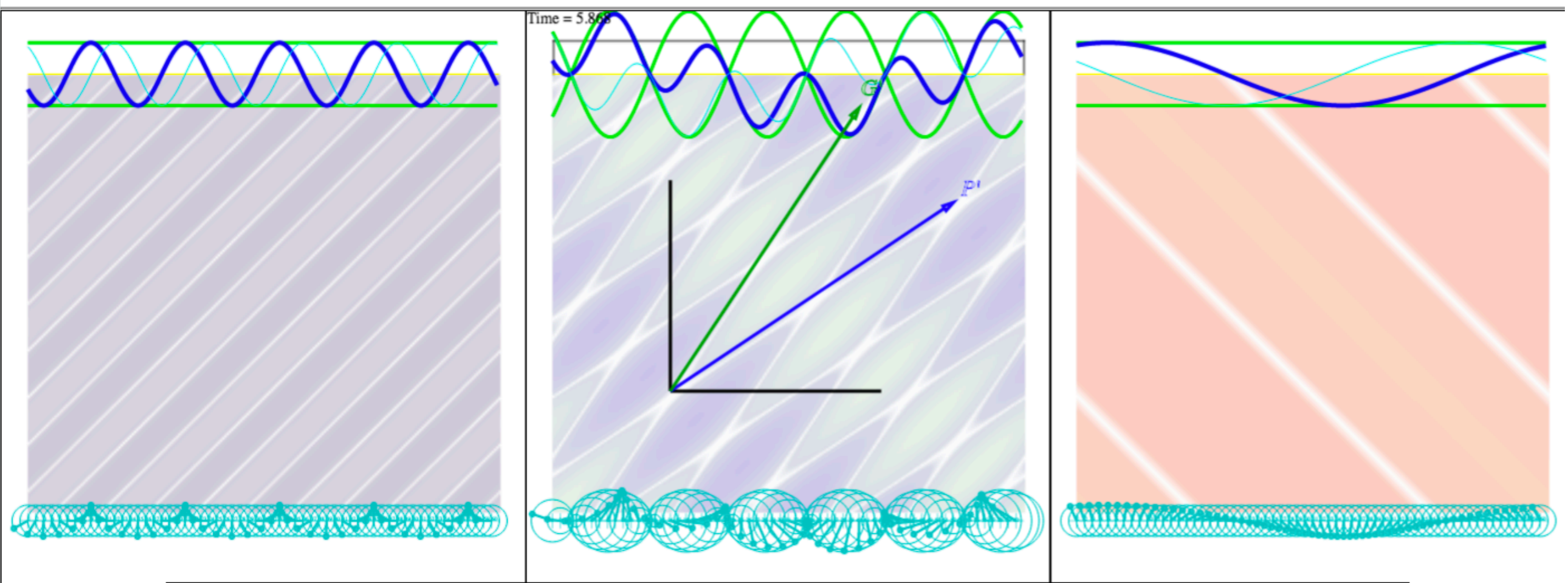
$$L = B \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix}$$

$$\approx \begin{pmatrix} -\frac{1}{2} \\ \frac{1}{2} \end{pmatrix}$$

$c \cdot \text{Wave Number } c \cdot \kappa' = c \cdot k/2\pi$   
(units:  $\nu_A = c \cdot \kappa_A$ )  
"waves per meter"

group	$b_{\text{Doppler RED}}$	$\frac{V_{\text{group}}}{c}$	$\frac{\nu_{\text{group}}}{\nu_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{\kappa_{\text{group}}}{\kappa_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{c}{V_{\text{group}}}$	$b_{\text{Doppler BLUE}}$
phase	$\frac{1}{b_{\text{Doppler BLUE}}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{\kappa_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{\nu_{\text{phase}}}{\nu_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{V_{\text{phase}}}{c}$	$\frac{1}{b_{\text{Doppler RED}}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\text{csc } \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{2}{5}=0.4$	$\frac{2}{3}=0.67$	$\frac{1}{1}=1.0$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{3}{2}=1.5$	$\frac{5}{2}=2.5$

(may contain errors)



This map has circle sector arc-area  $\sigma = 0.6435$

set to angle  $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

$$\begin{aligned} \sin(\sigma) &= 0.6000 &= \tanh(\rho) &= 3/5 \\ \tan(\sigma) &= 0.7500 &= \sinh(\rho) &= 3/4 \\ \sec(\sigma) &= 1.2500 &= \cosh(\rho) &= 5/4 \\ \cos(\sigma) &= 0.8000 &= \operatorname{sech}(\rho) &= 4/5 \\ \cot(\sigma) &= 1.3333 &= \operatorname{csch}(\rho) &= 4/3 \\ \csc(\sigma) &= 1.6667 &= \operatorname{coth}(\rho) &= 5/3 \end{aligned}$$

$$\cosh(\rho) + \sinh(\rho) = \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho}$$

$$\cosh(\rho) - \sinh(\rho) = \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho}$$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2} \quad \text{Half-Sum-}$$

$$\sinh(\rho) = \frac{e^{+\rho} - e^{-\rho}}{2} \quad \text{Half-Difference}$$

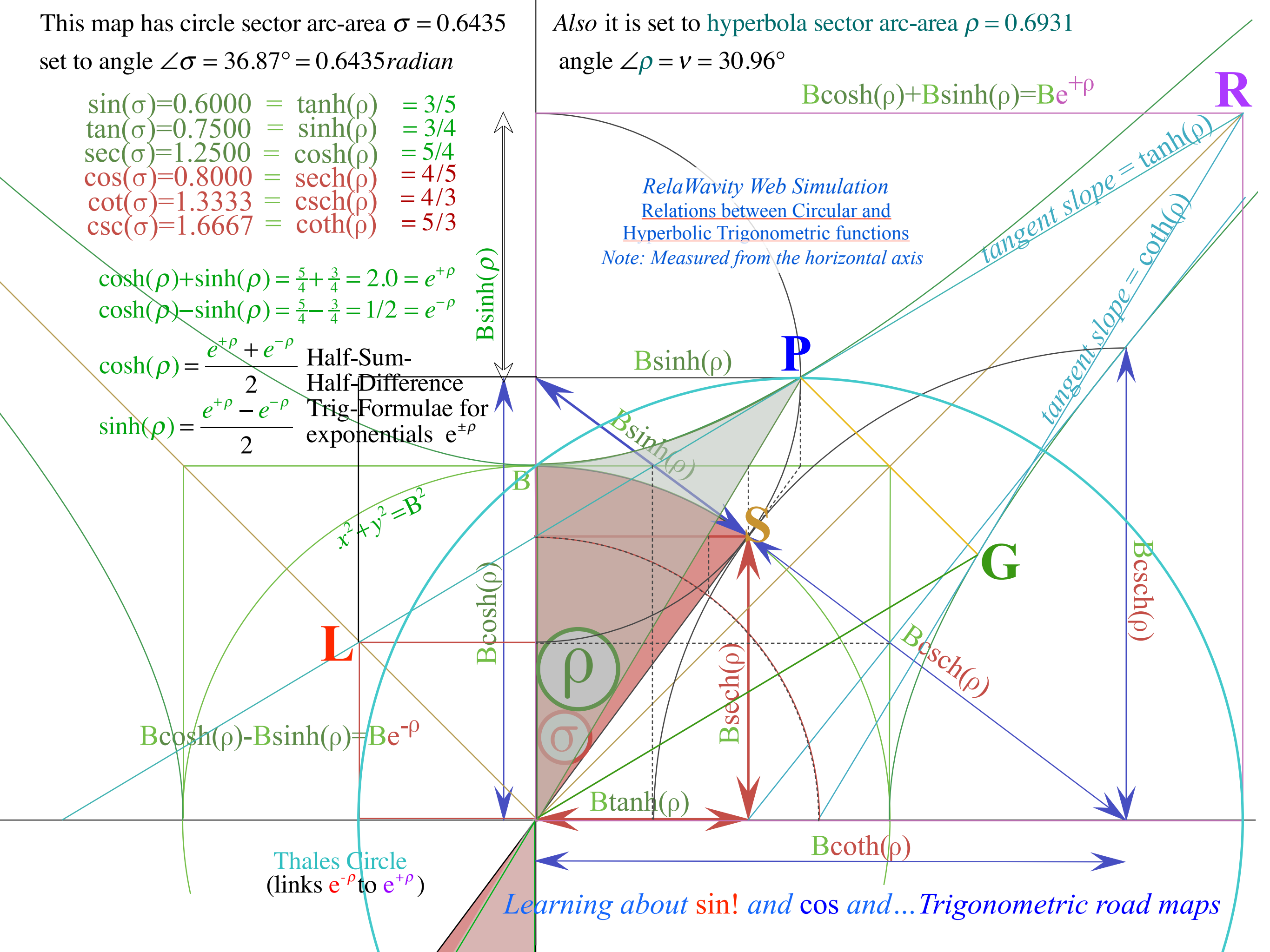
Trig-Formulae for  
exponentials  $e^{\pm\rho}$

Also it is set to hyperbola sector arc-area  $\rho = 0.6931$

angle  $\angle\rho = \nu = 30.96^\circ$

$$B\cosh(\rho) + B\sinh(\rho) = B e^{+\rho}$$

*RelaWavity Web Simulation*  
Relations between Circular and  
Hyperbolic Trigonometric functions  
*Note: Measured from the horizontal axis*



$$B\cosh(\rho) - B\sinh(\rho) = B e^{-\rho}$$

Thales Circle  
(links  $e^{-\rho}$  to  $e^{+\rho}$ )

Learning about **sin!** and **cos** and... Trigonometric road maps



Fig. 11 in text Relativity...

(a) Per-space-time  $(\nu', c\kappa')$  geometry of 2-CW vectors

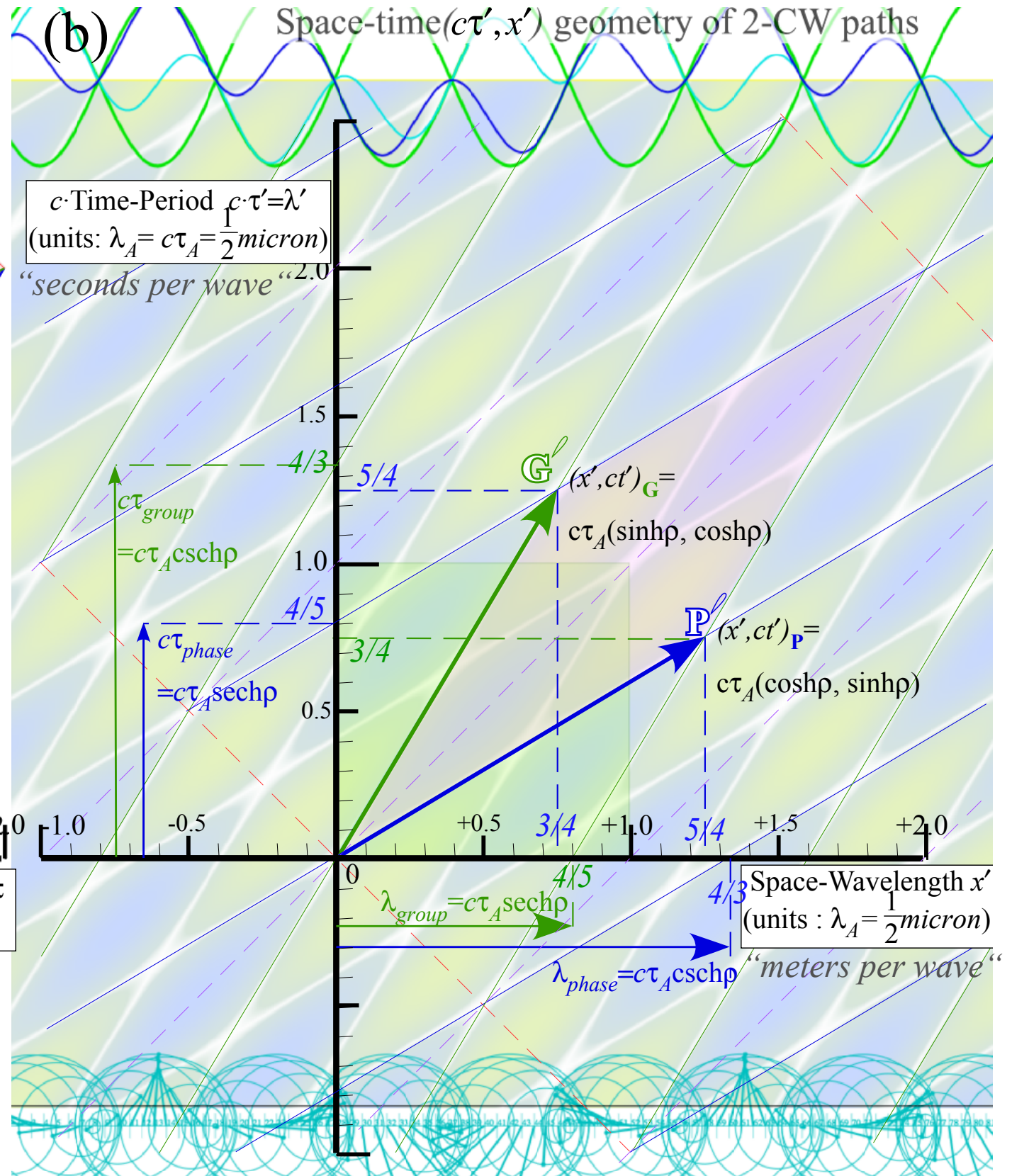
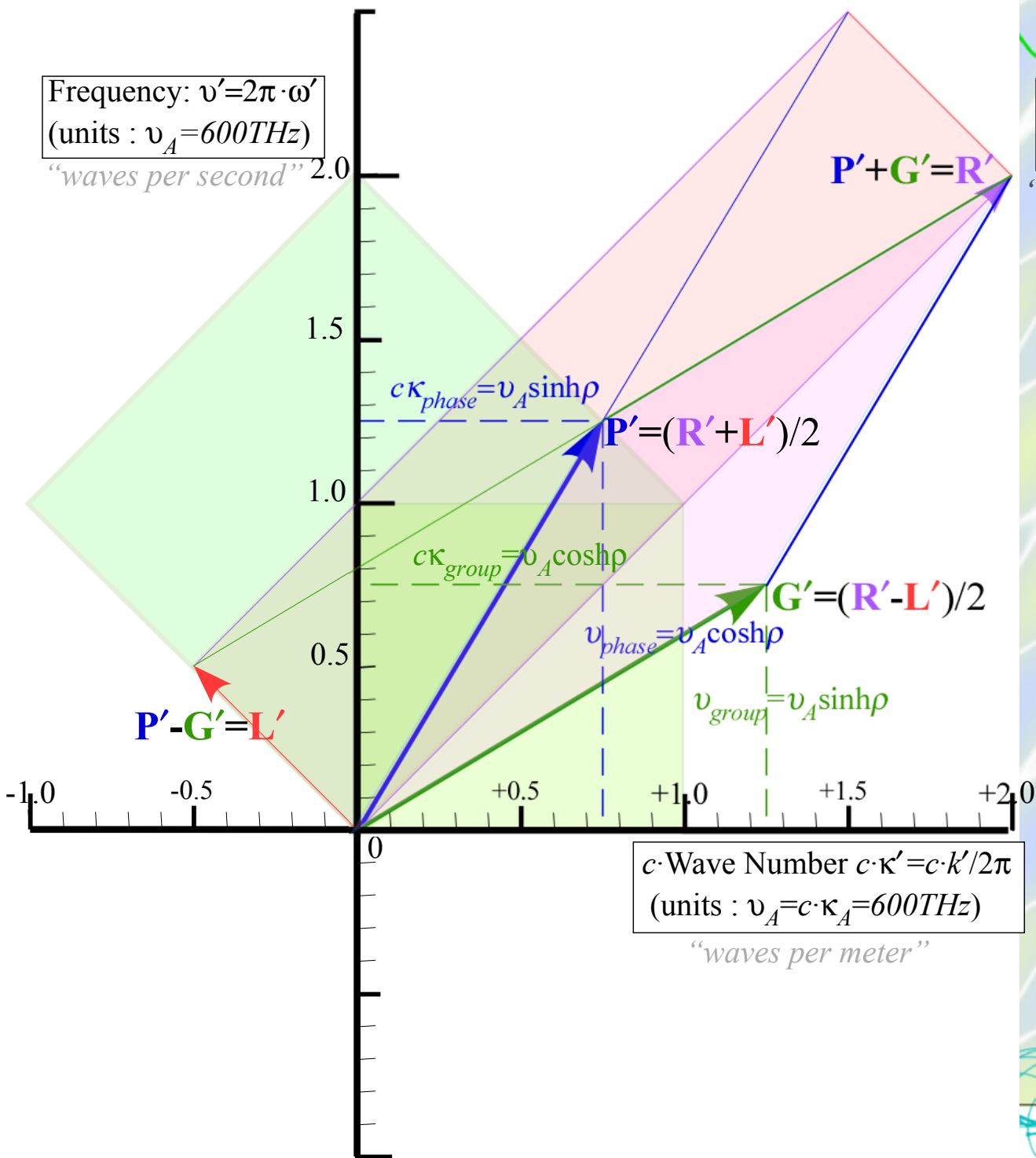
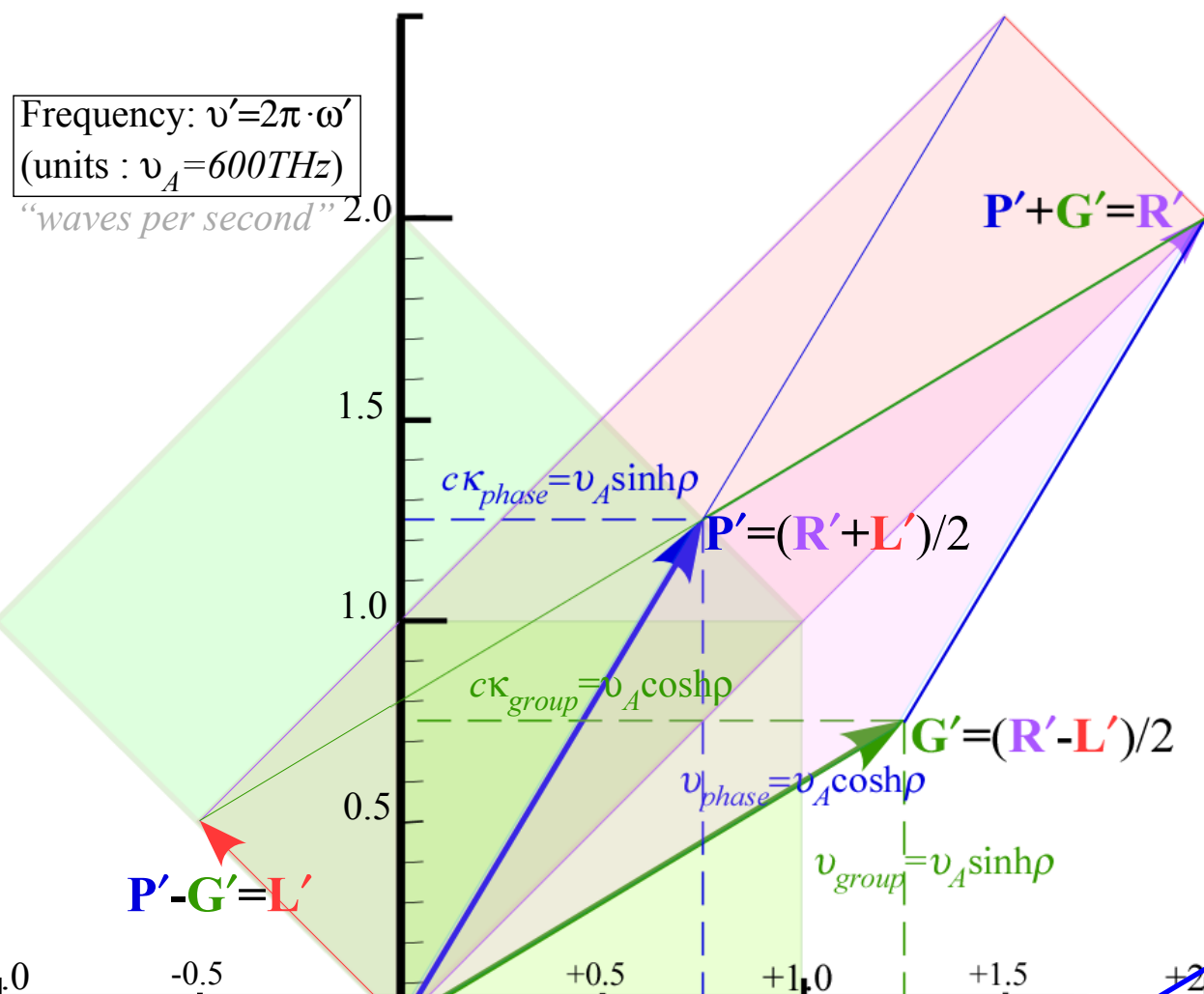


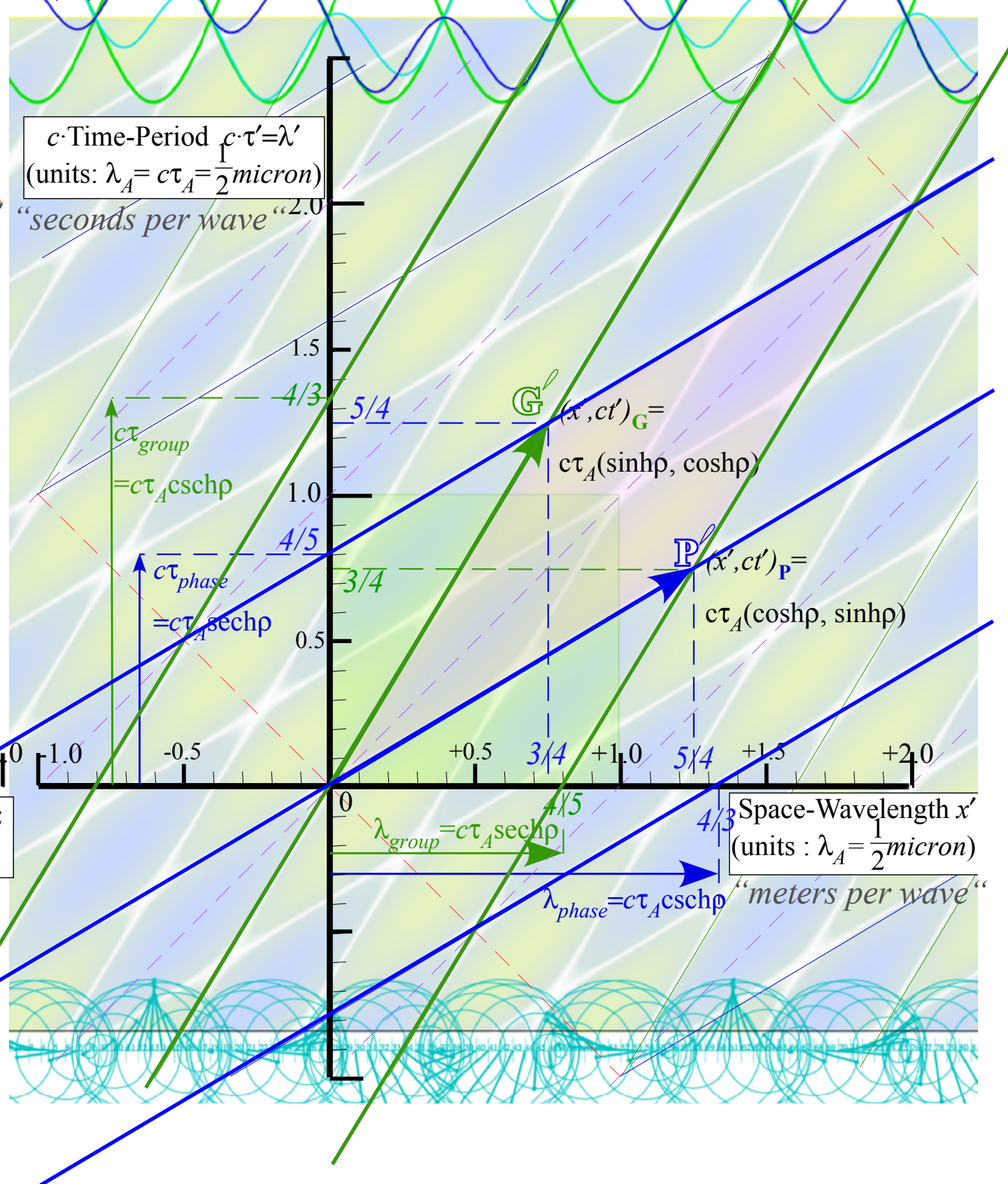


Fig. 4 in Ch.0 text introducing Relativity..

(a) Per-space-time ( $v', c\kappa'$ ) geometry of 2-CW vectors



(b) Space-time ( $c\tau', x'$ ) geometry of 2-CW paths



group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

# Lecture 30

## Wed. 12.05.2018

Review: Relativity  $\rho$  functions and plots vs.  $\rho$

Derive relativity parameters and Minkowski coordinates for  $\nu_R=2.5\text{THz}$  and  $\nu_L=0.5\text{THz}$

*Rapidity  $\rho$  related to stellar aberration angle  $\sigma$  and L. C. Epstein's approach to relativity*

*Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry*

*"Occams Sword" and summary of 16 parameter functions of  $\rho$  and  $\sigma$*

*Applications to optical waveguide, spherical waves, and accelerator radiation*

➔ Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

Relativistic action and Lagrangian-Hamiltonian relations

Poincaré' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization "photon" number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity in accelerated frames*

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c \text{)}$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \text{ (for } u \ll c \text{)}$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

At low speeds:

<i>group</i>	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
<i>phase</i>	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\forall$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

[RelaWavity Web Simulation - Relativistic Terms \(Short version\)](#)

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

At low speeds:

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

At low speeds:  $\Leftarrow$  for  $(u \ll c)$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$



# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

At low speeds:  
 $\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

time	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
space	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:  
 $\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy and momentum

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

# Using (some) wave parameters to develop relativistic quantum theory

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$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

So attach scale factor  $h$  to match units.

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{Doppler RED}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

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$$B = v_A$$

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At low speeds:

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$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor  $h$  to match units.

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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So attach scale factor  $h$  to match units.

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Lucky coincidences?? Cheap trick??  
...Try exact  $v_{phase}$  ...

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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[RelaWavity Web Simulation - Relativistic Terms \(Short version\)](#)



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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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(old-fashioned notation)

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Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

Einstein (1905)

(old-fashioned notation)

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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Max Planck  
1858-1947

# Using (some) wave parameters to develop relativistic quantum theory

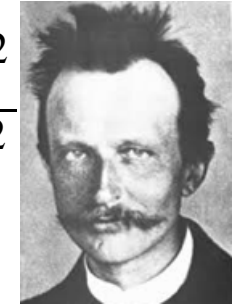
$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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Max Planck  
1858-1947

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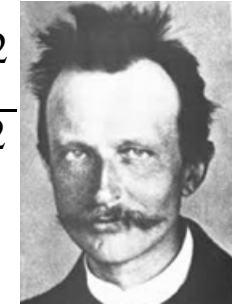
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# Using (some) wave parameters to develop relativistic quantum theory



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For more, visit the Pirelli Challenge Site  
[Quantized amplitude](#)

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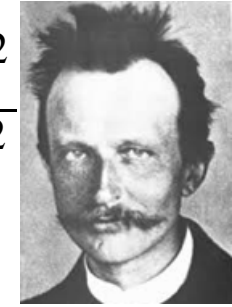
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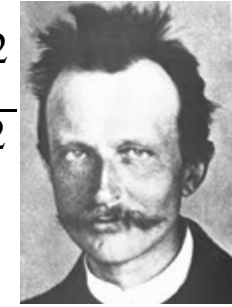
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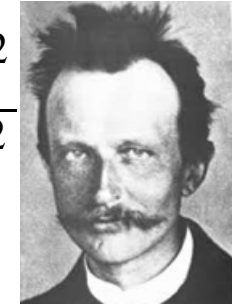
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Resolution and dirty secret:  $\mathbf{E}$ ,  $N$ , and  $v_{phase}$  are all frequencies!

So  $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hN v_{phase}$  is quadratic in  $v_{phase}$

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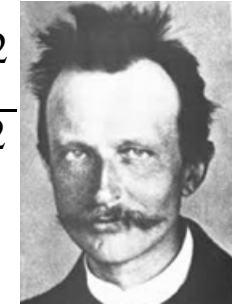
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Einstein (1905)

$$hK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}} \quad (\text{old-fashioned notation})$$

$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

$$\text{Momentum: } hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

DeBroglie (1921)

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
phase	$b_{Doppler BLUE}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$



Max Planck 1858-1947      Louis DeBroglie 1892-1987

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

for  $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$  or:  $hB = Mc^2$  (The famous  $Mc^2$  shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor  $h$  (or  $hN$ ) to match units.

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2$$

for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx Mu$$

~~Natural wave conspiracy~~  
~~Lucky coincidences??~~ ~~Expensive Cheap trick??~~  
...Try exact  $v_{phase}$  and  $K_{phase}$ ...

Need to replace  $h$  with  $hN$  to match e.m. energy density  $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hN v_{phase}$

$$hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

Einstein (1905)

This motivates the "particle" normalization  $\int \Psi^* \Psi dV = N$   $\Psi = \sqrt{\frac{\epsilon_0}{h\nu}} E$

$$h c K_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

group	$b_{\text{Doppler RED}}$	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{K_{\text{group}}}{K_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{V_{\text{phase}}}{c}$
phase	$b_{\text{Doppler BLUE}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{K_{\text{phase}}}{K_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$		
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$		
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$

$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}} \quad (\text{old-fashioned notation})$$

$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

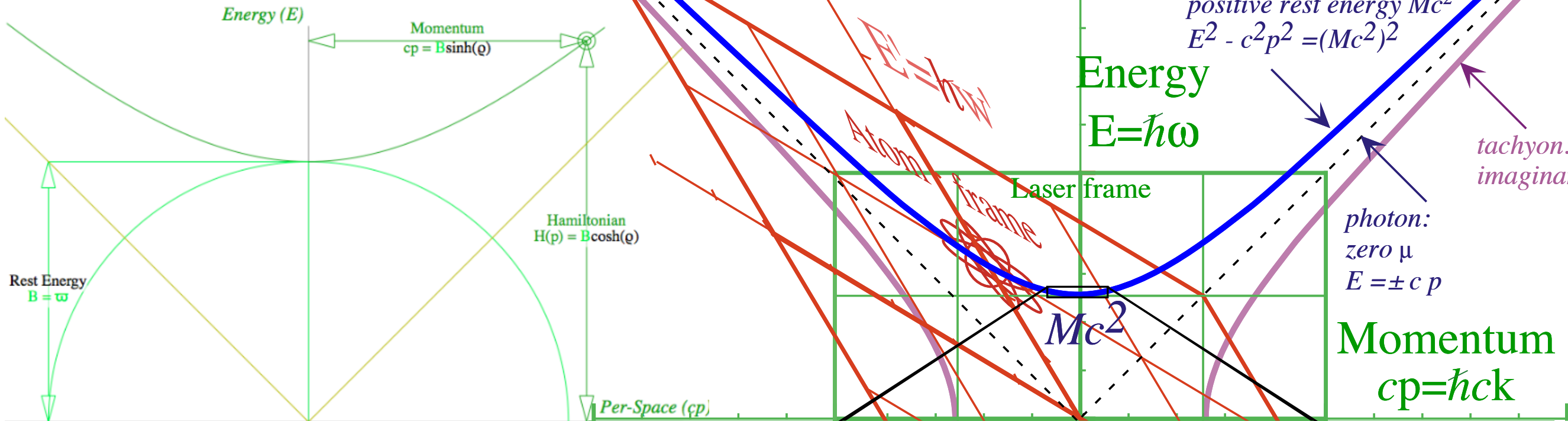
Momentum:  $hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$

DeBroglie (1921)



# Using (some) wave coordinates for relativistic quantum theory

## (a) Exact Einstein-Planck Dispersion



Mass (resting)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Energy

$$h\nu_{\text{phase}} = E = h\nu_A \cosh \rho$$

Momentum

$$hc\kappa_{\text{phase}} = cp = hc\kappa_A \sinh \rho = h\nu_A \sinh \rho$$

Energy versus Momentum

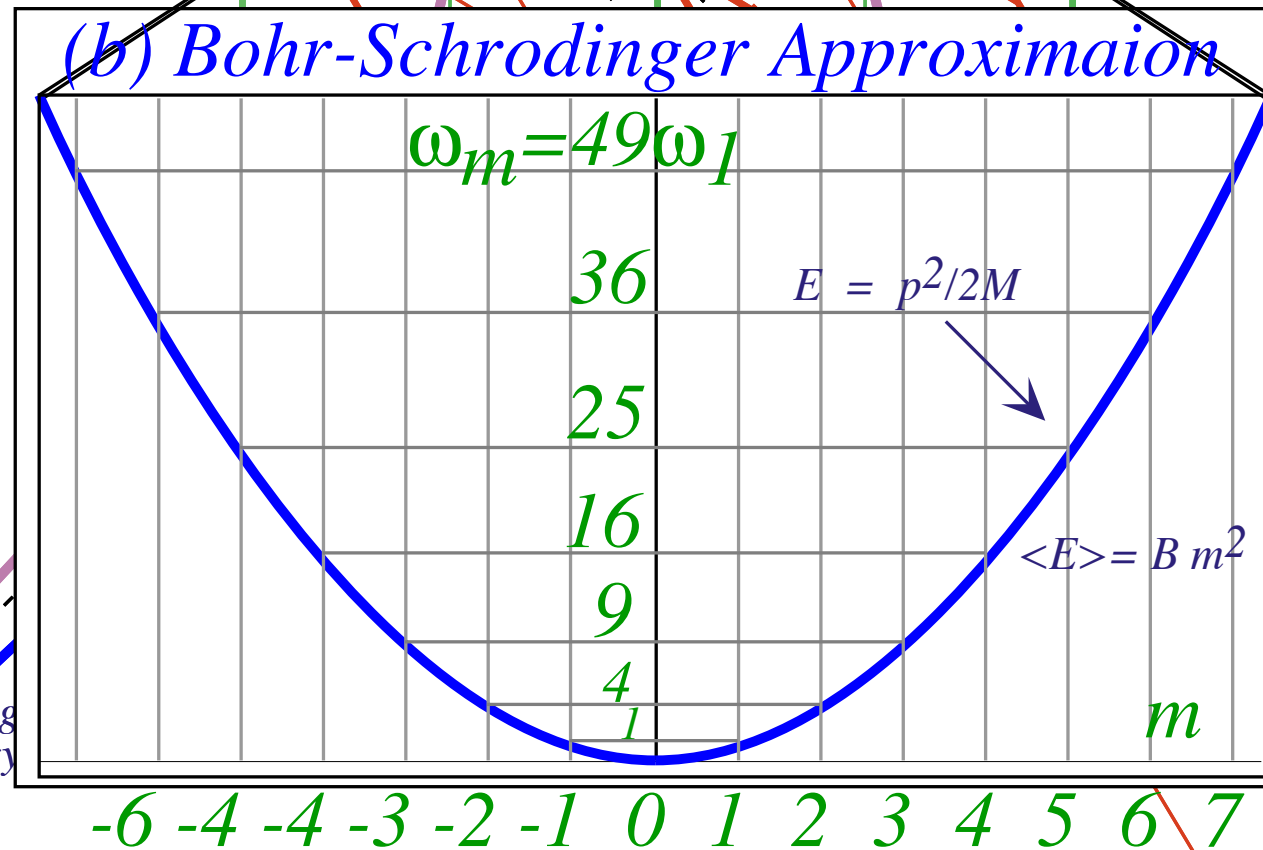
$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$



Neils Bohr  
1885-1962

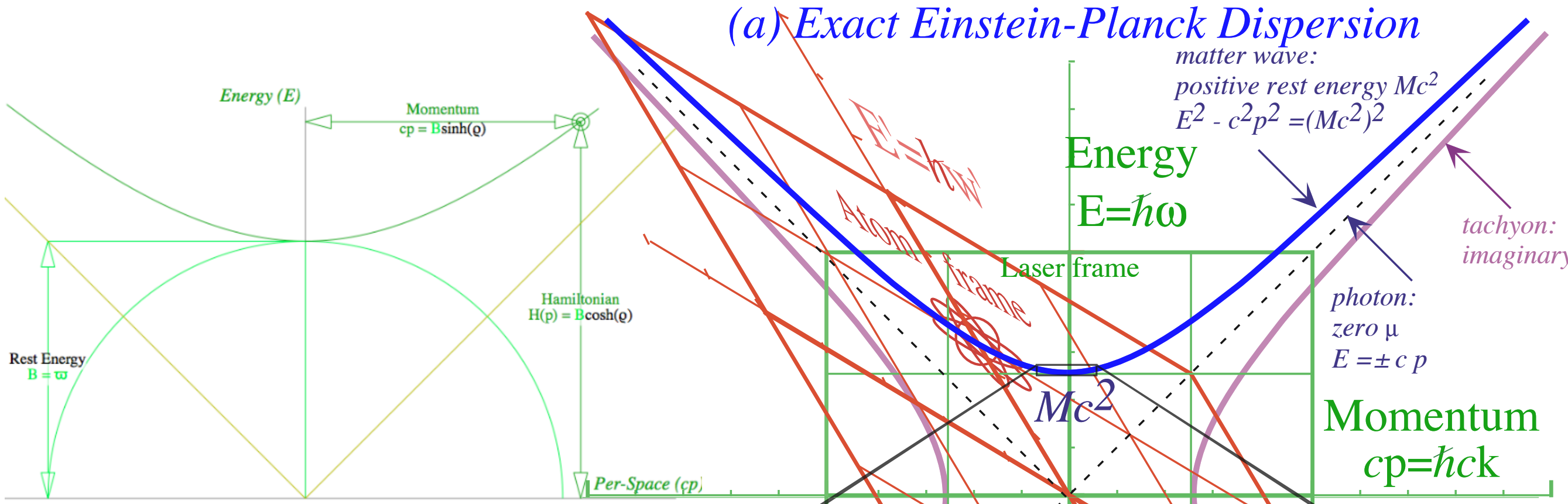
## (b) Bohr-Schrodinger Approximaion



Erwin Schrodinger  
1887-1961



# Using (some) wave coordinates for relativistic quantum theory



Mass (resting)  
 $hB = \hbar\omega_A = Mc^2 = \hbar ck_A$

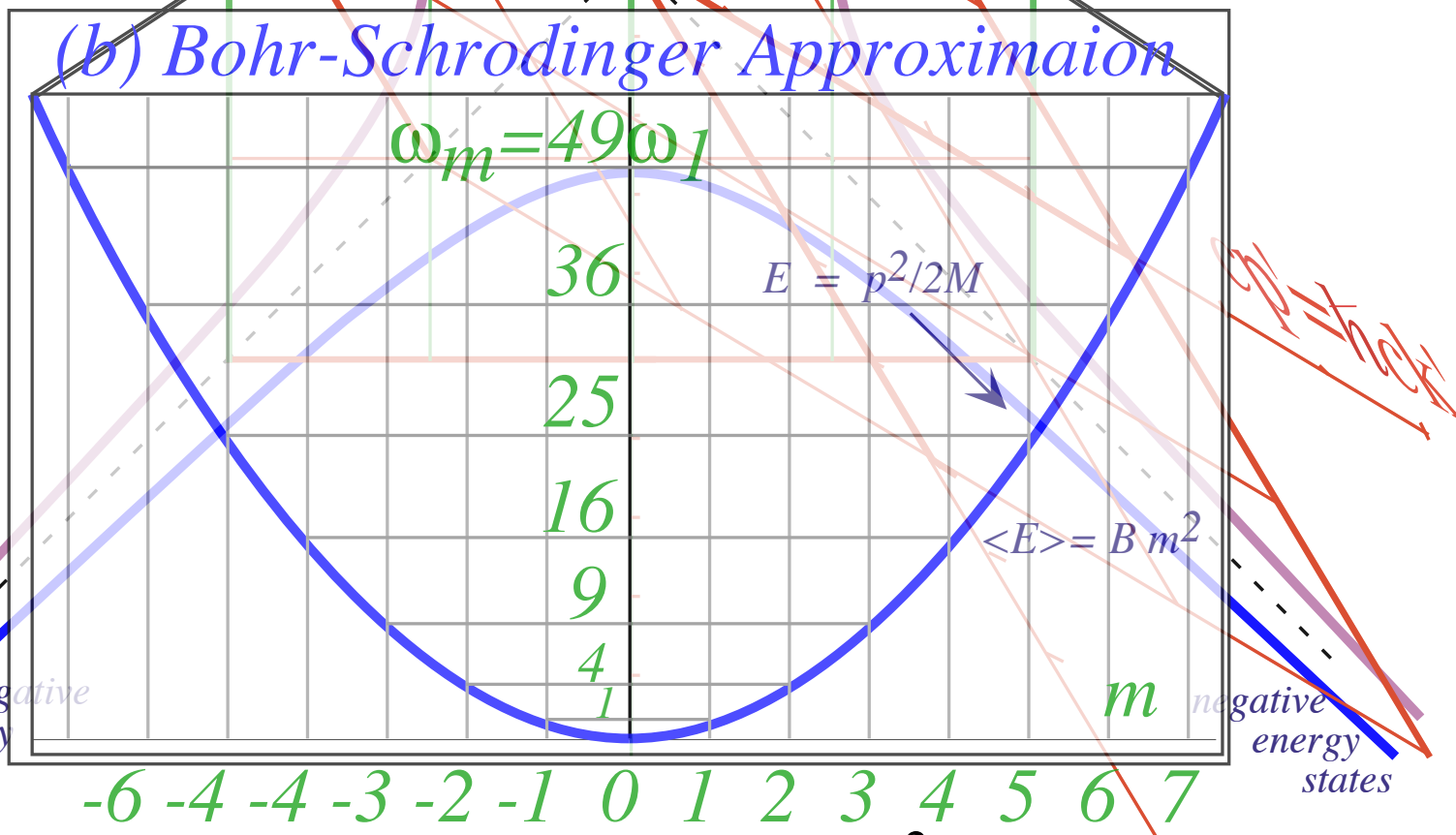
Energy  
 $\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$

Momentum  
 $\hbar ck_{phase} = cp = \hbar ck_A \sinh \rho = \hbar\omega_A \sinh \rho$

Energy versus Momentum  
 $E^2 = (Mc^2)^2 \cosh^2 \rho$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$

low speed approximation



## Relativity variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\nabla$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\operatorname{csc} \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>effects</i>	$b_{RED}^{Doppler}$	$V_{group}$	<i>past-future asymmetry</i> (off-diagonal Lorentz-transform)	<i>x-contraction</i> (Lorentz.) $\tau_{phase}$ -contraction	<i>t-dilation</i> (Einstein) $v_{phase}$ -dilation (on-diagonal Lorentz-transform)	<i>inverse asymmetry</i>	$V_{phase}$	$b_{BLUE}^{Doppler}$

## Relativistic quantum mechanics variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\nabla$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\operatorname{csc} \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{\beta}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{\beta}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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<i>functions</i>		$V_{group} = c \tanh \rho$	<i>momentum</i> $cp = Mc^2 \sinh \rho$	<i>-Lagrangian</i> $L = -Mc^2 \operatorname{sech} \rho$	<i>Hamiltonian</i> $H = Mc^2 \cosh \rho$	<i>DeBroglie</i> $\lambda = \alpha \operatorname{csch} \rho$	$V_{phase} = c \operatorname{coth} \rho$	

# Lecture 30

## Wed. 12.05.2018

Review: Relativity  $\rho$  functions and plots vs.  $\rho$

*Rapidity  $\rho$  related to stellar aberration angle  $\sigma$  and L. C. Epstein's approach to relativity*  
*Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry*  
*"Occams Sword" and summary of 16 parameter functions of  $\rho$  and  $\sigma$*   
*Applications to optical waveguide, spherical waves, and accelerator radiation*

Derivation of relativistic quantum mechanics

- ➔ What's the matter with mass? Shining some light on the Elephant in the room
- Relativistic action and Lagrangian-Hamiltonian relations
- Poincaré' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization "photon" number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity in accelerated frames*

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

$$= h\nu_{\text{phase}}$$

momentum:  $cp = Mc^2 \sinh \rho$

$$= hc\kappa_{\text{phase}}$$

velocity:  $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

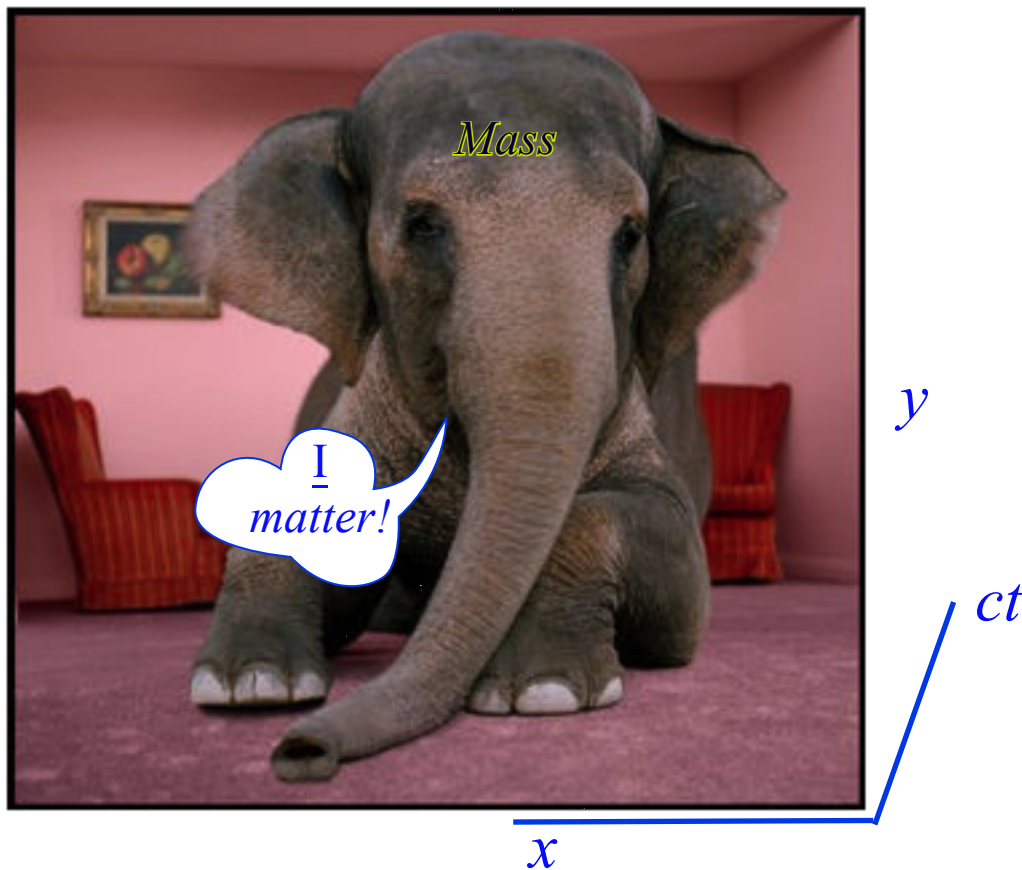
Rest Mass  $M_{\text{rest}}$  (Einstein's mass)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

- *What's the matter with Mass?*



*Shining some light on the elephant in the spacetime room*



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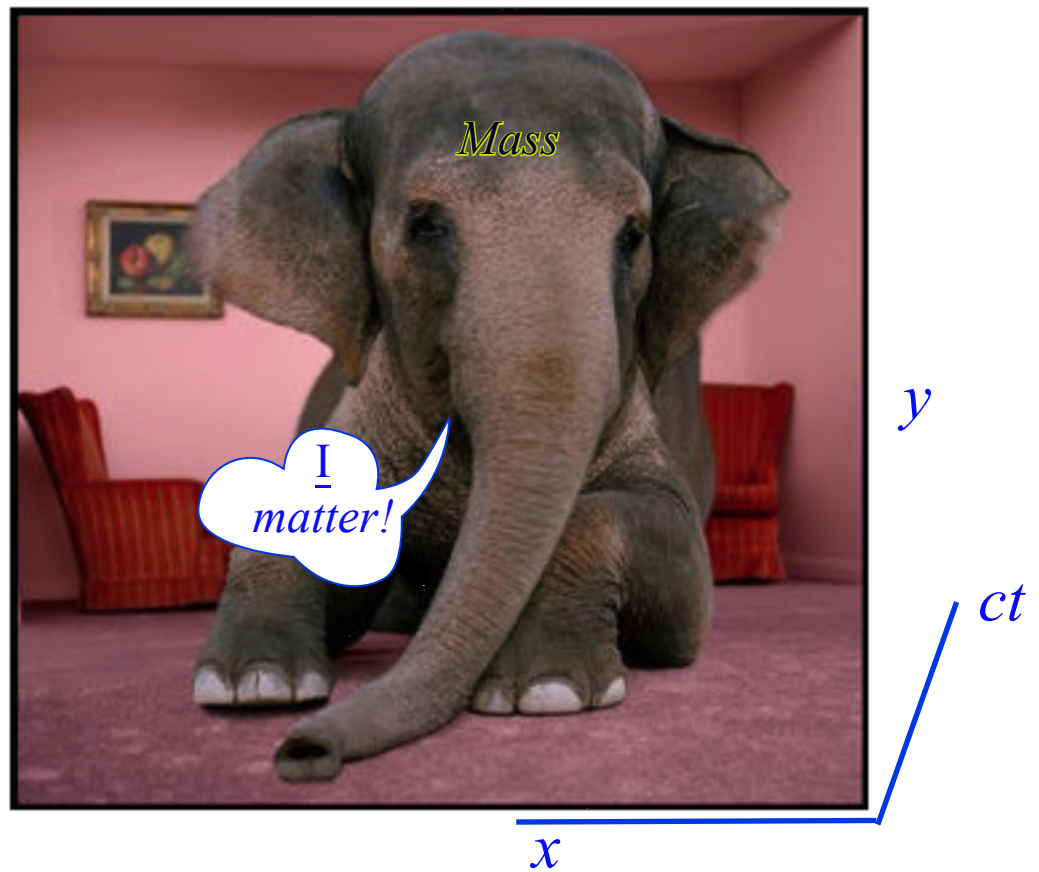
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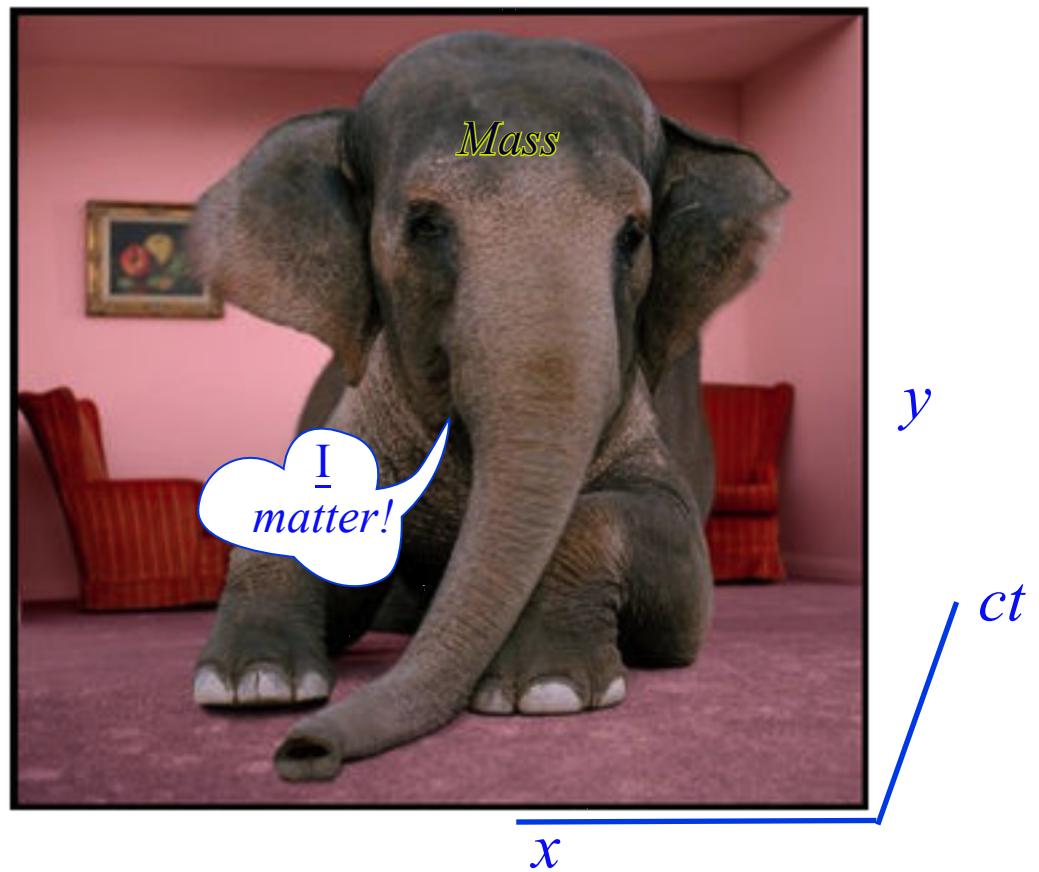
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- *What's the matter with Mass?*



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Rest Mass  $M_{rest}$  (Einstein's mass)

$$h\nu_A = hc\kappa_A = Mc^2$$

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Rest  
Mass

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$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \quad \text{Momentum Mass}$$

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Defines invariant hyperbola(s)

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Mass

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Limiting cases:  $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^\rho / 2$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}} \quad \text{Momentum  
Mass}$$



# Definition(s) of mass for relativity/quantum

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Mass}$$

Effective Mass  $M_{eff}$  (Newton's mass) Defined by ratio  $F/a = dp/du$  of relativistic force to acceleration.

# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

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# Definition(s) of mass for relativity/quantum

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$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho$$

# Definition(s) of mass for relativity/quantum

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$$= h\nu_{phase}$$

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$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho \quad \text{Effective Mass}$$

Limiting cases:  $M_{eff} \xrightarrow{u \rightarrow c} M_{rest} e^{3\rho/2}$

$$M_{eff} \xrightarrow{u \ll c} M_{rest}$$



# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

$$= h\nu_{phase}$$

Rest Mass  $M_{rest}$  (Einstein's mass)

Defines invariant hyperbola(s)

momentum:  $cp = Mc^2 \sinh \rho$

$$h\mathbf{B} = h\nu_{\mathbf{A}} = Mc^2 = hc\mathbf{K}_{\mathbf{A}}$$

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

$$= hc\mathbf{K}_{phase}$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} \quad \begin{array}{l} \text{Rest} \\ \text{Mass} \end{array}$$

velocity:  $u = c \tanh \rho = \frac{d\nu}{d\mathbf{K}}$

Momentum Mass  $M_{mom}$  (Galileo's mass) Defined by ratio  $p/u$  of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Limiting cases:  $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho/2}$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \quad \begin{array}{l} \text{Momentum} \\ \text{Mass} \end{array}$$

Effective Mass  $M_{eff}$  (Newton's mass) Defined by ratio  $F/a = dp/du$  of relativistic force to acceleration.

That is ratio of change  $dp = Mc \cosh \rho d\rho$  in momentum to change  $du = c \operatorname{sech}^2 \rho d\rho$  in group velocity.

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho \quad \begin{array}{l} \text{Effective} \\ \text{Mass} \end{array}$$

Limiting cases:  $M_{eff} \xrightarrow{u \rightarrow c} M_{rest} e^{3\rho/2}$

$$M_{eff} \xrightarrow{u \ll c} M_{rest}$$

More common derivation using group velocity:  $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\nu}{d\mathbf{K}}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}}$$

# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

$$= h\nu_{phase}$$

Rest Mass  $M_{rest}$  (Einstein's mass)

Defines invariant hyperbola(s)

momentum:  $cp = Mc^2 \sinh \rho$

$$h\mathbf{B} = h\nu_{\mathbf{A}} = Mc^2 = hc\mathbf{K}_{\mathbf{A}}$$

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

$$= hc\mathbf{K}_{phase}$$

Group velocity:  $u = c \tanh \rho = \frac{d\nu}{d\mathbf{K}}$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} \quad \text{Rest Mass}$$

Momentum Mass  $M_{mom}$  (Galileo's mass) Defined by ratio  $p/u$  of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Limiting cases:  $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho/2}$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \quad \text{Momentum Mass}$$

Effective Mass  $M_{eff}$  (Newton's mass) Defined by ratio  $F/a = dp/du$  of relativistic force to acceleration.

That is ratio of change  $dp = Mc \cosh \rho d\rho$  in momentum to change  $du = c \operatorname{sech}^2 \rho d\rho$  in group velocity.

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho \quad \text{Effective Mass}$$

Limiting cases:  $M_{eff} \xrightarrow{u \rightarrow c} M_{rest} e^{3\rho/2}$

$$M_{eff} \xrightarrow{u \ll c} M_{rest}$$

More common derivation using group velocity:  $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\nu}{d\mathbf{K}}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}} = M_{rest} \cosh^3 \rho \quad \text{Effective Mass}$$

general wave formula to accompany  $V_{group} = \frac{d\omega}{dk}$

# Definition(s) of mass for relativity/quantum

Rest Mass  $M_{rest}$  (Einstein's mass)

$$h\nu = h\nu_A = Mc^2 = hck_A$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hck_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

Momentum Mass  $M_{mom}$  (Galileo's mass) Defined by  $p/u$

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \quad \frac{\text{Momentum}}{\text{Mass}}$$

Effective Mass  $M_{eff}$  (Newton's mass) Defined by  $F/a = dp/du$

That is ratio of  $dp = Mc \cosh \rho d\rho$  to change  $du = c \operatorname{sech}^2 \rho d\rho$  in velocity

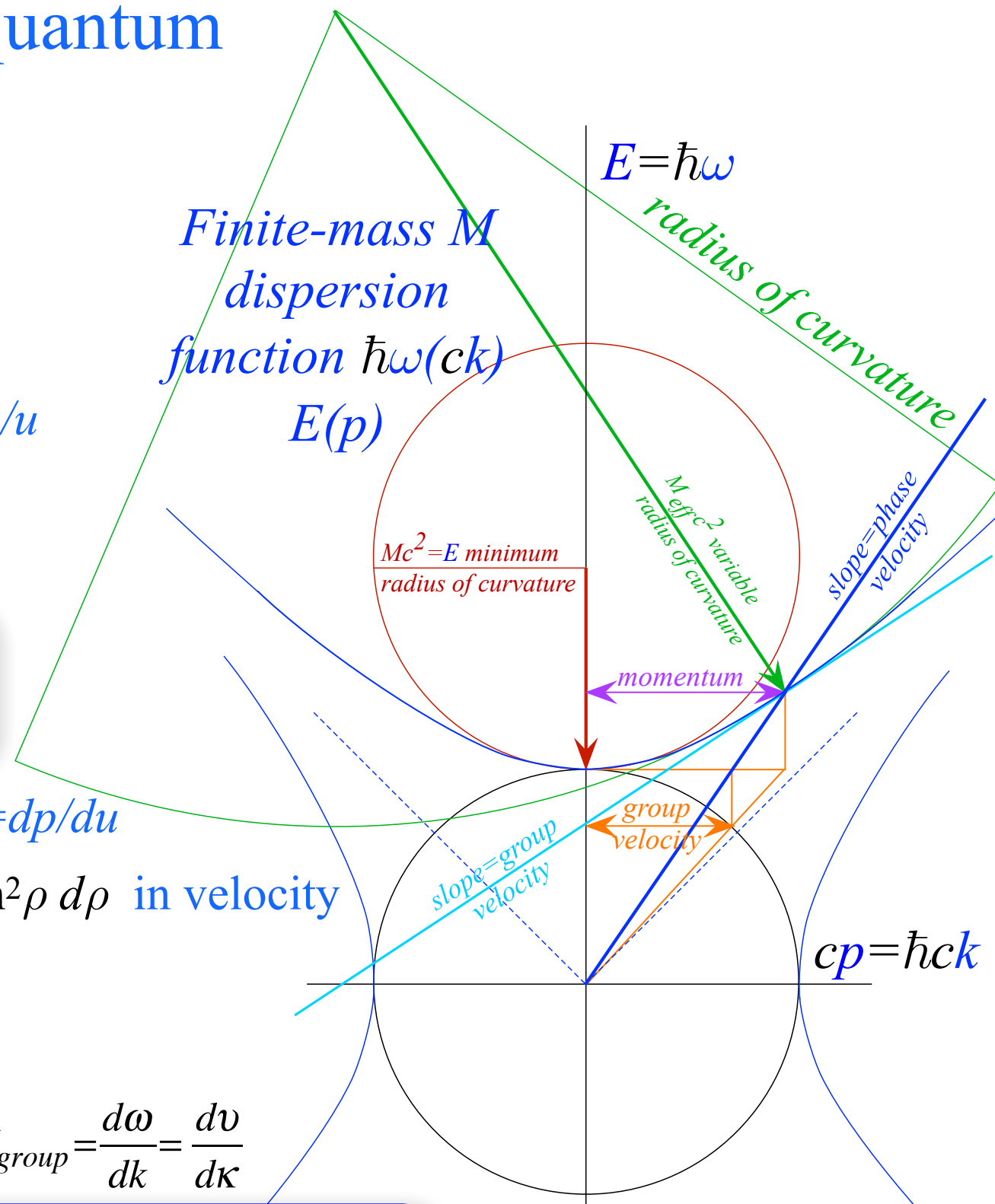
$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho \quad \frac{\text{Effective Mass}}$$

More common derivation using group velocity:  $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{dv}{dk}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}} = M_{rest} \cosh^3 \rho \quad \frac{\text{Effective Mass}}$$

general wave formula

to accompany  $V_{group} = \frac{d\omega}{dk}$



Effective mass is proportional to the radius of curvature of  $\omega(k)$  dispersion.

# Definition(s) of mass for relativity/quantum

How much mass does a  $\gamma$ -photon have?

Rest Mass (a)  $\gamma$ -rest mass:  $M_{rest}^{\gamma} = 0,$

Momentum Mass (b)  $\gamma$ -momentum mass:  $M_{mom}^{\gamma} = \frac{p}{c} = \frac{h\kappa}{c} = \frac{h\nu}{c^2},$

Effective Mass (c)  $\gamma$ -effective mass:  $M_{eff}^{\gamma} = \infty.$

Newton complained about his “corpuscles” of light having “fits” (going crazy).

(All this would be evidence of *triple Schizophrenia*.)

$$M_{mom}^{\gamma} = \frac{h\nu}{c^2} = \nu(1.2 \cdot 10^{-51}) \text{kg} \cdot \text{s} = 4.5 \cdot 10^{-36} \text{kg} \quad (\text{for: } \nu=600\text{THz})$$



# Lecture 30

## Wed. 12.05.2018

Review: Relativity  $\rho$  functions and plots vs.  $\rho$

*Rapidity  $\rho$  related to stellar aberration angle  $\sigma$  and L. C. Epstein's approach to relativity*  
*Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry*  
*"Occams Sword" and summary of 16 parameter functions of  $\rho$  and  $\sigma$*   
*Applications to optical waveguide, spherical waves, and accelerator radiation*

Derivation of relativistic quantum mechanics

- What's the matter with mass? Shining some light on the Elephant in the room
- ➔ Relativistic action and Lagrangian-Hamiltonian relations
- Poincaré' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization "photon" number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity in accelerated frames*

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# Relativistic action $S$ and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega \qquad \hbar \equiv \frac{h}{2\pi}$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor format →

$$\begin{aligned} \hbar\omega_A &= Mc^2 = \hbar ck_A \\ \hbar\omega_{phase} &= E = \hbar\omega_A \cosh \rho \\ \hbar ck_{phase} &= cp = \hbar\omega_A \sinh \rho \end{aligned}$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action $S$** and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$$

$$\hbar \equiv \frac{h}{2\pi}$$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor format →

$$\begin{aligned} \hbar\omega_A &= Mc^2 = \hbar c k_A \\ \hbar\omega_{phase} &= E = \hbar\omega_A \cosh \rho \\ \hbar c k_{phase} &= cp = \hbar\omega_A \sinh \rho \end{aligned}$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action $S$** and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \qquad \hbar \equiv \frac{h}{2\pi}$$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor format →

$$\begin{aligned} \hbar\omega_A &= Mc^2 = \hbar ck_A \\ \hbar\omega_{phase} &= E = \hbar\omega_A \cosh \rho \\ \hbar ck_{phase} &= cp = \hbar\omega_A \sinh \rho \end{aligned}$$



# Relativistic **action $S$** and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian*  $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor format →

$$\begin{aligned} \hbar \omega_A &= Mc^2 = \hbar c k_A \\ \hbar \omega_{phase} &= E = \hbar \omega_A \cosh \rho \\ \hbar c k_{phase} &= cp = \hbar \omega_A \sinh \rho \end{aligned}$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action $S$** and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian*  $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

*Legendre transformation*

Use *Group velocity* :  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor format →

$$\begin{aligned} \hbar \omega_A &= Mc^2 = \hbar c k_A \\ \hbar \omega_{phase} &= E = \hbar \omega_A \cosh \rho \\ \hbar c k_{phase} &= cp = \hbar \omega_A \sinh \rho \end{aligned} \quad \hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action $S$** and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian*  $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

*Legendre transformation*

Use Group velocity:  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

Prior wave relations

← linear Hz  
format

angular phasor  
format →

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian H = E*

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity :  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

$$L \text{ is : } Mc^2 \frac{-1}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

Prior wave relations

← linear Hz format      angular phasor format →

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \equiv \frac{h}{2\pi}$$



# Relativistic **action $S$** and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian*  $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity:  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

Note:  $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Compare *Lagrangian*  $L$

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

Prior wave relations

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

← linear Hz format

angular phasor →

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

format

format

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian H = E*

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity:  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

Note:  $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Compare *Lagrangian L*

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

with *Hamiltonian H = E*

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

## Prior wave relations

$$\begin{aligned} \hbar \nu_A &= Mc^2 = \hbar c \kappa_A \\ \hbar \nu_{phase} &= E = \hbar \nu_A \cosh \rho \\ \hbar c \kappa_{phase} &= cp = \hbar \nu_A \sinh \rho \end{aligned}$$

← linear Hz  
format

angular phasor  
format →

$$\begin{aligned} \hbar \omega_A &= Mc^2 = \hbar c \kappa_A \\ \hbar \omega_{phase} &= E = \hbar \omega_A \cosh \rho \\ \hbar c \kappa_{phase} &= cp = \hbar \omega_A \sinh \rho \end{aligned}$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic action $S$ and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian*  $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity:  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

Note:  $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Also:  $cp = Mc^2 \sinh \rho$

Compare *Lagrangian*  $L$

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

with *Hamiltonian*  $H = E$

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

Prior wave relations  $\hbar = h/2\pi$

← linear Hz format      angular phasor format →

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian H = E*

$$\frac{dS}{dt} = L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p u - H = L$$

Legendre transformation

Use Group velocity:  $u = \frac{dx}{dt} = c \tanh \rho = c \sin \sigma$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$$

$$= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

Note:  $Mc u = Mc^2 \tanh \rho$

$$= Mc^2 \sin \sigma$$

Also:  $cp = Mc^2 \sinh \rho$

$$= \hbar ck = Mc^2 \tan \sigma$$

Including stellar angle  $\sigma$

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian H = E*

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define *Action S = \hbar \Phi*

$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

Prior wave relations

← linear Hz format

angular phasor format →

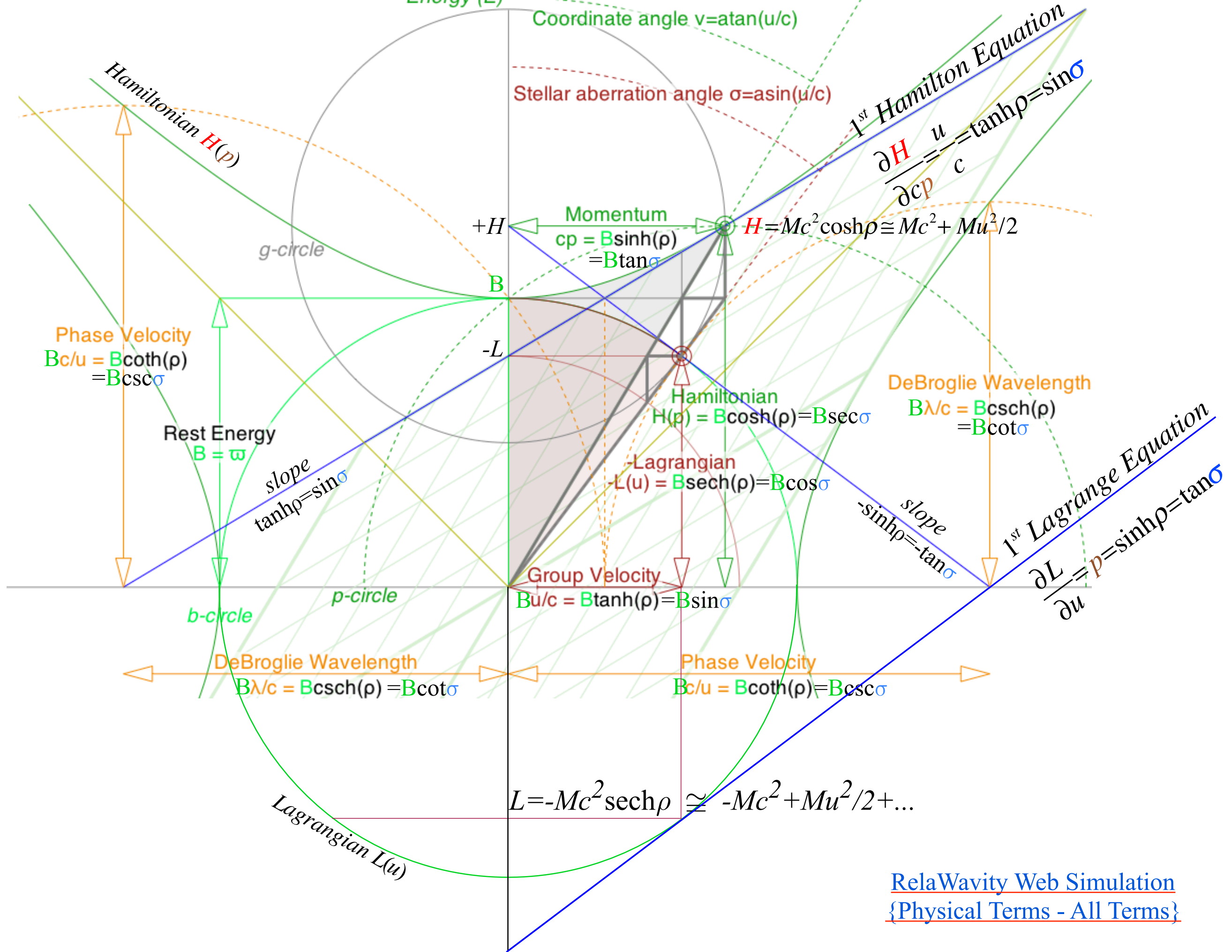
$$\hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c \kappa_{phase} = cp = \hbar \omega_A \sinh \rho$$

$$\hbar \equiv \frac{h}{2\pi}$$





# Lecture 30

## Wed. 12.05.2018

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➔ Poincaré' and Hamilton-Jacobi equations

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Compton recoil related to rocket velocity formula

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# Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian H = E*

$$\frac{dS}{dt} = L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian H = E*

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define *Action S = \hbar \Phi*

## Prior wave relations

$$\begin{aligned} \hbar \nu_A &= Mc^2 = \hbar c \kappa_A \\ \hbar \nu_{phase} &= E = \hbar \nu_A \cosh \rho \\ \hbar c \kappa_{phase} &= cp = \hbar \nu_A \sinh \rho \end{aligned}$$

← linear Hz  
format

angular phasor  
format

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$$dS \equiv Ldt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt$$

*Poincare Invariant action differential*

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

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*Legendre transformation*

Use Group velocity:  $u = \frac{dx}{dt} = c \tanh \rho$

$$dS \equiv L dt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt$$

*Poincare Invariant action differential*

$$\frac{\partial S}{\partial x} = p$$

$$\frac{\partial S}{\partial t} = -H$$

*Hamilton-Jacobi equations*

Compare Lagrangian  $L$

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

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Define Action  $S = \hbar \Phi$

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Comparing 2<sup>nd</sup>-quantization "photon" number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity in accelerated frames*

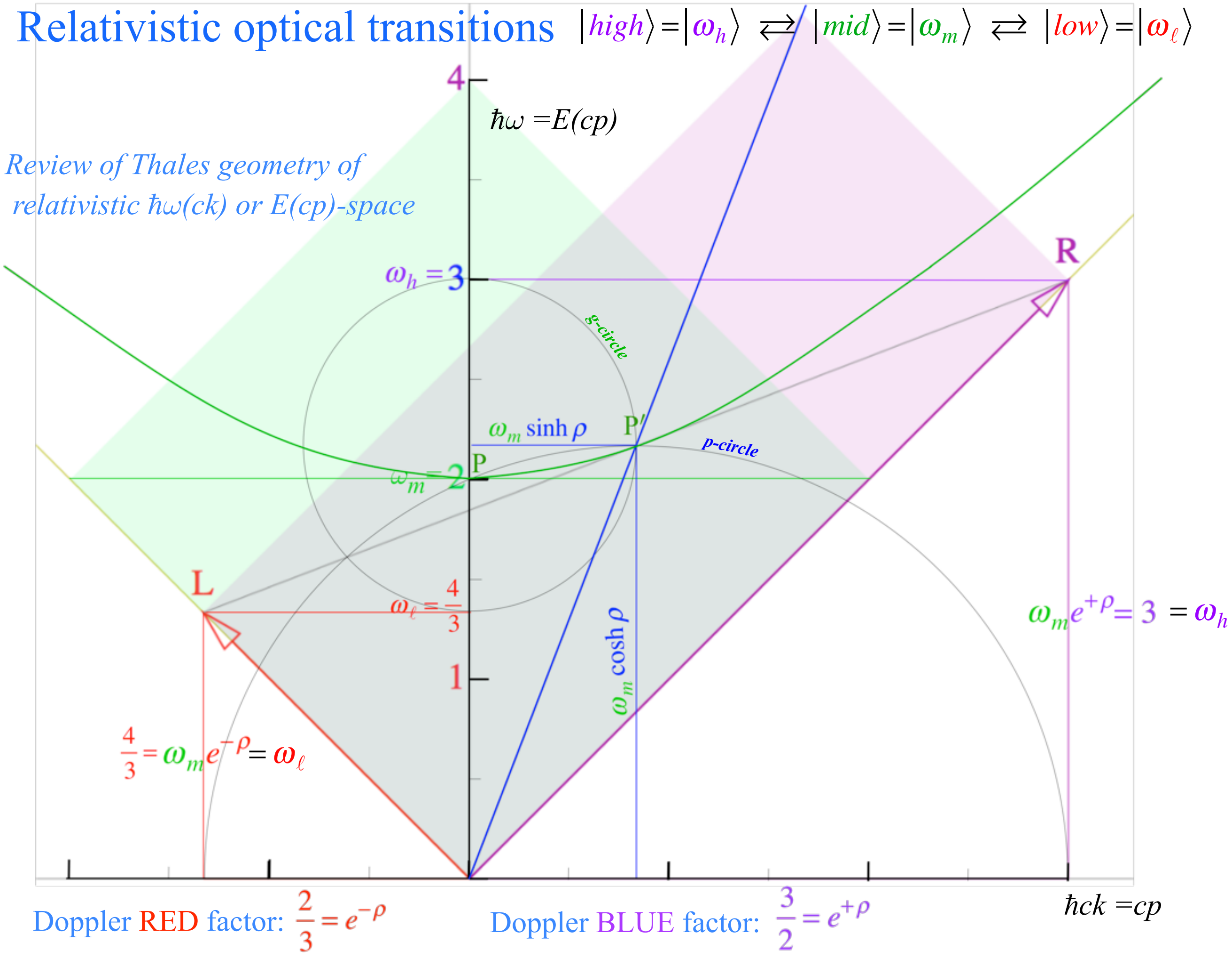
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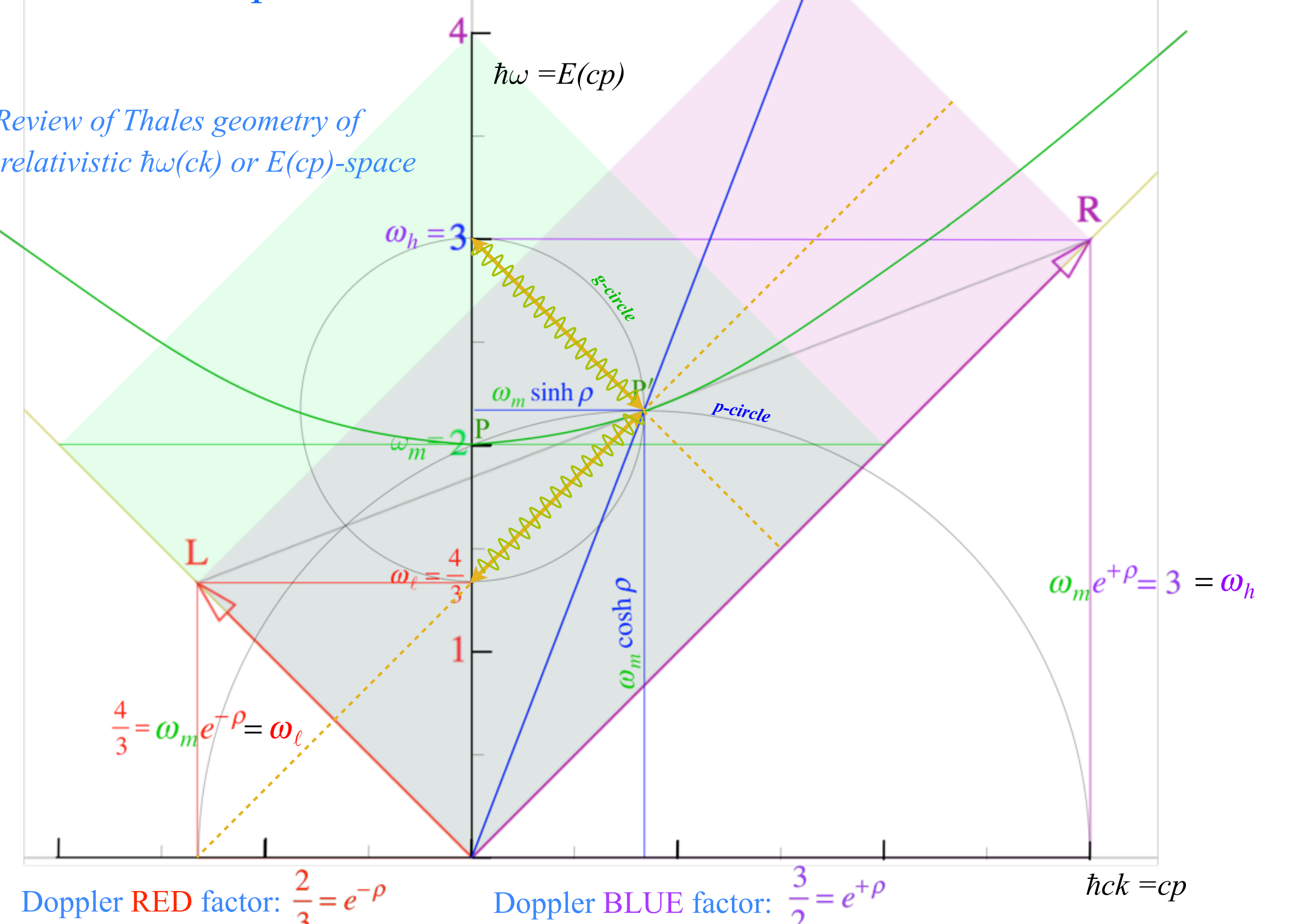
# Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

*Review of Thales geometry of relativistic  $\hbar\omega(ck)$  or  $E(cp)$ -space*



Relativistic optical transitions  $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

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Review of Thales geometry of relativistic  $\hbar\omega(ck)$  or  $E(cp)$ -space

Initial stationary

BLUE  $K_h$  thing  $\omega_h = M_h c^2$

transitions to

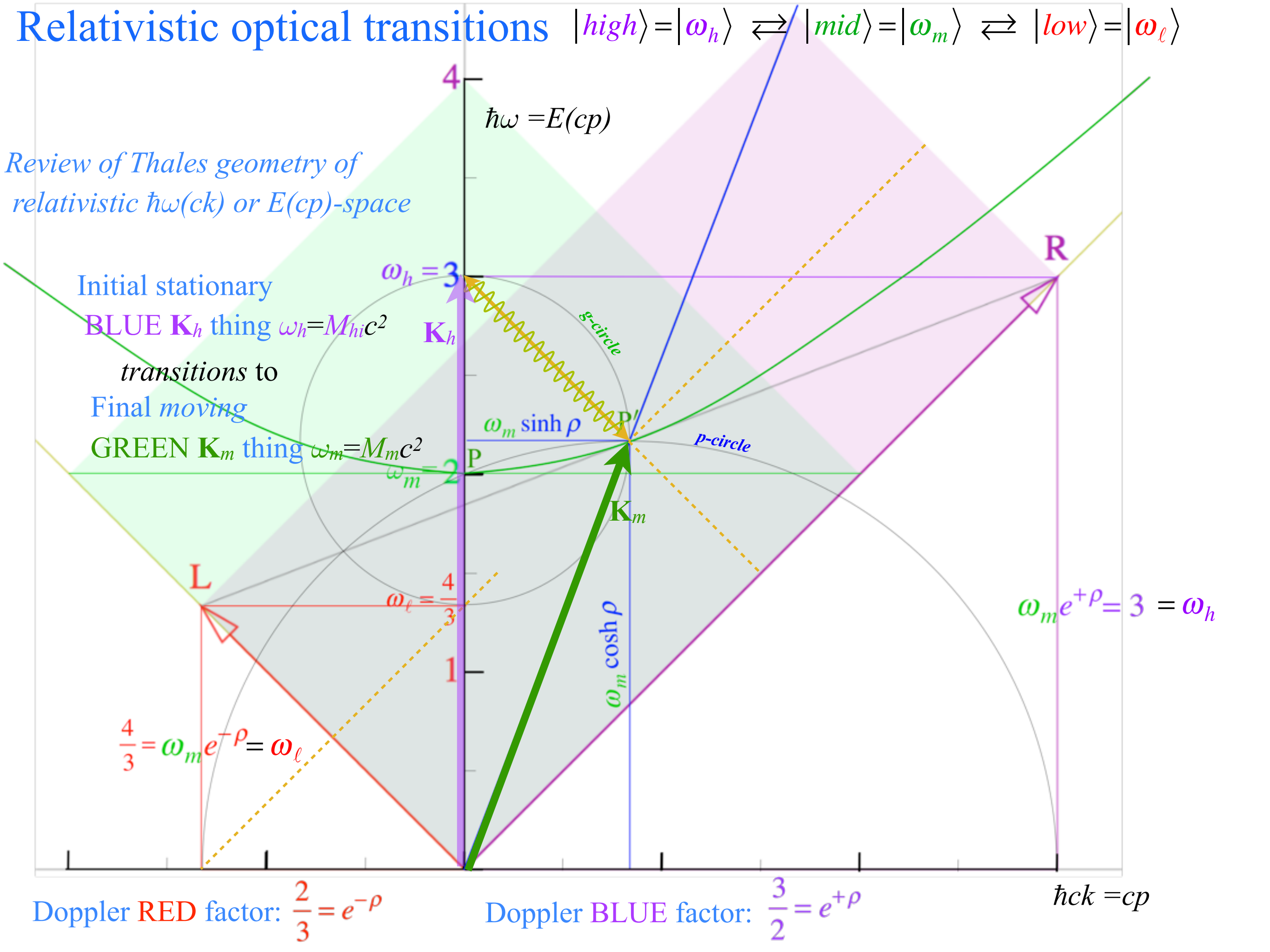
Final moving

GREEN  $K_m$  thing  $\omega_m = M_m c^2$

Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

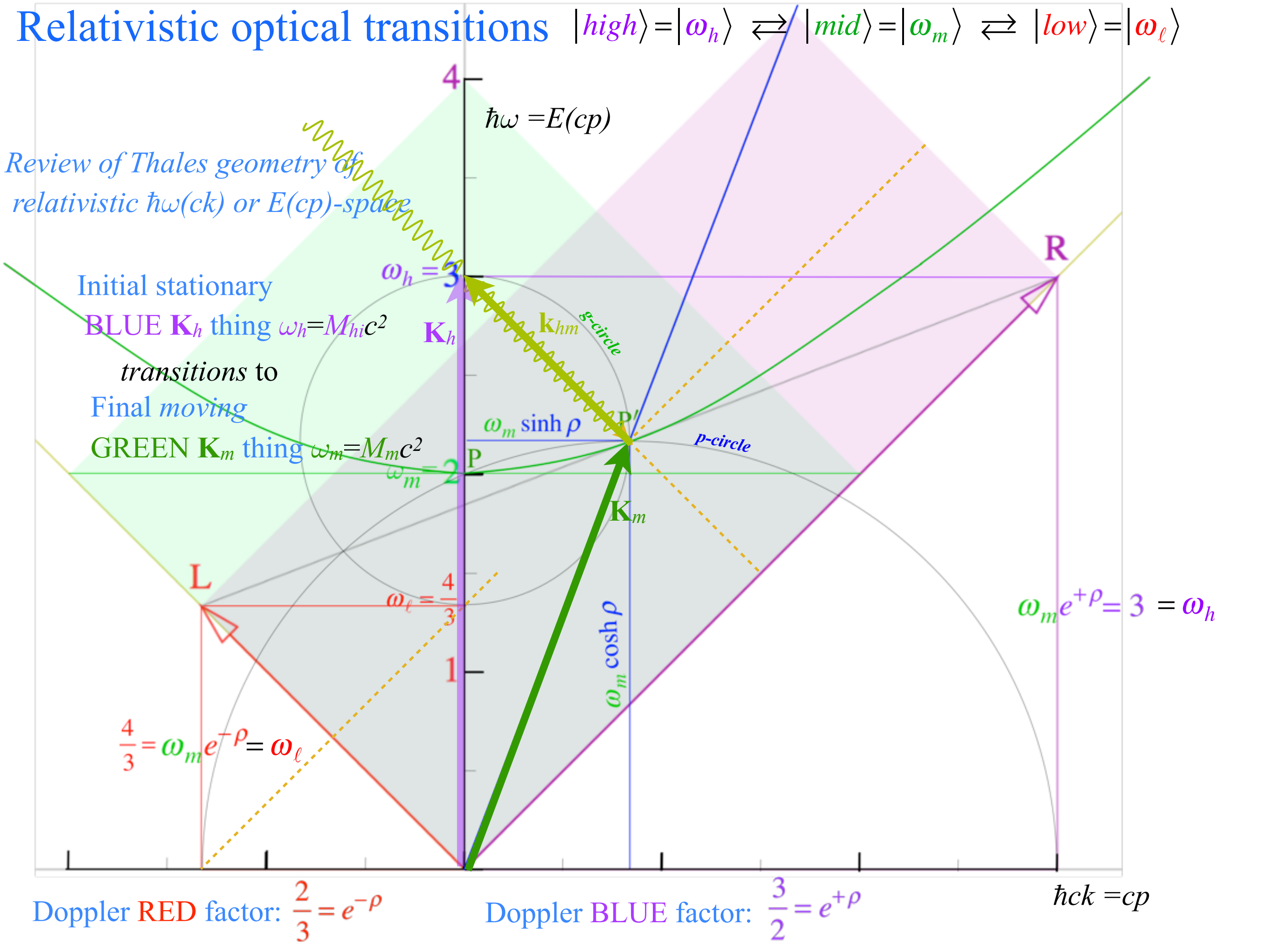
Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$



Relativistic optical transitions  $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

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- ➔ Feynman diagram geometry
  - Compton recoil related to rocket velocity formula
  - Comparing 2<sup>nd</sup>-quantization "photon" number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

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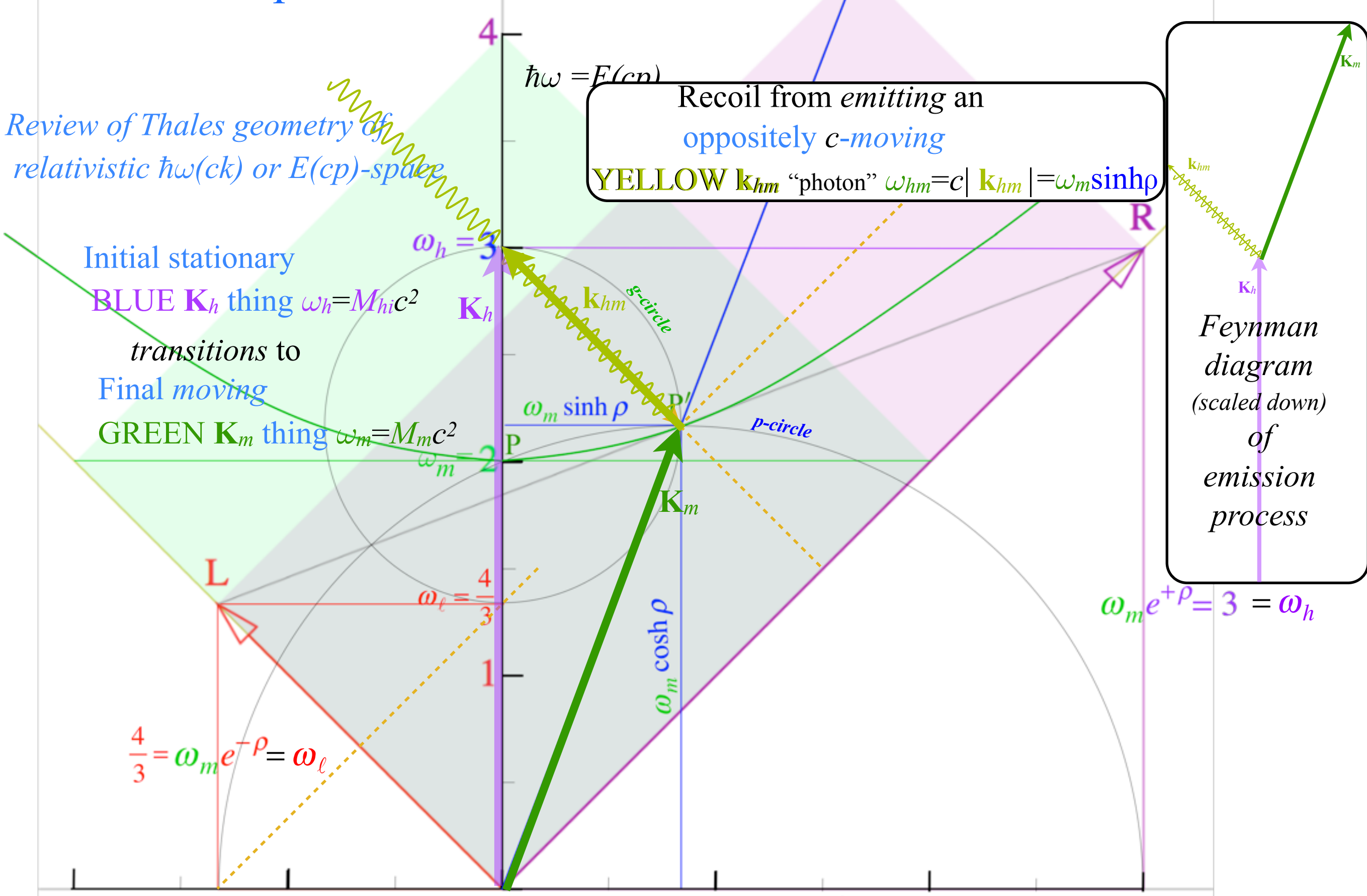
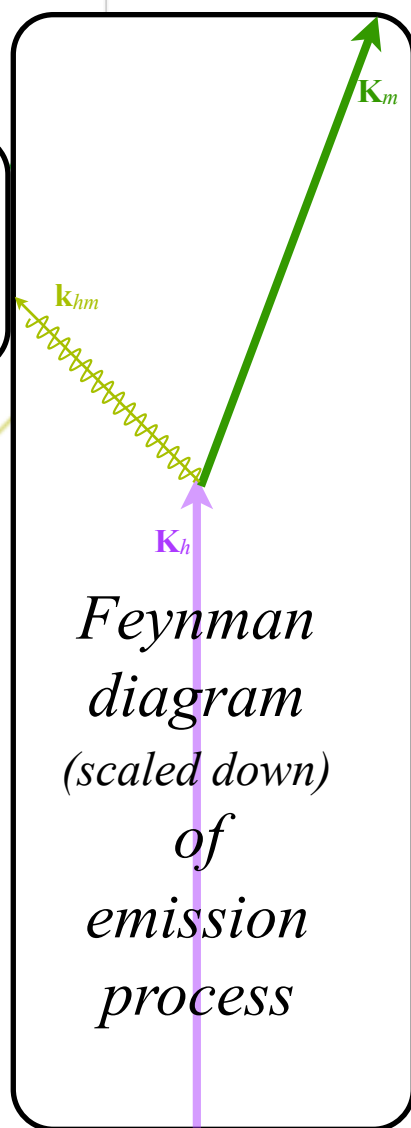


Relativistic optical transitions  $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

*Review of Thales geometry of relativistic  $\hbar\omega(cp)$  or  $E(cp)$ -space*

Initial stationary  
**BLUE  $K_h$  thing**  $\omega_h = M_h c^2$   
 transitions to  
 Final moving  
**GREEN  $K_m$  thing**  $\omega_m = M_m c^2$

Recoil from emitting an oppositely  $c$ -moving **YELLOW  $k_{hm}$**  "photon"  $\omega_{hm} = c |k_{hm}| = \omega_m \sinh \rho$



Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

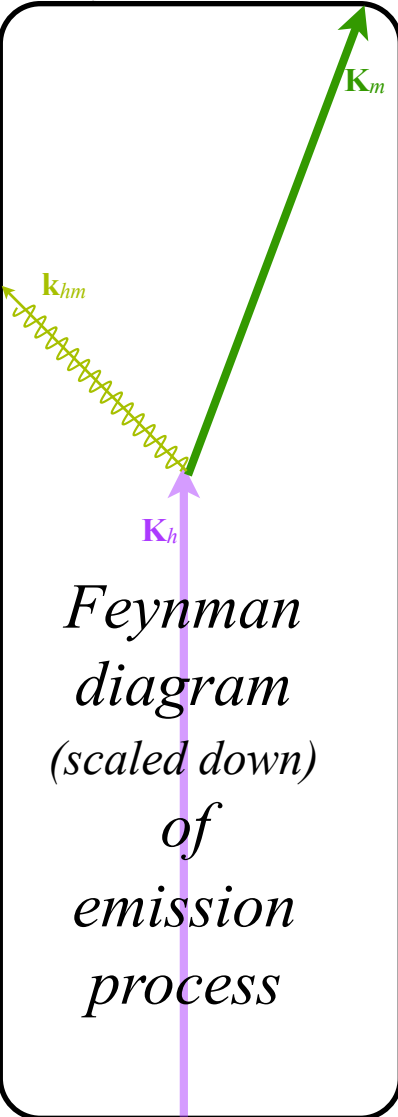
$\hbar ck = cp$

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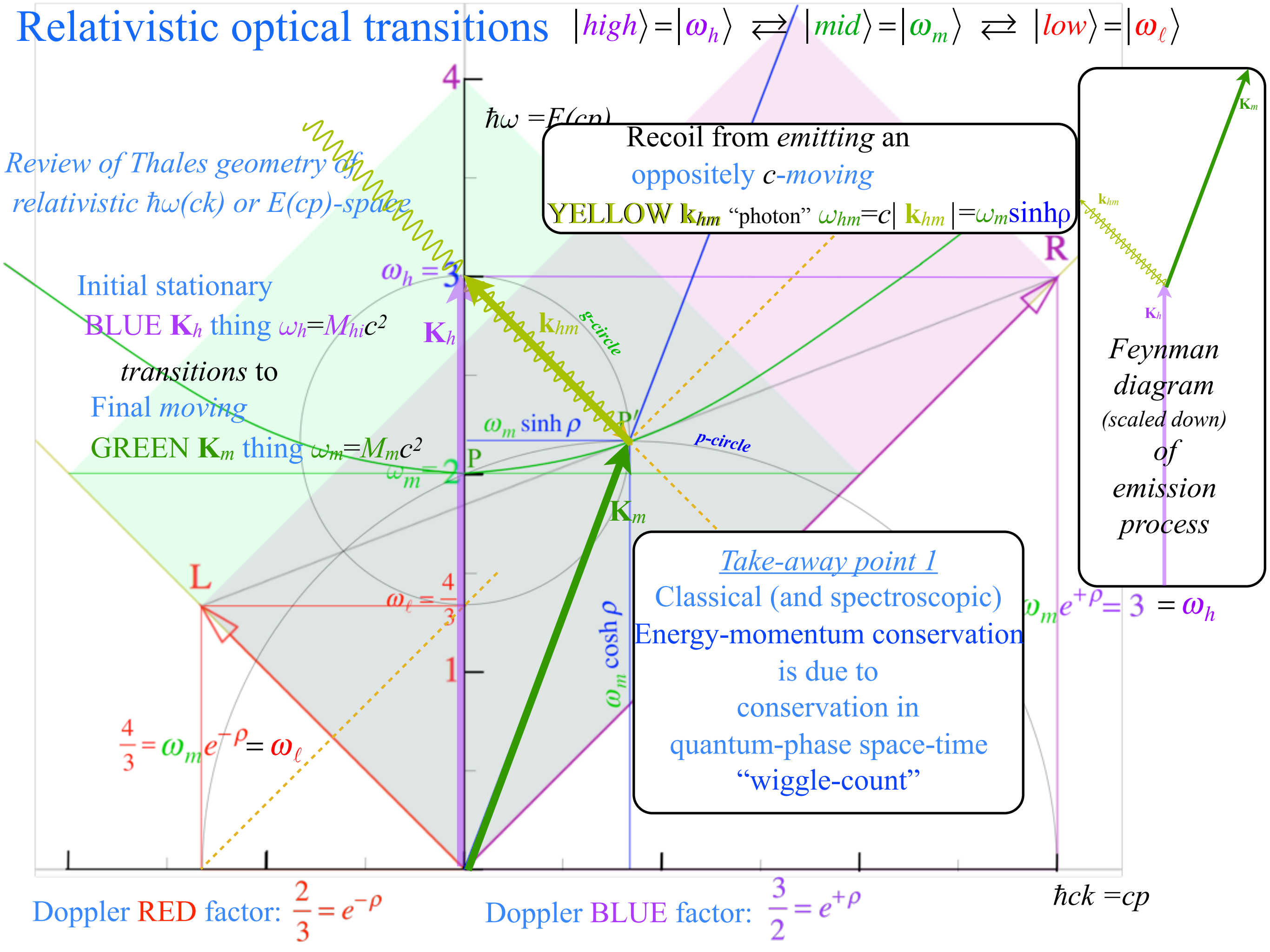


Take-away point 1  
 Classical (and spectroscopic) Energy-momentum conservation is due to conservation in quantum-phase space-time "wiggle-count"

Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$



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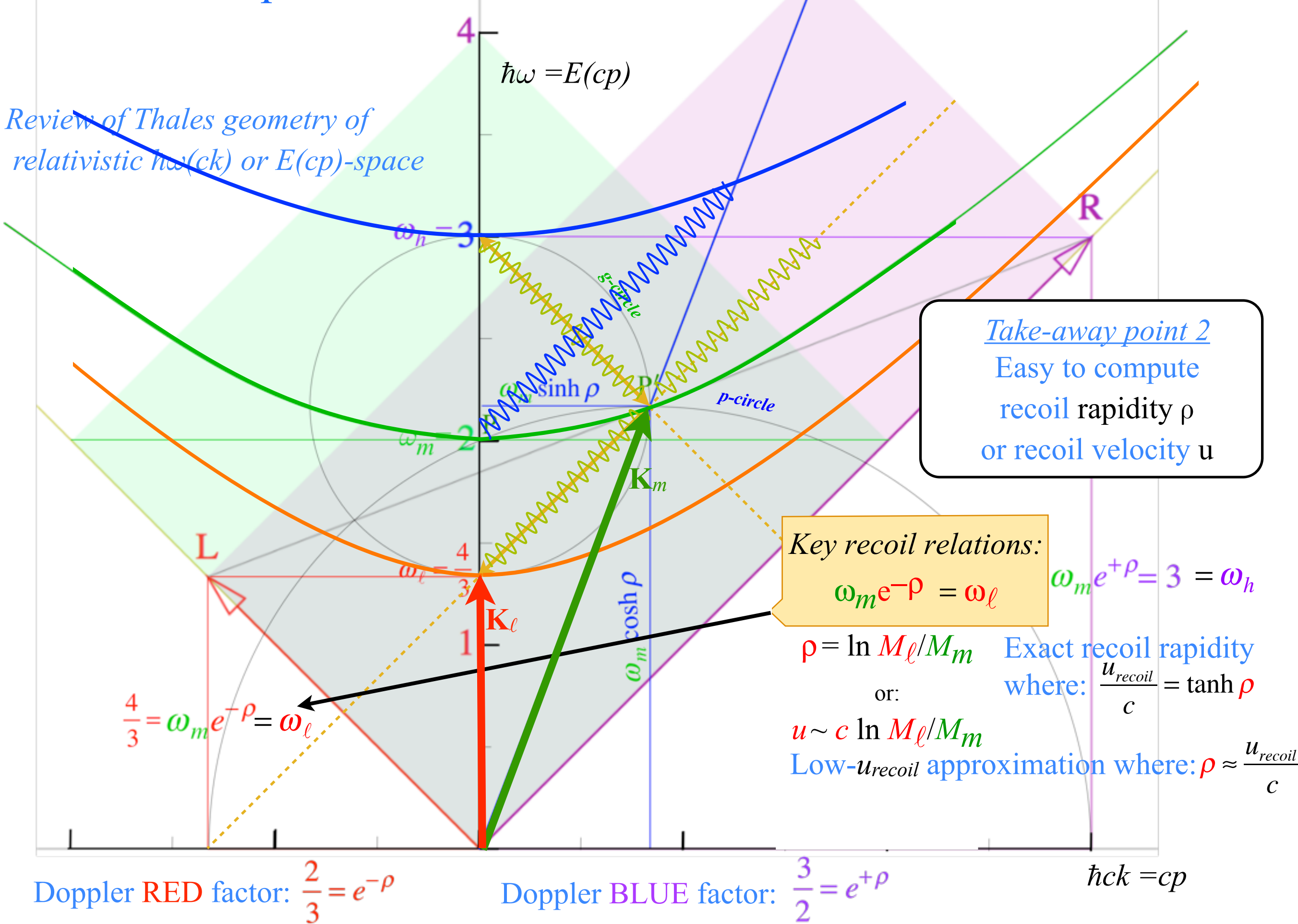
- ➔ Compton recoil related to rocket velocity formula  
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Take-away point 2  
Easy to compute recoil rapidity  $\rho$  or recoil velocity  $u$

Key recoil relations:  
 $\omega_m e^{-\rho} = \omega_l$   
 $\omega_m e^{+\rho} = \omega_h$

$\rho = \ln M_l/M_m$  Exact recoil rapidity where:  $\frac{u_{recoil}}{c} = \tanh \rho$

or:  
 $u \sim c \ln M_l/M_m$  Low- $u_{recoil}$  approximation where:  $\rho \approx \frac{u_{recoil}}{c}$

Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

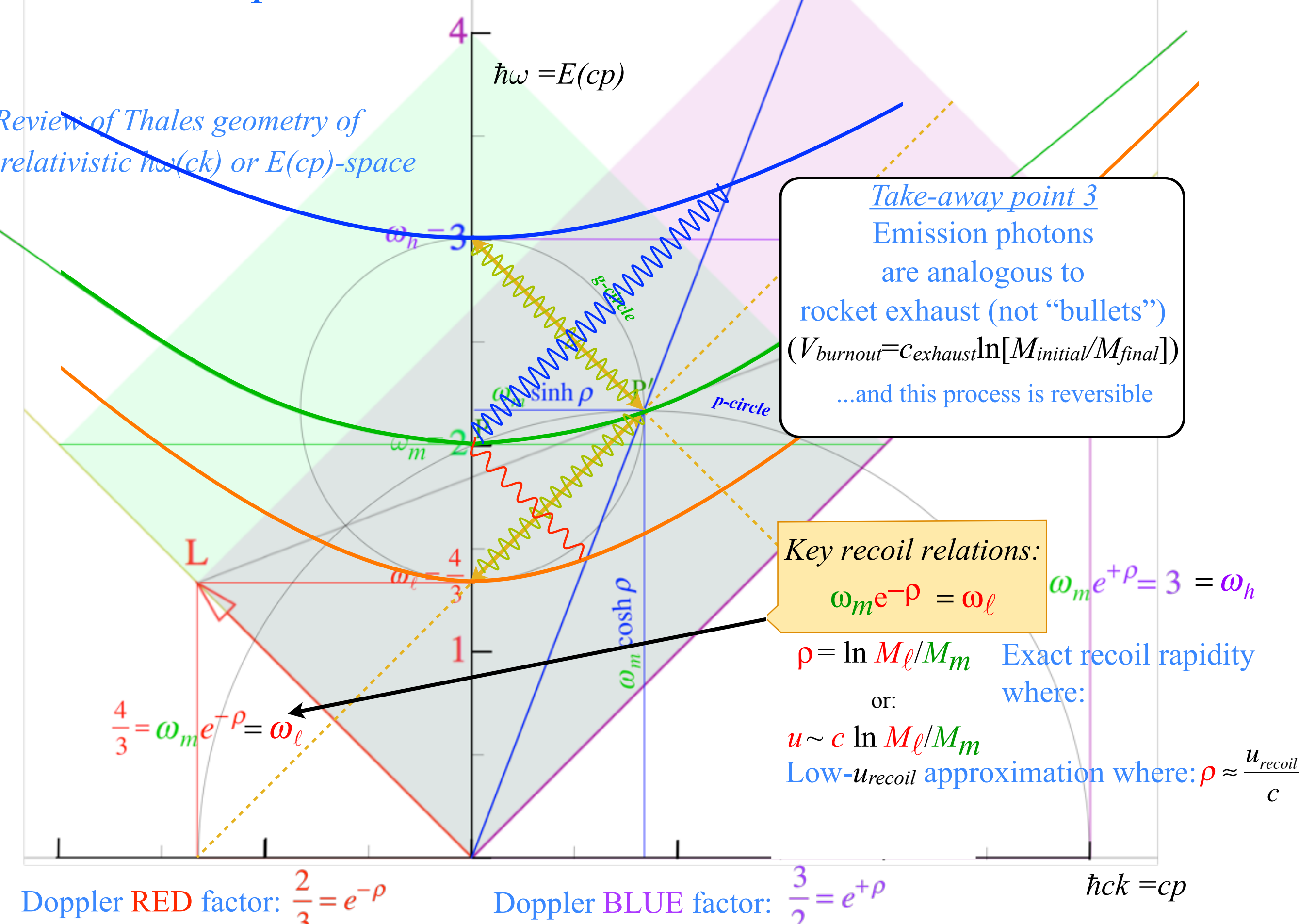
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Take-away point 3  
 Emission photons are analogous to rocket exhaust (not “bullets”) ( $V_{burnout} = C_{exhaust} \ln[M_{initial}/M_{final}]$ ) ...and this process is reversible

Key recoil relations:

$\omega_m e^{-\rho} = \omega_l$        $\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_l/M_m$       Exact recoil rapidity where:

or:  
 $u \sim c \ln M_l/M_m$   
 Low- $u_{recoil}$  approximation where:  $\rho \approx \frac{u_{recoil}}{c}$

Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

$(p, q)$  - coordinates

rest frequency:      rapidity:

$$\omega_q = \omega_m e^{q\rho} \qquad \rho_p = p\rho$$

$$P_{p,q} = (ck_{p,q}, \omega_{p,q})$$

$$= \omega_m e^{q\rho} (\sinh p\rho, \cosh p\rho)$$

All-rational-fraction lattice  
defined by discrete sub-group  
of Lorentz Poincare Group  
(Feynman path integrals defined  
by group transformations)

+2

+1

0

-1

-2

$L = \text{lefthand shift power}$   
 $\omega_L = \omega_m e^{L\rho}$

-2

-1

0

+1

+2

$R = \text{righthand shift power}$   
 $\omega_R = \omega_m e^{R\rho}$

$(p, q) - (R, L)$   
coordinate

transformations:

$$p = \frac{R-L}{2}, \quad q = \frac{R+L}{2}$$

$$R = p+q, \quad L = q-p$$

Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

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# 2<sup>nd</sup> Quantization:

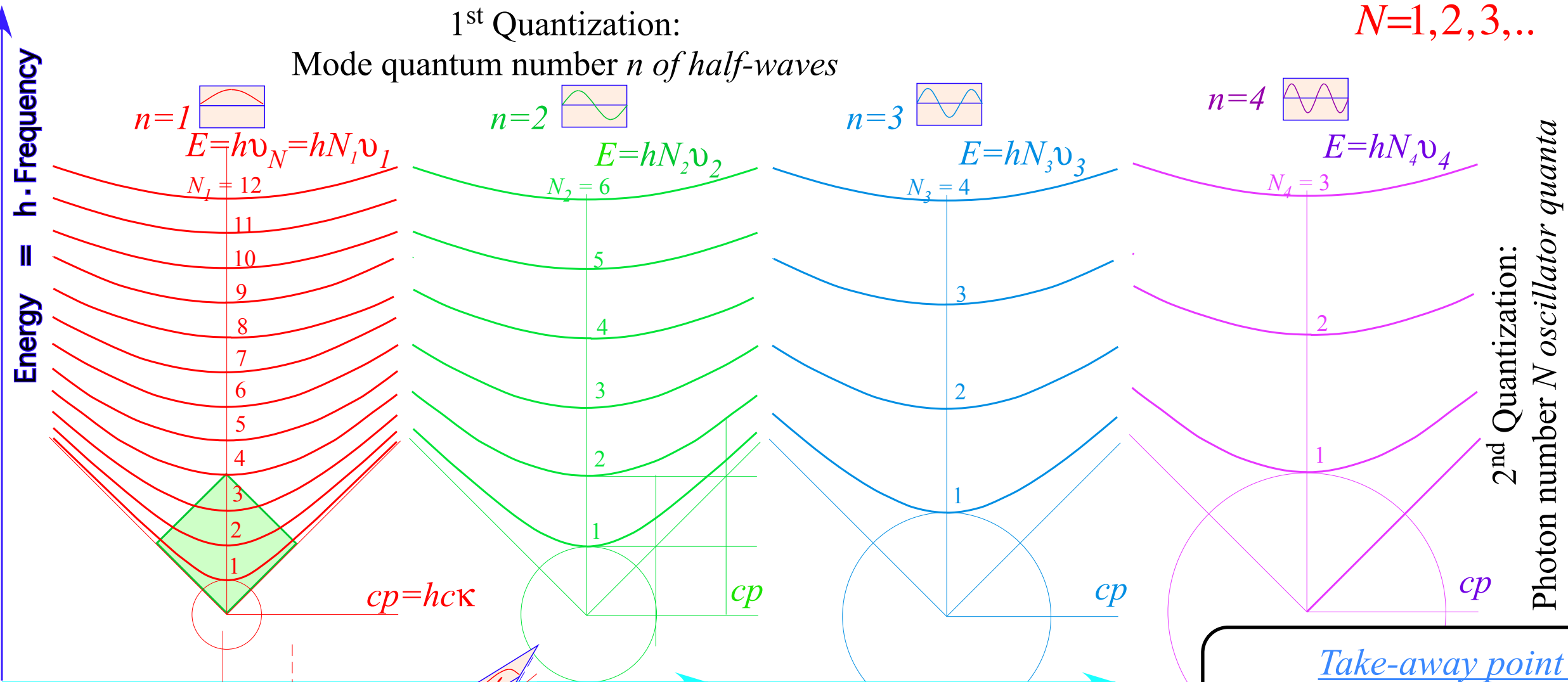
$h\nu$  is actually  $hN\nu$

$(h\nu_{phase} = E = h\nu_A \cosh \rho)$  is actually  $(hN\nu_{phase} = E_N = hN\nu_A \cosh \rho)$  with quantum numbers

$N=1,2,3,\dots$

1<sup>st</sup> Quantization:

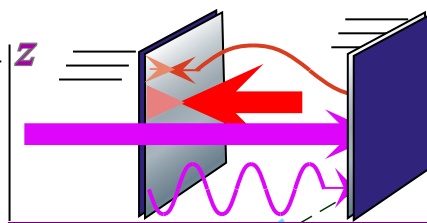
Mode quantum number  $n$  of half-waves



$c \cdot \text{Momentum} = hc \cdot \text{Wavenumber}$

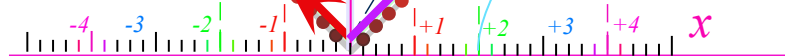
Boosted wave mode

Boosted cavity wave has invariant mode number  $n$  photon number  $N_n$



Lorentz contracted cavity length  $L=3.2$   
Proper length  $l=4.0$

Pirelli Challenge Site  
Quantized amplitude



Take-away point 4  
Cavity quantum electrodynamics (CQED) and spectra are analogous to molecular rovibronic dynamics with rotation-vibration algebra replaced by Lorentz-Poincare-Dirac algebra (and geometry!)

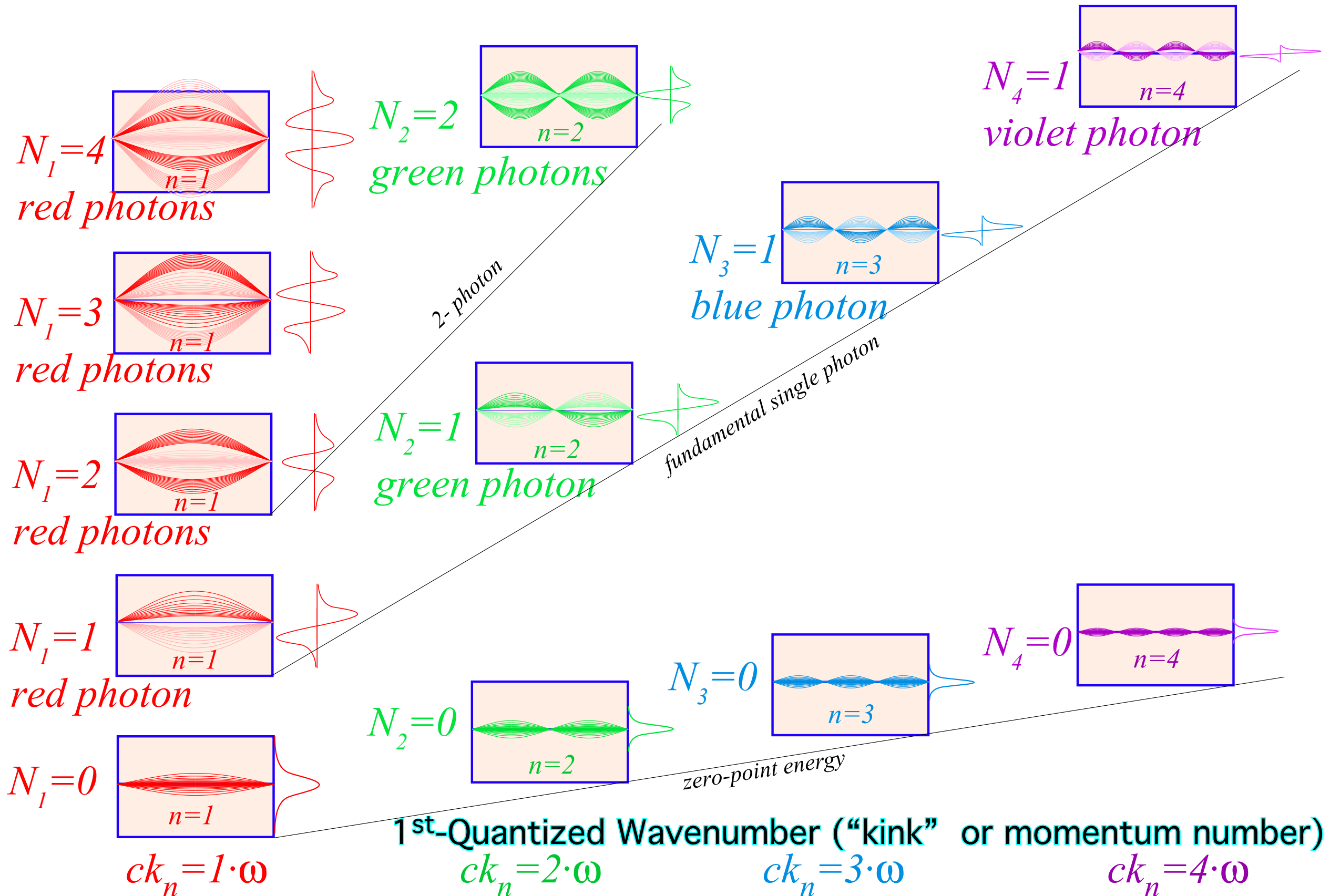


# 2<sup>nd</sup> Quantization:

$h\nu$  is actually  $hN\nu$

$(h\nu_{phase}=E=h\nu_A \cosh \rho)$  is actually  $(hN\nu_{phase}=E_N=hN\nu_A \cosh \rho \quad (N=1,2,..) )$

2<sup>nd</sup>-Quantized Amplitude (“photon” number)



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# Acceleration by chirping laser pairs

## Varying acceleration (General case)

From Lect. 35 ModPhys (2012)

Varying local acceleration  $\rho = \rho(\tau)$

$$u = \frac{dx}{dt} = c \tanh(\tau)$$

$$\frac{dt}{d\tau} = \cosh \rho(\tau)$$

$$\frac{dx}{d\tau} = \frac{dx}{dt} \frac{dt}{d\tau} = c \tanh \rho(\tau) \cosh \rho(\tau) = c \sinh \rho(\tau)$$

$$ct = c \int \cosh \rho(\tau) d\tau$$

$$x = c \int \sinh \rho(\tau) d\tau$$

Constant local acceleration  $\rho = \frac{g\tau}{c}$  "Einstein Elevator"

$$ct = c \int \cosh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \sinh \frac{g\tau}{c}$$

$$x = c \int \sinh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \cosh \frac{g\tau}{c}$$

Previous examples involved constant velocity

Constant velocity  $\rho = \rho_0 = \text{const.}$  "Lorentz transformation"

$$ct = c \int \cosh \rho_0 d\tau = c\tau \cosh \rho_0$$

$$x = c \int \sinh \rho_0 d\tau = c\tau \sinh \rho_0$$

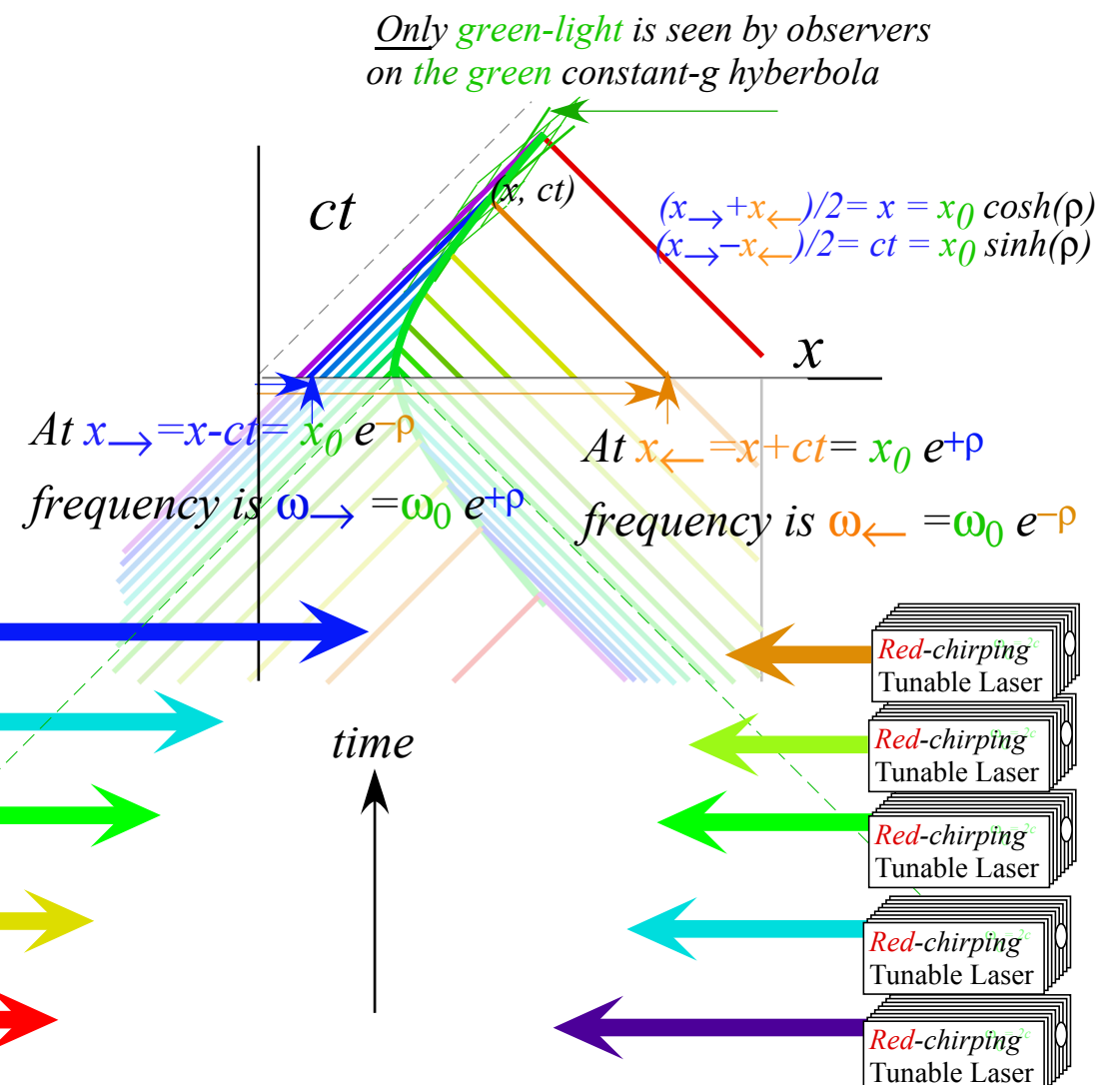
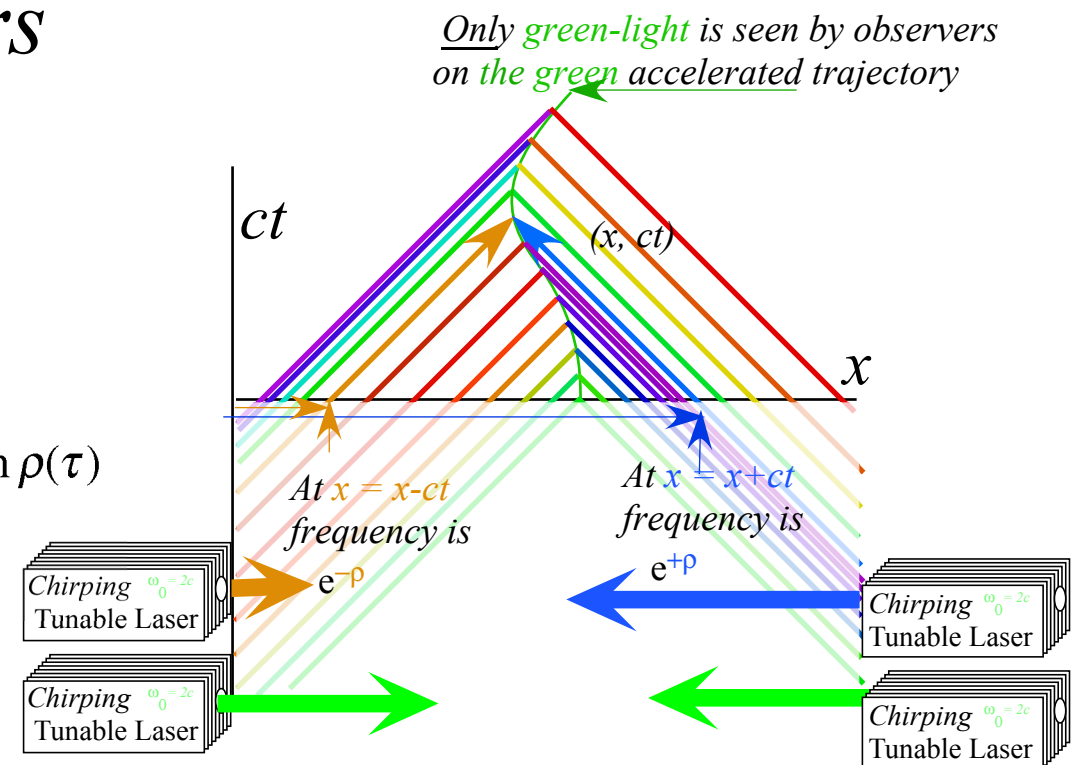
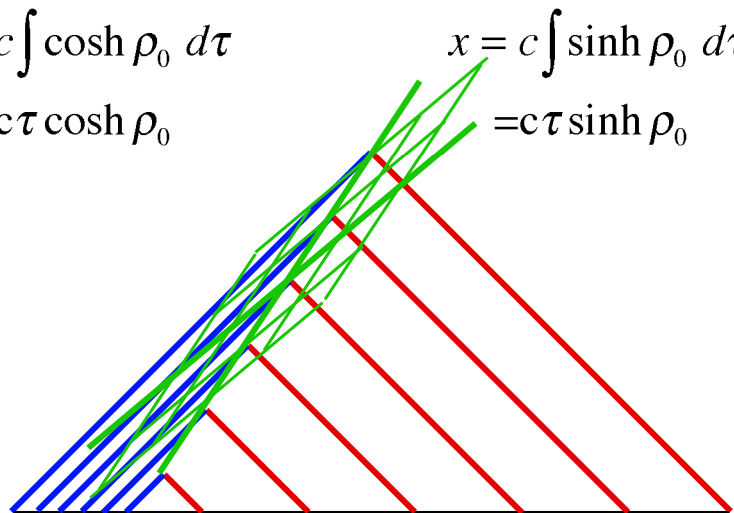
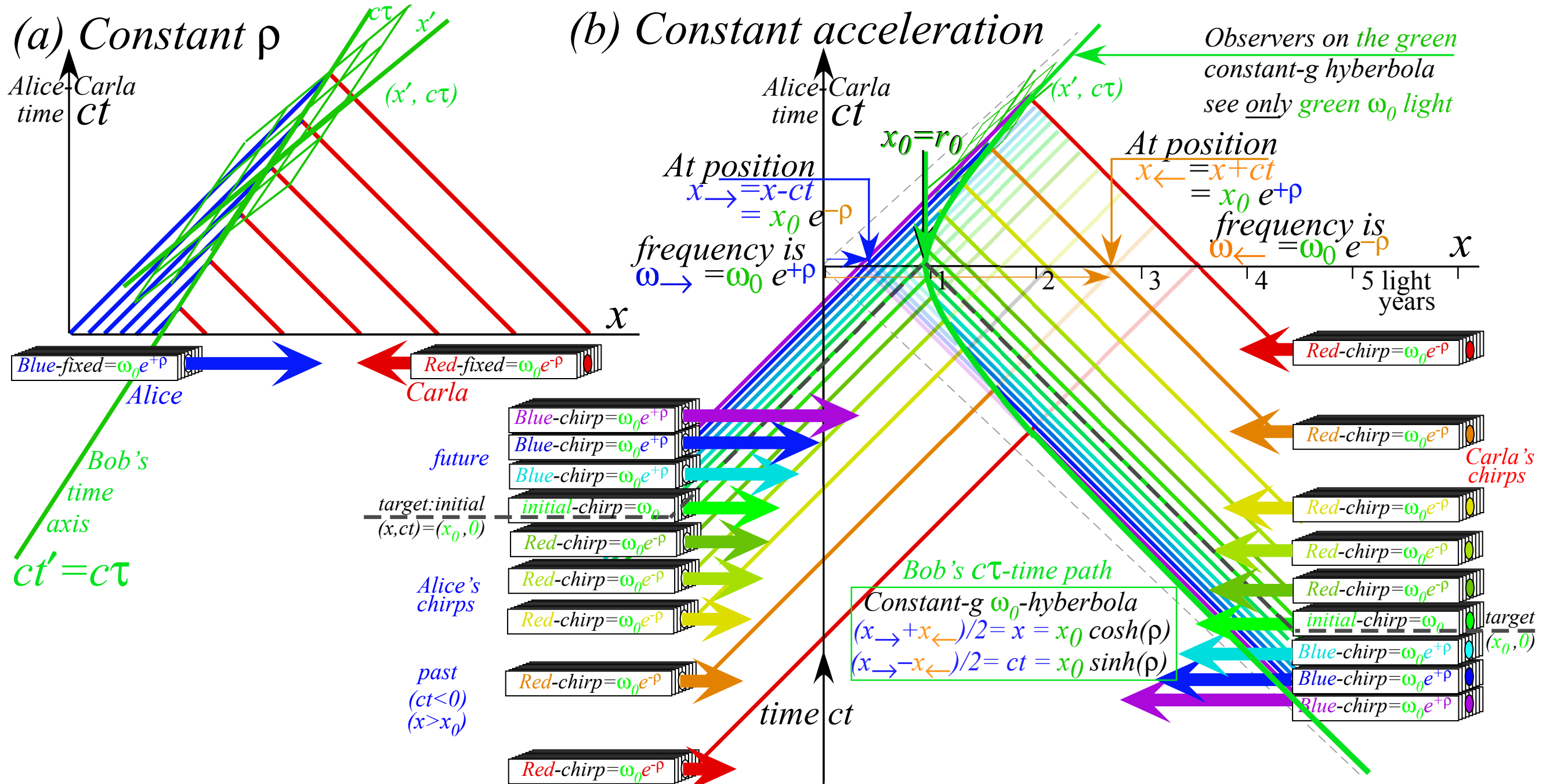
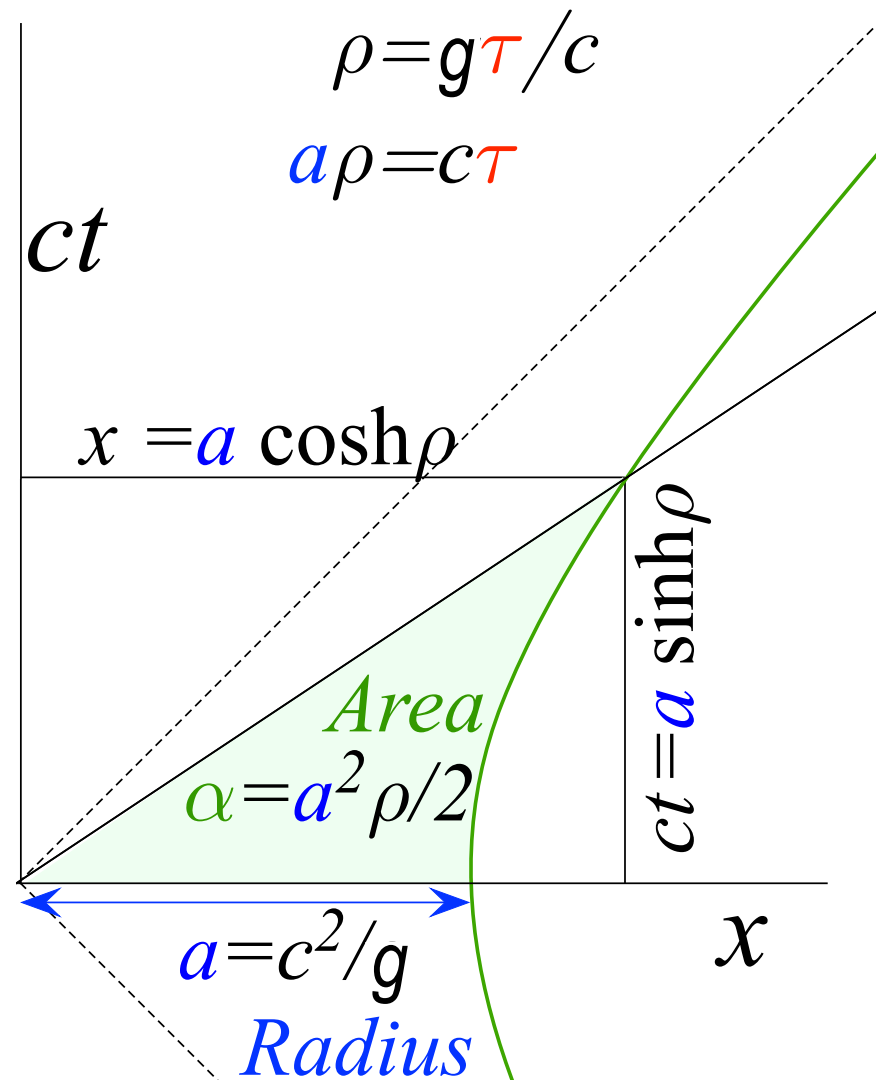


Fig. 8.1 Optical wave frames by red-and-blue-chirped lasers (a) Varying acceleration (b) Constant g





(a) Constant acceleration  $g$   
 Rapidity  $\rho$  vs proper time  $\tau$



(b) Traveler paths of acceleration  $g_q$

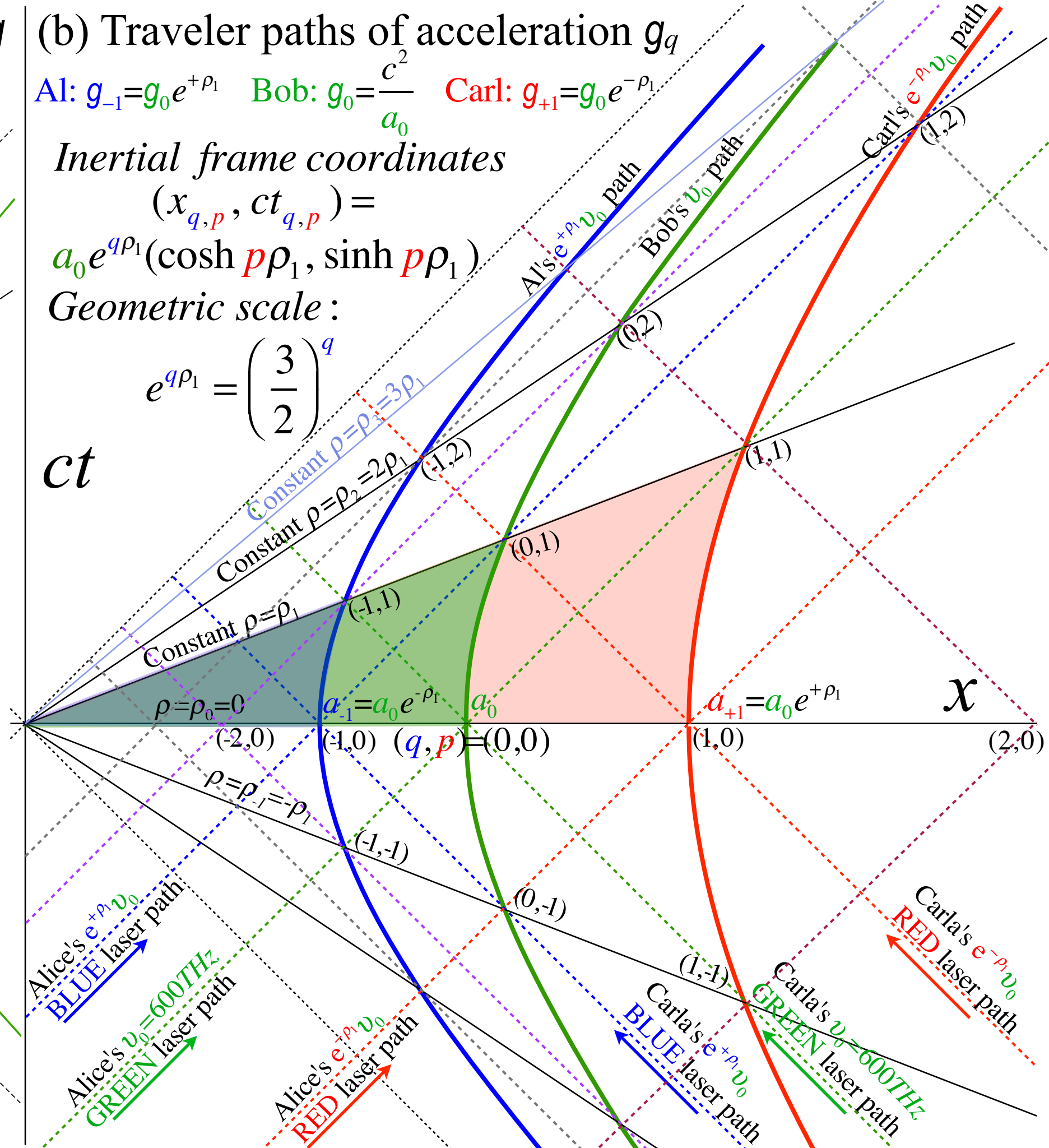
Al:  $g_{-1} = g_0 e^{+\rho_1}$     Bob:  $g_0 = \frac{c^2}{a_0}$     Carl:  $g_{+1} = g_0 e^{-\rho_1}$

*Inertial frame coordinates*

$(x_{q,p}, ct_{q,p}) =$   
 $a_0 e^{q\rho_1} (\cosh p\rho_1, \sinh p\rho_1)$

*Geometric scale:*

$e^{q\rho_1} = \left(\frac{3}{2}\right)^q$



# Lecture 31

## Thur. 12.10.2015

Review: Relativity  $\rho$  functions      Two famous ones      Extremes and plot vs.  $\rho$   
Doppler jeopardy      Geometric mean and Relativistic hyperbolas  
Animation of  $e^{\rho}=2$  spacetime and per-spacetime plots

*Rapidity*  $\rho$  related to *stellar aberration angle*  $\sigma$  and L. C. Epstein's approach to relativity

Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry

“Occams Sword” and summary of 16 parameter functions of  $\rho$  and  $\sigma$

Applications to optical waveguide, spherical waves, and accelerator radiation

*Learning about sin! and COS and...*

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

Relativistic action and Lagrangian-Hamiltonian relations

Poincaré' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization “photon” number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

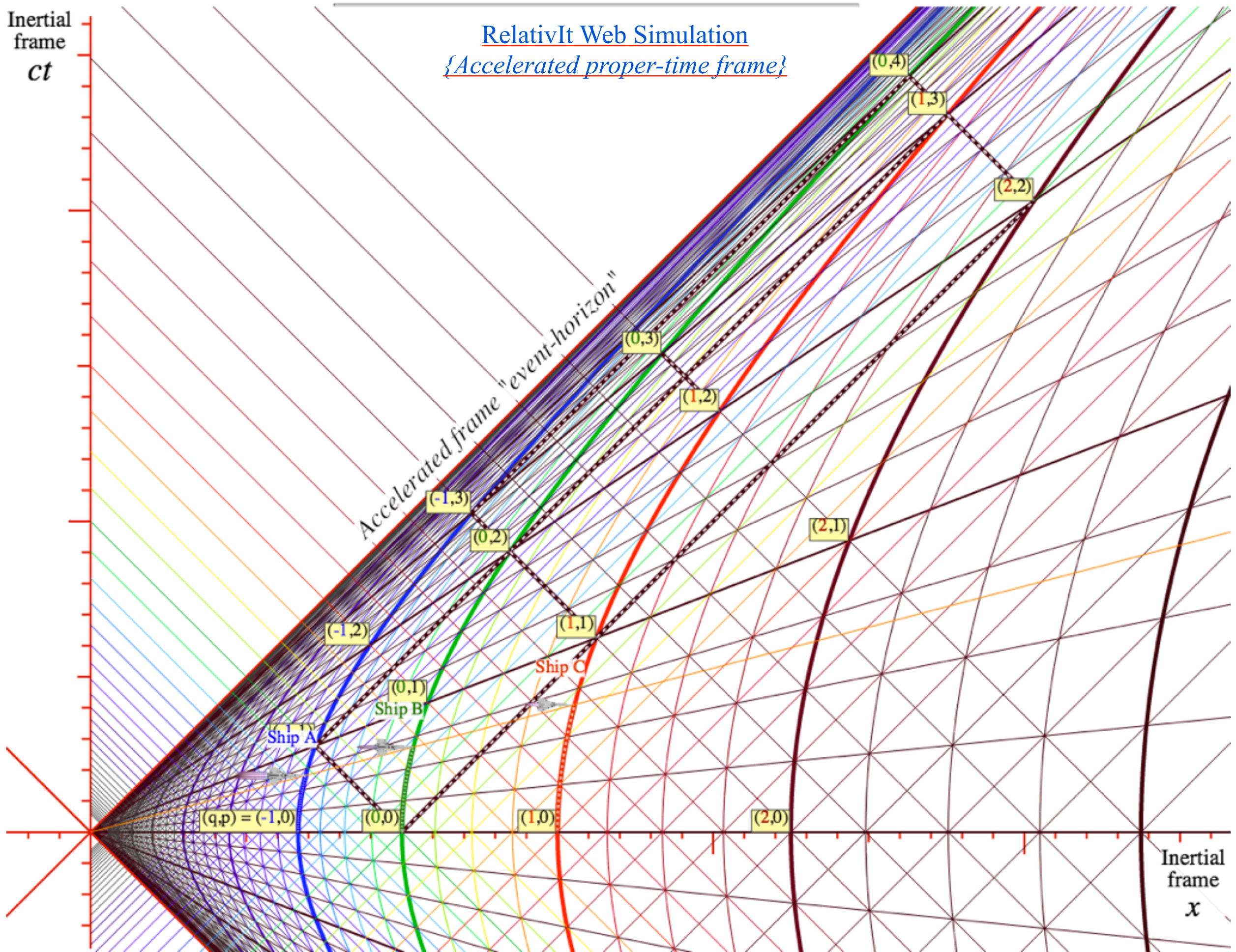
*Relativity* in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

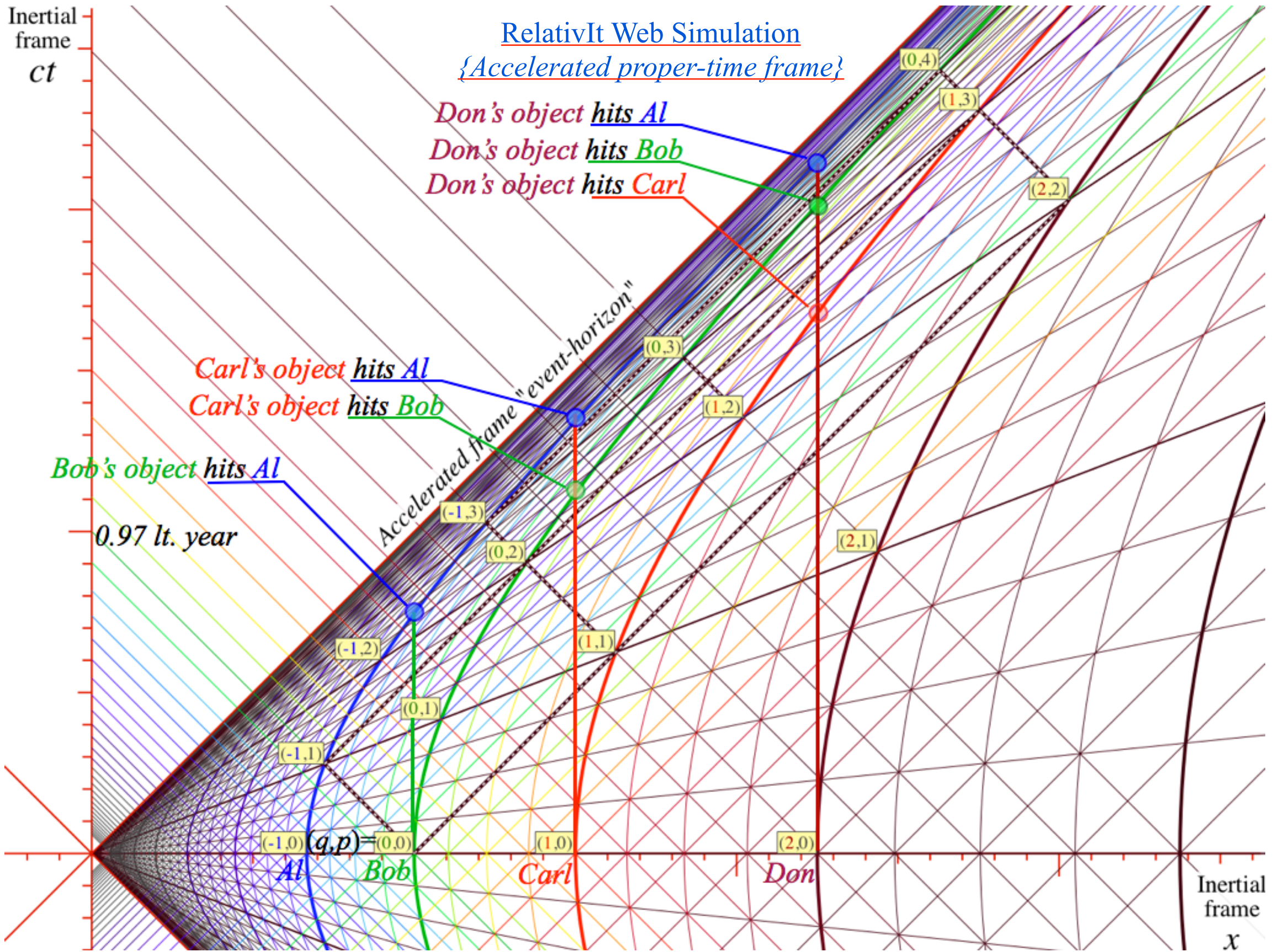
Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

➔ Animation of mechanics and metrology of constant- $g$  grid



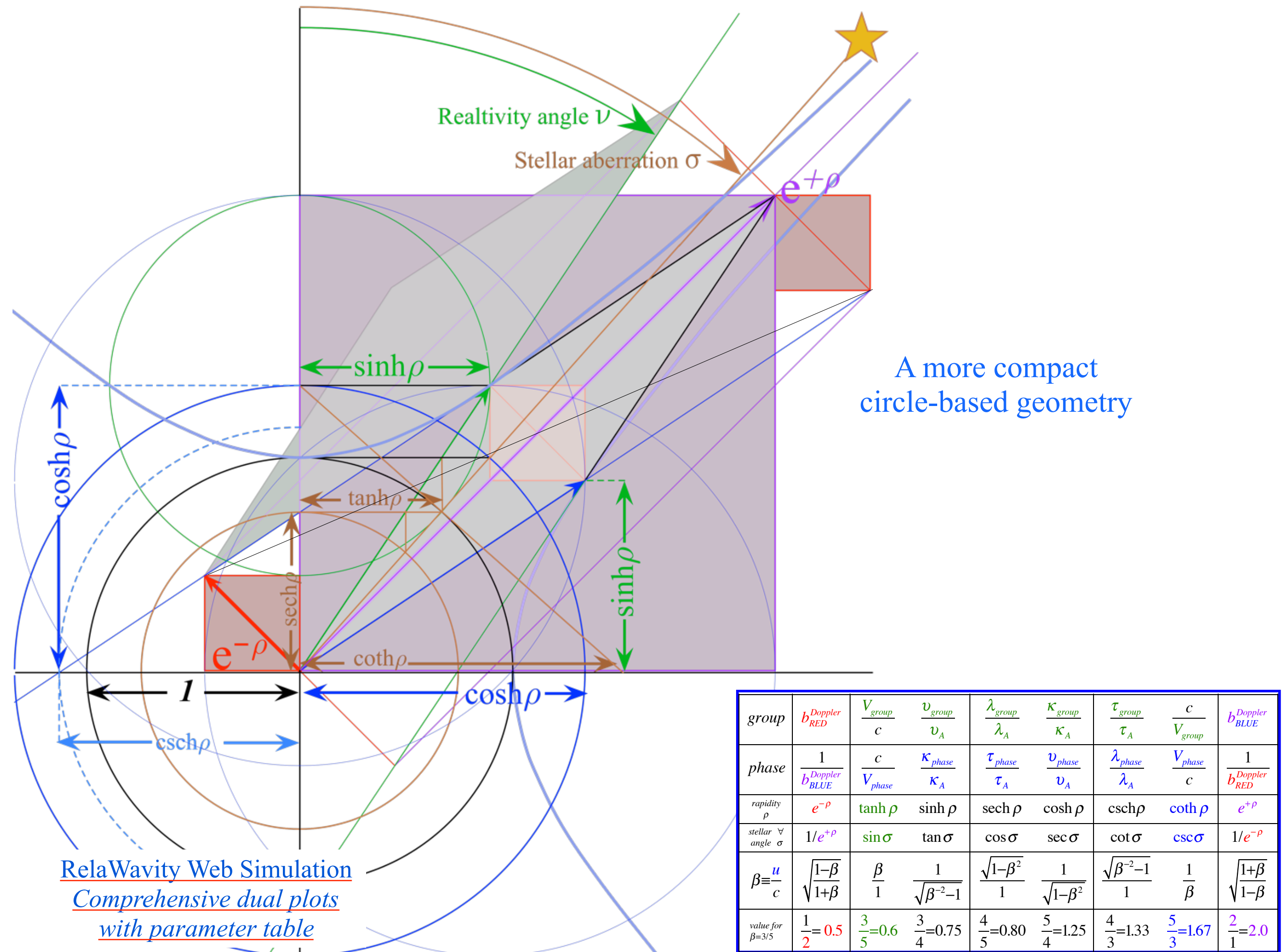












A more compact circle-based geometry

RelaWavity Web Simulation  
Comprehensive dual plots  
with parameter table

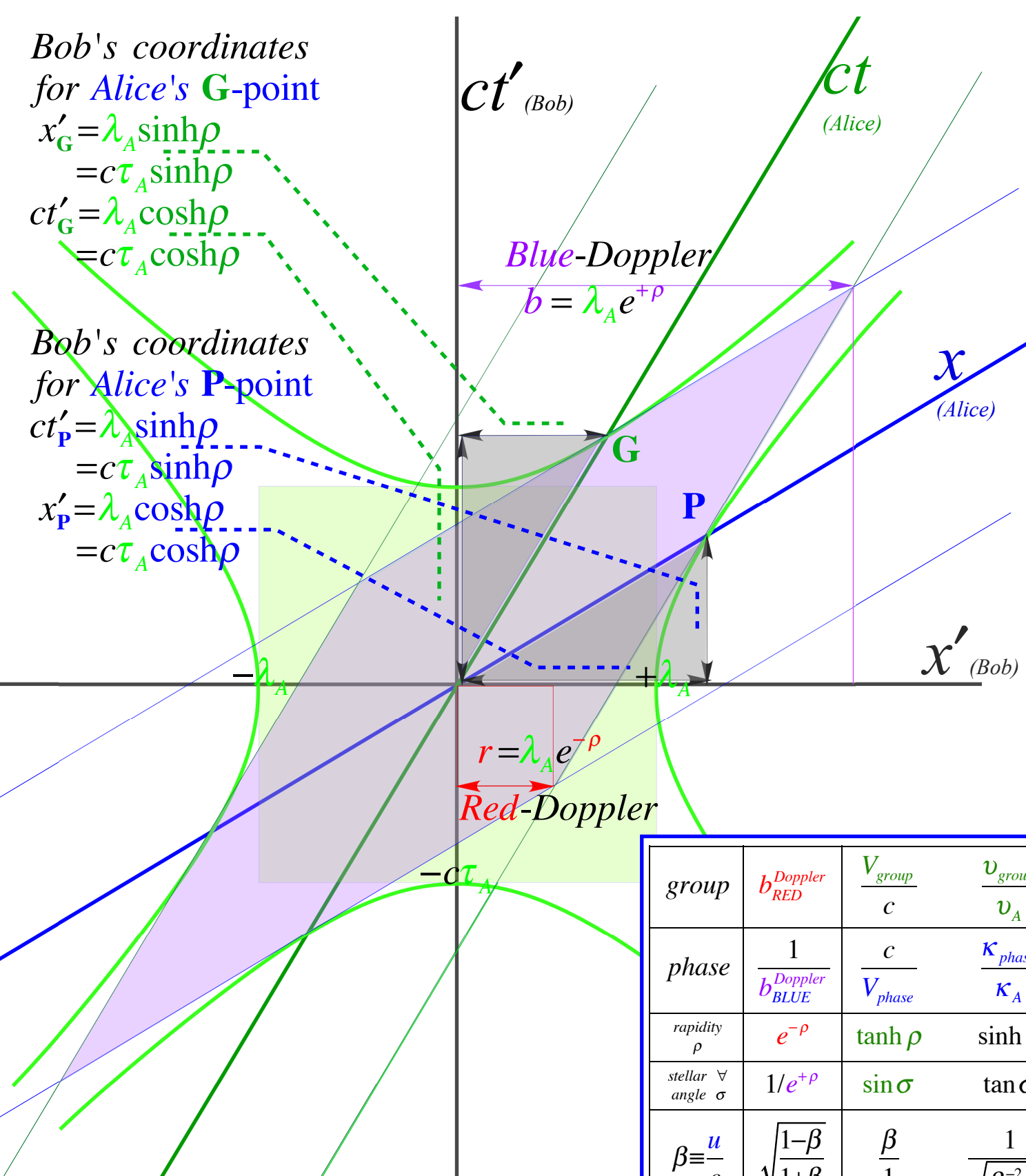
group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\forall$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Bob's coordinates  
for Alice's **G**-point

$$\begin{aligned} x'_G &= \lambda_A \sinh \rho \\ &= c\tau_A \sinh \rho \\ ct'_G &= \lambda_A \cosh \rho \\ &= c\tau_A \cosh \rho \end{aligned}$$

Bob's coordinates  
for Alice's **P**-point

$$\begin{aligned} ct'_P &= \lambda_A \sinh \rho \\ &= c\tau_A \sinh \rho \\ x'_P &= \lambda_A \cosh \rho \\ &= c\tau_A \cosh \rho \end{aligned}$$



Space-time parameters

$$\begin{aligned} \lambda_{phase} &= \lambda_A \operatorname{csch} \rho \\ \lambda_{group} &= \lambda_A \operatorname{sech} \rho \\ c\tau_{phase} &= c\tau_A \operatorname{sech} \rho \\ c\tau_{group} &= c\tau_A \operatorname{csch} \rho \end{aligned}$$

Per-space-time parameters

$$\begin{aligned} cK_{phase} &= cK_A \sinh \rho \\ cK_{group} &= cK_A \cosh \rho \\ v_{phase} &= v_A \cosh \rho \\ v_{group} &= v_A \sinh \rho \end{aligned}$$

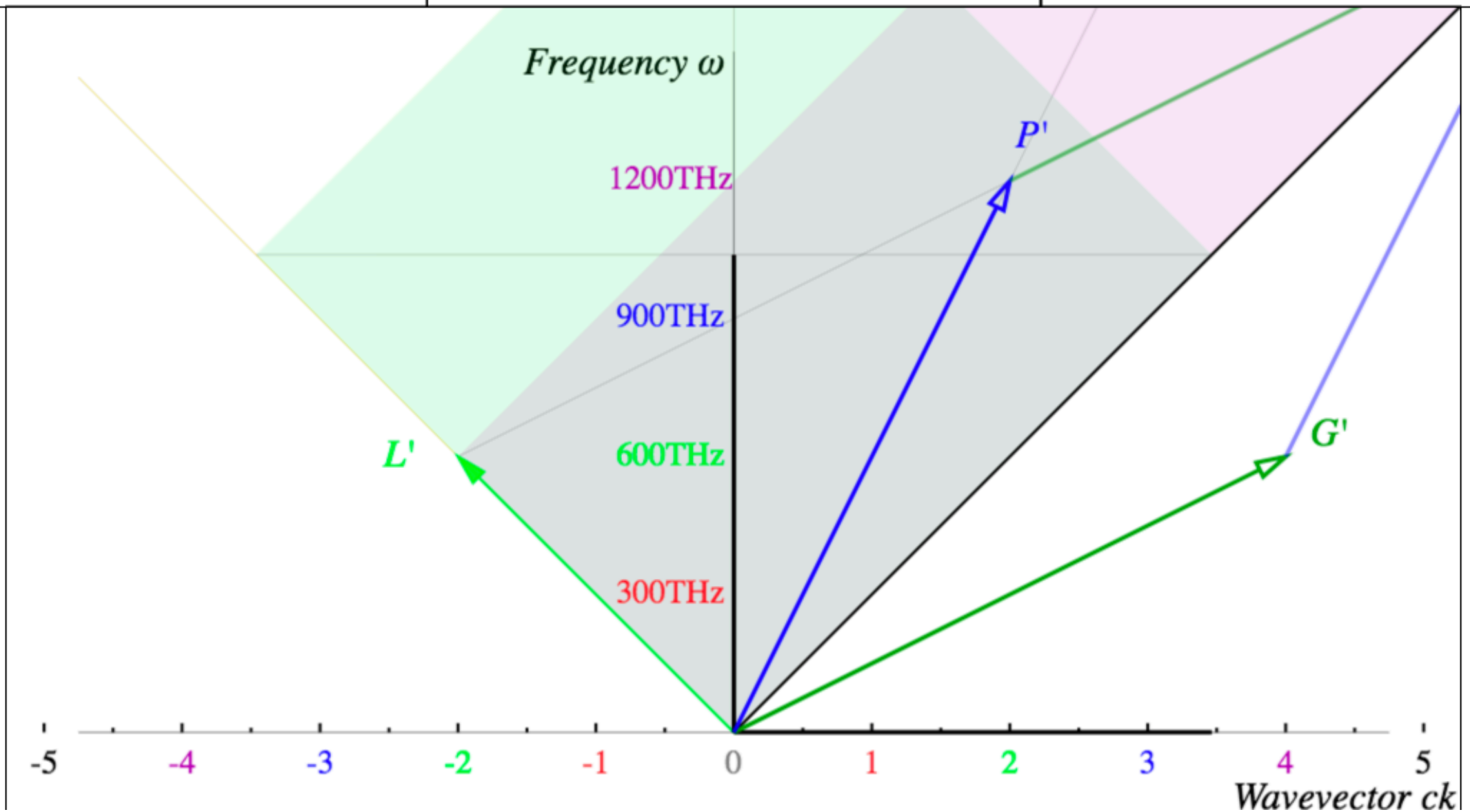
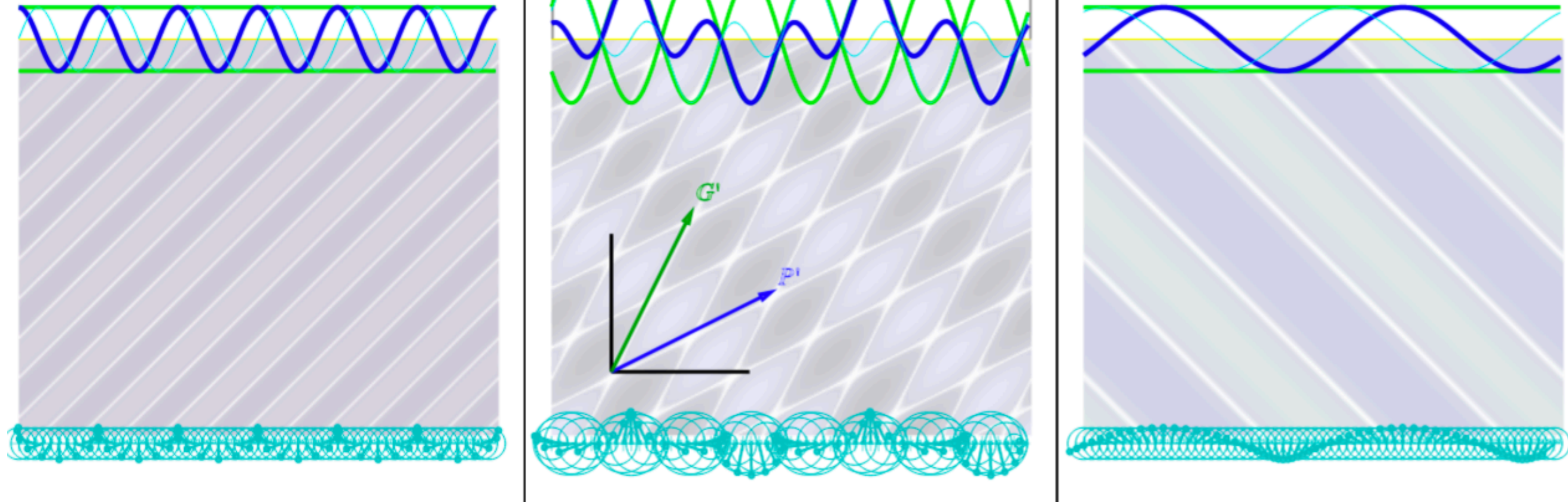
RelaWavity Web Simulations

[Comprehensive dual plots](#)  
[with parameter table](#)

[ct' vs x' with parameter table](#)

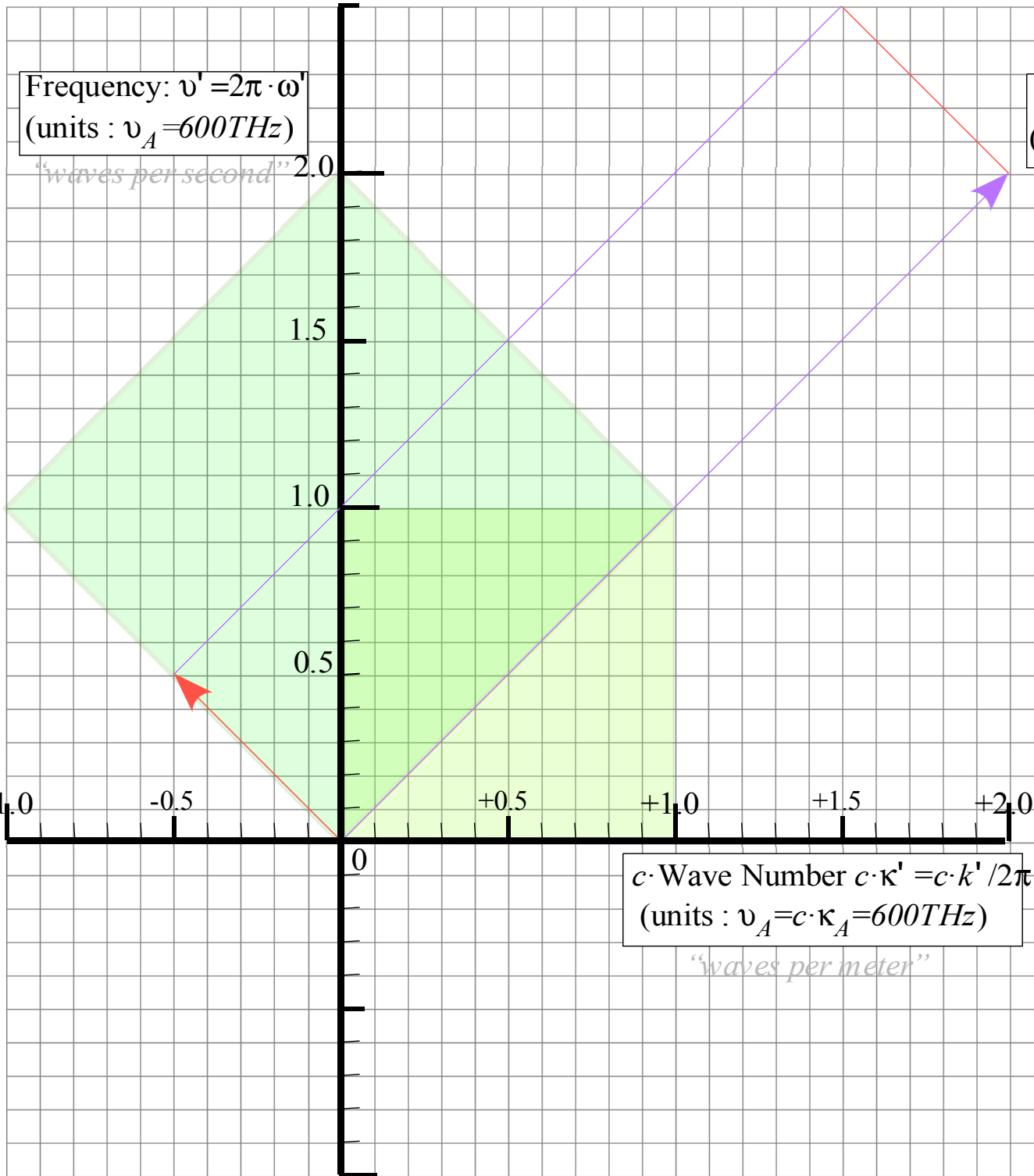
group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
effects	$b_{RED}^{Doppler}$	$V_{group}$	past-future asymmetry (off-diagonal Lorentz-transform)	$x$ -contraction <sup>(Lorentz)</sup> $\tau_{phase}$ -contraction	$t$ -dilation <sup>(Einstein)</sup> $v_{phase}$ -dilation (on-diagonal Lorentz-transform)	inverse asymmetry	$V_{phase}$	$b_{BLUE}^{Doppler}$







Per-space-time ( $\nu'$ ,  $c\kappa'$ ) geometry of 2-CW vectors



Space-time ( $c\tau'$ ,  $x'$ ) geometry of 2-CW paths

